



Overview of High Energy Experiments

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2024.7.17-18 @ VSON2024, Quy Nhon

Who am I?

Professor in Kyoto University, Japan



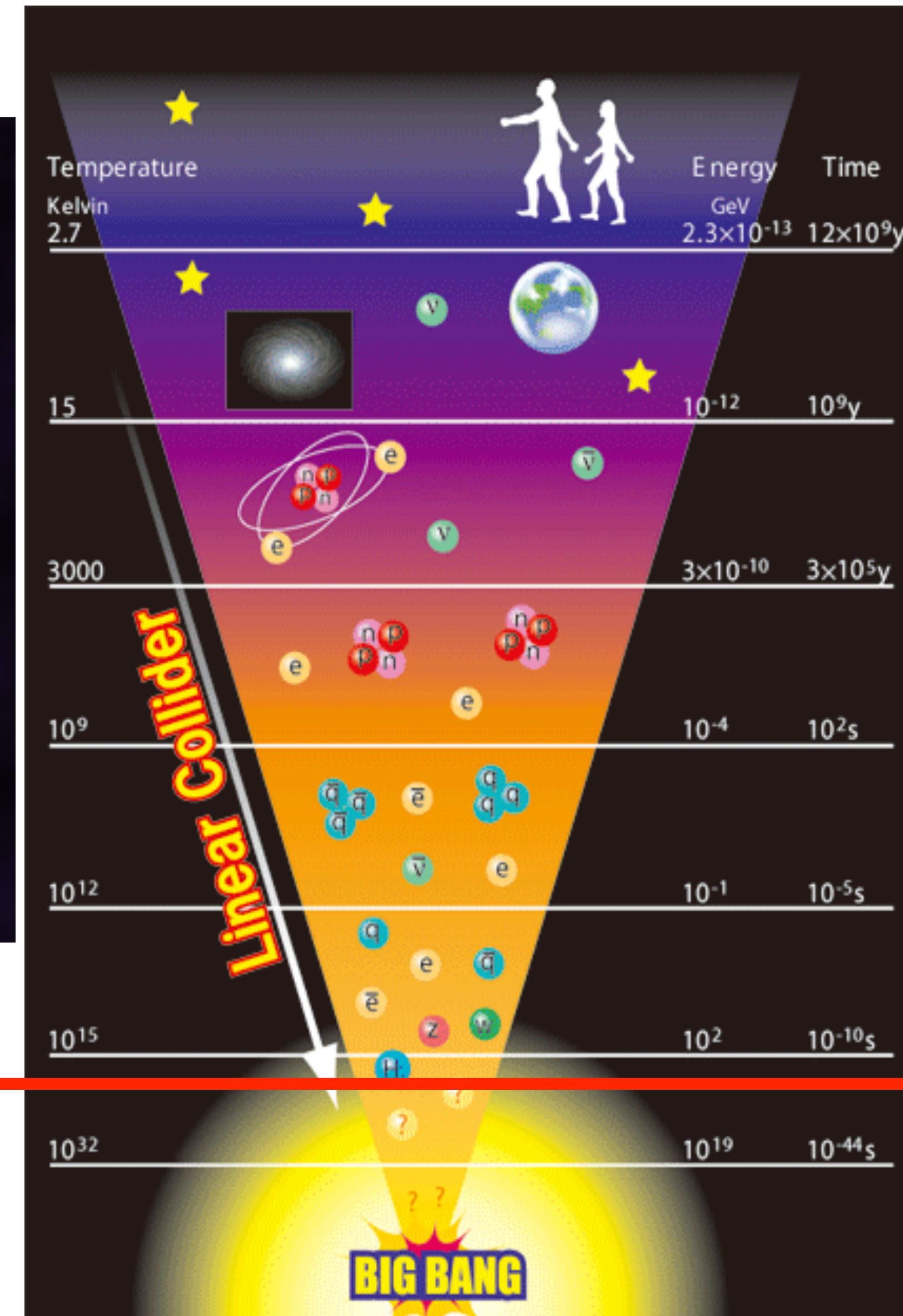
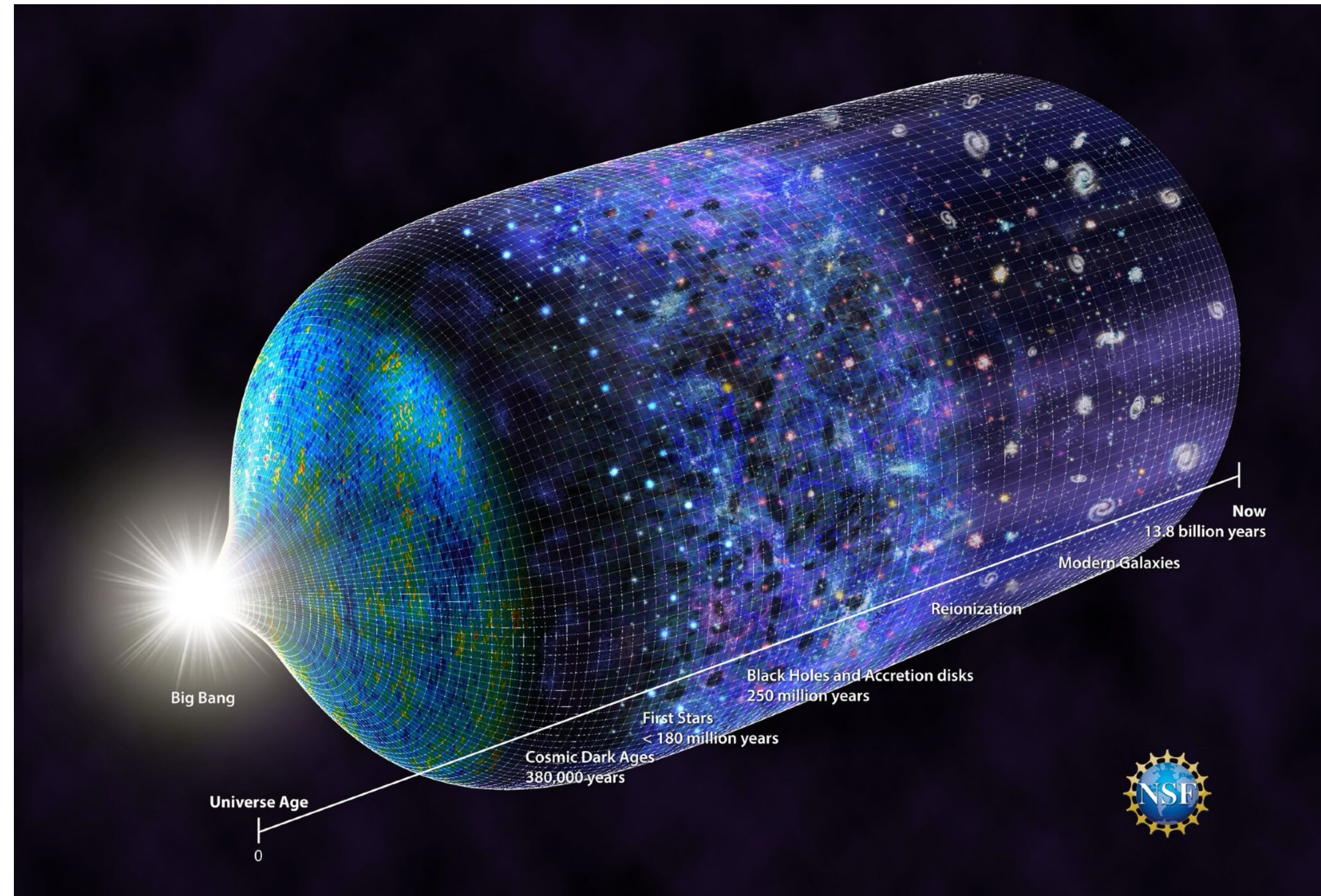
- 1995: Ph.D.
- 1995-1997: JSPS Fellow at Fermilab, US
 - working for the Kaon experiments.
- 1997-1999: Fermi Fellow at U. Chicago, US.
 - working for the Collider experiment at Fermilab (CDF)
- 1999-now: Faculty members in Kyoto U.
 - working for neutrino experiments: K2K, Super-K, SciBooNE, T2K, NINJA, Hyper-K, AXEL

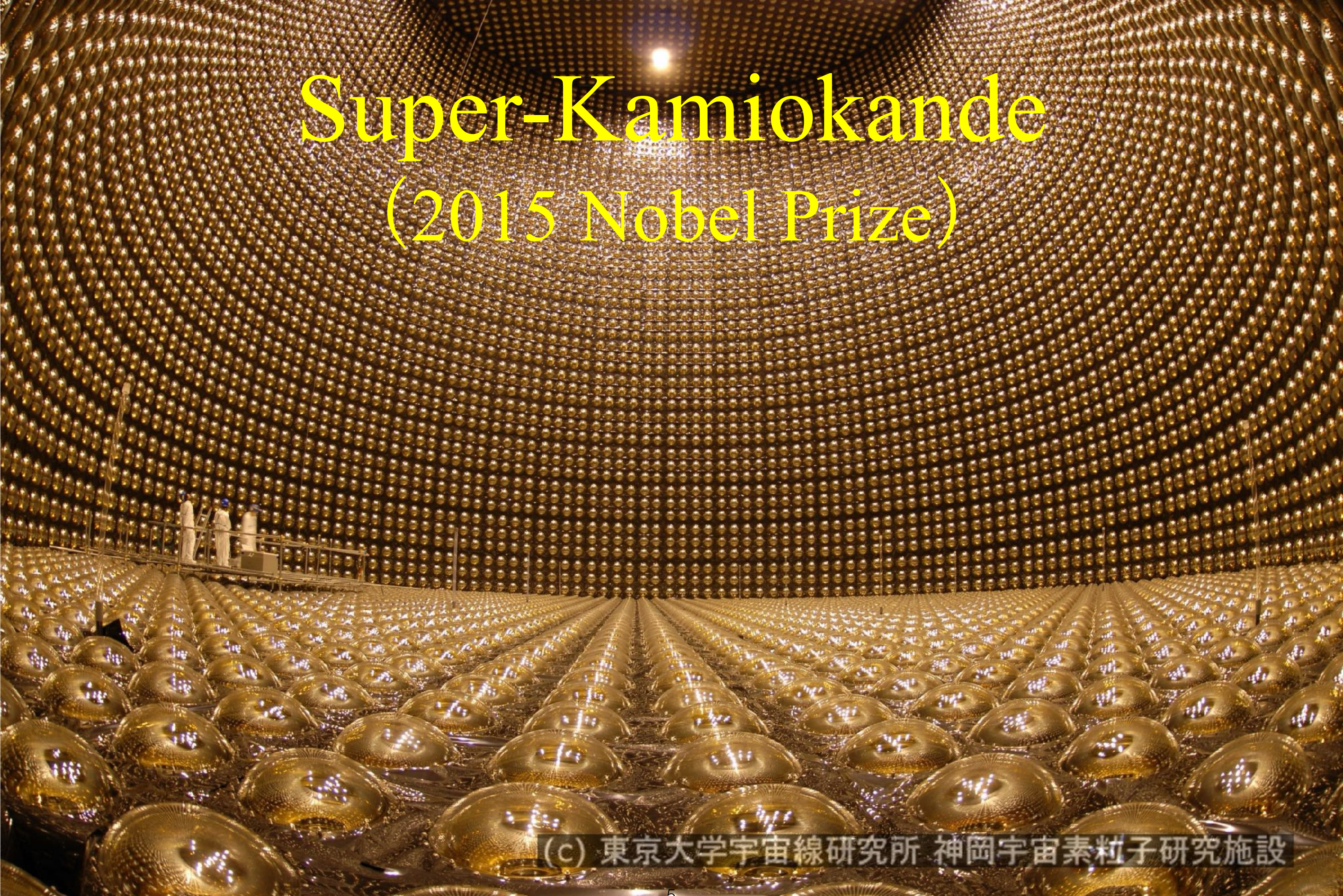
High Energy Physics

- **Study the ultimate building blocks of the universe and matter.**
 - What is matter made of? What exists in the universe? What are the properties of those components? How do they interact?
 - Standard Model of Particle Physics: quarks, leptons, gauge particles, Higgs boson.
 - Unknown entities: dark matter, dark energy, other unknown particles, unexpected phenomena.
 - Theoretical guidance: supersymmetric particles, extra dimensions
- Studies that question the meaning of time and even space (spacetime).
 - Origin of mass, space-reversal symmetry, time-reversal symmetry breaking.
- (Personal view) Differences with other disciplines and fields
 - Eliminates the complexity (diversity) of many-body systems and explores the underlying physical laws and views of physics.

Explore frontiers!

After the birth of the universe (after the Big Bang), to a world of 10^{-10} (0.1 nano) seconds.

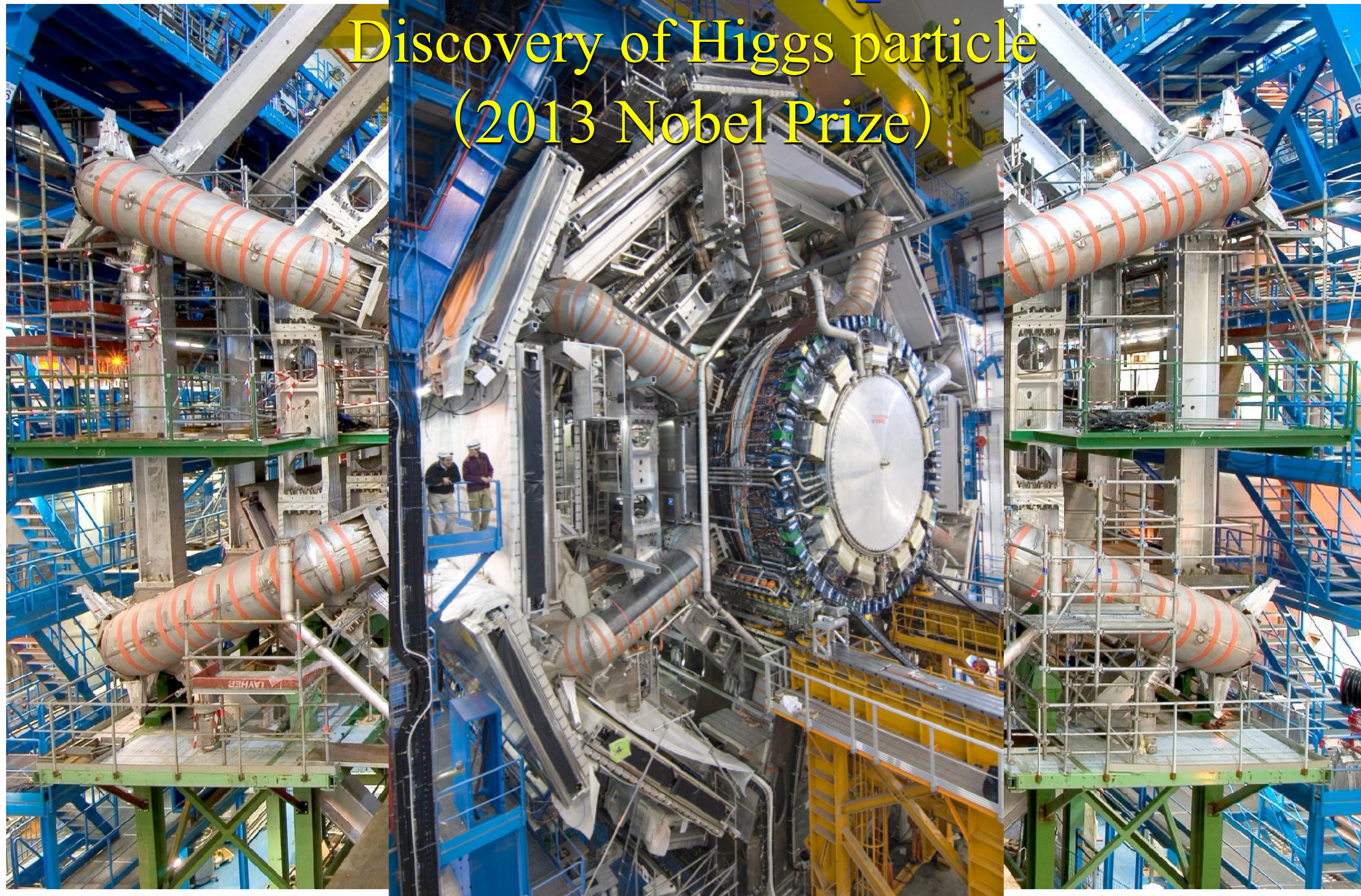




Super-Kamiokande (2015 Nobel Prize)

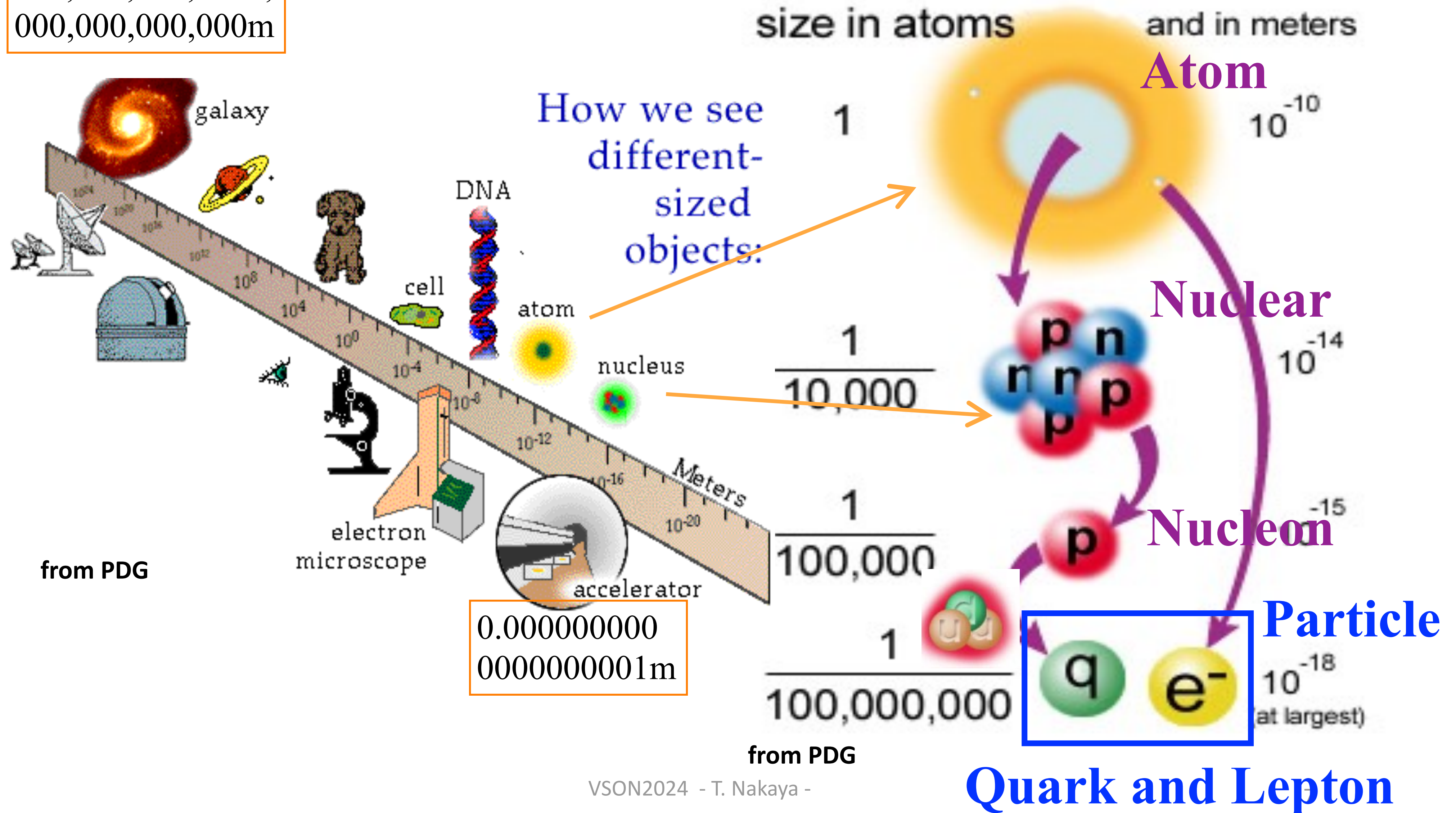
(c) 東京大学宇宙線研究所 神岡宇宙素粒子研究施設

LHC/ATLAS Experiment



Particle Physics

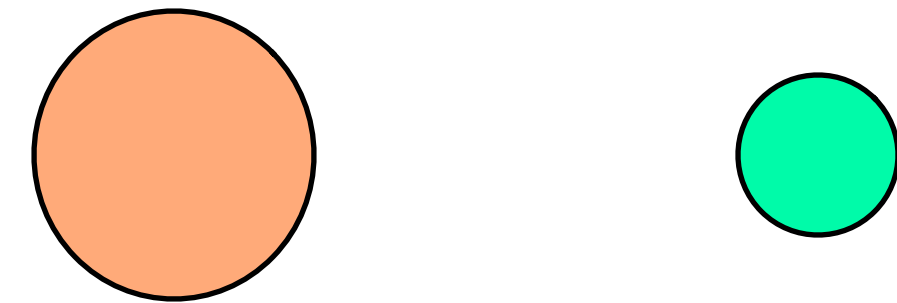
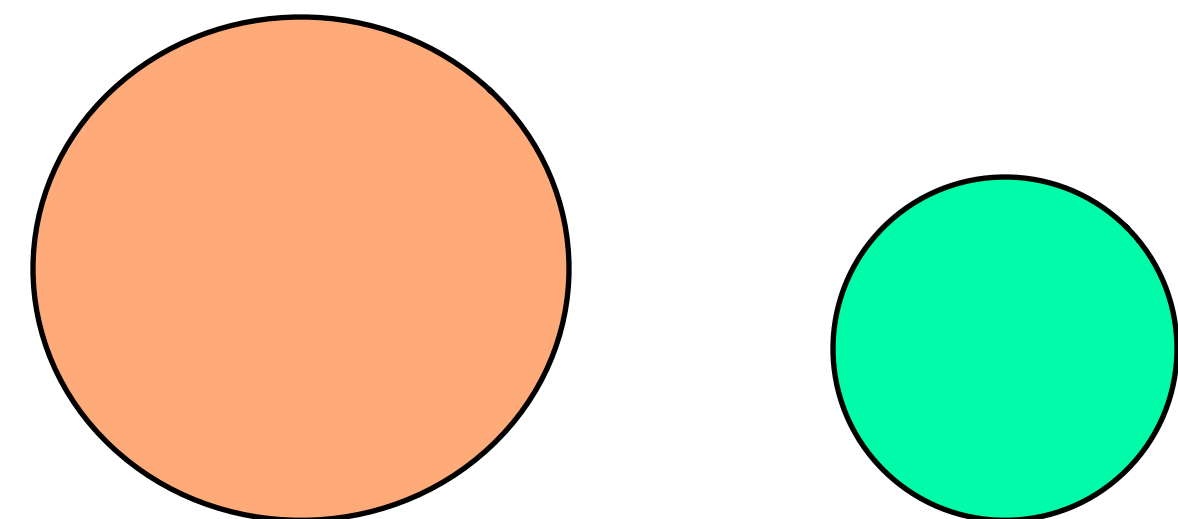
100,000,000,0000,
000,000,000,000m



Quark and Lepton

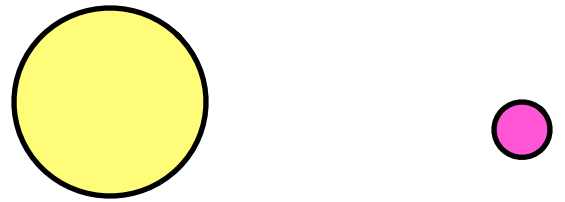
× 2 (Antiparticles)

Quark



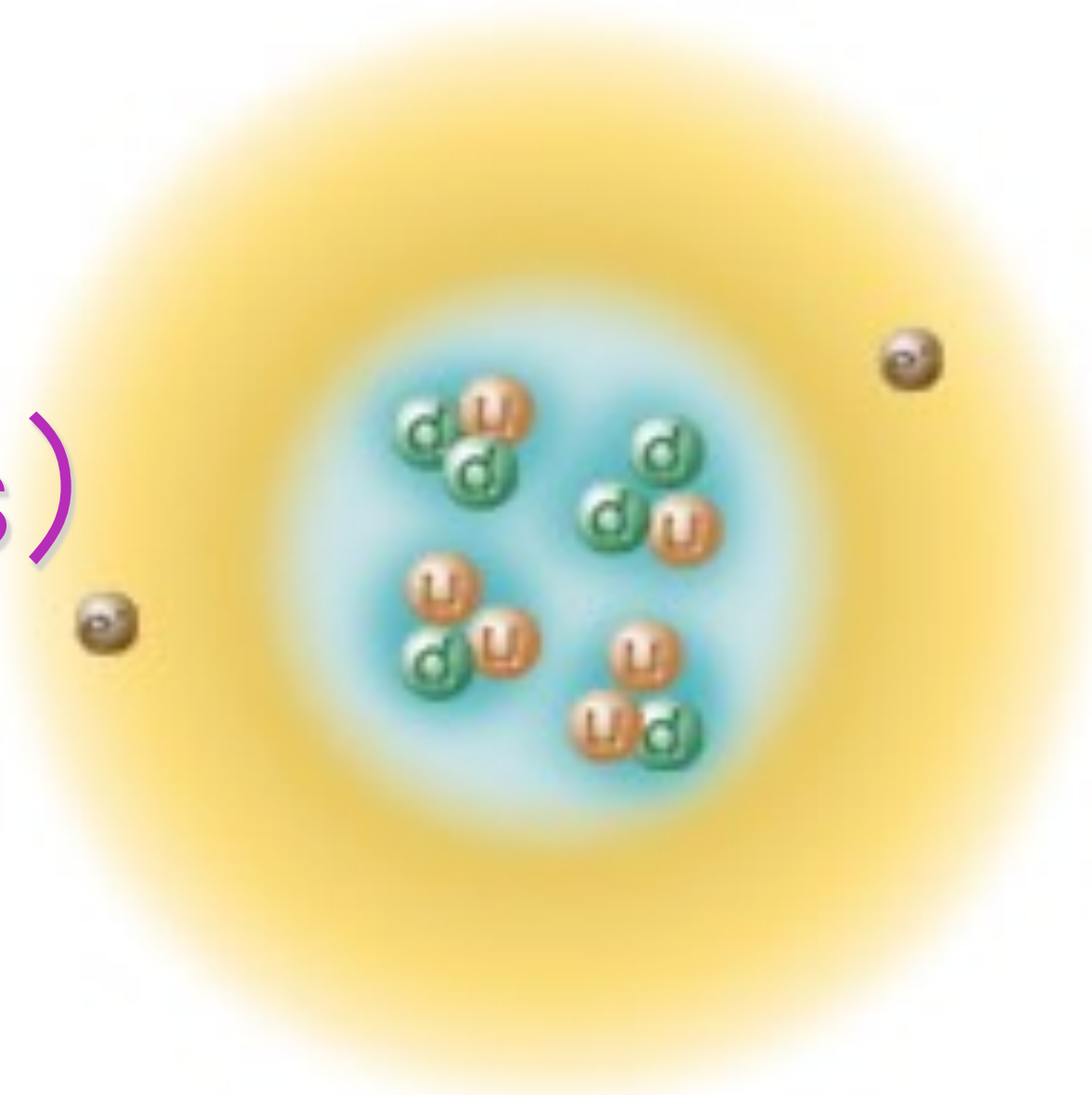
uud(proton)

