

Hyper-Kamiokande and proton decay

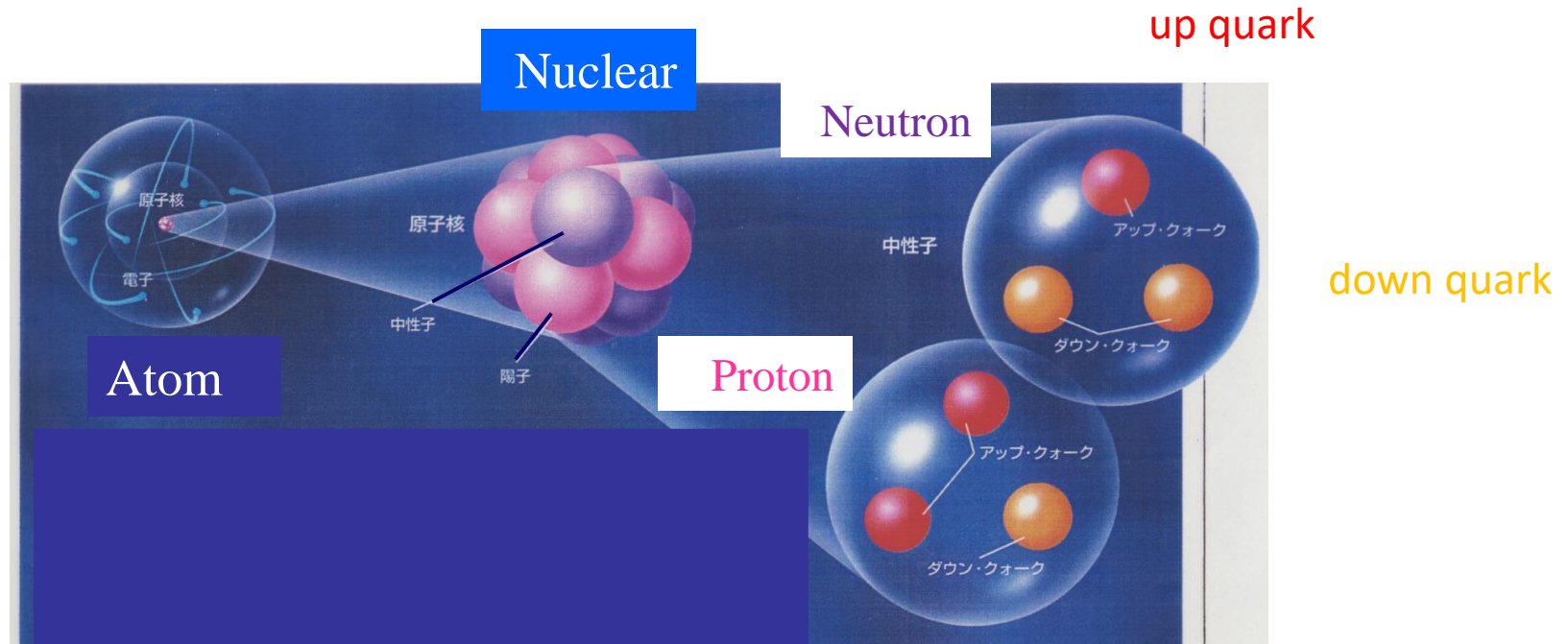
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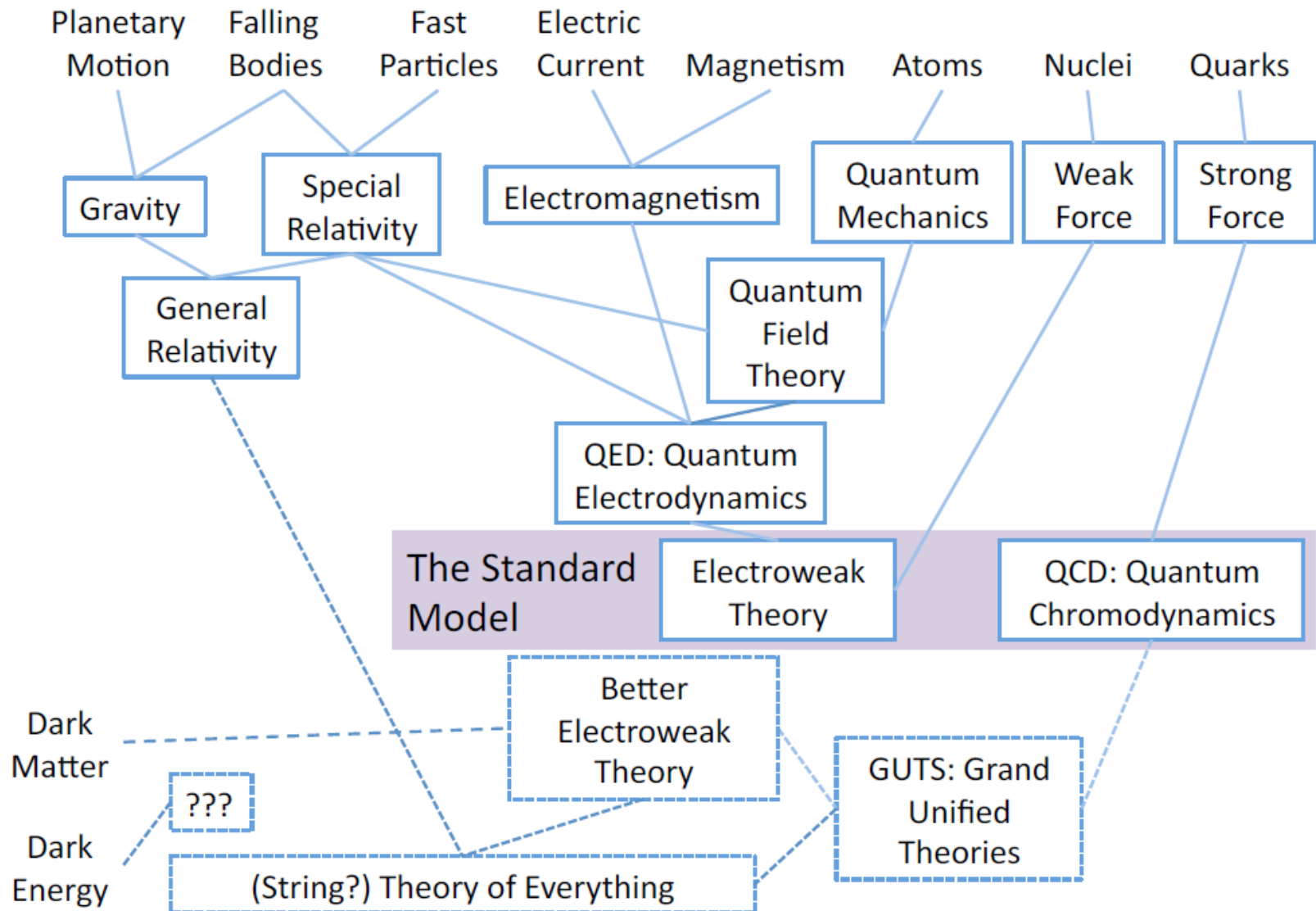
1. Does Proton Decay ?