Lepton

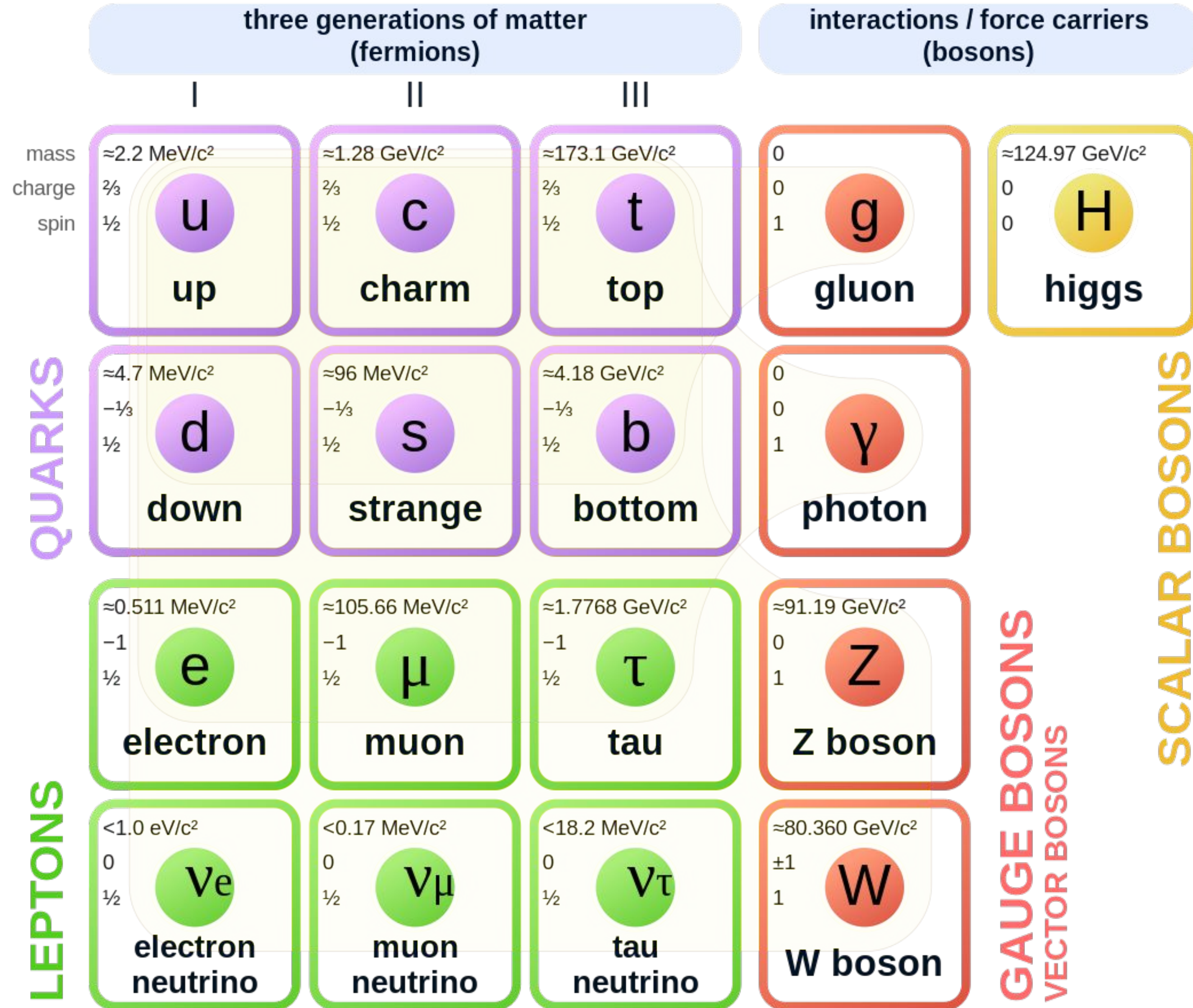


Electron

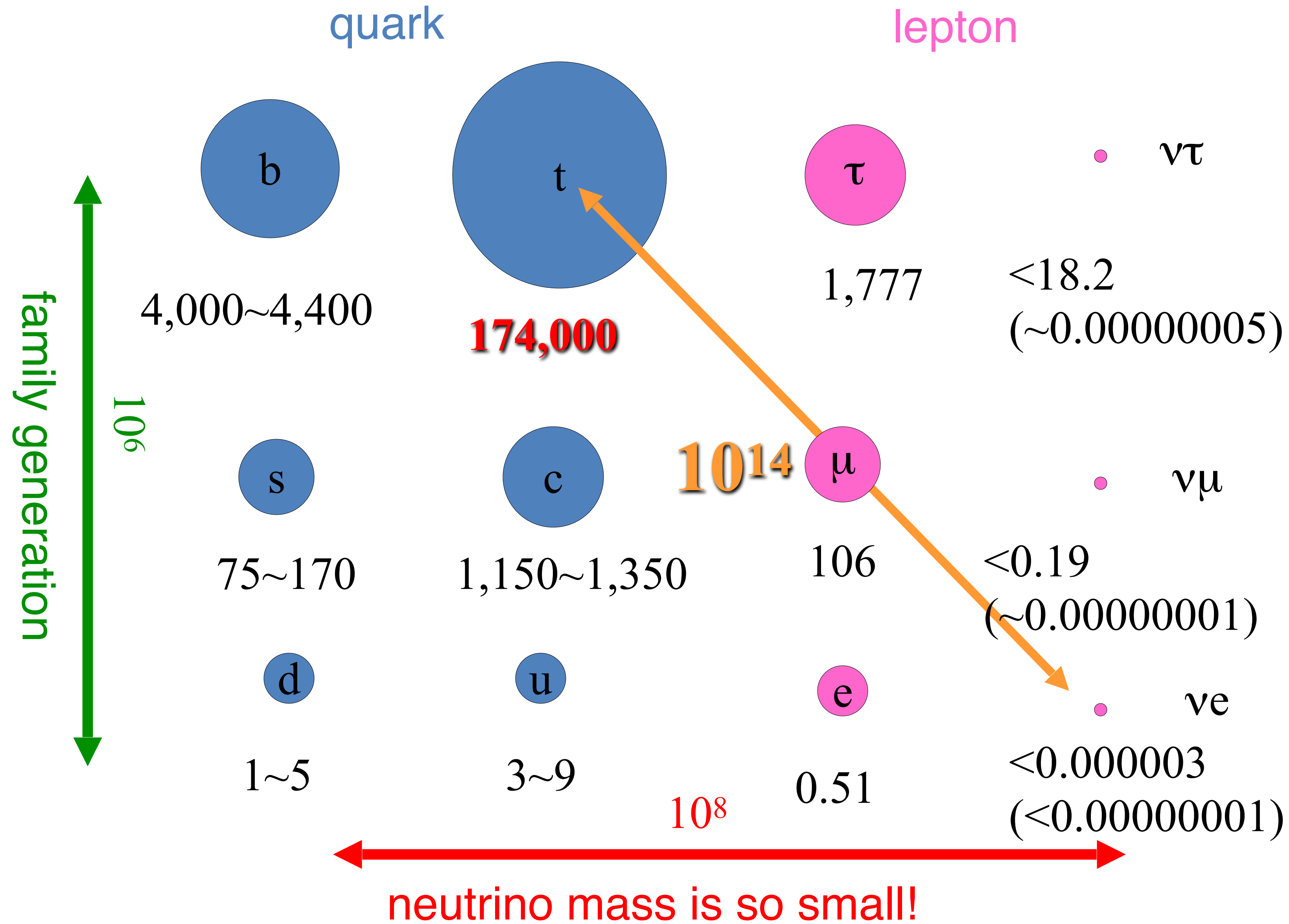
Neutrino



Standard Model of Elementary Particles



Mass (MeV) : $m_{\Psi\bar{\Psi}}$



Fundamental Questions!

not solved yet

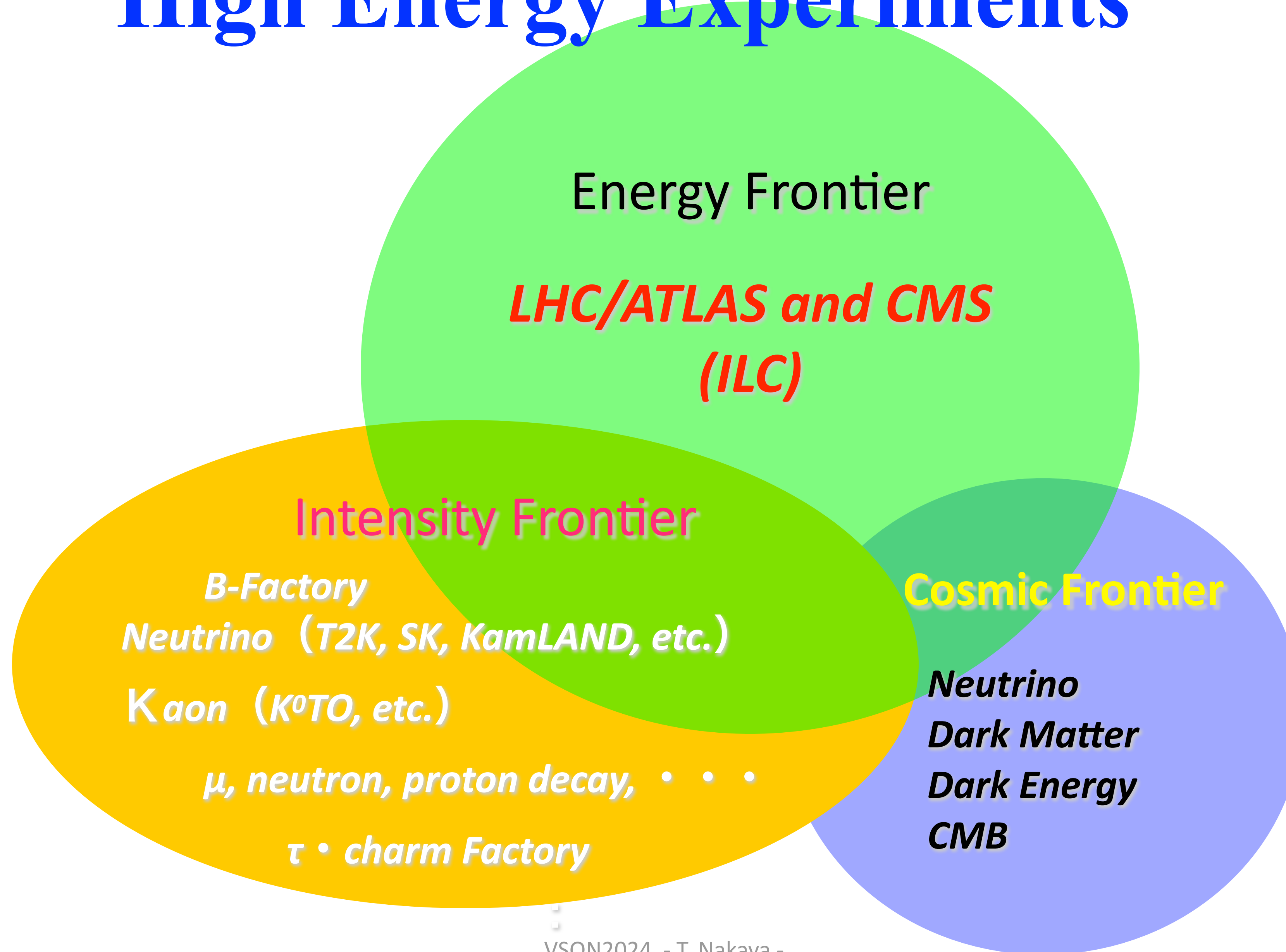
- What is dark matter?
- Where are antimatter? How did they disappear?
- The accelerating universe?
- Why is neutrino mass so small?
- How are the forces are unified?
- Generation structure. Why are the masses of particles are so different?

High Energy Physics and Experiments

Three categories

- Energy Frontier
 - Study particles at high energy (\sim TeV)
- Intensity Frontier
 - Study particles with high precision and look for rare processes
- Cosmic Frontier
 - Observe our universe and look for footprints made by past particles.

High Energy Experiments



What experiments are there?

- CERN (Europe)
 - **LHC/ATLAS, CMS**, LHCb, FASAR, milliQan, NA62, DsTau, (CEPE)
- PSI (Switzerland)
 - MEG II
- Fermilab (US)
 - Muon g-2, Mu2e, Neutrino experiments: DUNE, NOvA, SBN, ICARUS, etc.
- KEK/J-PARC (Japan)
 - **Super-KEKB/Belle II**, J-PARC/T2K, KoTO, COMET, muon g-2, JSNS², Super-K, Hyper-K, KAGLA, KamLAND, (ILC)
- IHEP (China)
 - BESSIII, JUNO, (CEPC)
- *Cosmology and Astro-particle experiments*

1. Energy Frontier



LHC

Large Hadron Collider
(Superconducting Accelerator)

CERN and LHC

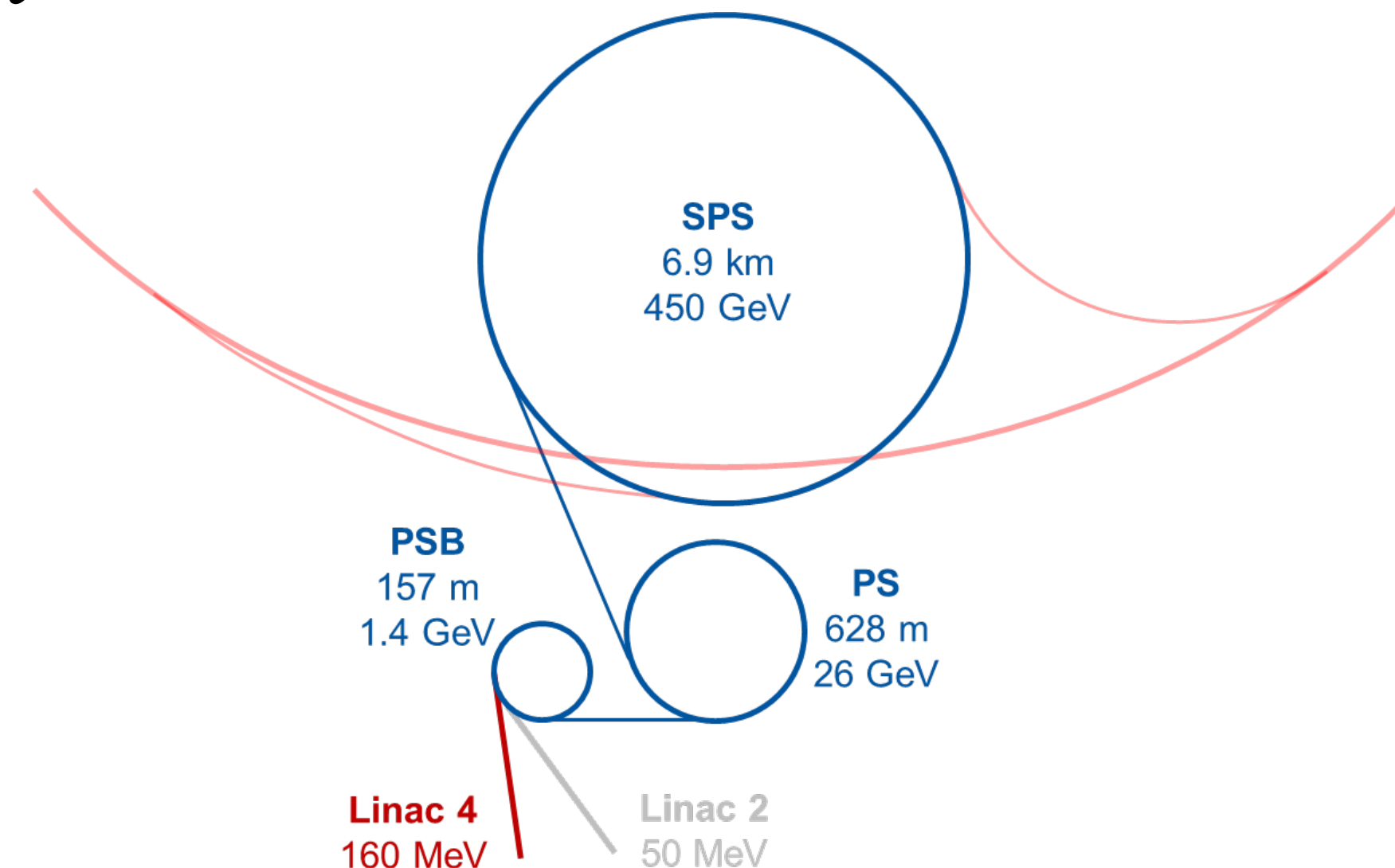
Circumference 27km
(Why does it require 27km?
Magnetic field: 8T)



Circular accelerator

- **Proton Synchrotron**

- Protons go round and round
 - Inject proton beam periodically
- Acceleration mode : Increasing energy with every lap
 - LHC: 450 GeV \rightarrow 6,800 GeV (6.8 TeV)
 - Increasing magnetic field synchronized with energy
- Colliding mode : constant energy
 - How many turns per second?
 - Circumference 27km

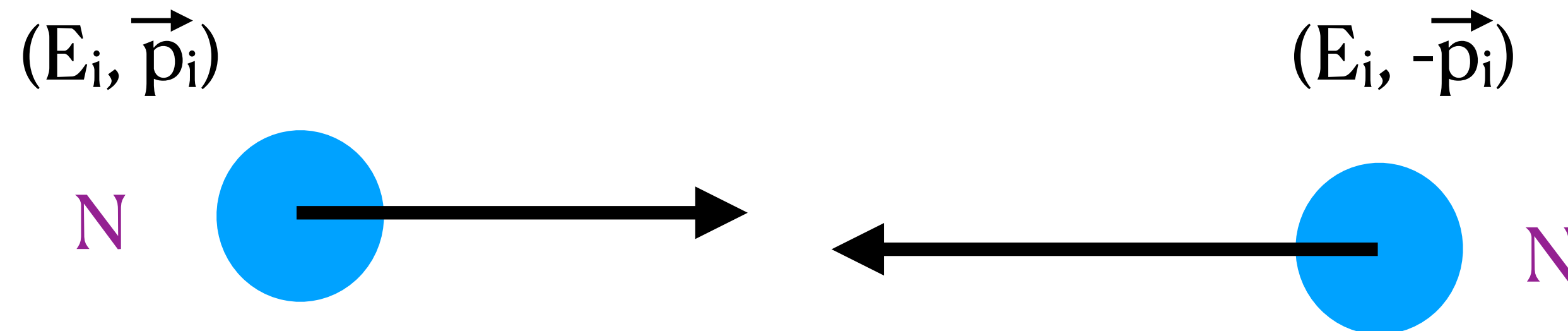


Collider experiment

- Collider experiment

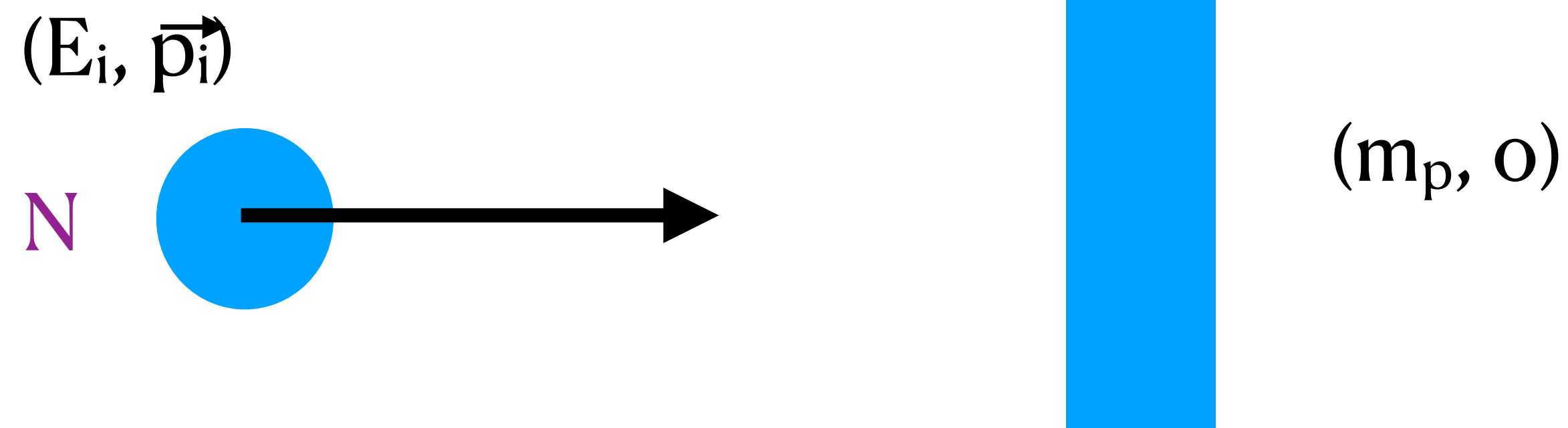
- Effective use of beam energy

How heavy subatomic particles can be made?

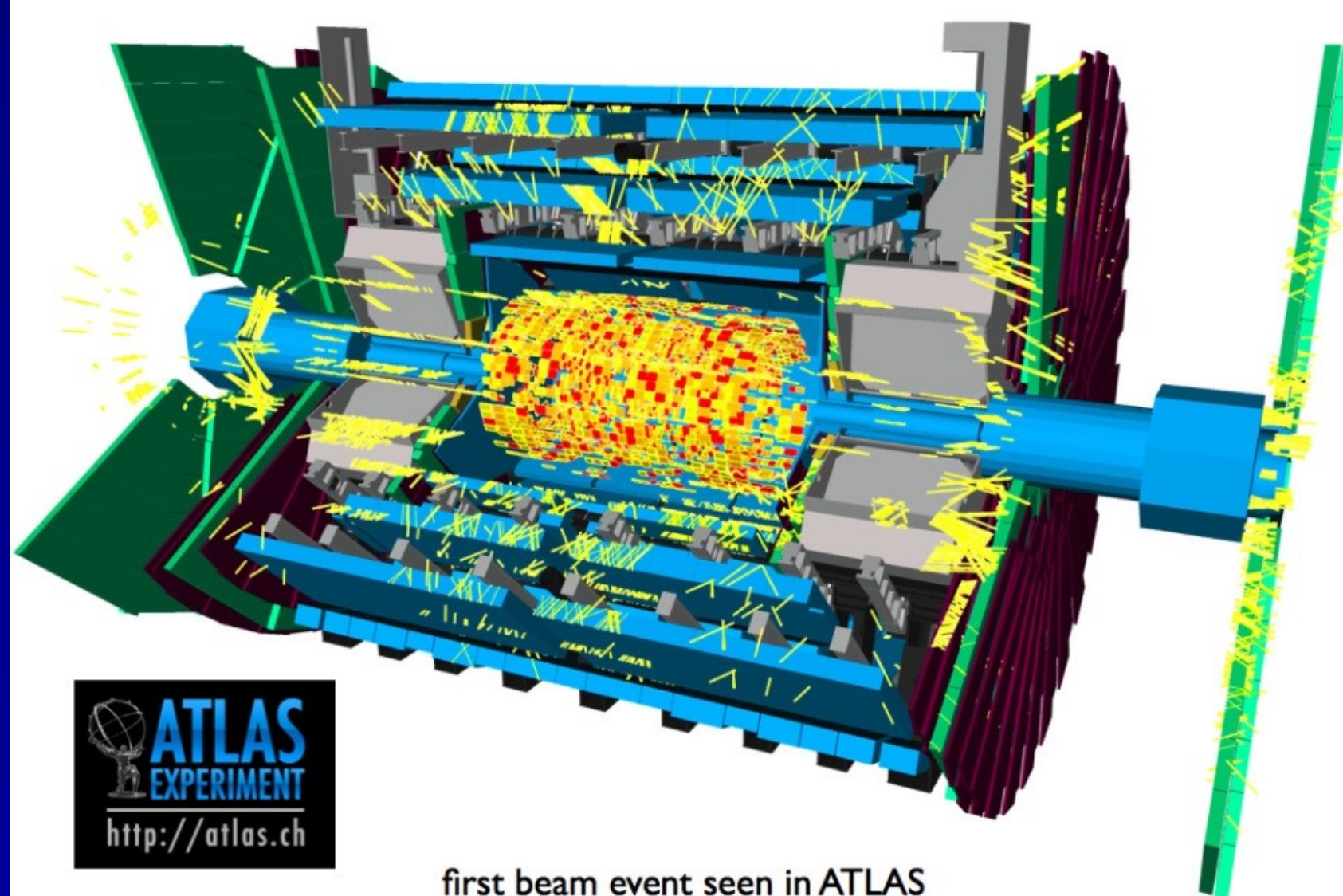


- Fixed target experiment

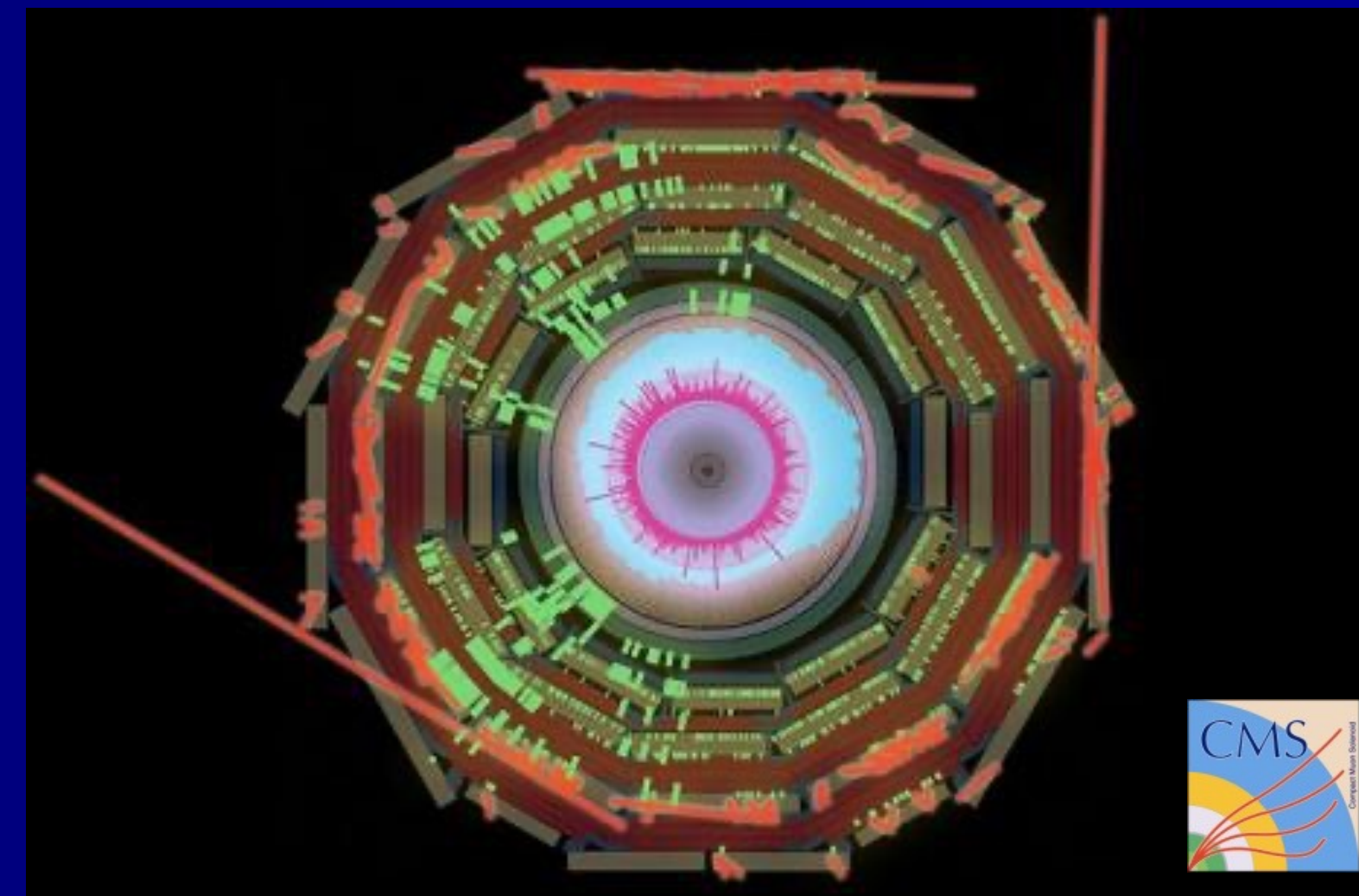
- Effective use of beam flux

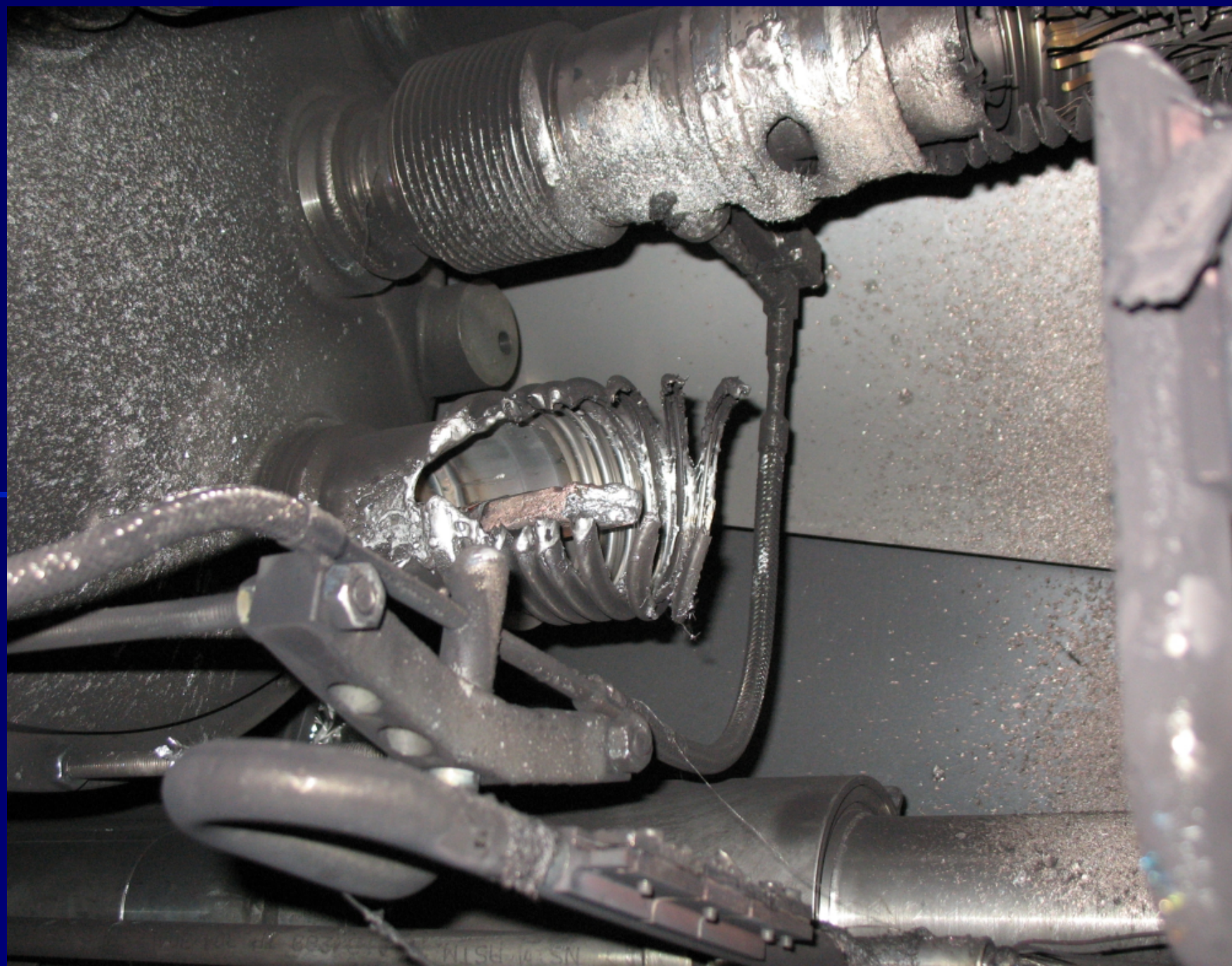


16 years ago
(2008/9/10)

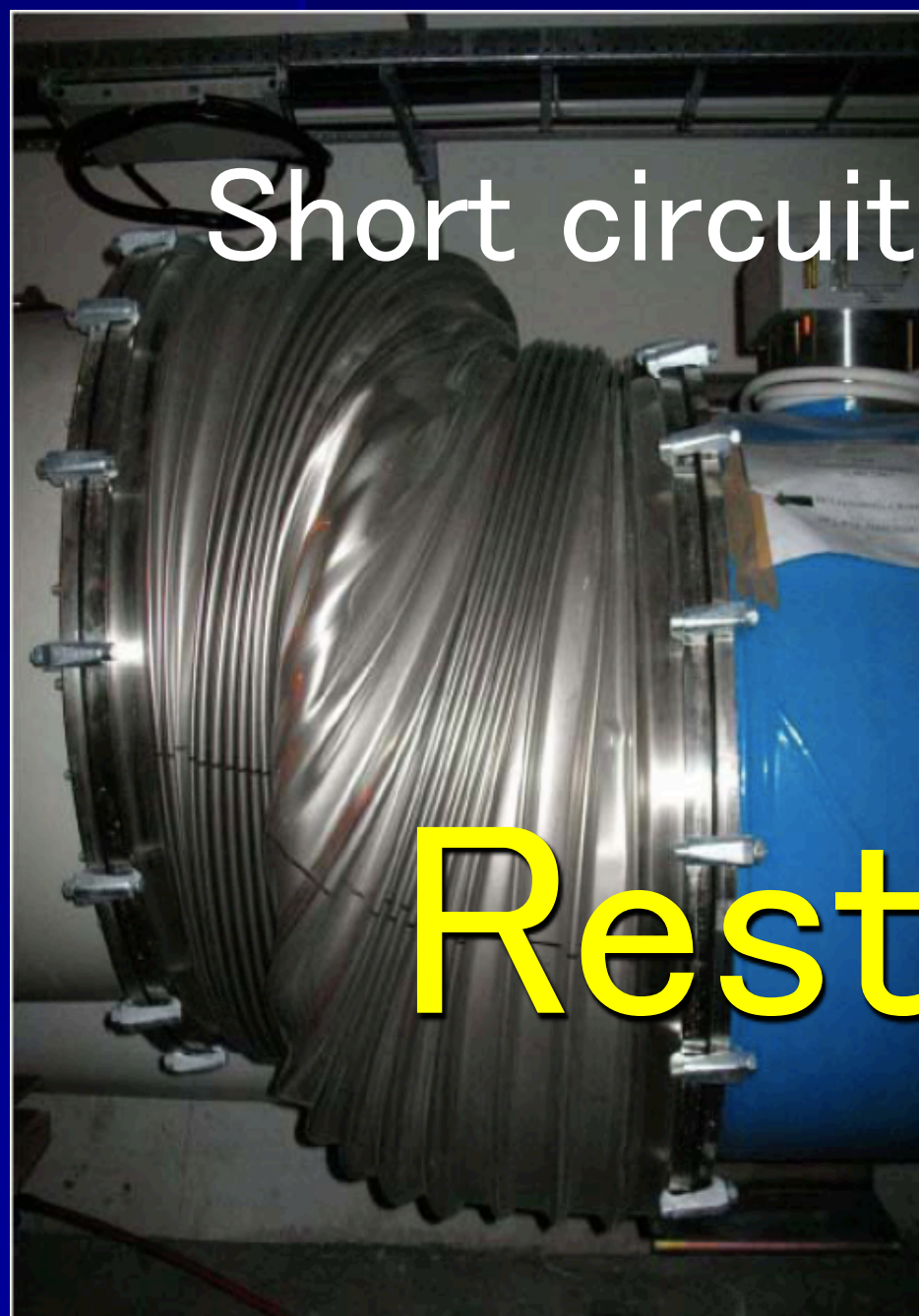


first beam event seen in ATLAS





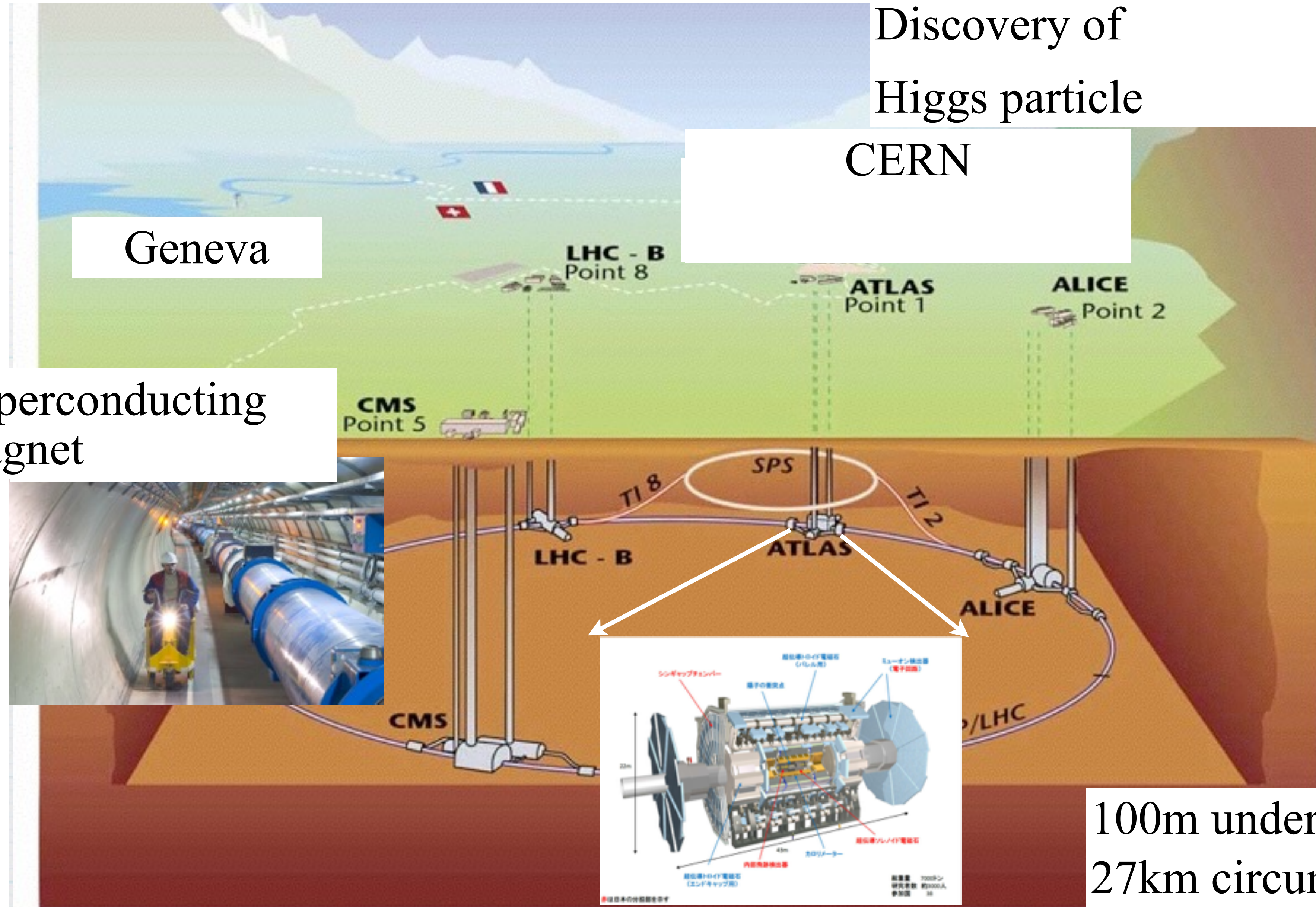
**9 days after
(2008/9/19)**



Short circuit → superconducting collapse
→ liquid helium boiling and explosion

Restart in 2010

LHC/ATLAS/CMS Proton-Proton Collider experiments



Discovery of Higgs particle

CERN

Geneva

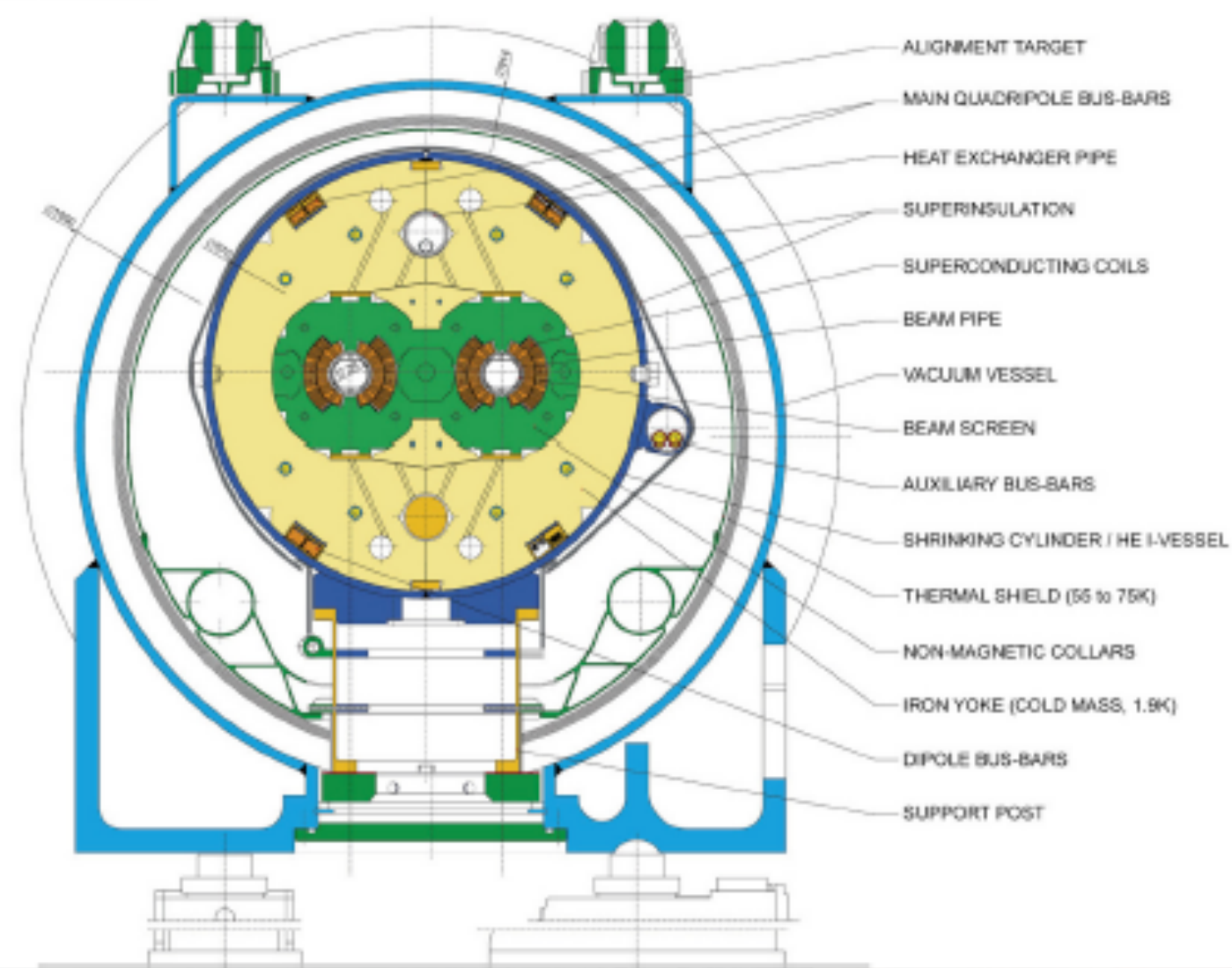
Superconducting magnet

100m underground
27km circumference

Large Hadron Collider (LHC) at CERN

- 14 TeV proton-proton collider
- 27 km circumference
- 14 years construction
- US 5B dollar for construction

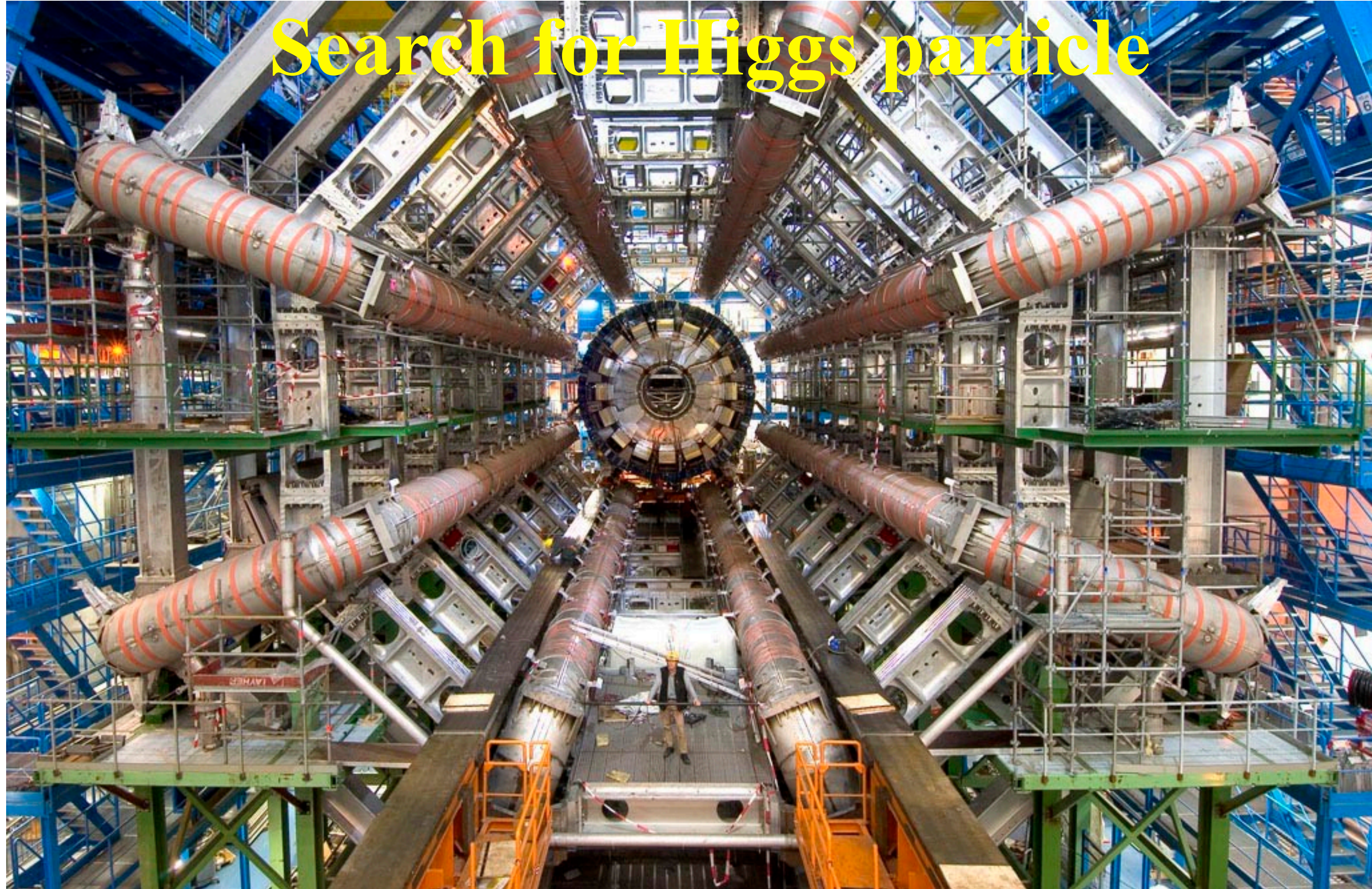
LHC DIPOLE : STANDARD CROSS-SECTION



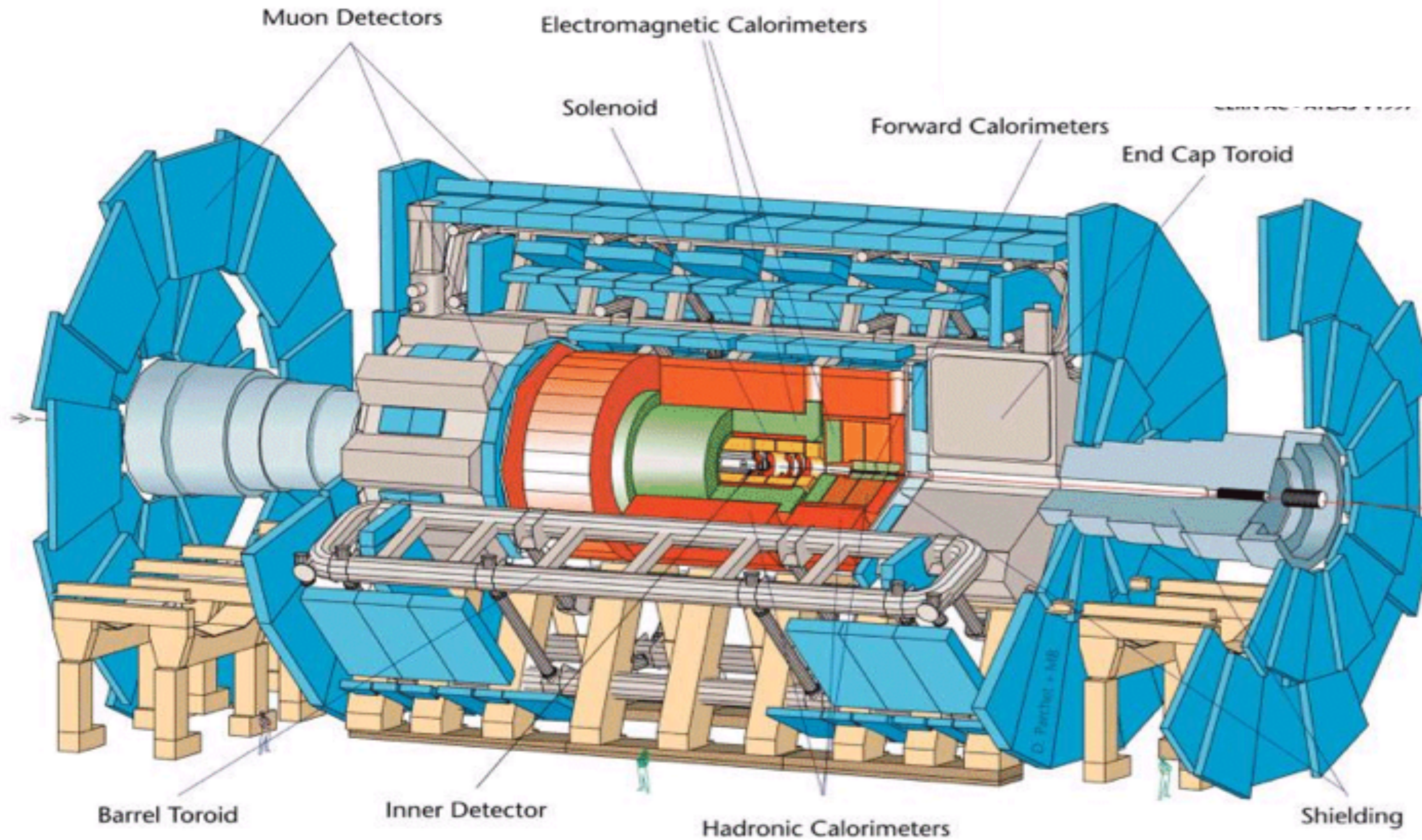
- 1.9 K cooling by Helium
- 8.3 T magnetic field with 1232 magnets

LHC/ATLAS experiments

Search for Higgs particle

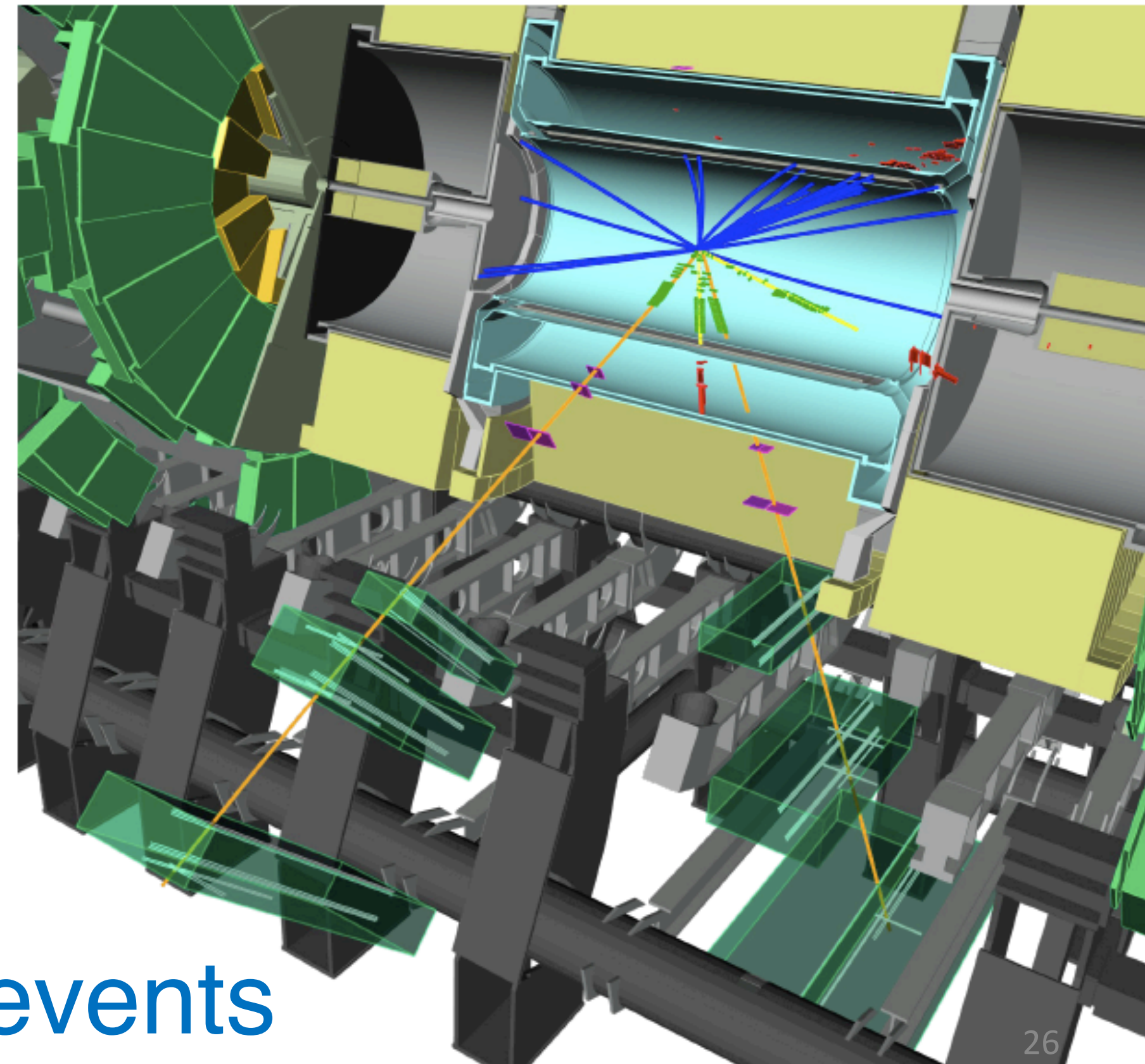


ATLAS detector



ATLAS detector

- Multi-purpose detector
 - Hight 25 m, Length 44m, Weight 7,000 ton
 - 1M signal channels