- Nucleus consists of protons and neutrons.

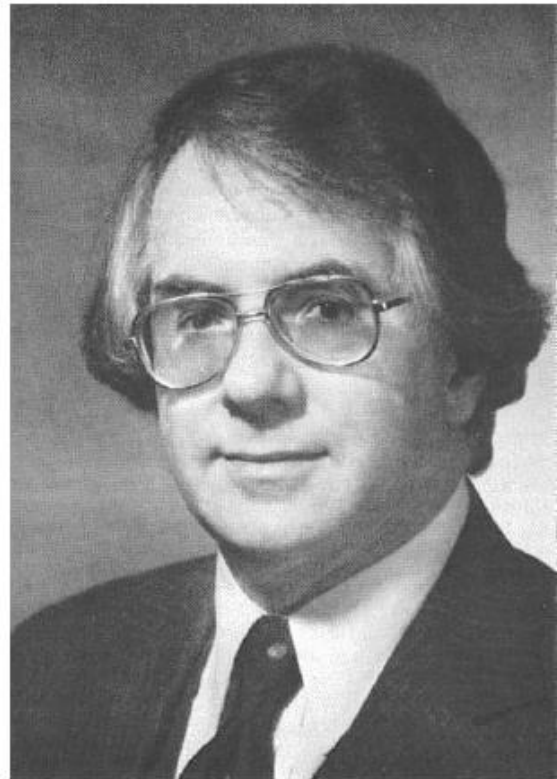


- It is well known that neutron decays spontaneously as β -decay: $n \rightarrow p + e^- + \bar{\nu}_e$
 - Note that $M_n > M_p$.
- People thought proton is stable because of baryon number conservation.
 - n,p has baryon number 1.
 - We have never observed phenomena with baryon number violation.
 - Proton is the lightest baryon in the world.
 - 1929: Weyl suggests absolute stability of proton
- But is it really true ?

Can we explain everything by a single theory ?



$SU(5)$ by Georgi and Glashow (1974)



Grand Unified Theories

Assume $SU(3) \otimes SU(2) \otimes U(1)$ is part of a larger symmetry group

QCD
E.g. $SU(3)$

EW

$$\bar{5} = \begin{pmatrix} \bar{d}_g \\ \bar{d}_r \\ \bar{d}_b \\ e^- \\ -\nu_e \end{pmatrix}_L \quad 10 = \begin{pmatrix} 0 & \bar{u}_b & -\bar{u}_r & -u_g & -d_g \\ & 0 & \bar{u}_g & -u_r & d_r \\ & & 0 & -u_b & -d_b \\ & & & 0 & -e^+ \\ & & & & 0 \end{pmatrix}_L$$

$$24 = \begin{pmatrix} G_{11} - \frac{2B}{\sqrt{30}} & G_{12} & G_{13} \\ G_{21} & G_{22} - \frac{2B}{\sqrt{30}} & G_{23} \\ G_{31} & G_{32} & G_{33} - \frac{2B}{\sqrt{30}} \end{pmatrix} \quad \begin{pmatrix} \bar{X}_1 & \bar{Y}_1 \\ \bar{X}_2 & \bar{Y}_2 \\ \bar{X}_3 & \bar{Y}_3 \end{pmatrix}$$

$$\begin{pmatrix} X_1 & X_2 & X_3 \\ Y_1 & Y_2 & Y_3 \end{pmatrix} \quad \begin{pmatrix} \frac{W^3}{\sqrt{2}} + \frac{3B}{\sqrt{30}} & W^+ \\ W^- & -\frac{W^3}{\sqrt{2}} + \frac{3B}{\sqrt{30}} \end{pmatrix}$$

Transition

generators

between lepton and quark !