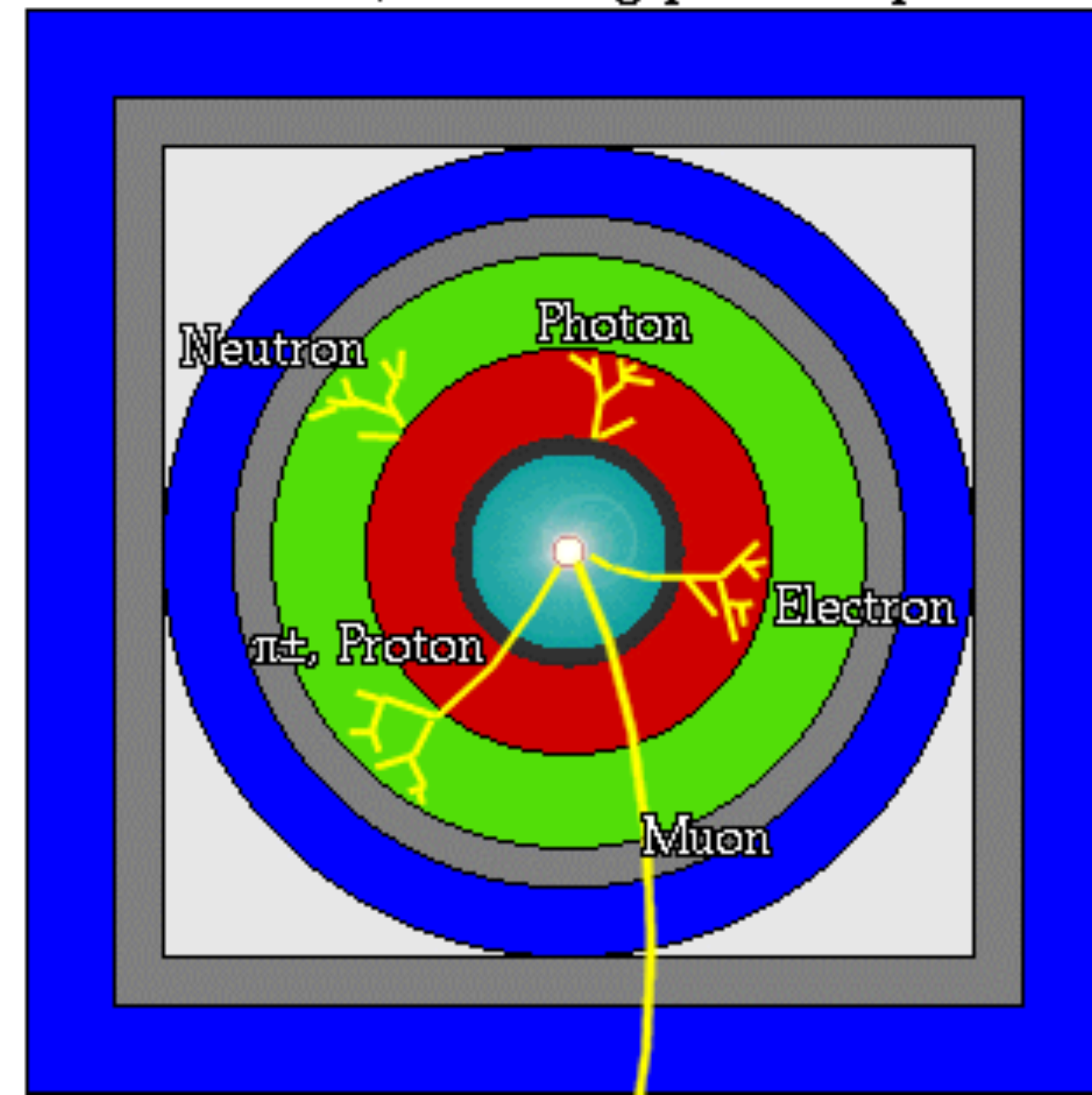
$H \rightarrow Z Z^* \rightarrow 4 l$ events

Why is multi-purpose detector?

- Particles observed
 - Electron
 - Photon
 - Proton (uud), π^+ (ud), K^+ (us)
 - Neutron (udd)
 - Muon
 - ?? Neutrino ??

A detector cross-section, showing particle paths

- Beam Pipe (center)
- Tracking Chamber
- Magnet Coil
- E-M Calorimeter
- Hadron Calorimeter
- Magnetized Iron
- Muon Chambers

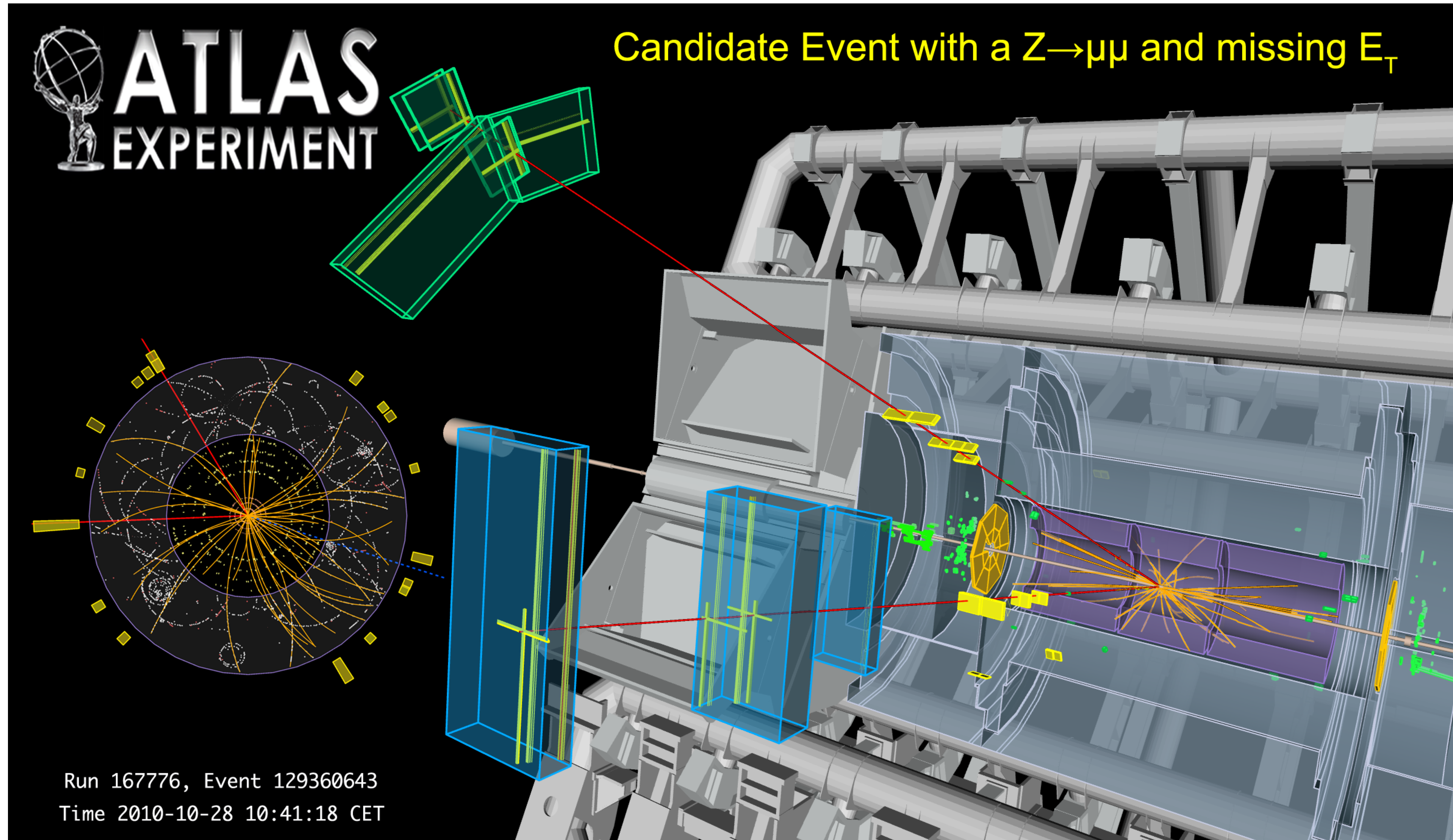


Principle of Particle measurements

- Tracking the particle trajectories under a magnetic field, and measuring their position and momentum.
- Next, photons and electrons are stopped and their energy and position are measured.
- Hadrons (mainly protons and pions) are then stopped and their energy and position measured.
- Finally, charged particles not stopped are muons.
- Non-observed particles are neutrinos, which are identified by missing energy.

Candidate for $ZZ \rightarrow \mu\mu\nu\nu$

- $m_{\mu\mu}$ 94 GeV, $E_{\tau}^{\text{miss}} = 161$ GeV

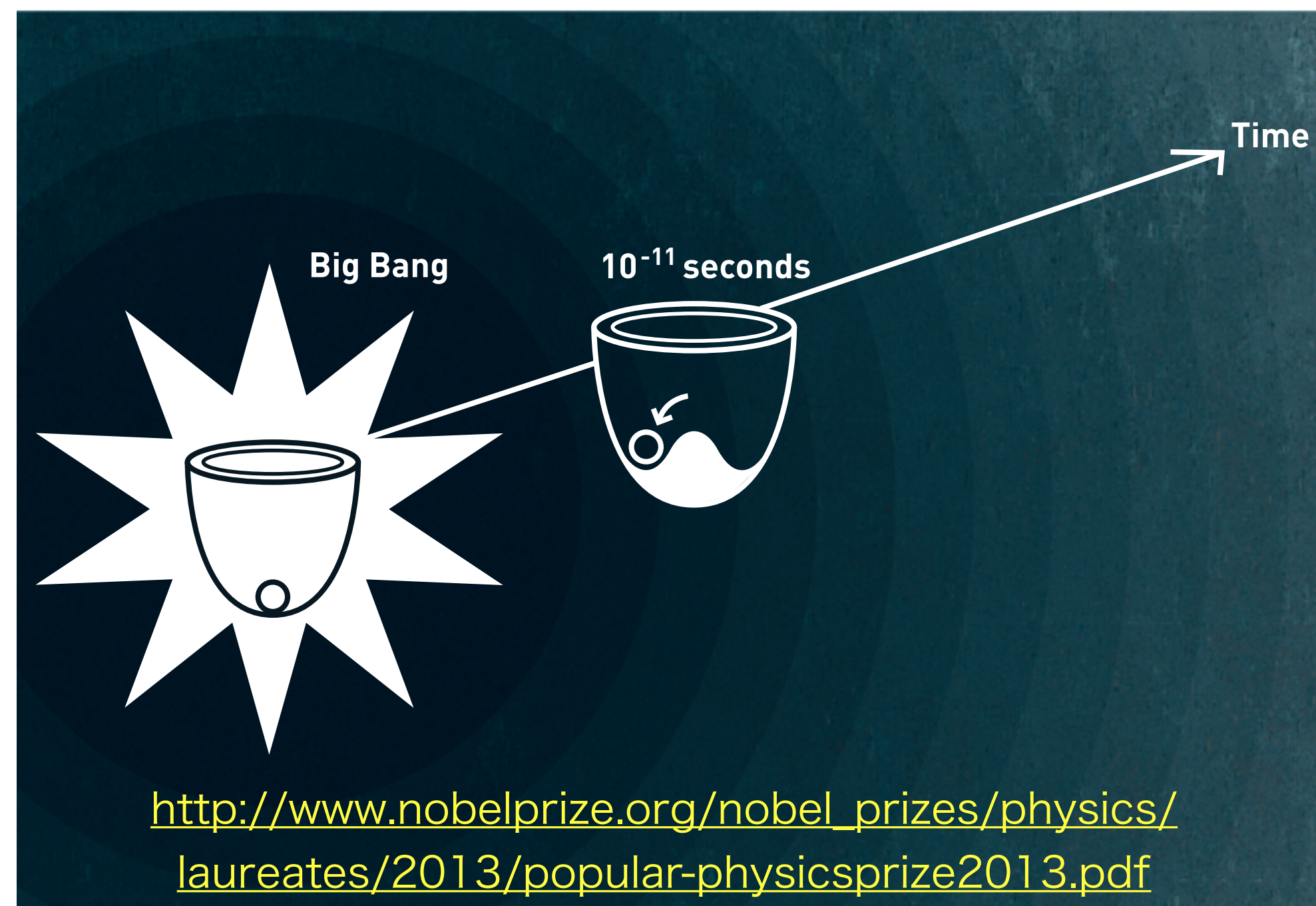


Physics by LHC at the TeV energy region

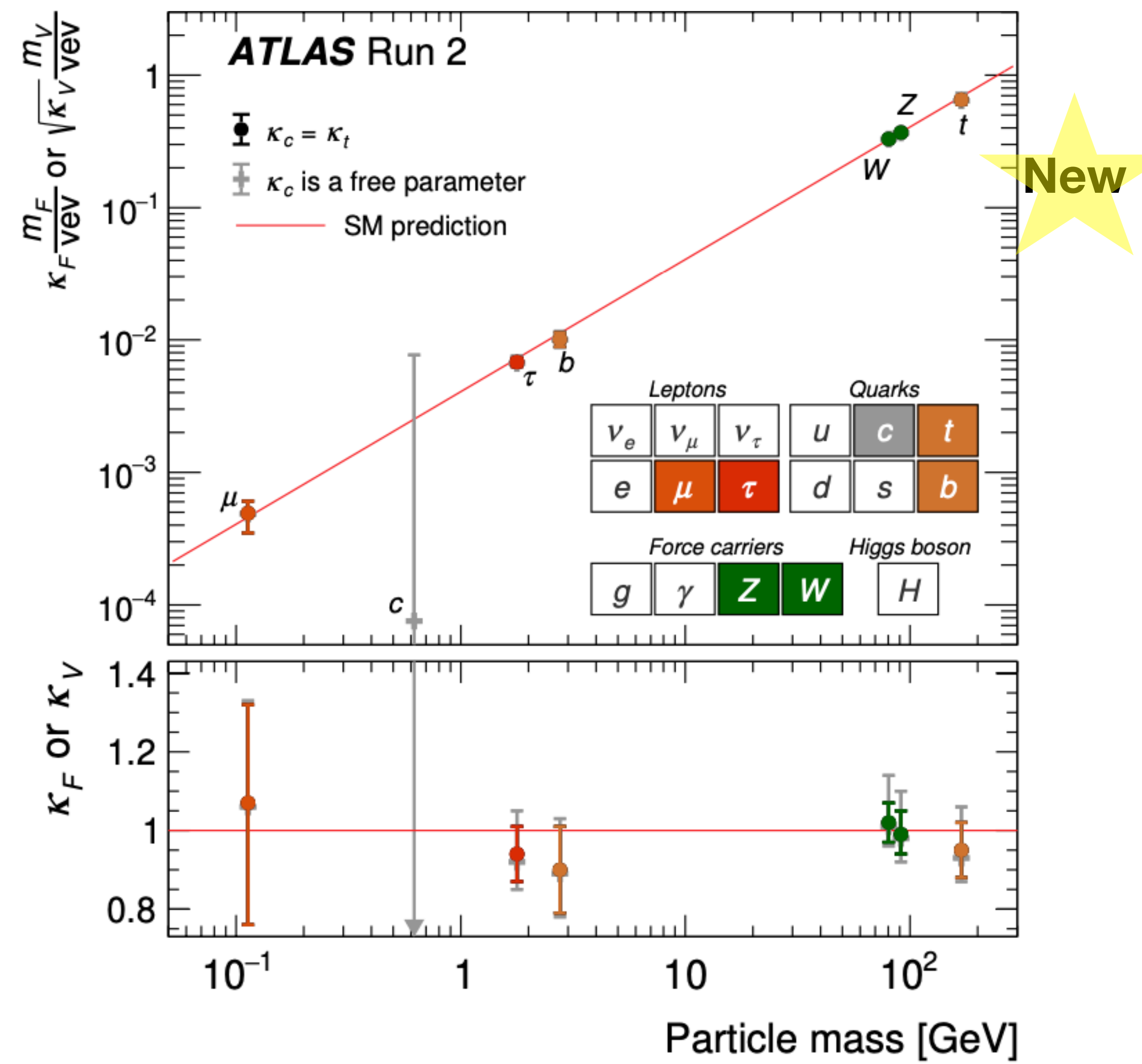
- Discovering the Higgs boson to understand the origin of mass due to spontaneous symmetry breaking.
- Search for supersymmetry particles
 - Linking space-time symmetry and internal symmetry of particles.
- Look for any new phenomena.

Higgs boson

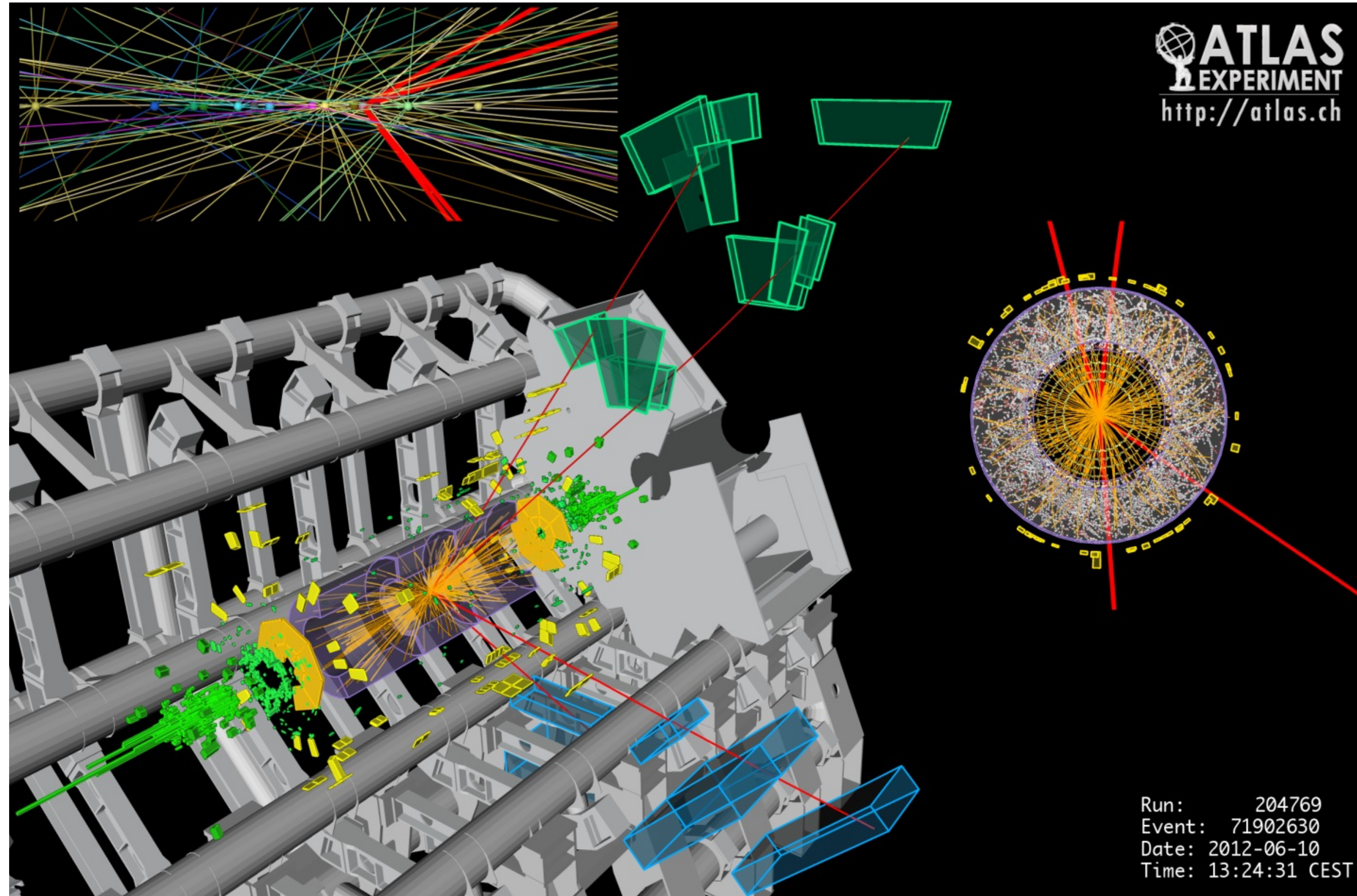
- Particles (Higgs field condensate) as the source of the phase transition in the vacuum
 - Phase transition: the energy of the vacuum changes.
 - Particle is given (inertial) mass.
 - Coupling with the Higgs boson is related to the mass of the coupling particle



Coupling with the Higgs boson with the mass of the coupling particle

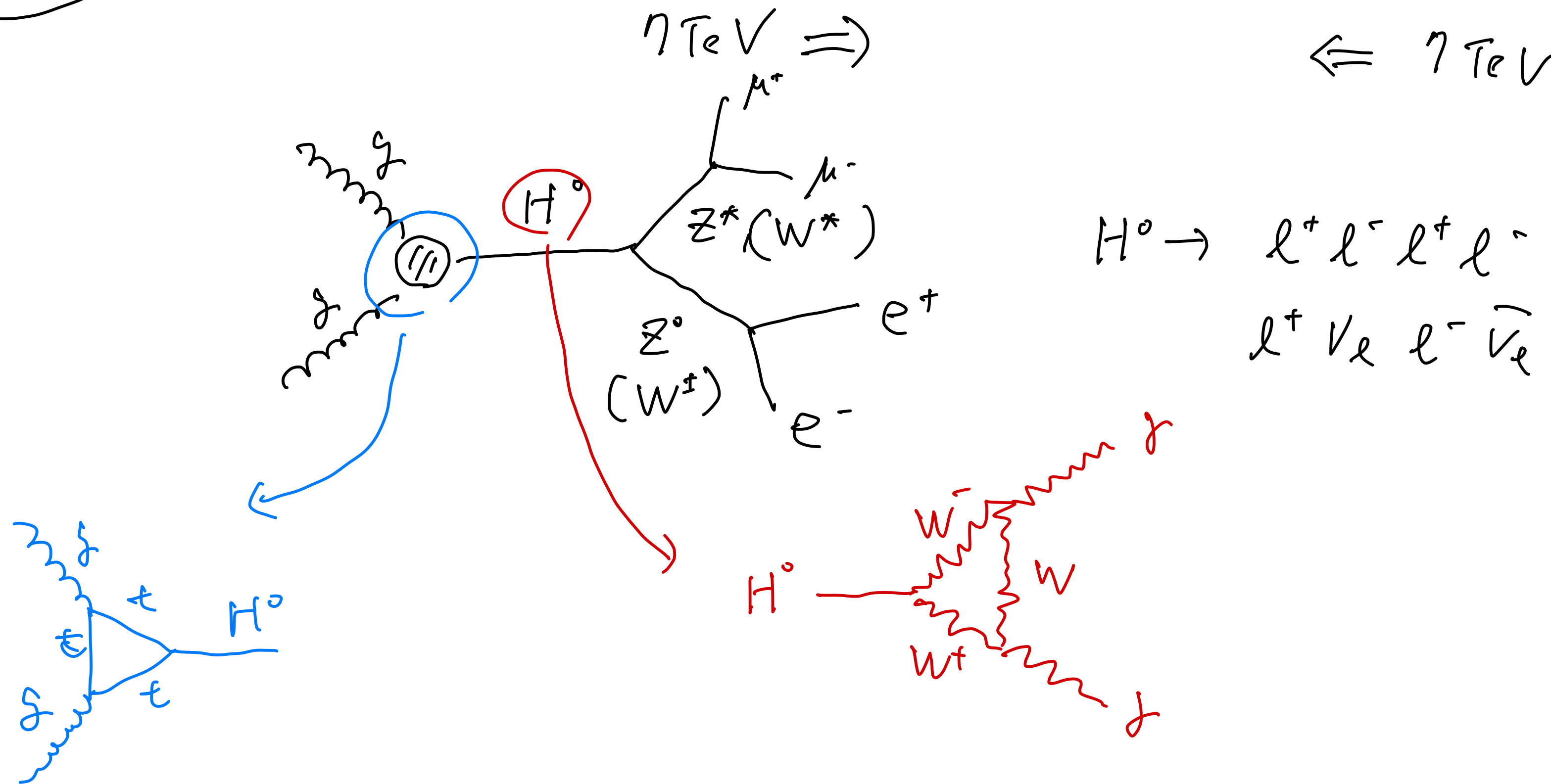
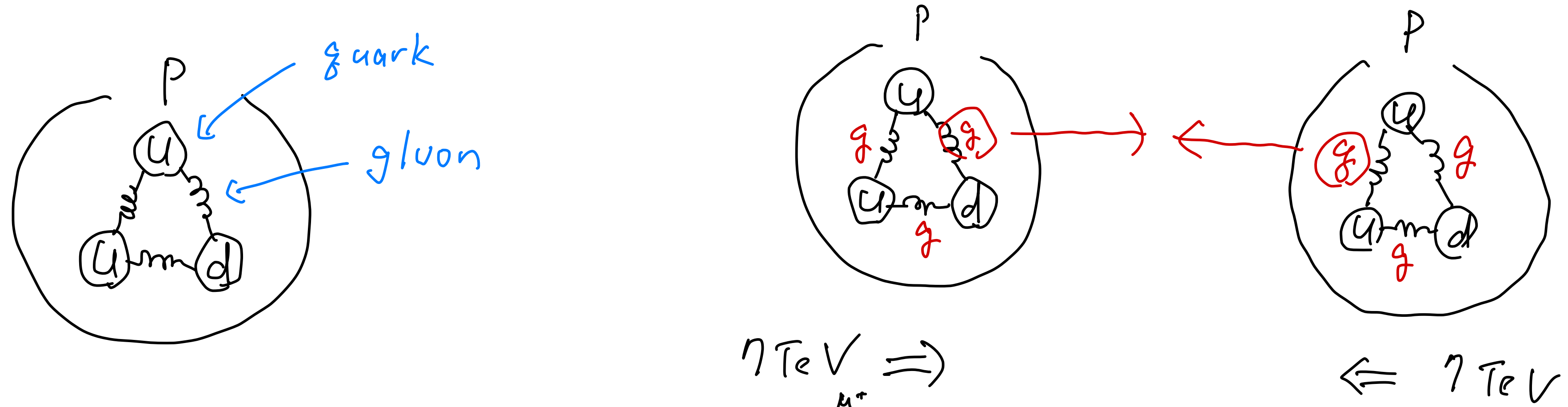


How to detect Higgs bosons



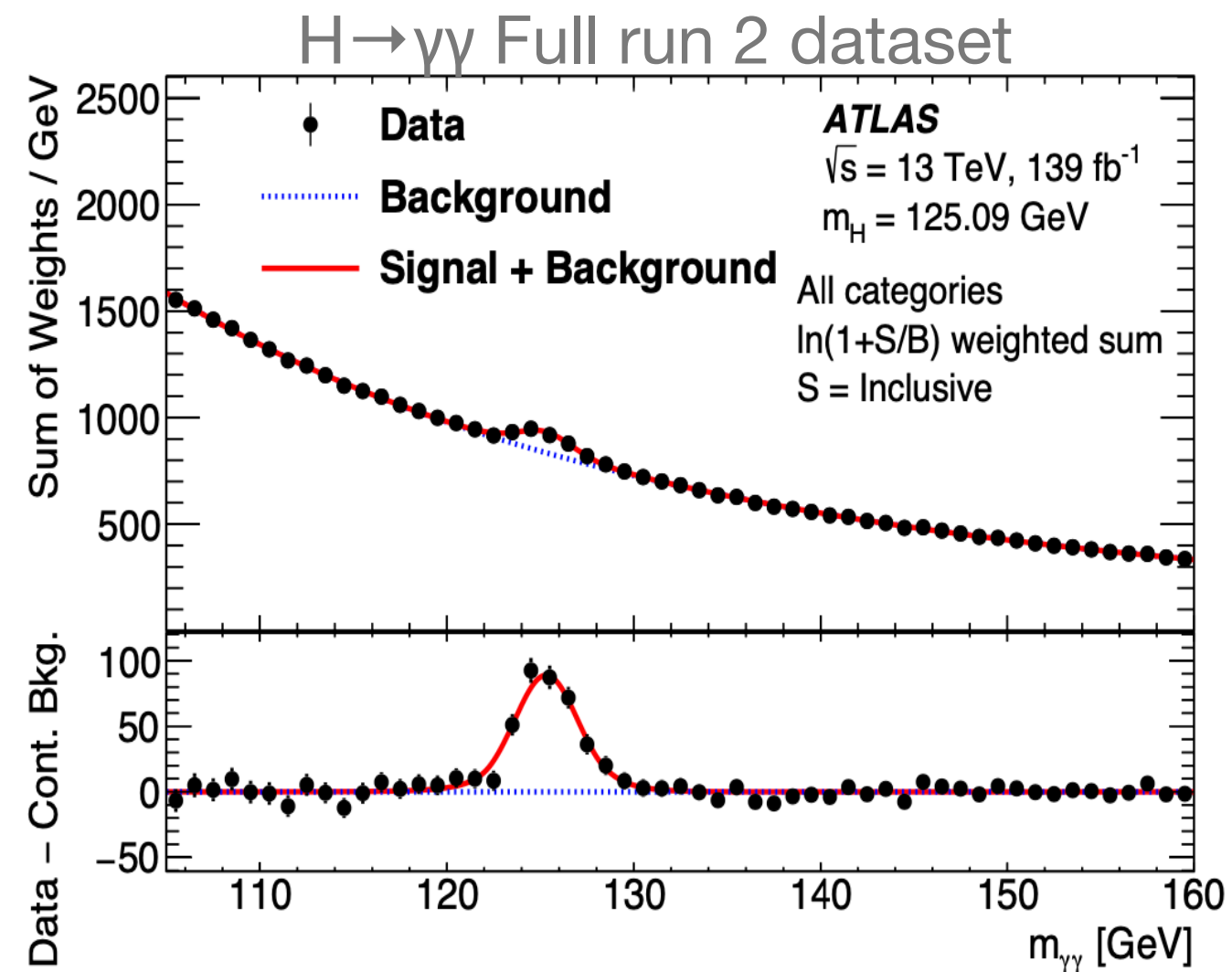
<https://atlas.cern/updates/atlas-feature/higgs-boson>

Proton - Proton collisions

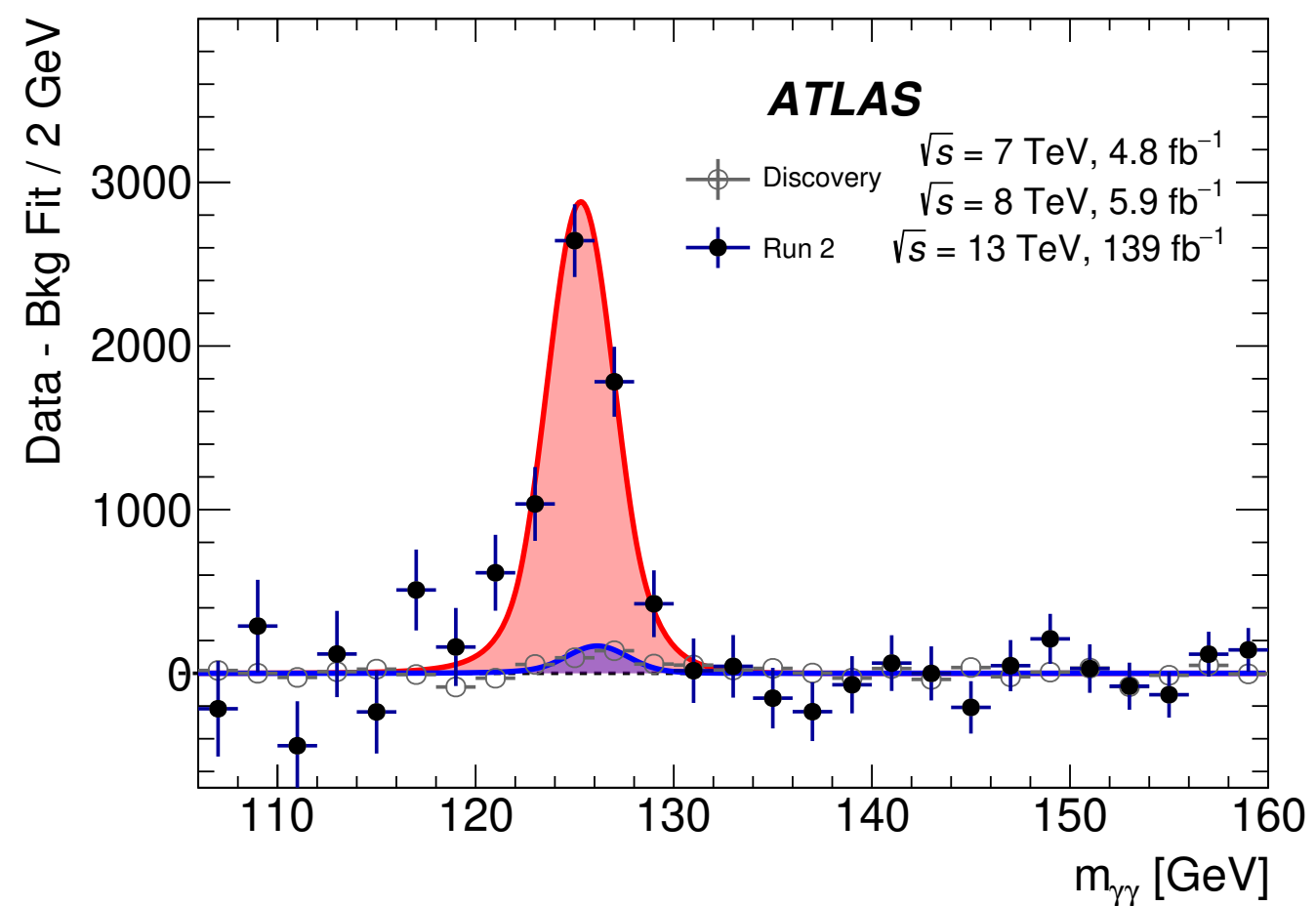


Detecting Higgs bosons

$$H \rightarrow \gamma \gamma$$



Comparison with discovery dataset in 2012



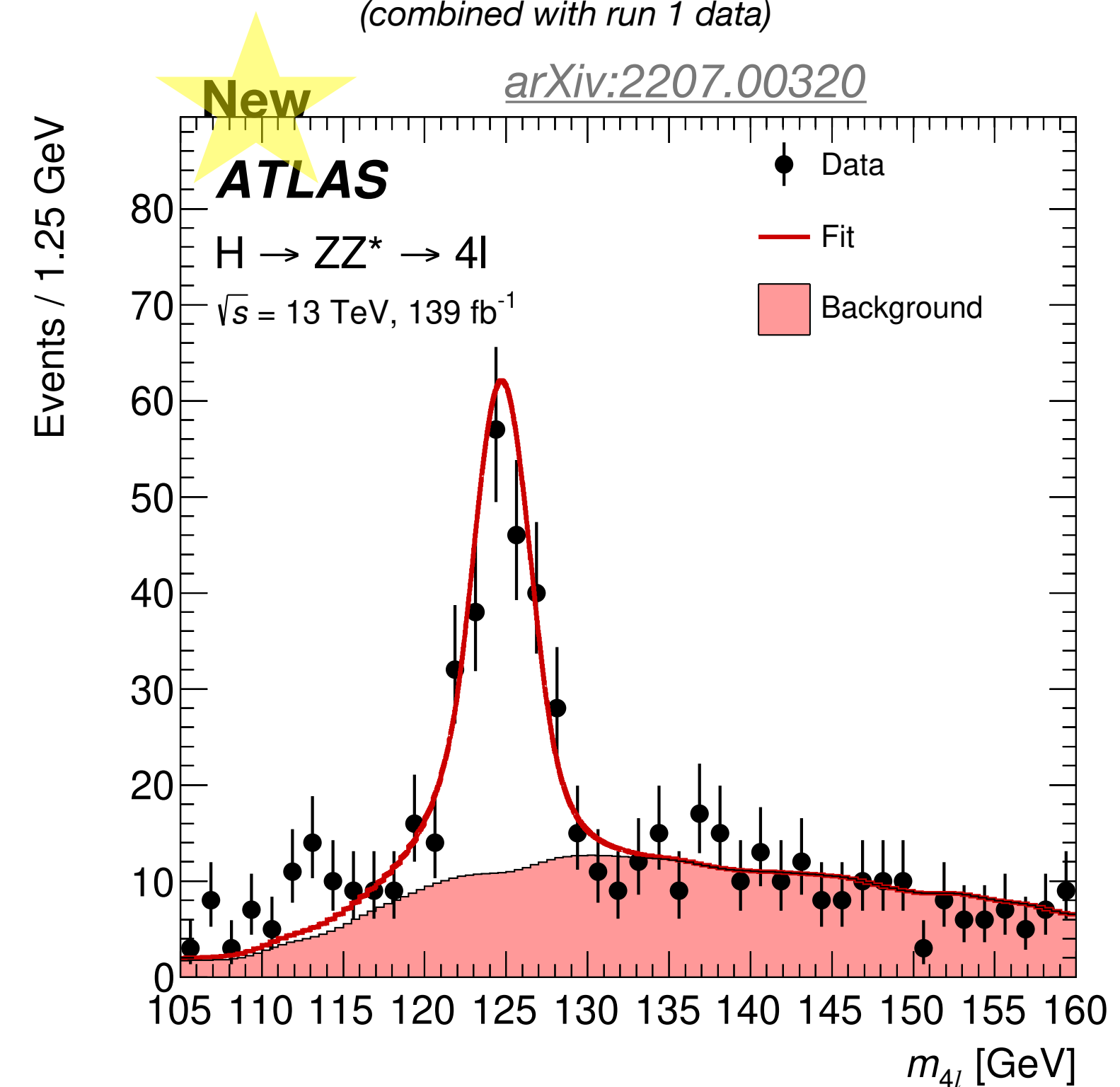
$$H \rightarrow ee\mu\mu, \mu\mu\mu\mu, eeee$$

Precise mass measurement using H $\rightarrow 4l$

Event-by-event resolution, DNN for S/B separation, precise muon and electron momentum calibration

$$m_H = 124.94 \pm 0.17(\text{stat.}) \pm 0.03(\text{syst.}) \text{ GeV}$$

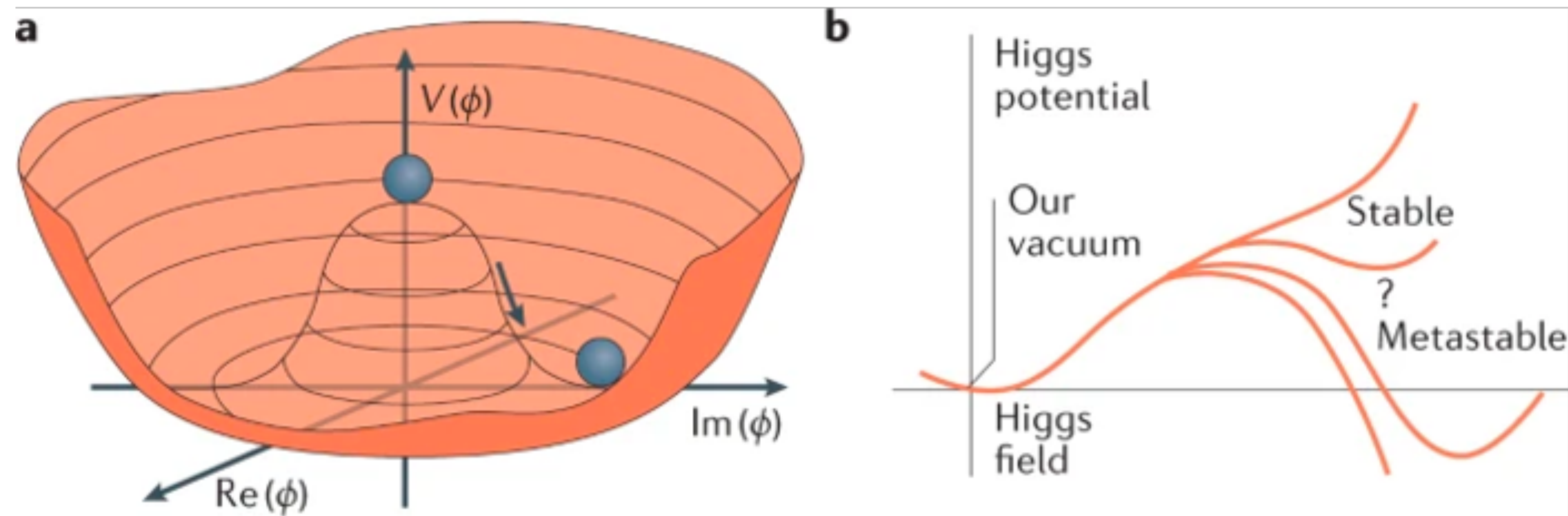
(combined with run 1 data)



What next with Higgs bosons

- Higgs potential
 - detecting a pair production of two Higgs bosons

$$V(\phi) = \frac{1}{2}m_H^2\phi^2 + \sqrt{\lambda/2}m_H\phi^3 + \frac{1}{4}\lambda\phi^4$$



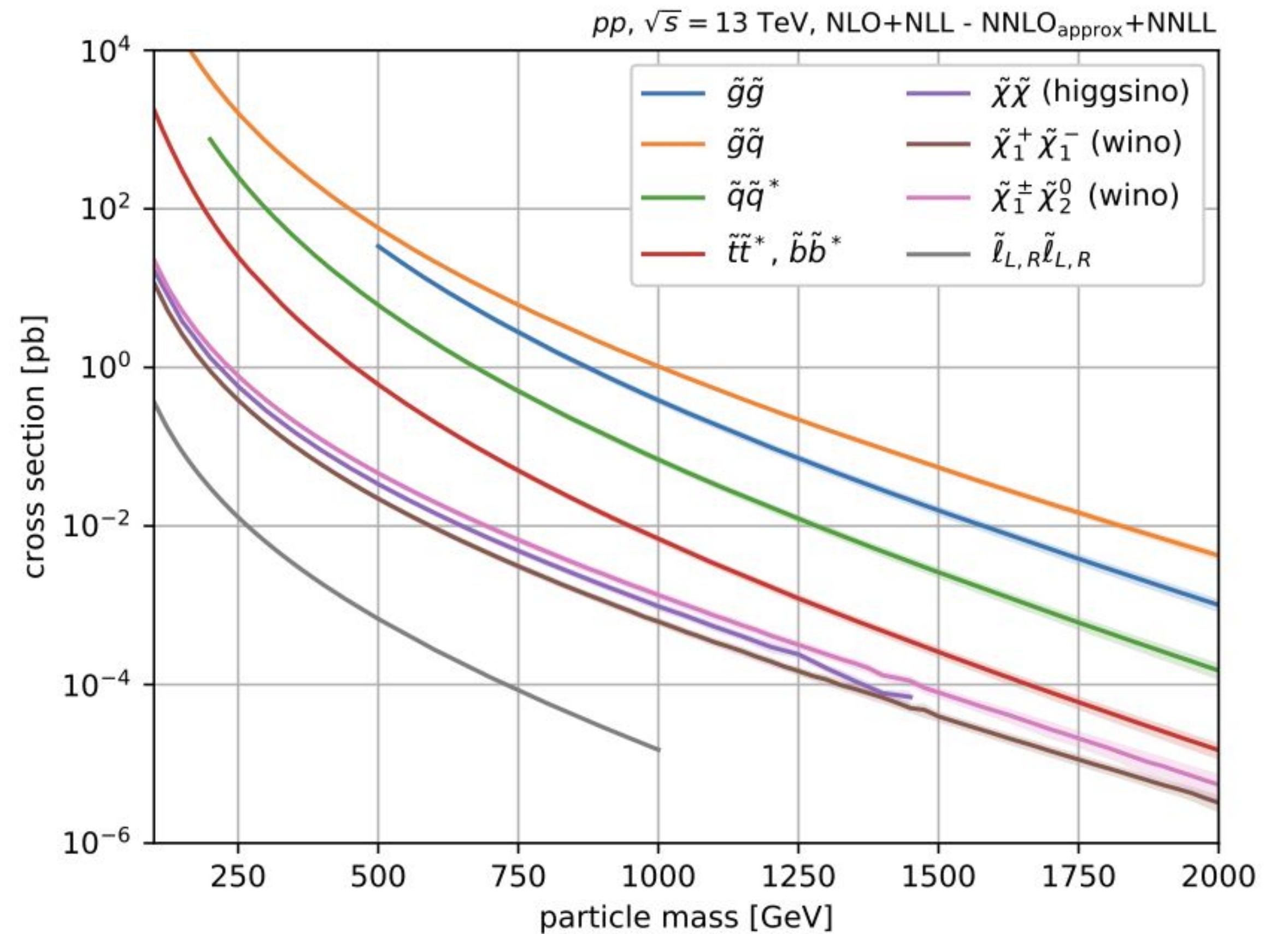
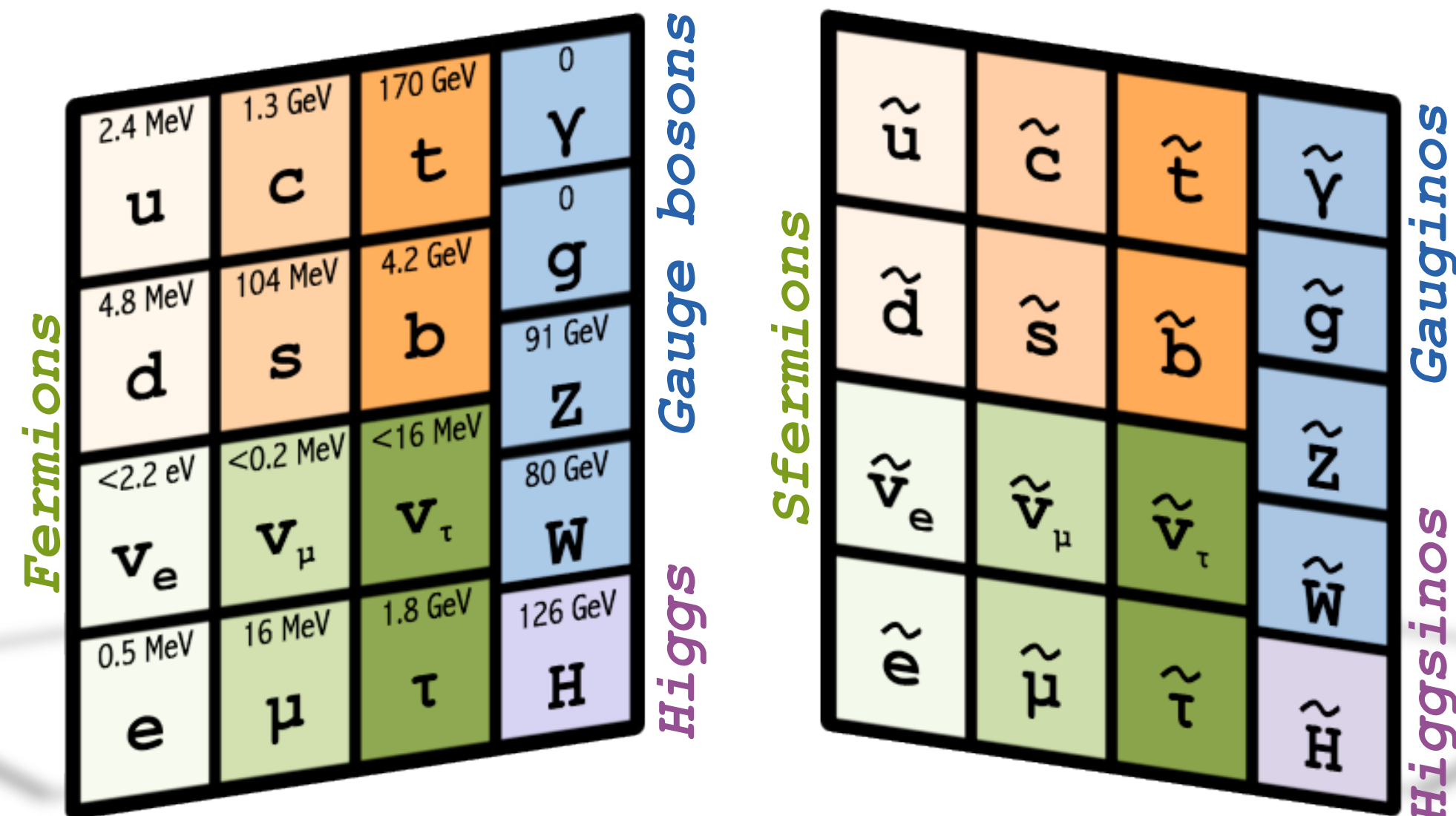
Beyond Higgs bosons

- **Supersymmetry (SUSY)**
 - Symmetry between Bose particles (spin integers) and Fermi particles (spin half-integers)
 - Example of Bose particle: photon
 - Fermi particles e.g. electrons
 - Are dark matter particles SUSY particles?
- **New Physics**
 - Extra dimensions (>4)
 - How many dimensions is space, are dimensions beyond 4 compactified?
 - Leptoquarks, Heavy majorana neutrinos, etc..

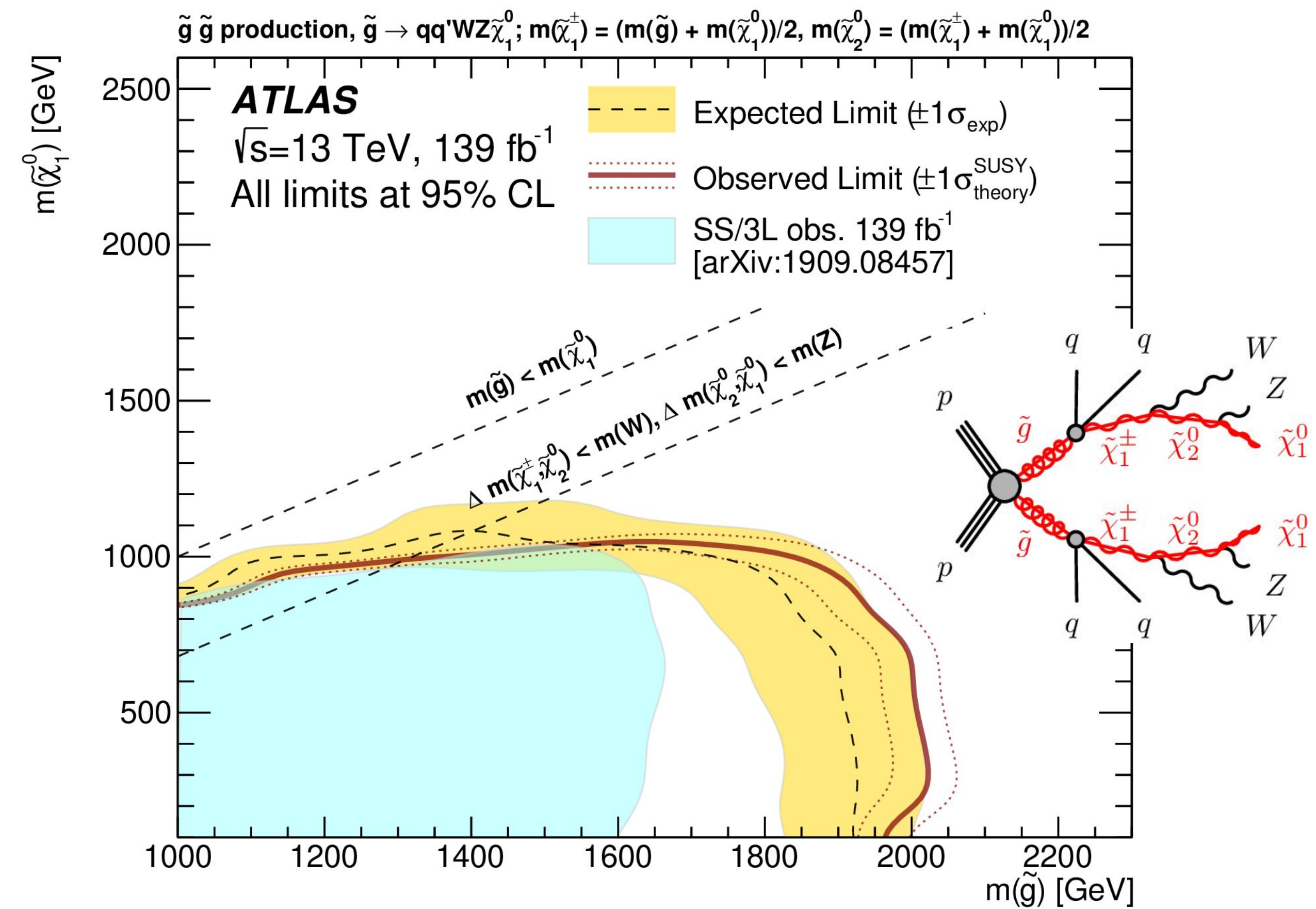
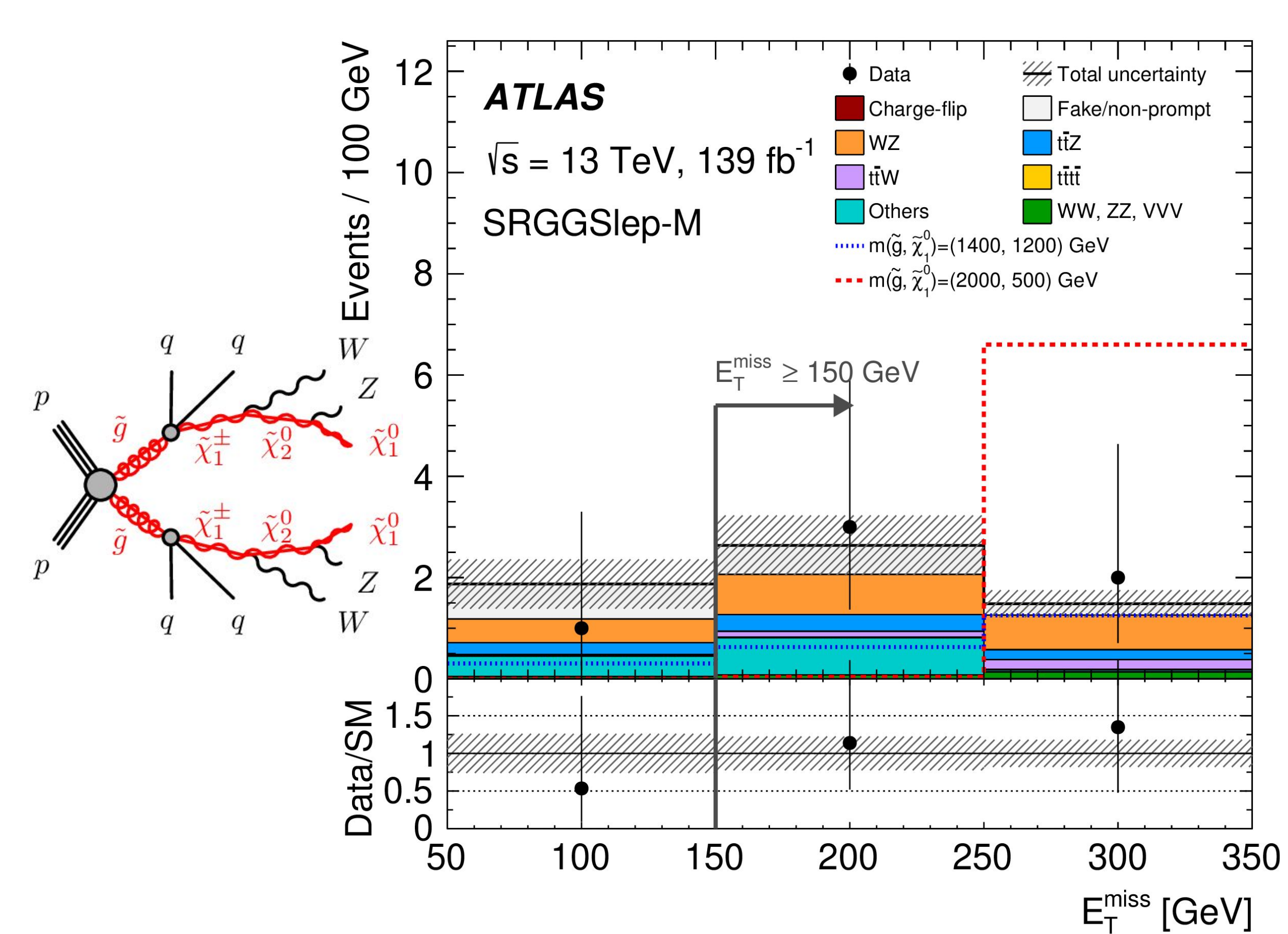
SUSY