Consequences:

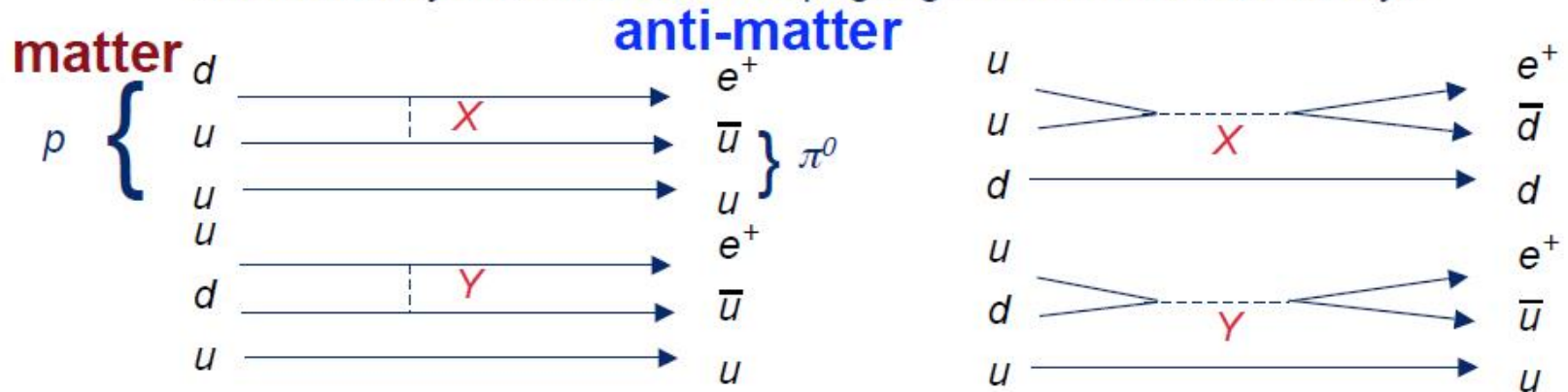
- ◆ Single (unified) coupling
- ◆ Charge quantization: $Q_d = Q_e/3$, $Q_u = -2Q_d \Rightarrow Q_p = -Q_e$
- ◆ New gauge interactions (X, Y bosons) \Rightarrow proton decay

Proton Decay in SU(5)

By Georgi and Glashow (1974)

- Decay mechanisms

dominated by the dimension=6 op. gauge boson mediated decays



($X^{\pm 4/3}, Y^{\pm 1/3}$: new gauge bosons)

- Predictions

$$\tau_B(p \rightarrow e^+ \pi^0) = 4 \times 10^{29 \pm 1.7} \text{ years}, \quad B(p \rightarrow e^+ \pi^0) \approx 40 \sim 60 \%$$

$p \rightarrow e^+ \pi^0$ became the most famous and popular decay mode.

2. How to find proton decay

- Watch a proton for very long time ($> 10^{30}$ years).
 - Age of the universe: $\sim 10^{10}$ years
 - Obviously impossible.

OR

- Watch many protons for (relatively) short time.
 - Lifetime τ : $N(t) = N(t=0)\exp(-t/\tau)$
 - **Need huge detector !**

Proton Decay Detectors

IMB

H_2O

Kamiokande

H_2O

H_2O

Super-Kamiokande

Fe

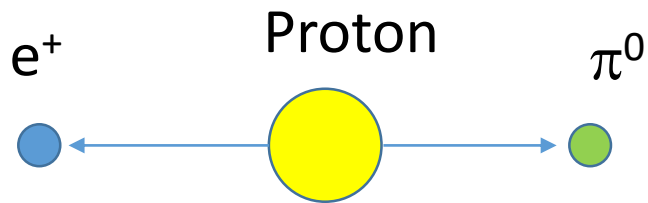
Soudan

Higher sensitivity in Water Cherenkov Detectors

2. Dominant decay mode:

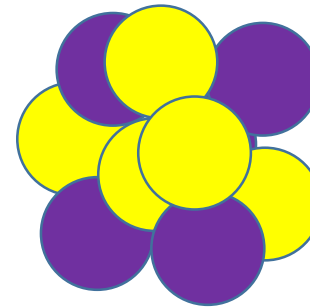
$$p \rightarrow e^+ \pi^0$$

What's important for $p \rightarrow e^+ \pi^0$?



In “free” proton case, e^+ and π^0 emit in back-to-back. Energy corresponding to proton mass is fully used.

Nucleus

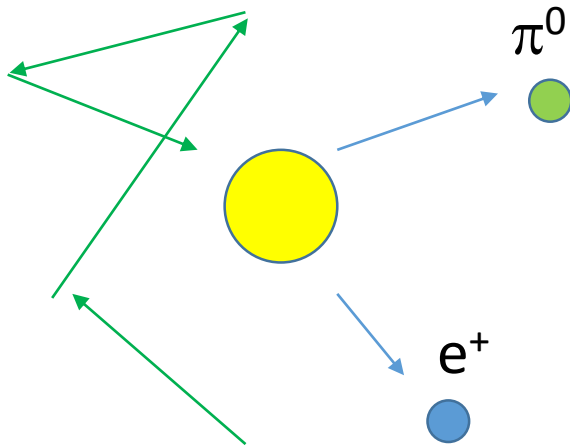


What happens if a bound proton in nucleus decays ?

Inefficiencies and uncertainties of proton decay search come from nuclear effect !