- SUSY searches covering vast regions of SUSY phase space
- Typical main signals in final states are
 - Strong SUSY: 2 same-sign or 3 Leptons
 - Electroweakino: with taus
 - Higgsinos: with b-jets and photons

© Forces \leftrightarrow Matter



Missing energy with 2 same-sign or 3 leptons



ATLAS SUSY Searches* - 95% CL Lower Limits

March 2023

ATLAS Preliminary

$\sqrt{s} = 13$ TeV

Model	Signature	$\int \mathcal{L} dt$ [fb $^{-1}$]	Mass limit	Reference					
Inclusive Searches	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0 e, μ mono-jet	2-6 jets 1-3 jets	E_T^{miss} E_T^{miss}	139	\tilde{q} [1x, 8x Degen.] 1.0 1.85 \tilde{q} [8x Degen.] 0.9	$m(\tilde{\chi}_1^0) < 400$ GeV $m(\tilde{q}) - m(\tilde{\chi}_1^0) = 5$ GeV	2101.14293 2102.10874	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0 e, μ	2-6 jets	E_T^{miss}	139	\tilde{g} 2.3 \tilde{g} Forbidden 1.15-1.95	$m(\tilde{\chi}_1^0) = 0$ GeV $m(\tilde{\chi}_1^0) = 1000$ GeV	2010.14293 2010.14293	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}W\tilde{\chi}_1^0$	1 e, μ	2-6 jets	E_T^{miss}	139	\tilde{g} 2.2	$m(\tilde{\chi}_1^0) < 600$ GeV	2101.01629	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell)\tilde{\chi}_1^0$	$ee, \mu\mu$	2 jets	E_T^{miss}	139	\tilde{g} 2.2	$m(\tilde{\chi}_1^0) < 700$ GeV	2204.13072	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$	0 e, μ SS e, μ	7-11 jets 6 jets	E_T^{miss} E_T^{miss}	139 139	\tilde{g} 1.15 1.97	$m(\tilde{\chi}_1^0) < 600$ GeV $m(\tilde{g}) - m(\tilde{\chi}_1^0) = 200$ GeV	2008.06032 1909.08457	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 e, μ SS e, μ	3 b 6 jets	E_T^{miss} E_T^{miss}	139 139	\tilde{g} 1.25 2.45	$m(\tilde{\chi}_1^0) < 500$ GeV $m(\tilde{g}) - m(\tilde{\chi}_1^0) = 300$ GeV	2211.08028 1909.08457	
	3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1$	0 e, μ	2 b	E_T^{miss}	139	\tilde{b}_1 1.255 \tilde{b}_1 0.68	$m(\tilde{\chi}_1^0) < 400$ GeV 10 GeV $< \Delta m(\tilde{b}_1, \tilde{\chi}_1^0) < 20$ GeV	2101.12527 2101.12527
$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_2^0 \rightarrow bh\tilde{\chi}_1^0$		0 e, μ 2 τ	6 b 2 b	E_T^{miss} E_T^{miss}	139 139	\tilde{b}_1 Forbidden 0.23-1.35 \tilde{b}_1 0.13-0.85	$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 100$ GeV $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 0$ GeV	1908.03122 2103.08189	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$		0-1 e, μ	≥ 1 jet	E_T^{miss}	139	\tilde{t}_1 1.25	$m(\tilde{\chi}_1^0) = 1$ GeV	2004.14060, 2012.03799	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$		1 e, μ	3 jets/1 b	E_T^{miss}	139	\tilde{t}_1 Forbidden 0.65	$m(\tilde{\chi}_1^0) = 500$ GeV	2012.03799	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{\tau}_1 b\nu, \tilde{\tau}_1 \rightarrow \tau\tilde{G}$		1-2 τ	2 jets/1 b	E_T^{miss}	139	\tilde{t}_1 Forbidden 1.4	$m(\tilde{\tau}_1) = 800$ GeV	2108.07665	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0 / \tilde{c}\tilde{c}, \tilde{c} \rightarrow c\tilde{\chi}_1^0$		0 e, μ 0 e, μ	2 c mono-jet	E_T^{miss} E_T^{miss}	36.1 139	\tilde{c} 0.85 \tilde{t}_1 0.55	$m(\tilde{\chi}_1^0) = 0$ GeV $m(\tilde{t}_1, \tilde{c}) - m(\tilde{\chi}_1^0) = 5$ GeV	1805.01649 2102.10874	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow Z/h\tilde{\chi}_1^0$		1-2 e, μ	1-4 b	E_T^{miss}	139	\tilde{t}_1 0.067-1.18	$m(\tilde{\chi}_2^0) = 500$ GeV	2006.05880	
$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$		3 e, μ	1 b	E_T^{miss}	139	\tilde{t}_2 Forbidden 0.86	$m(\tilde{\chi}_1^0) = 360$ GeV, $m(\tilde{t}_1) - m(\tilde{\chi}_1^0) = 40$ GeV	2006.05880	
EW direct		$\tilde{\chi}_1^+\tilde{\chi}_2^0$ via WZ	Multiple ℓ /jets $ee, \mu\mu$	≥ 1 jet	E_T^{miss} E_T^{miss}	139 139	$\tilde{\chi}_1^+/\tilde{\chi}_2^0$ 0.96 $\tilde{\chi}_1^+/\tilde{\chi}_2^0$ 0.205	$m(\tilde{\chi}_1^0) = 0$, wino-bino $m(\tilde{\chi}_1^+) - m(\tilde{\chi}_1^0) = 5$ GeV, wino-bino	2106.01676, 2108.07586 1911.12606
		$\tilde{\chi}_1^+\tilde{\chi}_1^+$ via WW	2 e, μ		E_T^{miss}	139	$\tilde{\chi}_1^+$ 0.42	$m(\tilde{\chi}_1^0) = 0$, wino-bino	1908.08215
	$\tilde{\chi}_1^+\tilde{\chi}_2^0$ via Wh	Multiple ℓ /jets		E_T^{miss}	139	$\tilde{\chi}_1^+/\tilde{\chi}_2^0$ Forbidden 1.06	$m(\tilde{\chi}_1^0) = 70$ GeV, wino-bino	2004.10894, 2108.07586	
	$\tilde{\chi}_1^+\tilde{\chi}_1^+$ via $\tilde{\ell}_L/\tilde{\nu}$	2 e, μ		E_T^{miss}	139	$\tilde{\chi}_1^+$ 1.0	$m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^+) + m(\tilde{\chi}_1^0))$	1908.08215	
	$\tilde{\tau}\tilde{\tau}, \tilde{\tau} \rightarrow \tau\tilde{\chi}_1^0$	2 τ		E_T^{miss}	139	$\tilde{\tau}$ [$\tilde{\tau}_L, \tilde{\tau}_{R,L}$] 0.16-0.3 0.12-0.39	$m(\tilde{\chi}_1^0) = 0$	1911.06660	
	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 e, μ $ee, \mu\mu$	0 jets ≥ 1 jet	E_T^{miss} E_T^{miss}	139 139	$\tilde{\ell}$ 0.7 $\tilde{\ell}$ 0.256	$m(\tilde{\chi}_1^0) = 0$ $m(\tilde{\ell}) - m(\tilde{\chi}_1^0) = 10$ GeV	1908.08215 1911.12606	
	$\tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G}$	0 e, μ 4 e, μ 0 e, μ 2 e, μ	≥ 3 b 0 jets ≥ 2 large jets ≥ 2 jets	E_T^{miss} E_T^{miss} E_T^{miss} E_T^{miss}	36.1 139 139 139	\tilde{H} 0.13-0.23 0.29-0.88 \tilde{H} 0.55 0.45-0.93 \tilde{H} 0.77	$\text{BR}(\tilde{\chi}_1^0 \rightarrow h\tilde{G}) = 1$ $\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = 1$ $\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = 1$ $\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = \text{BR}(\tilde{\chi}_1^0 \rightarrow h\tilde{G}) = 0.5$	1806.04030 2103.11684 2108.07586 2204.13072	
	Long-lived particles	Direct $\tilde{\chi}_1^+\tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^+$	Disapp. trk	1 jet	E_T^{miss}	139	$\tilde{\chi}_1^+$ 0.66 $\tilde{\chi}_1^+$ 0.21	Pure Wino Pure higgsino	2201.02472 2201.02472
Stable \tilde{g} R-hadron		pixel dE/dx		E_T^{miss}	139	\tilde{g} 2.05		2205.06013	
Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$		pixel dE/dx		E_T^{miss}	139	\tilde{g} [$\tau(\tilde{g}) = 10$ ns] 2.2	$m(\tilde{\chi}_1^0) = 100$ GeV	2205.06013	
$\tilde{\ell}\tilde{\ell}, \tilde{\ell} \rightarrow \ell\tilde{G}$		Displ. lep		E_T^{miss}	139	$\tilde{\ell}, \tilde{\mu}$ 0.7 $\tilde{\tau}$ 0.34 $\tilde{\tau}$ 0.36	$\tau(\tilde{\ell}) = 0.1$ ns $\tau(\tilde{\ell}) = 0.1$ ns $\tau(\tilde{\ell}) = 10$ ns	2011.07812 2011.07812 2205.06013	
RPV	$\tilde{\chi}_1^+\tilde{\chi}_1^+/\tilde{\chi}_1^0, \tilde{\chi}_1^+ \rightarrow Z\ell \rightarrow \ell\ell\ell$	3 e, μ		E_T^{miss}	139	$\tilde{\chi}_1^+/\tilde{\chi}_1^0$ [BR(Z τ)=1, BR(Z e)=1] 0.625 1.05	Pure Wino	2011.10543	
	$\tilde{\chi}_1^+\tilde{\chi}_1^+/\tilde{\chi}_2^0 \rightarrow WW/Z\ell\ell\ell\nu\nu$	4 e, μ	0 jets	E_T^{miss}	139	$\tilde{\chi}_1^+/\tilde{\chi}_2^0$ [$\lambda_{333} \neq 0, \lambda_{12k} \neq 0$] 0.95 1.55	$m(\tilde{\chi}_1^0) = 200$ GeV	2103.11684	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq\tilde{q}$	4-5 large jets		E_T^{miss}	36.1	\tilde{g} [$m(\tilde{\chi}_1^0) = 200$ GeV, 1100 GeV] 1.3 1.9	Large λ'_{112}	1804.03568	
	$\tilde{u}, \tilde{t} \rightarrow t\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow tbs$	Multiple		E_T^{miss}	36.1	\tilde{t} [$\lambda'_{323} = 2e-4, 1e-2$] 0.55 1.05	$m(\tilde{\chi}_1^0) = 200$ GeV, bino-like	ATLAS-CONF-2018-003	
	$\tilde{u}, \tilde{t} \rightarrow b\tilde{\chi}_1^+, \tilde{\chi}_1^+ \rightarrow bbs$	$\geq 4b$		E_T^{miss}	139	\tilde{t} Forbidden 0.95	$m(\tilde{\chi}_1^+) = 500$ GeV	2010.01015	
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	2 jets + 2 b		E_T^{miss}	36.7	\tilde{t}_1 [qq, bs] 0.42 0.61		1710.07171	
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow q\ell$	2 e, μ 1 μ	2 b DV	E_T^{miss} E_T^{miss}	36.1 136	\tilde{t}_1 0.4-1.45 \tilde{t}_1 [$1e-10 < \lambda'_{23k} < 1e-8, 3e-10 < \lambda'_{23k} < 3e-9$] 1.0 1.6	$\text{BR}(\tilde{t}_1 \rightarrow b\ell/b\mu) > 20\%$ $\text{BR}(\tilde{t}_1 \rightarrow q\mu) = 100\%, \cos\theta_i = 1$	1710.05544 2003.11956	
$\tilde{\chi}_1^+/\tilde{\chi}_2^0/\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow tbs, \tilde{\chi}_1^+ \rightarrow bbs$	1-2 e, μ	≥ 6 jets		E_T^{miss}	139	$\tilde{\chi}_1^0$ 0.2-0.32	Pure higgsino	2106.09609	

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10⁻¹

1

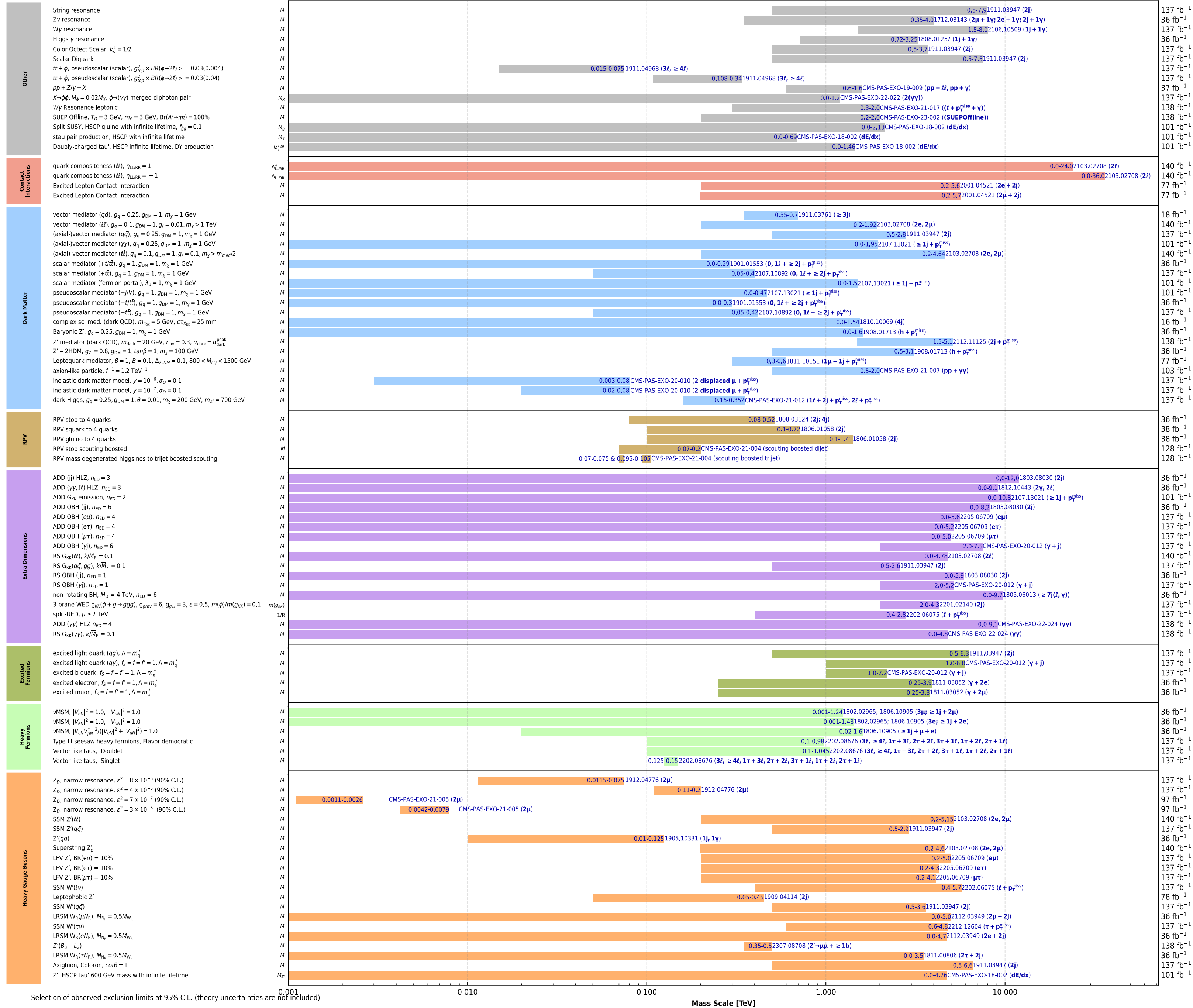
Mass scale [TeV]



Overview of CMS EXO results

CMS preliminary

March 2024



Selection of observed exclusion limits at 95% C.L. (theory uncertainties are not included).

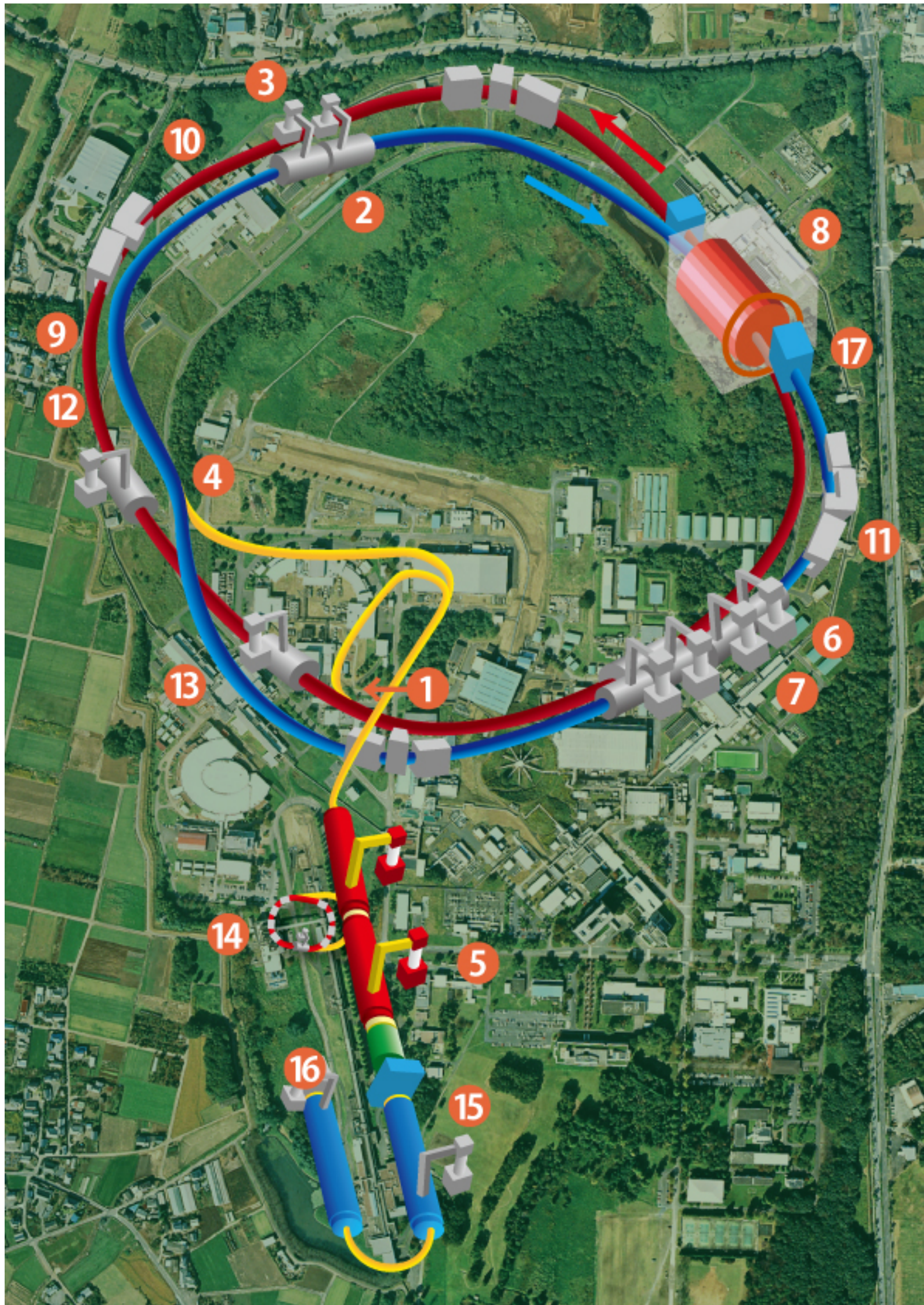
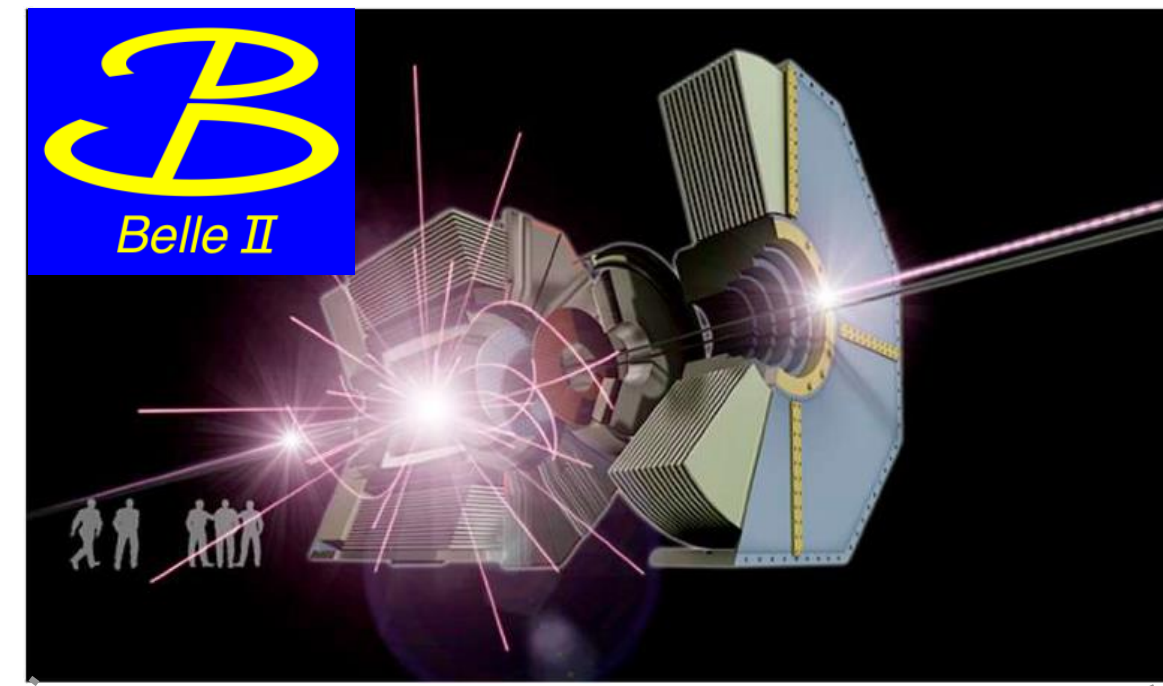
Mass Scale [TeV]

2. Intensity Frontier

What experiments are there?

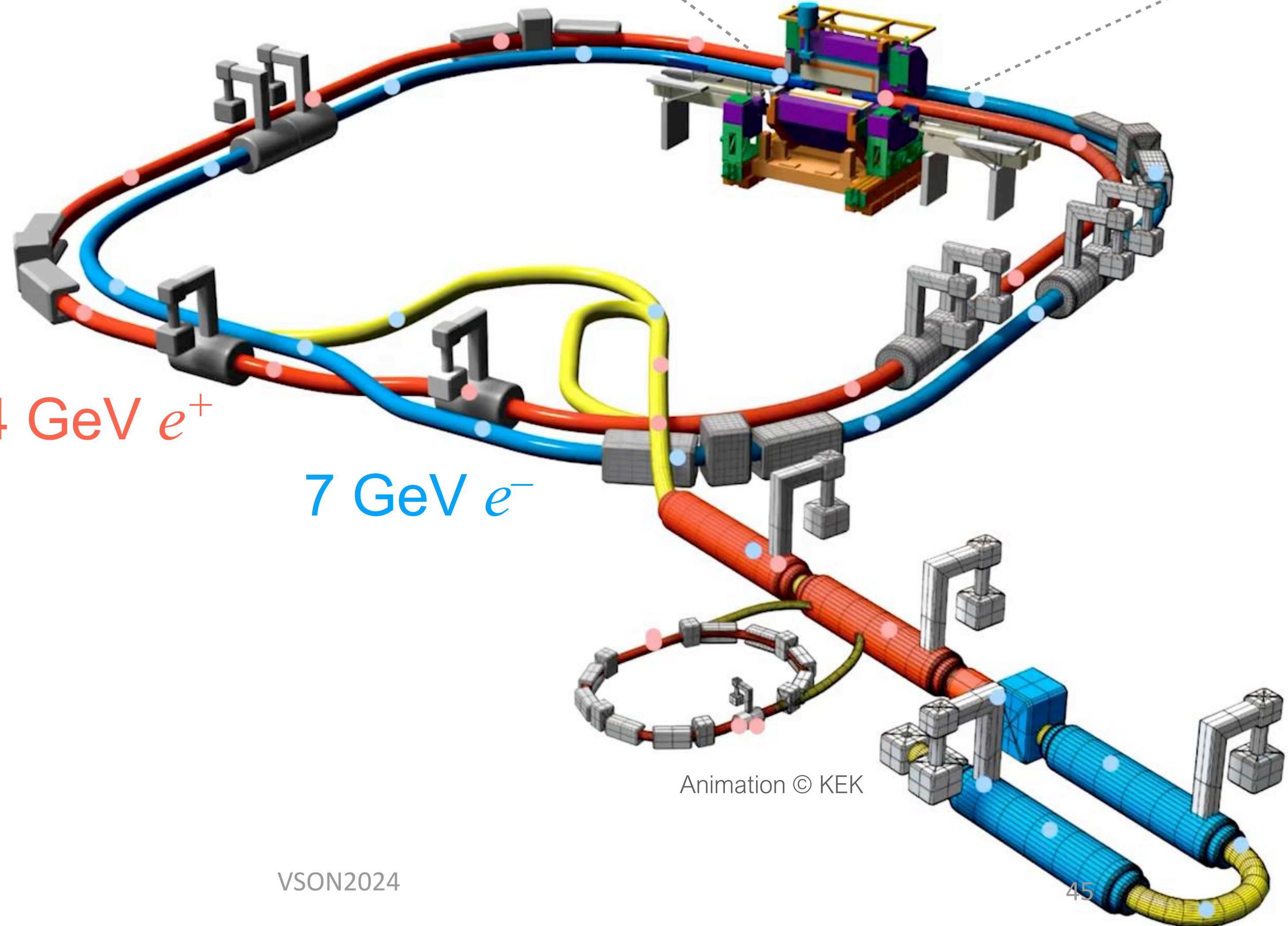
- CERN (Europe)
 - **LHC/ATLAS, CMS**, LHCb, FASAR, milliQan, NA62, DsTau, (CEPE)
- PSI (Switzerland)
 - MEG II
- Fermilab (US)
 - Muon g-2, Mu2e, Neutrino experiments: DUNE, NOvA, SBN, ICARUS, etc.
- KEK/J-PARC (Japan)
 - **Super-KEKB/Belle II**, J-PARC/T2K, KoTO, COMET, muon g-2, JSNS², Super-K, Hyper-K, KAGLA, KamLAND, (ILC)
- IHEP (China)
 - BESSIII, JUNO, (CEPC)
- *Cosmology and Astro-particle experiments*

B-factory



4 GeV e^+

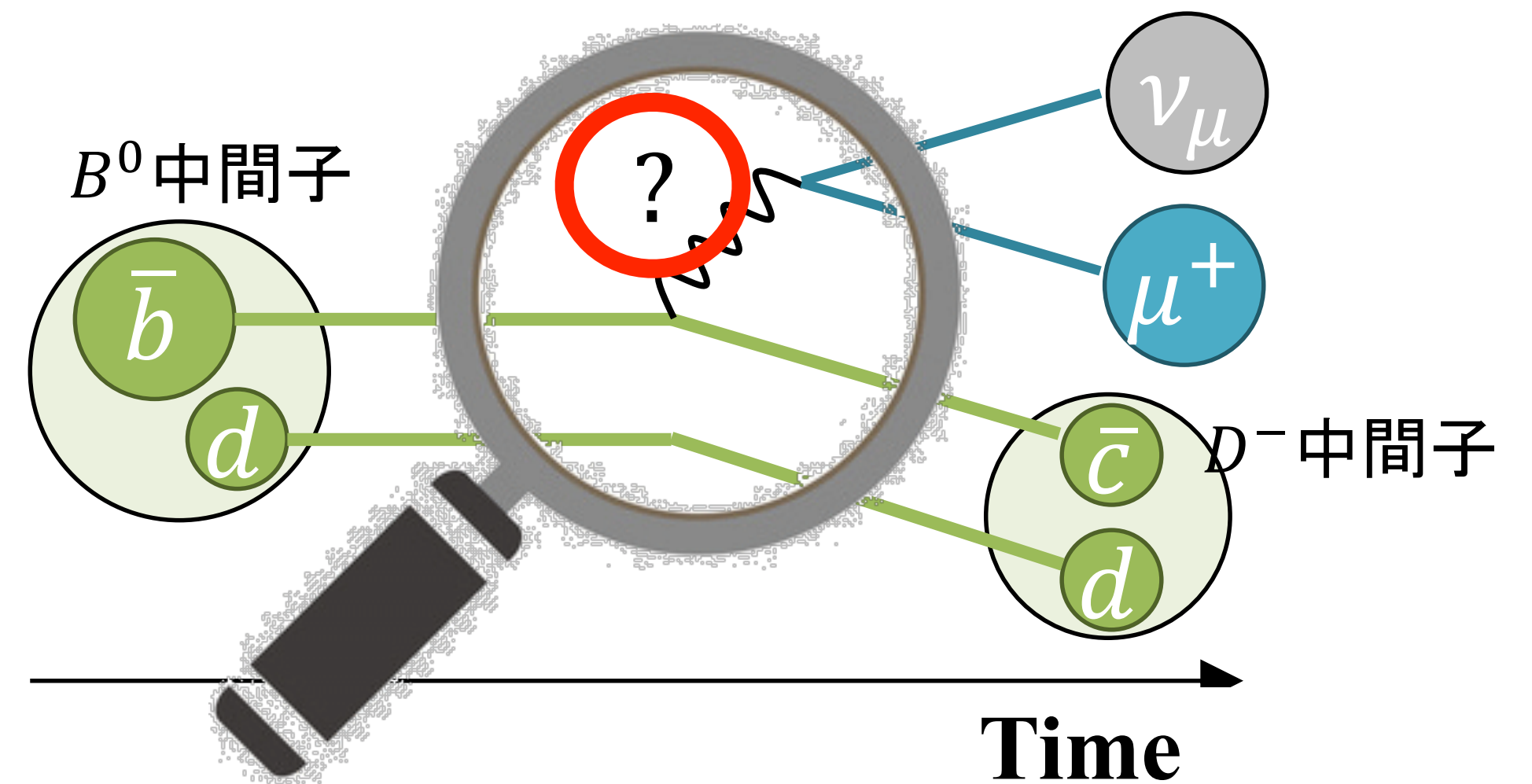
7 GeV e^-



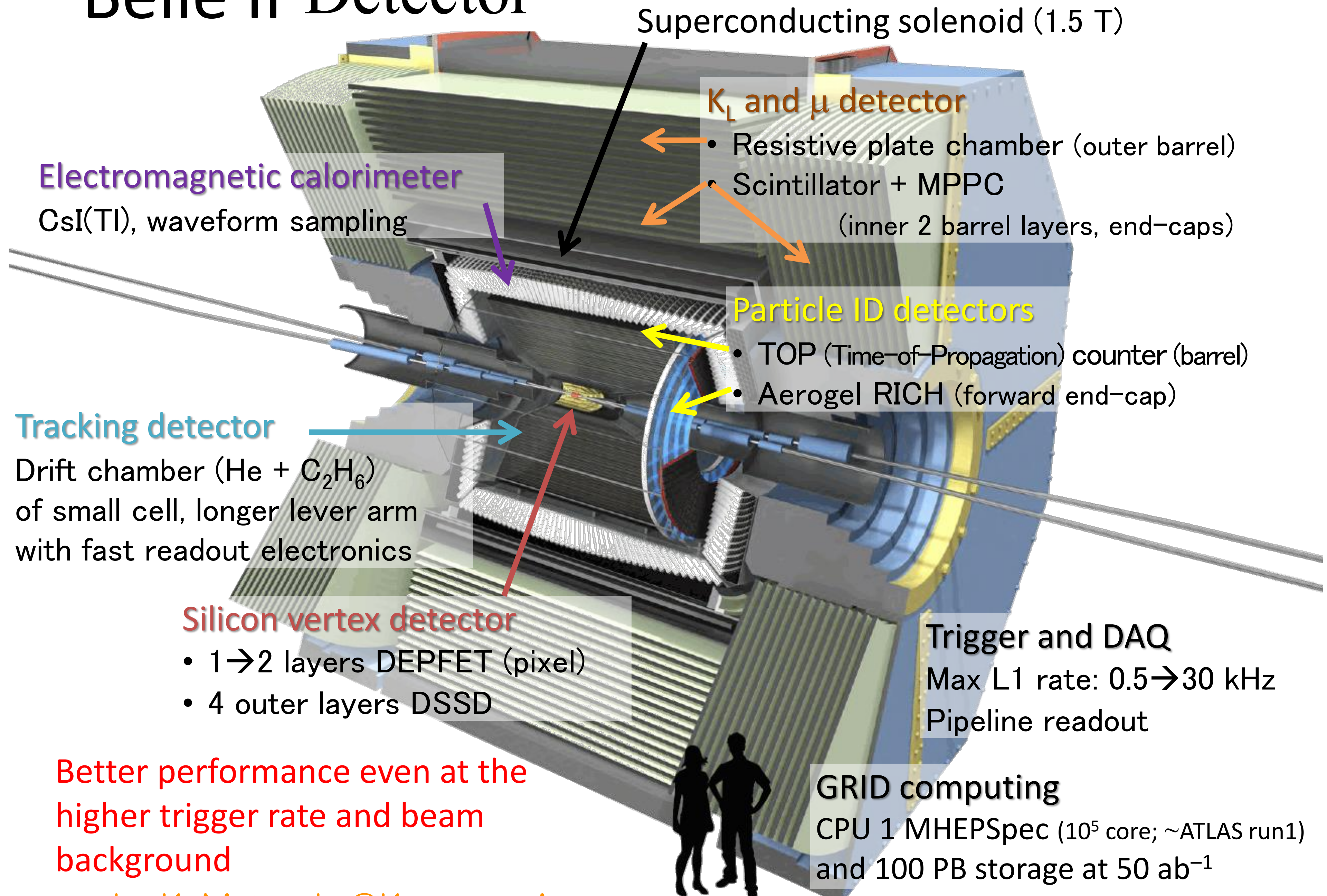
Animation © KEK

Basic ideas

- Quantum effects generate high-energy phenomena with a certain probability as intermediate states
- The intermediate state is indirectly probed by precise measurements because the probability including high-energy phenomena is rare.



Belle II Detector



Superconducting solenoid (1.5 T)

K_L and μ detector

- Resistive plate chamber (outer barrel)
- Scintillator + MPPC (inner 2 barrel layers, end-caps)

Electromagnetic calorimeter

CsI(Tl), waveform sampling

Particle ID detectors

- TOP (Time-of-Propagation) counter (barrel)
- Aerogel RICH (forward end-cap)

Tracking detector

Drift chamber (He + C₂H₆)
of small cell, longer lever arm
with fast readout electronics

Silicon vertex detector

- 1→2 layers DEPFET (pixel)
- 4 outer layers DSSD

Trigger and DAQ

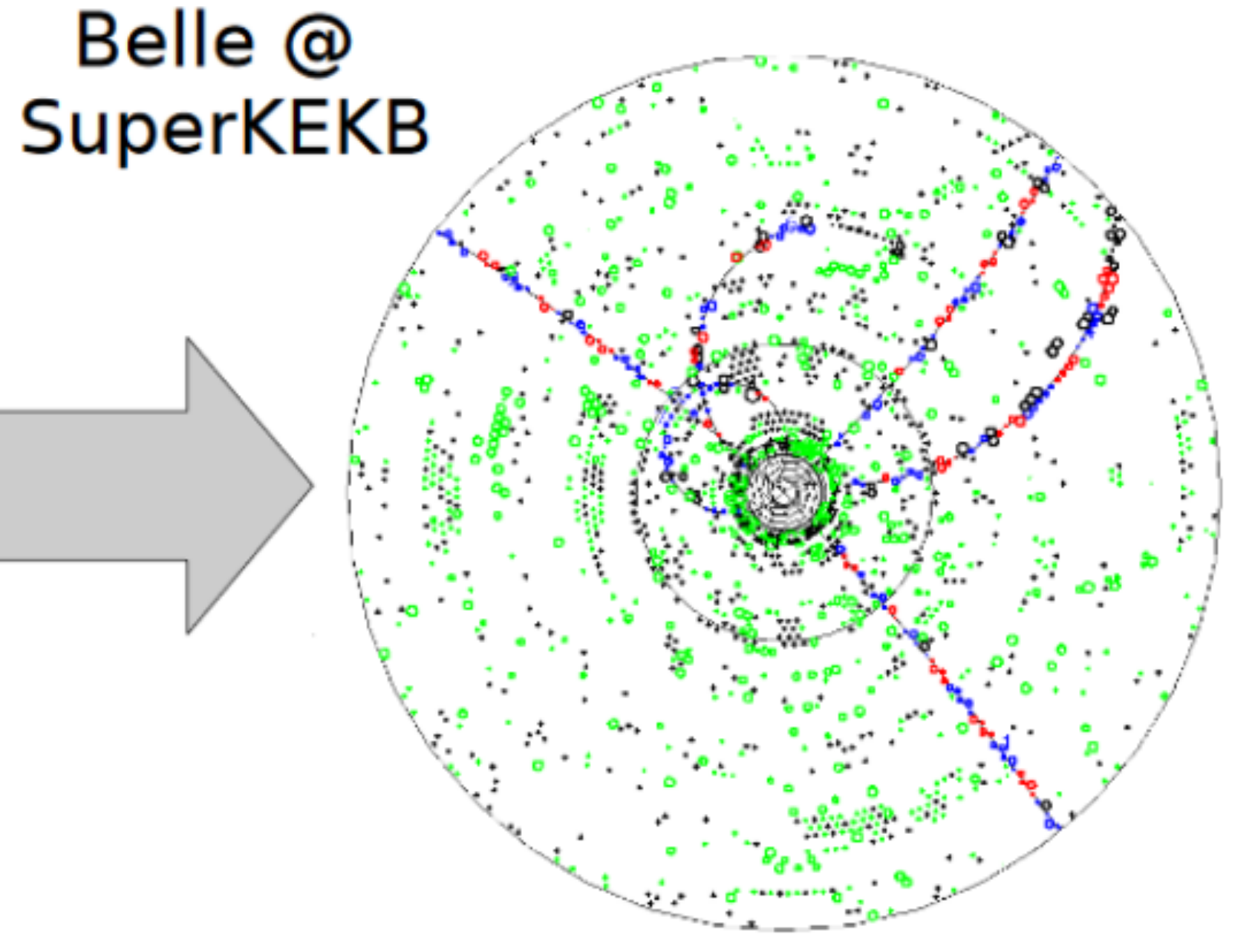
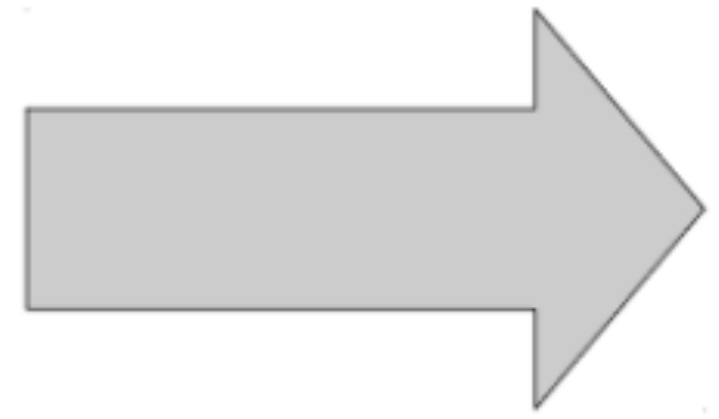
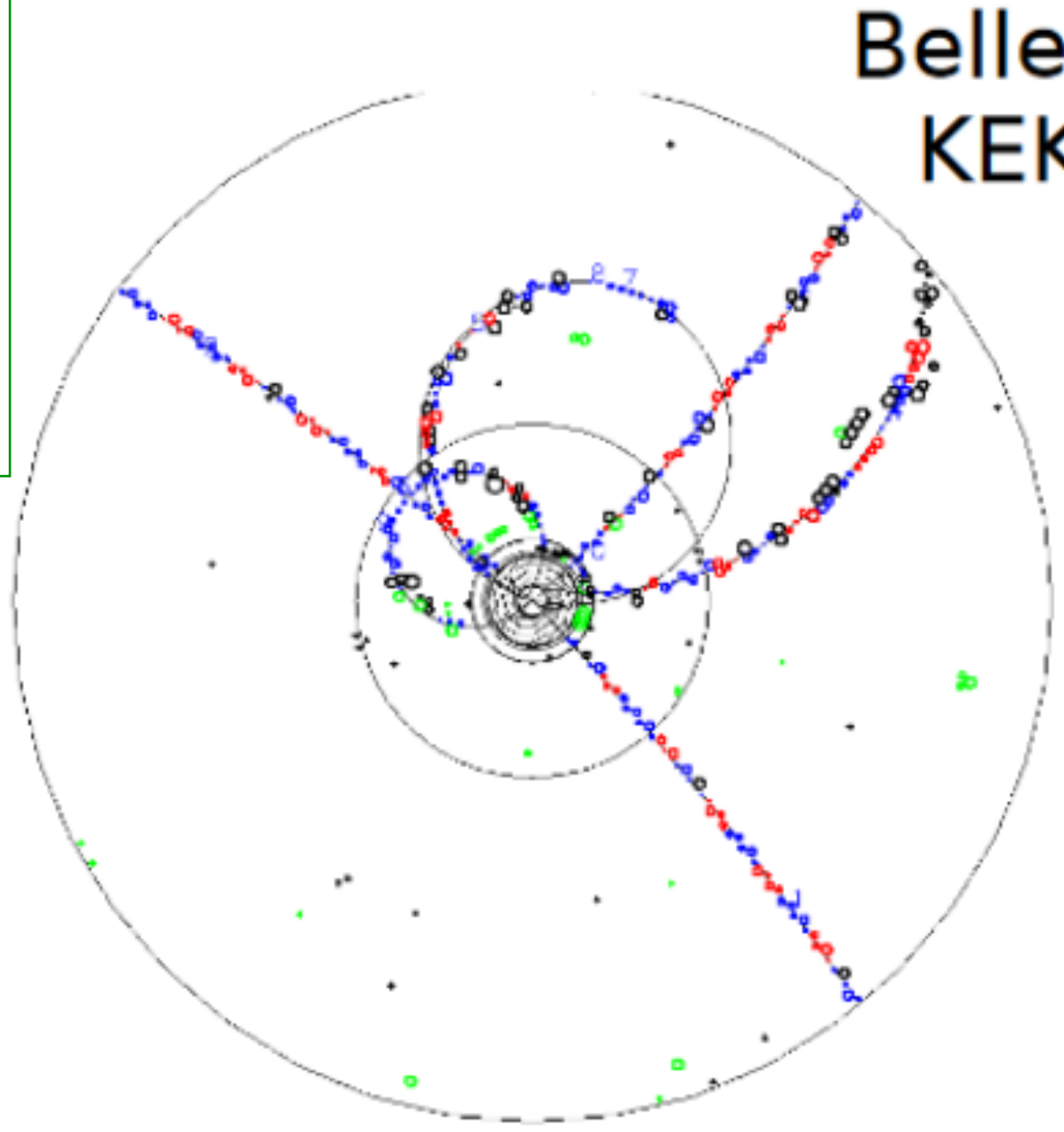
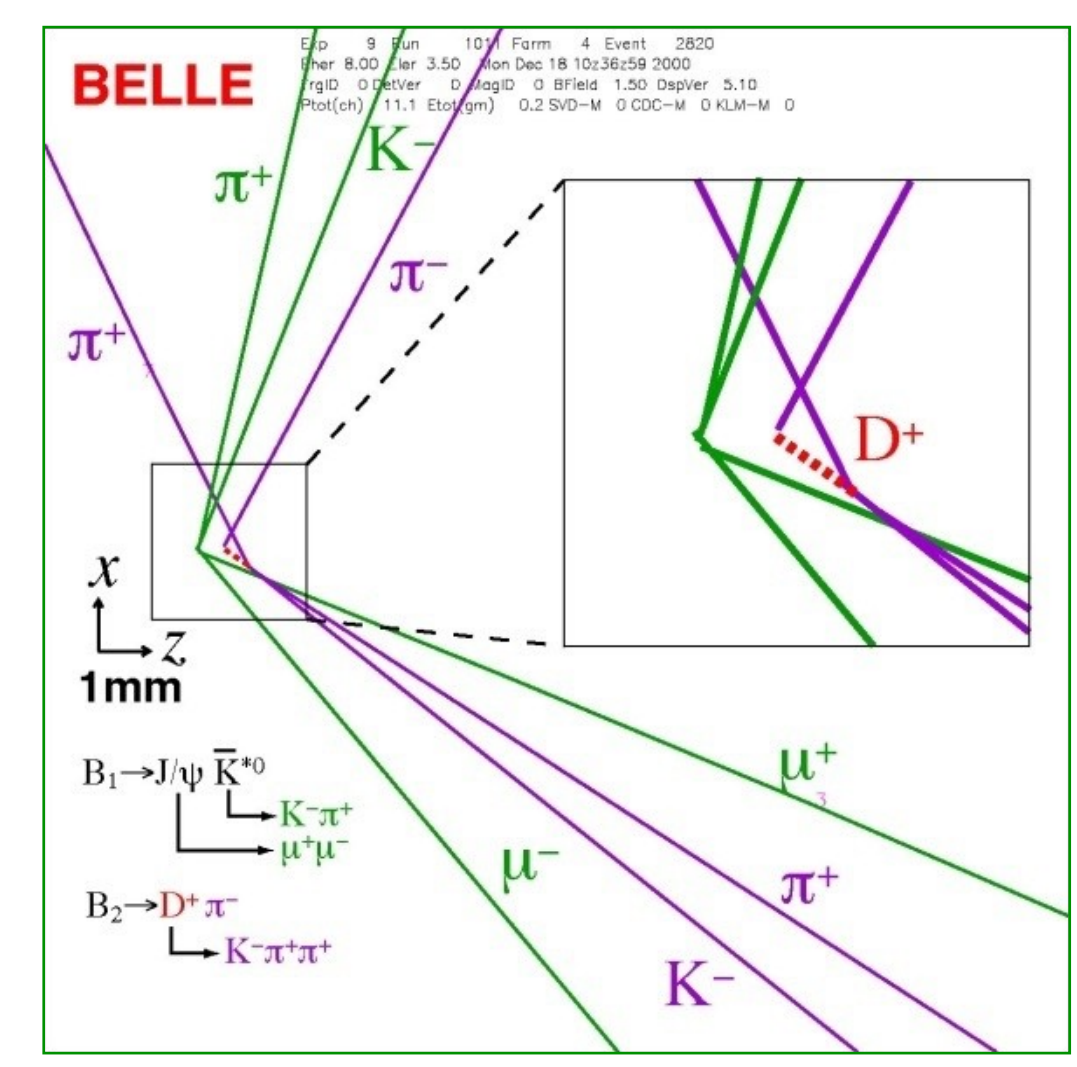
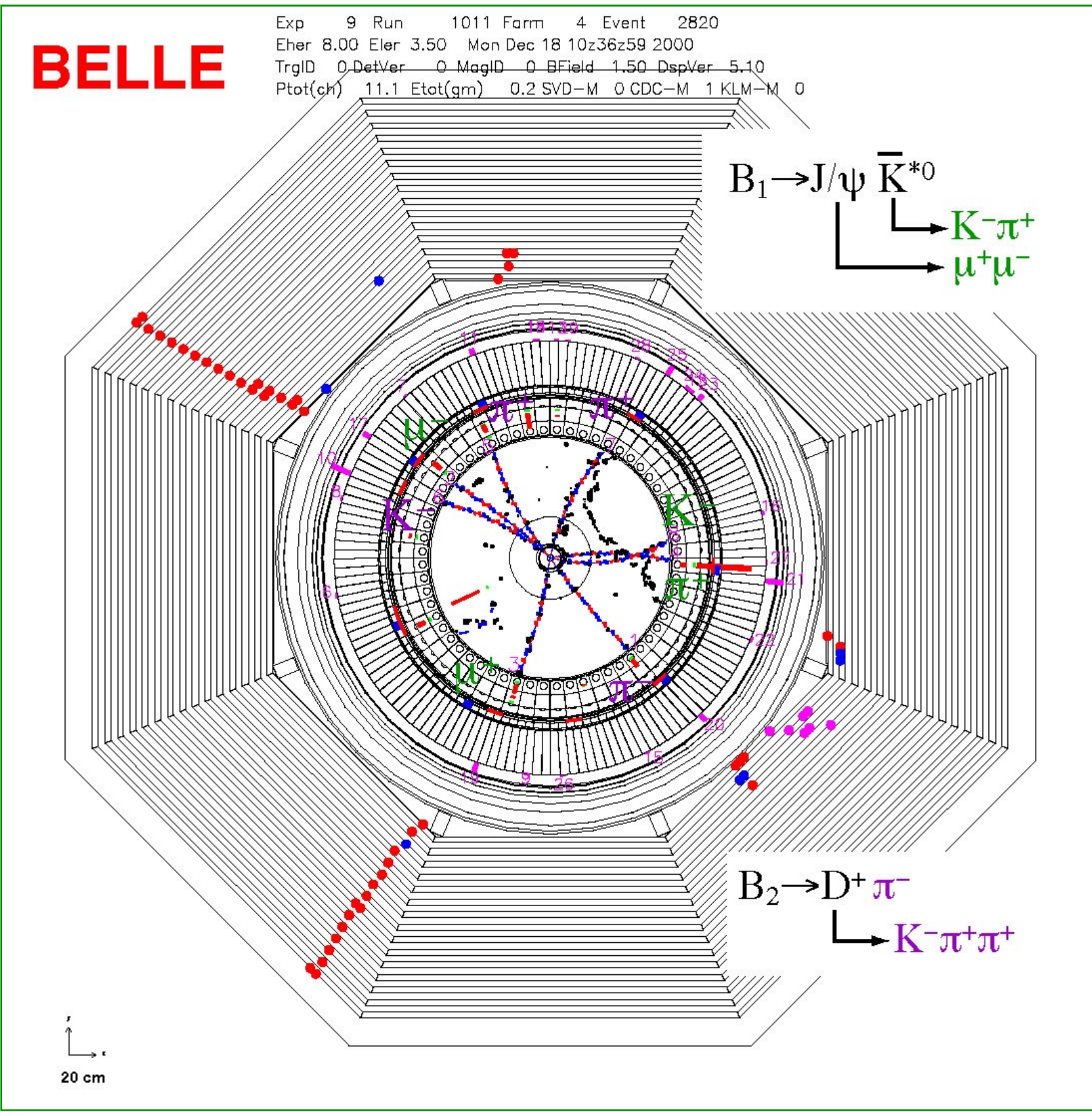
Max L1 rate: 0.5→30 kHz
Pipeline readout

**Better performance even at the
higher trigger rate and beam
background**

GRID computing

CPU 1 MHEPSpec (10⁵ core; ~ATLAS run1)
and 100 PB storage at 50 ab⁻¹

by K. Matsuoka@Kyoto seminar

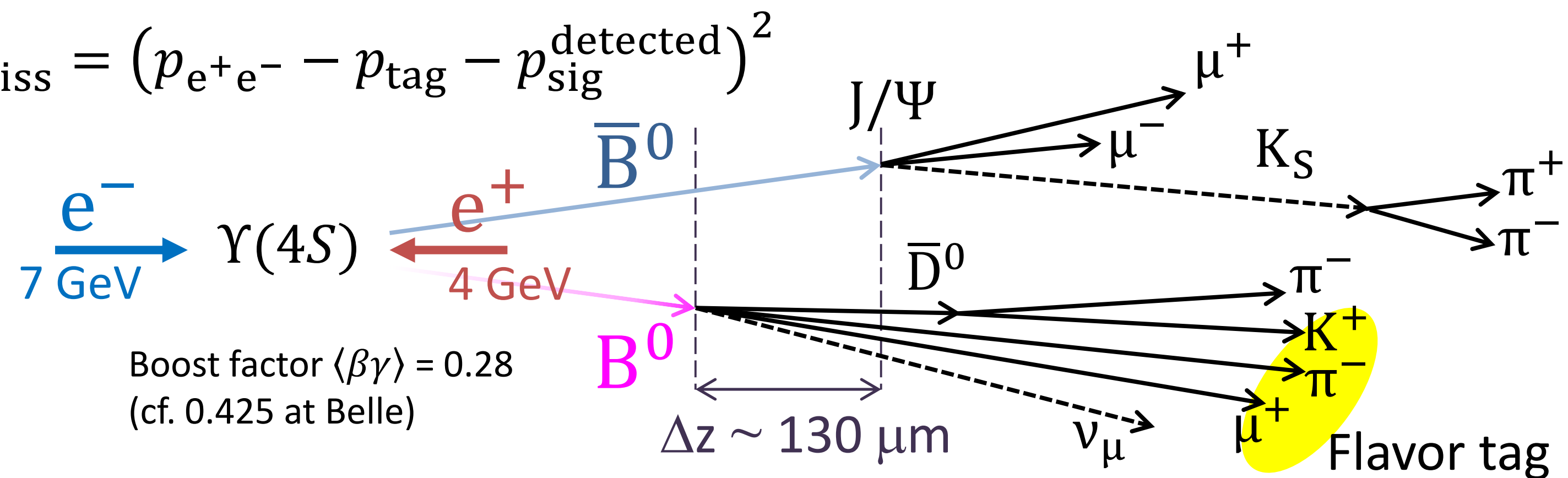


The same event with simulated background at $8 \times 10^{35} \text{ cm}^{-2}\text{s}^{-1}$