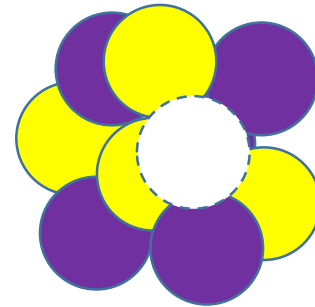
Key 1: Proton never stops in nucleus

- Protons don't exist locally in nucleus. It is always moving in the nuclear potential (Fermi motion, $p_f \sim 225 \text{ MeV}/c$).



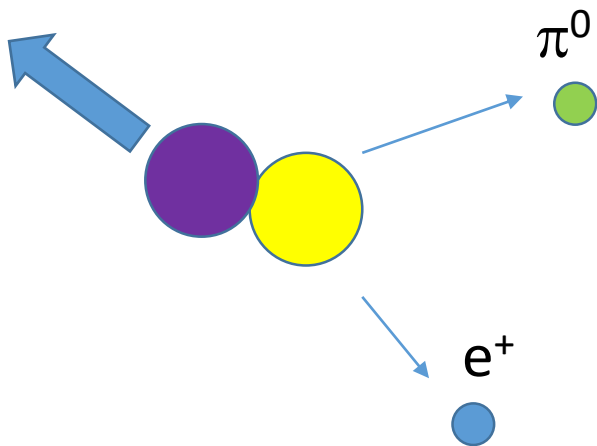
Key 2: Binding energy

- Energy corresponding proton mass should be used for compensating its binding energy (s-state: $\sim 40 \text{ MeV}$, p-state: 15 MeV in Oxygen).



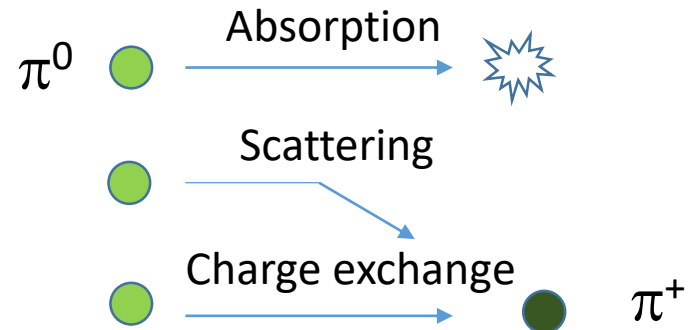
Key 3: Proton strongly binding to other nucleus

- ~ 20% protons are strongly binding to other nucleon which also bring energy when the proton decays (correlated decay)



Key 4: π interacts in nucleus

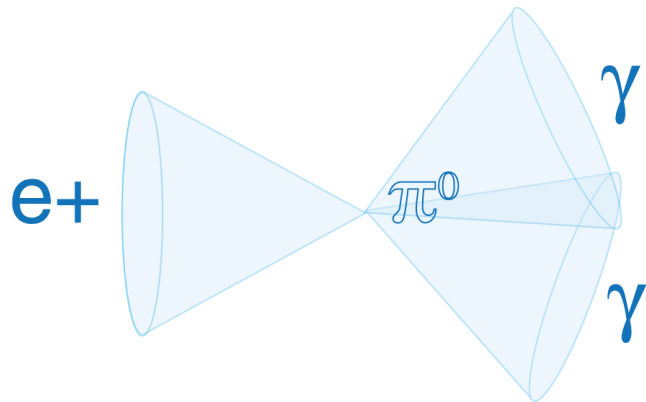
- Mesons (π , K, e.t.c.) in decay products are affected in nuclear interactions before exiting nucleus.