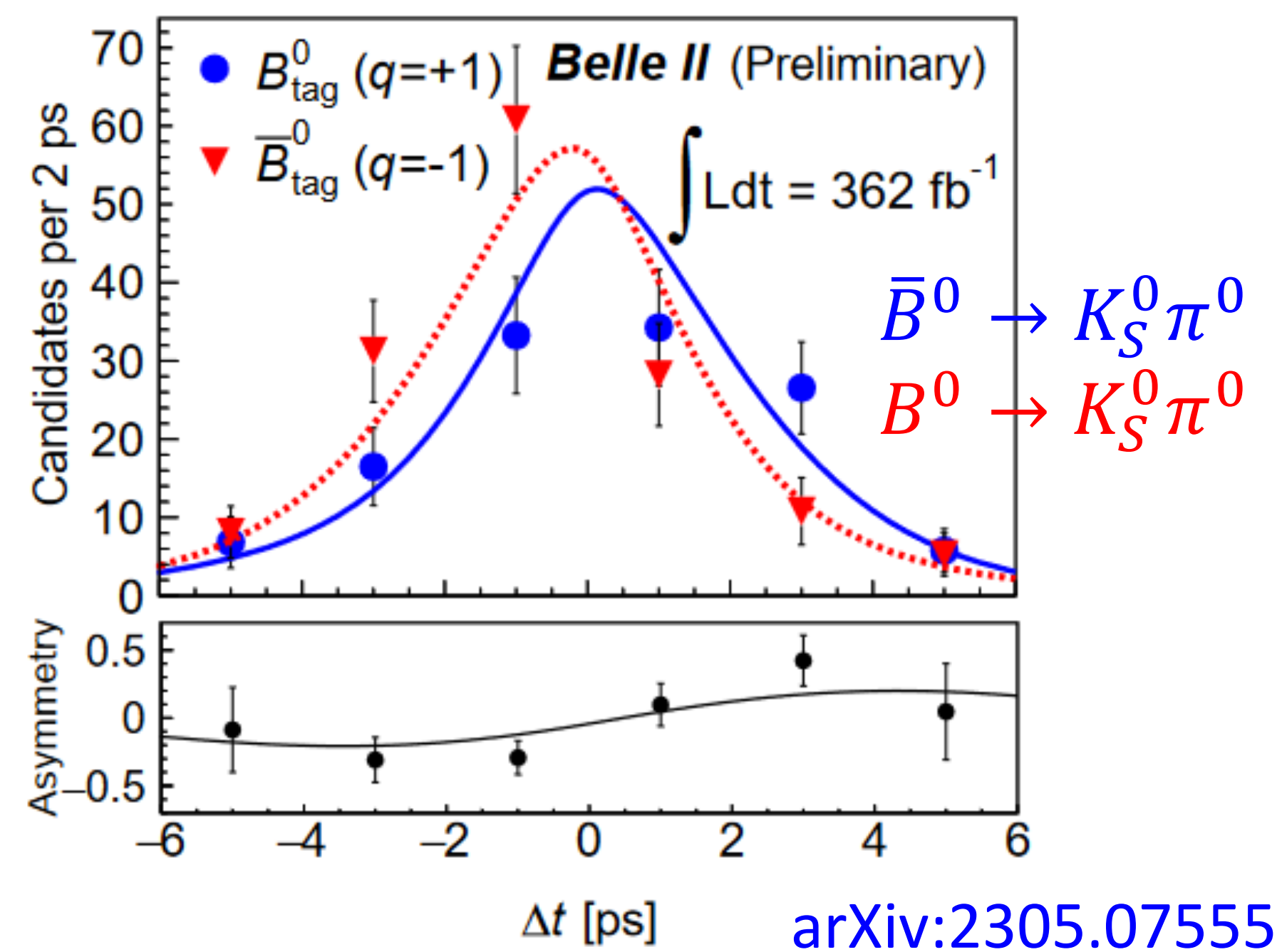
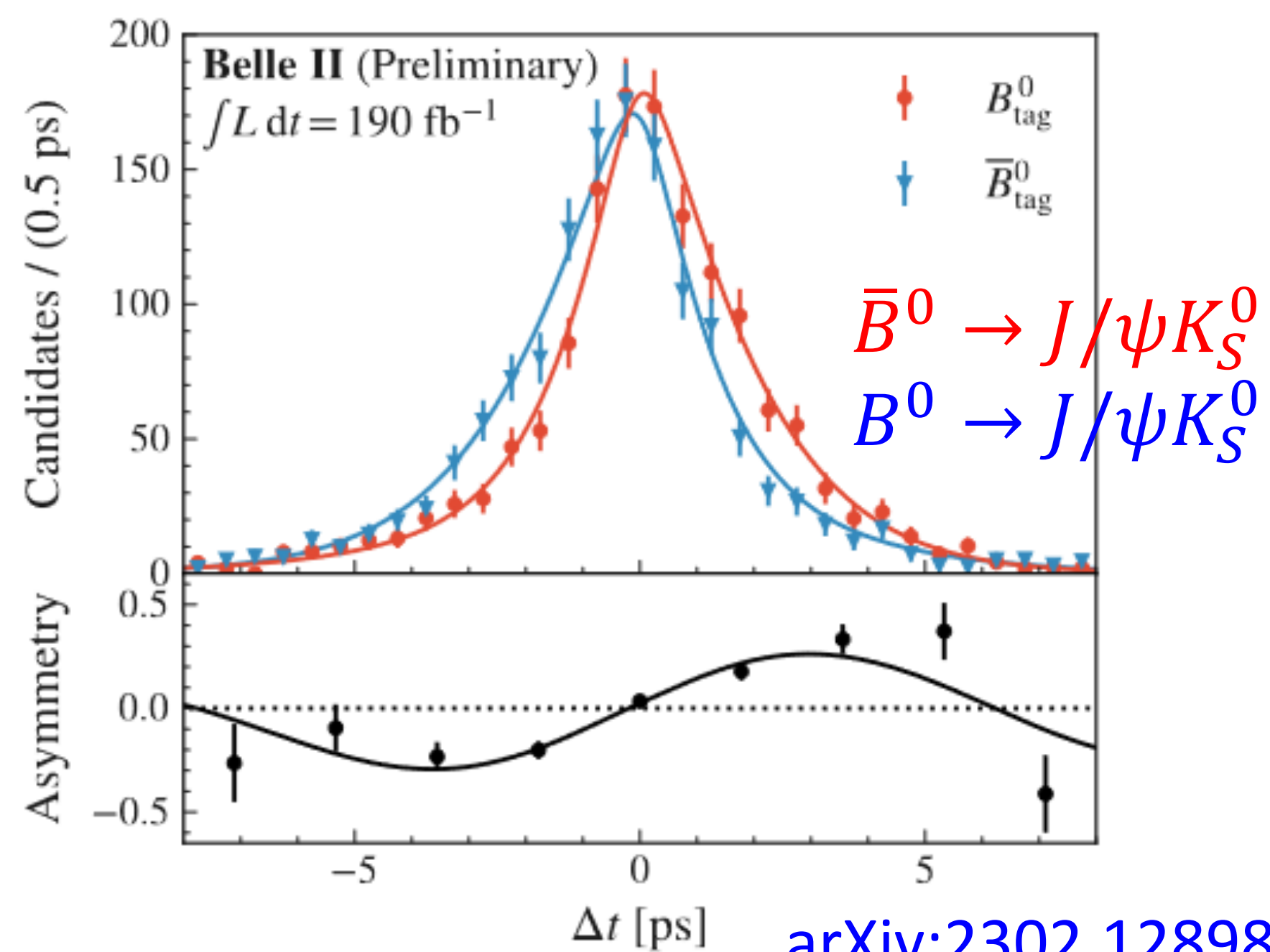
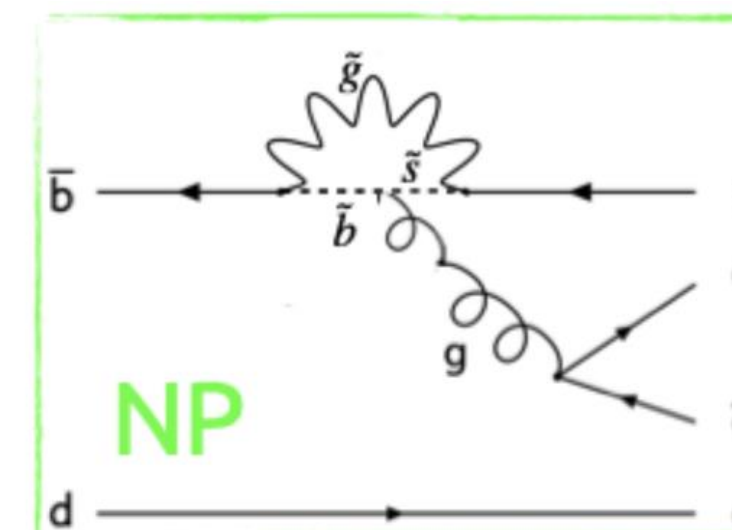
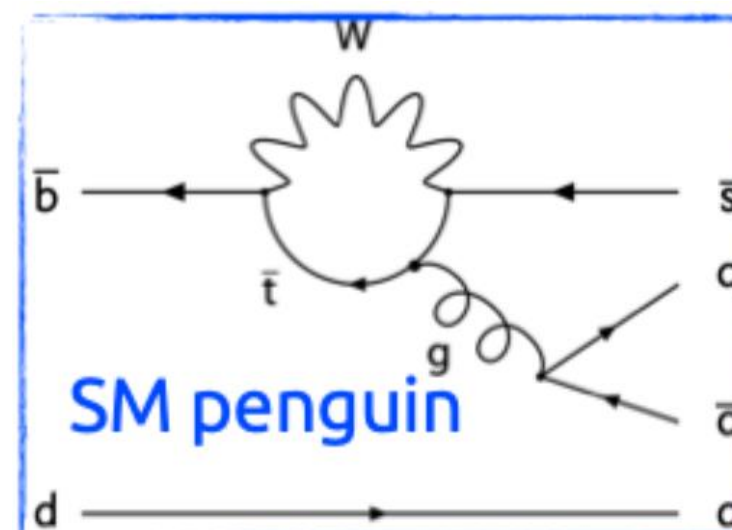
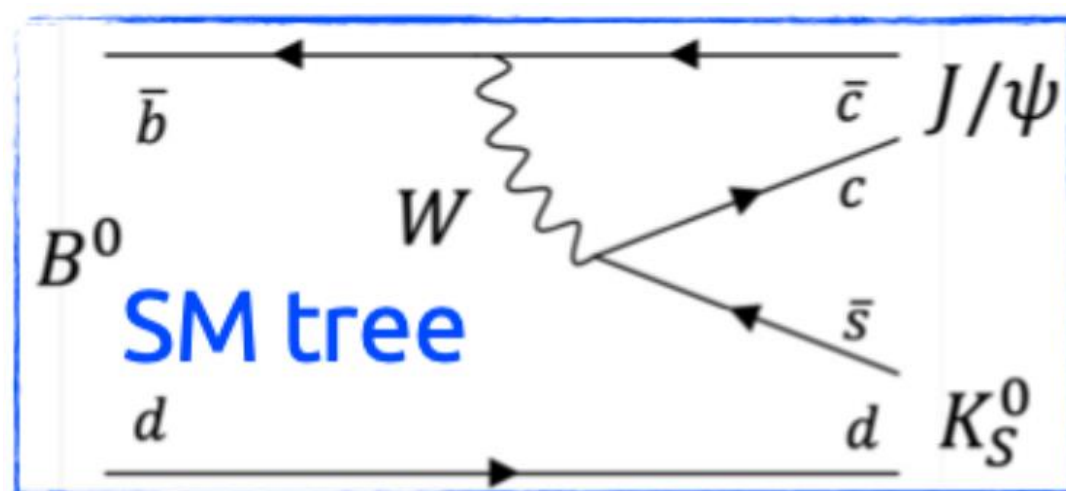
Super-KEKB/Belle II Basic

- ✓ e^+e^- collisions at (or around) $\Upsilon(4S)$
 - Well-known initial state kinematics
 - $B\bar{B}$ production from $\Upsilon(4S)$ without extra energy
 - No event pile-up
- ✓ Hermetic Belle II detector capable of detecting charged particles and reconstructing neutrals (γ, π^0, K_L^0 , etc) with high efficiencies.
- Tagging one of the B 's to infer the other B flavor and momentum.
 - Powerful S/N separation

$$m_{\text{miss}}^2 = (p_{e^+e^-} - p_{\text{tag}} - p_{\text{sig}}^{\text{detected}})^2$$



CP Violation in B decays



Belle II: $S = 0.720 \pm 0.062 \pm 0.016$

Belle: $S = 0.667 \pm 0.023 \pm 0.012$

($S \approx \sin 2\phi_1$ in this mode)

by K. Matsuoka@Kyoto seminar

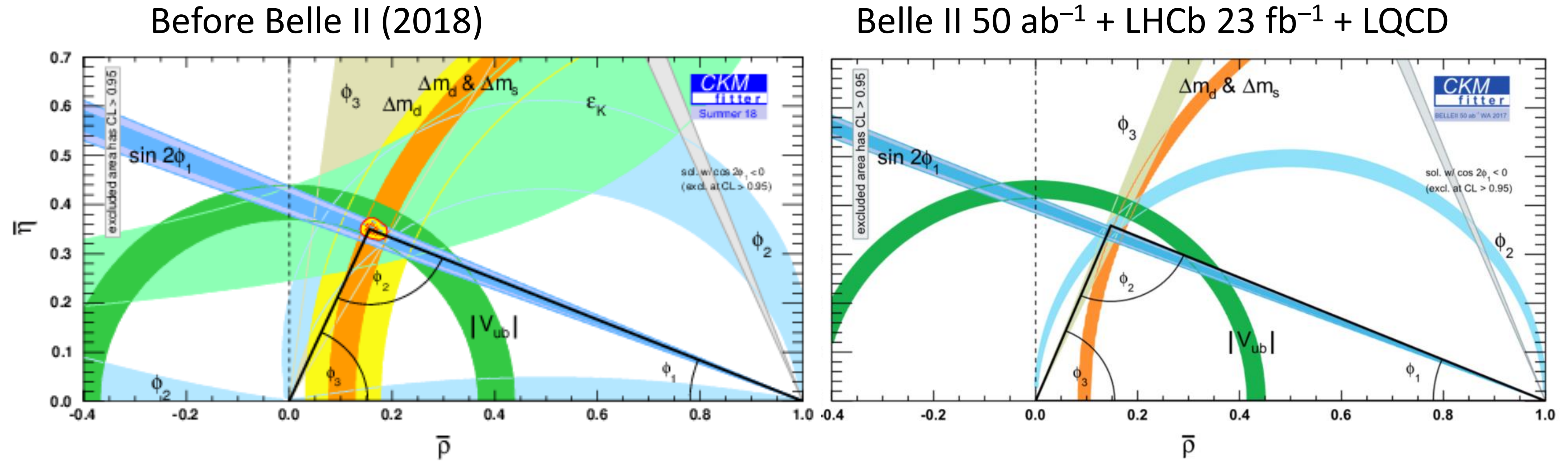
Belleの約60%のデータで凌駕する精度

Belle II: $S = 0.75 \pm_{-0.23}^{+0.20} \pm 0.04$

Belle: $S = 0.67 \pm 0.31 \pm 0.08$

Precision measurements

Fig. 226 of [Prog. Theor. Exp. Phys. 2019, 123C01](#)



The Standard Model has been tested with $\sim 10\%$ precision.

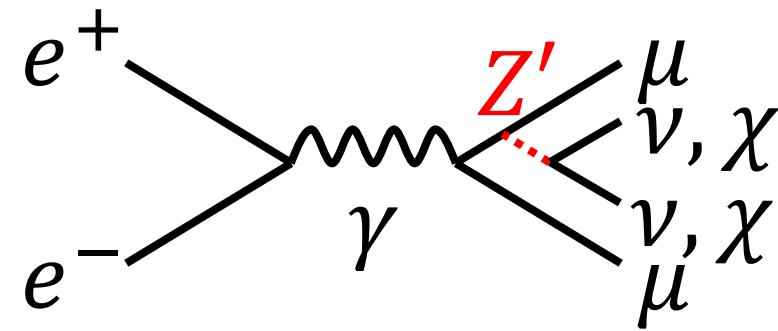
→ Search for non-standard effects that can only appear as small corrections to the Standard Model.

Access higher energy scales via quantum effects than are directly reachable at current or future colliders.

e.g. $\Lambda < \sim 1000$ TeV in B^0 mixing ($\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{1}{\Lambda^2} \mathcal{O}_{\Delta F=2}$) [\[arXiv:1302.0661\]](#)

Search for Dark particles

- Z' or LFV Z' to invisible

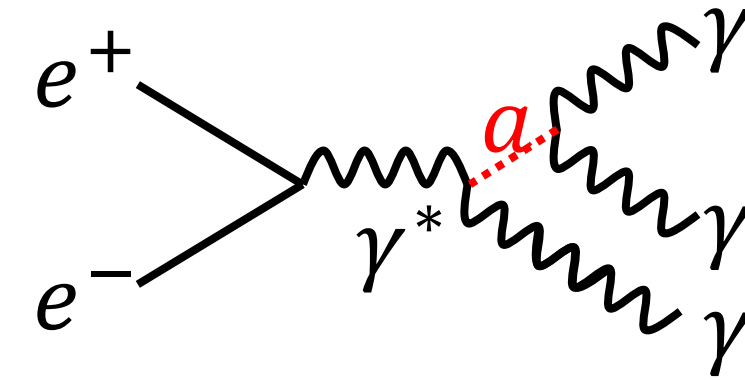


[PRL 124, 141801 \(2020\)](#)

... 1st Belle II physics paper

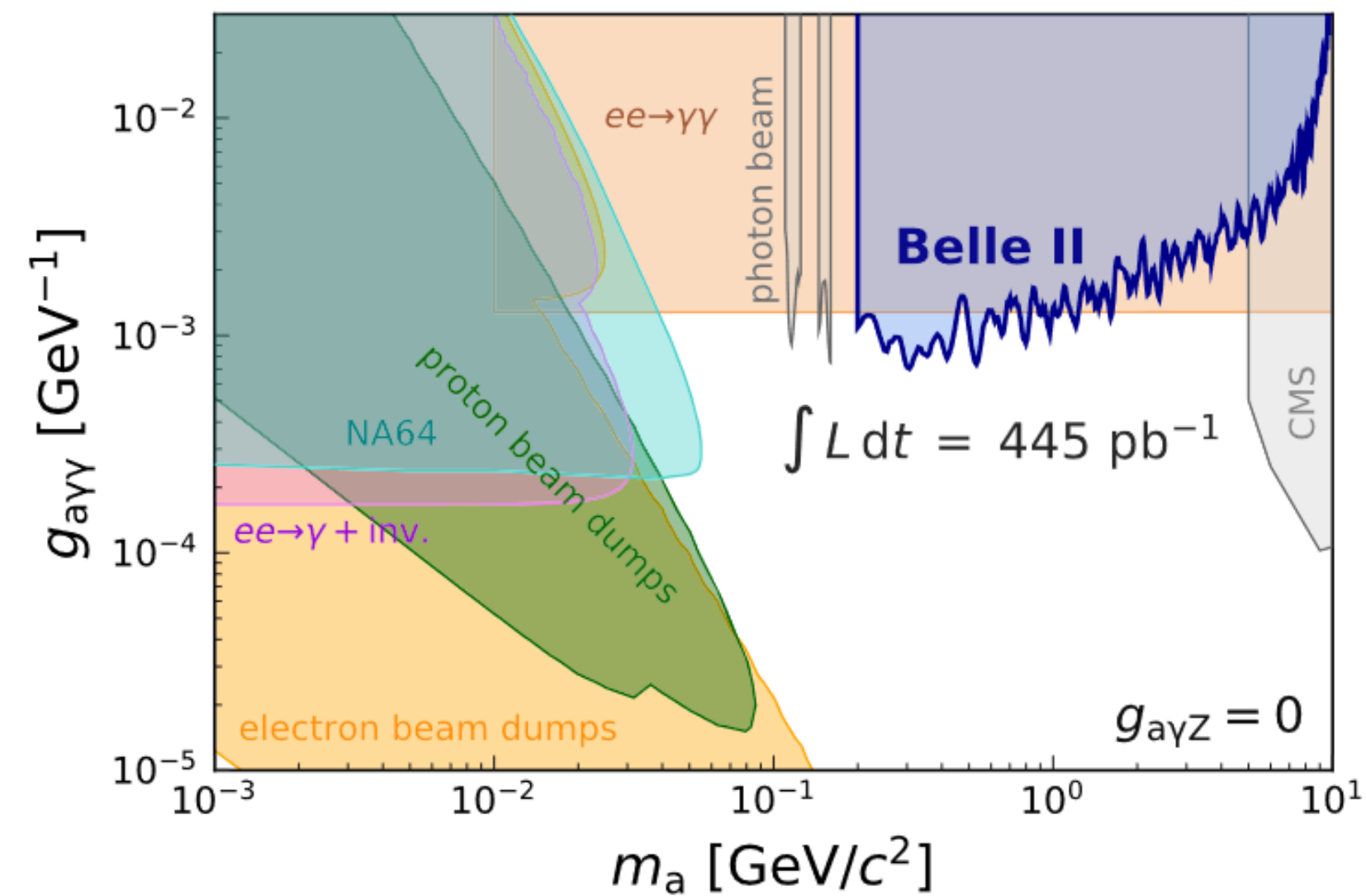
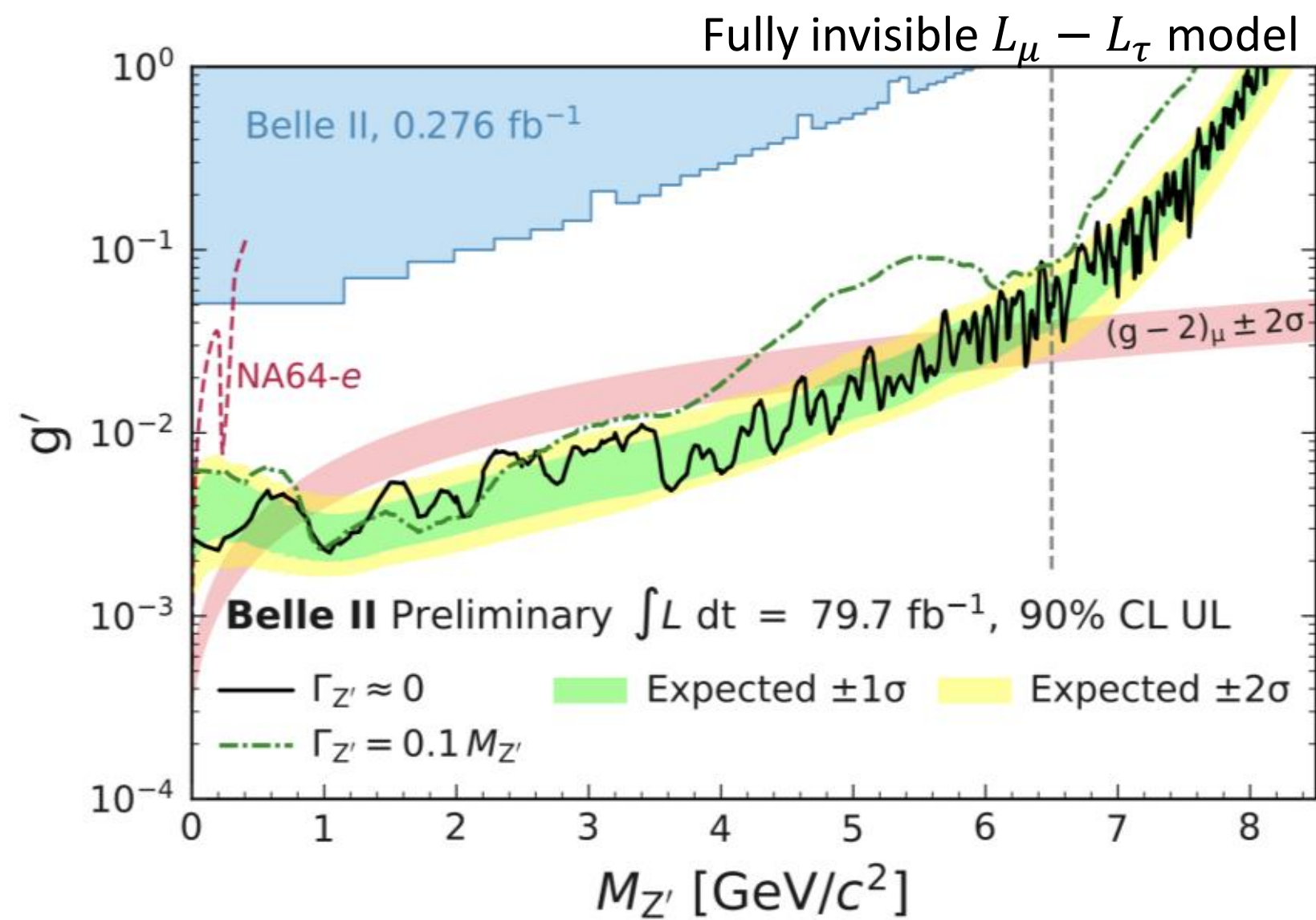
[arXiv:2212.03066](#) ... update, 2022

- Axion Like Particle



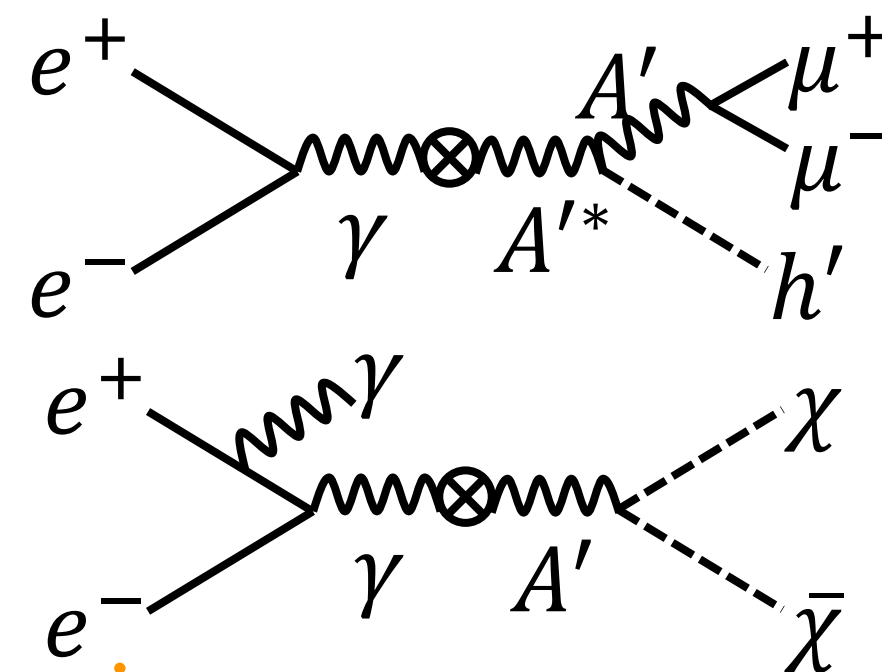
[PRL 125, 161806 \(2020\)](#)

... 2nd Belle II physics paper



- Dark higgstrahlung
[PRL 130, 071804 \(2023\)](#)

- Invisible dark photon



- Visible dark photon
- $Z' \rightarrow \mu\mu$
- Inelastic dark matter
- Dark scalar