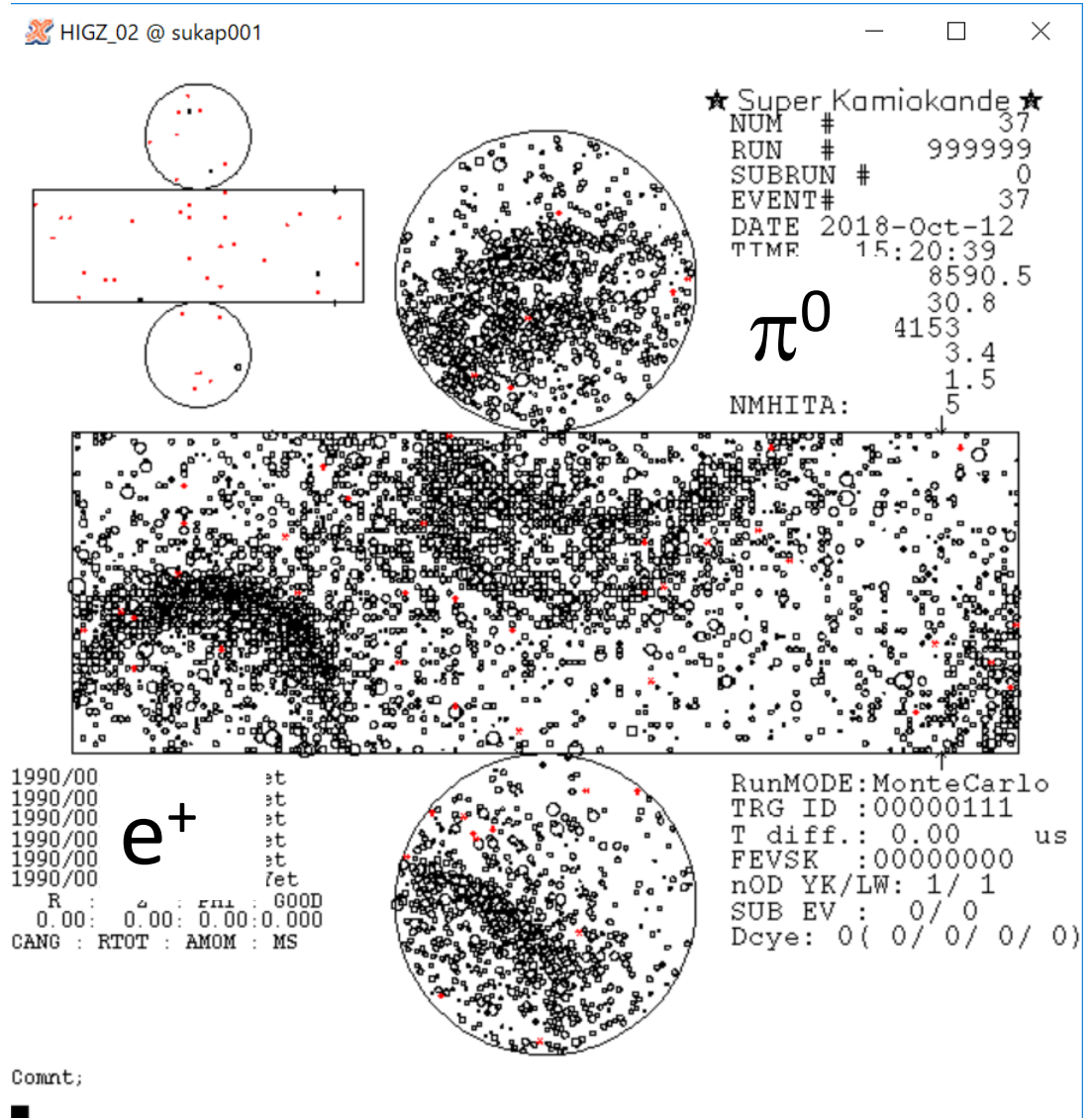
Why water is used for proton decay search ?

- Easy to construct larger detector.
 - Much cheaper than iron.
 - You can find large water tank everywhere (common technology).
- High efficiency and low uncertainty.
 - H_2O has two hydrogens which are not affected by nuclear effect . They are regarded as “free” proton.
 - ✓ Bound proton: $\sim 200 \text{ MeV}/c \sim 0.2c = 6 \times 10^7 \text{ m/s}$
 - ✓ velocity of molecular in liquid $\sim 10^2 \text{ m/s}$
 - Free protons contribute high selection efficiency and low uncertainty.

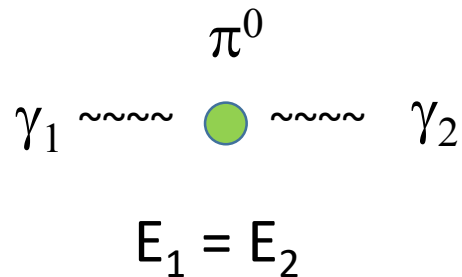
How look like $p \rightarrow e^+ \pi^0$ in SK ?



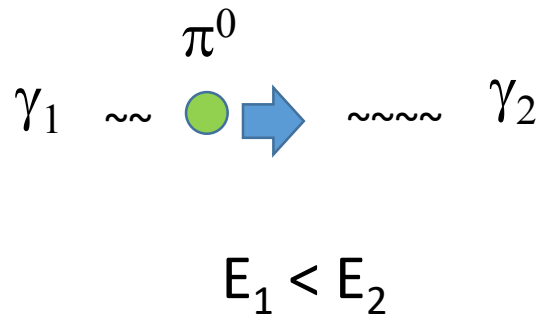
Three e-like rings
should be observed.



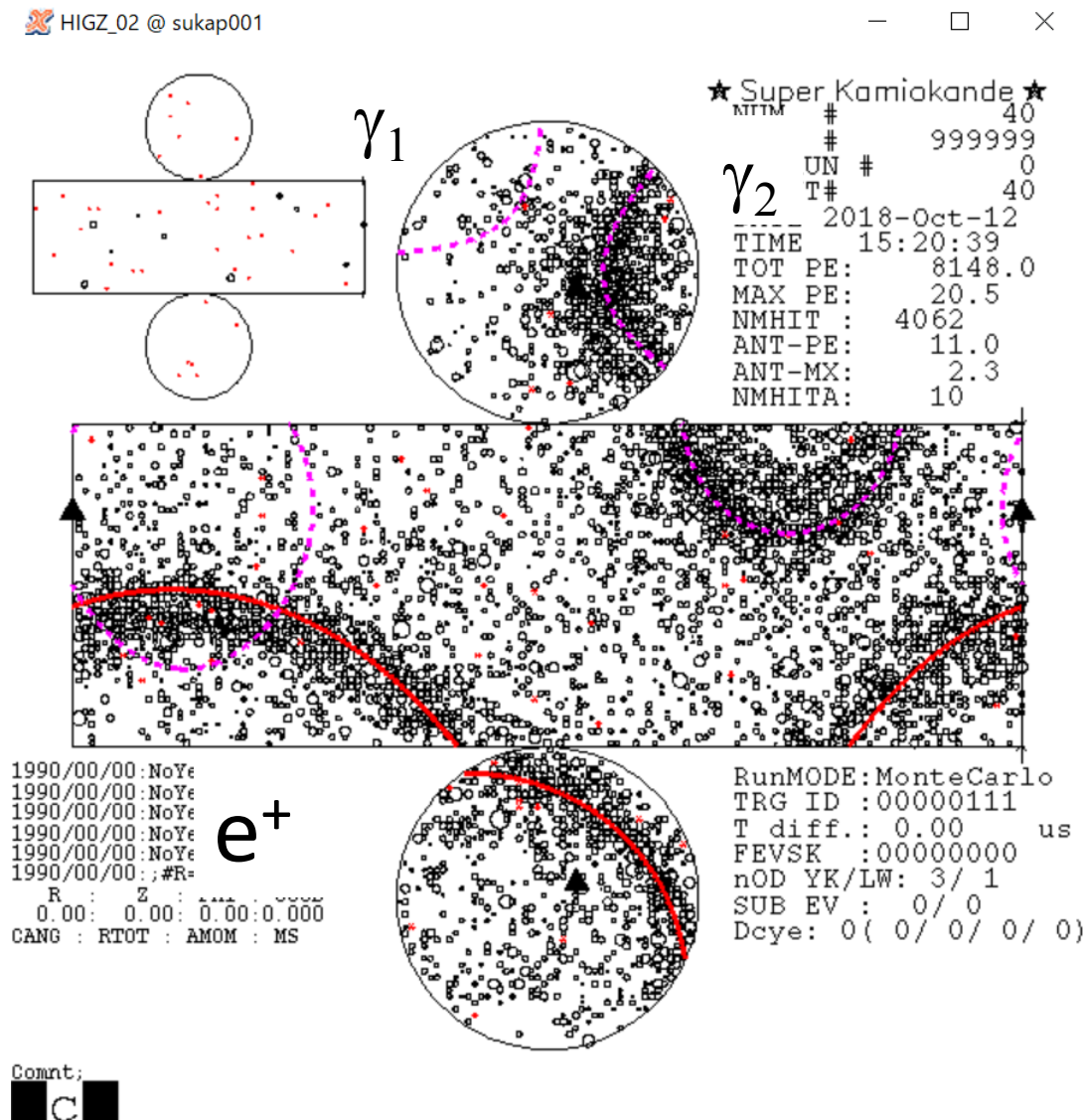
Stopped π^0 case

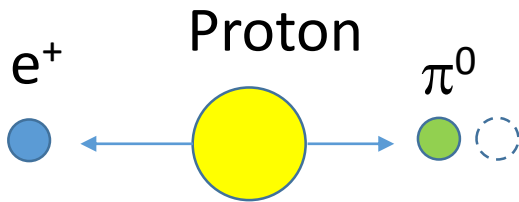


If a γ is emitted π^0 direction

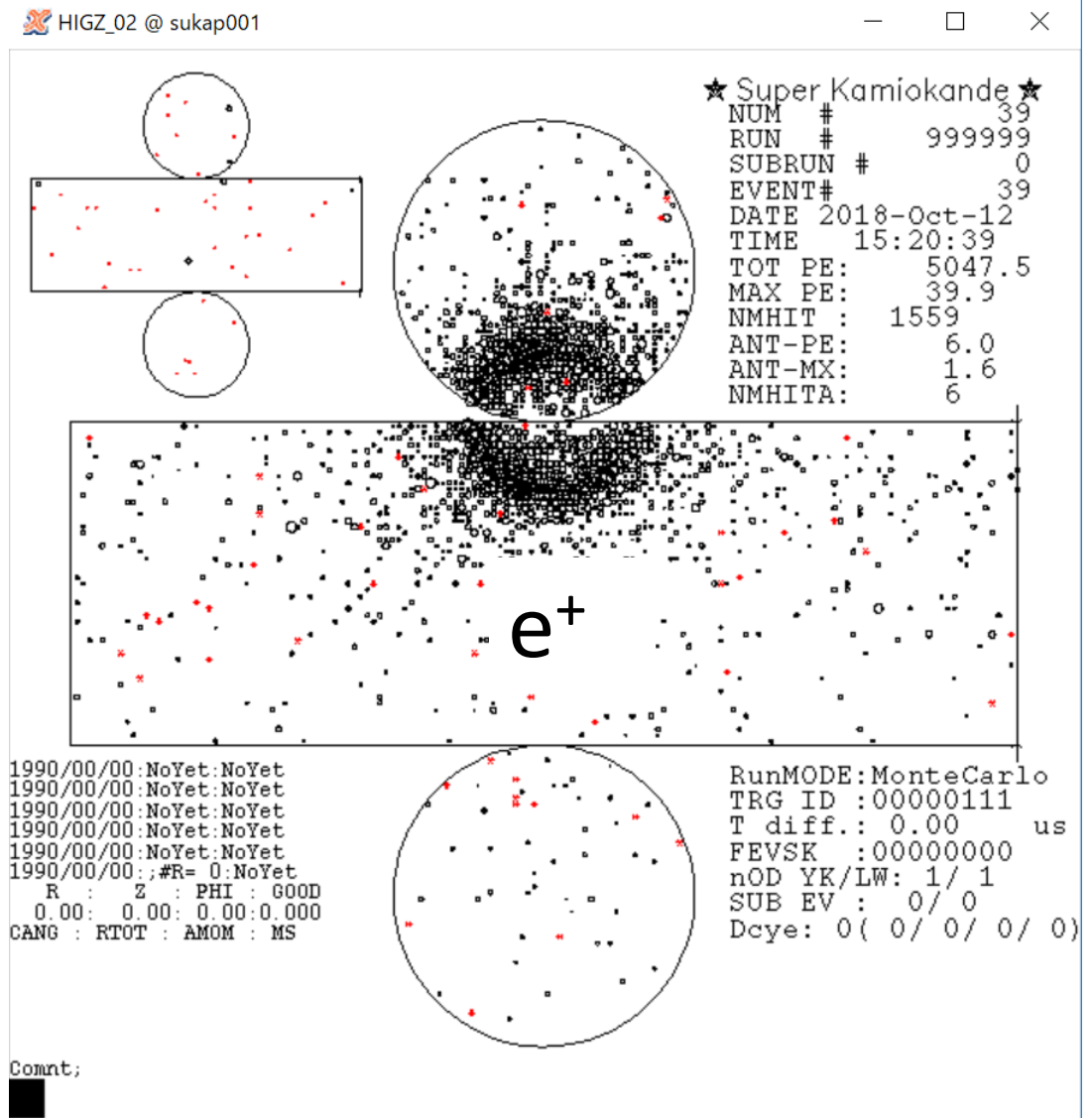


Sometimes one γ is failed to reconstruct and observed only two rings.

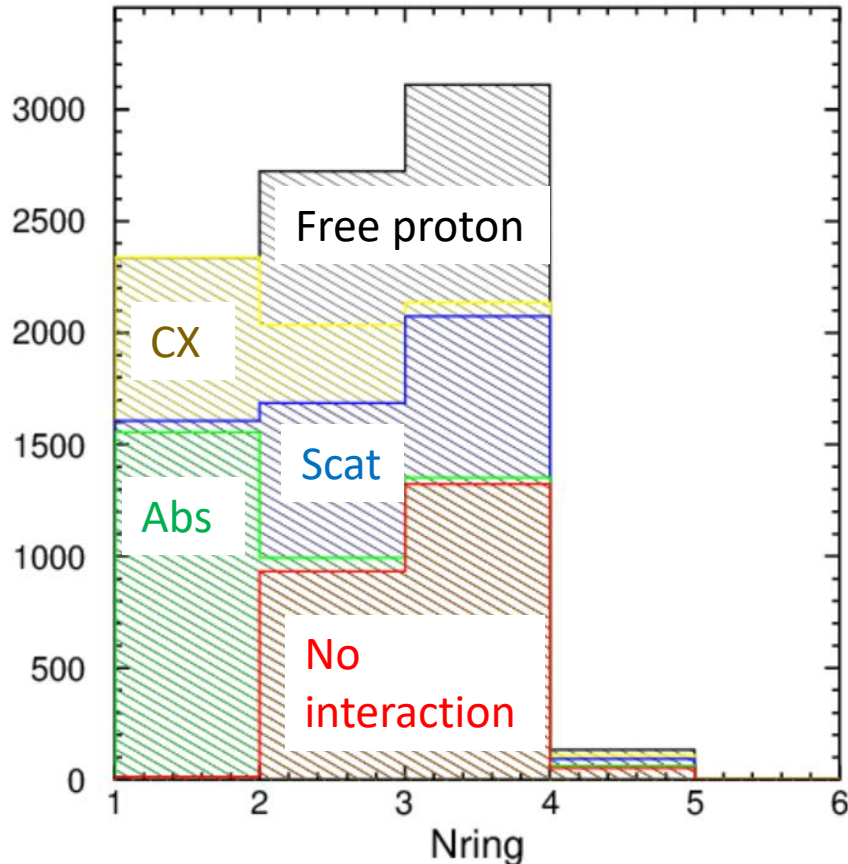




If π^0 is absorbed before exiting nucleus, only e^+ is observed (one ring).



Observed number of ring for $p \rightarrow e^+ \pi^0$



Free proton: H in H₂O

No interaction in Nucleus

Abs: π^0 absorption in Nucleus

Scat: scattered

CX: charge exchange

($\pi^0 \rightarrow \pi^\pm$, below threshold)

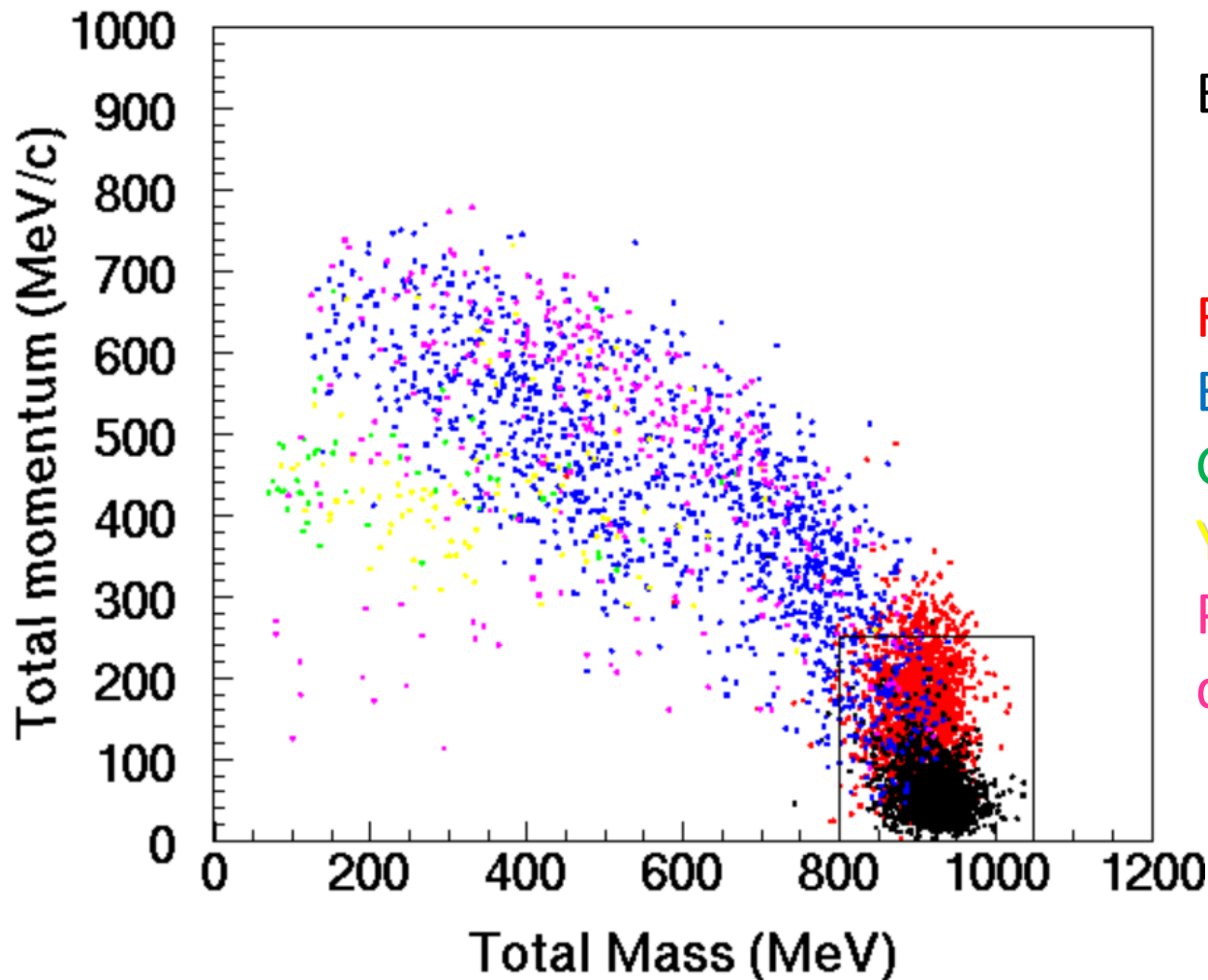
Choose 2 or 3 rings.

Selection criteria for $p \rightarrow e^+ \pi^0$

1. Event vertex should be located 2 m inward from the tank wall (fiducial volume cut, 22.5kton).
2. 2 or 3 ring event.
3. All ring should be e-like (Particle Identification).
4. No Michel electrons.
5. Reconstruct π^0 mass for 3 ring events. It should be $85 < M_{\pi^0} < 185 \text{ MeV}/c^2$
6. Reconstruct total mass and momentum should be $800 < M_{\text{tot}} < 1050 \text{ MeV}/c^2$, $P_{\text{tot}} < 250 \text{ MeV}/c$.

Total mass vs Total momentum for $p \rightarrow e^+ \pi^0$

- Selection efficiency $\sim 40\%$
- Inefficiency is dominated by unavoidable physics processes.



Black: Free proton

Bound proton

Red: No π interaction

Blue: π scatter

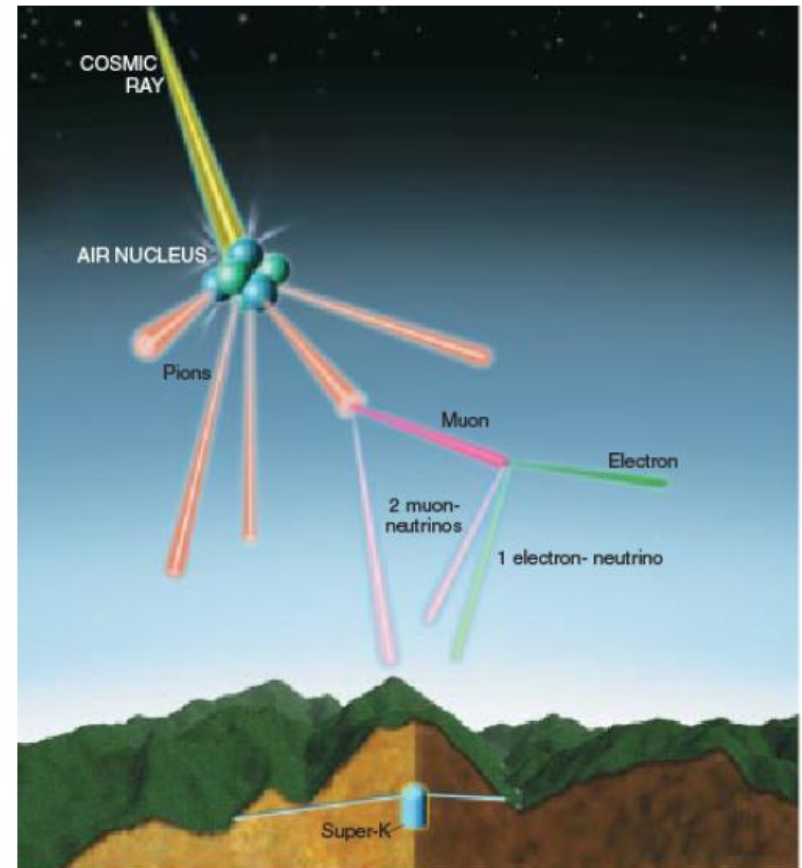
Green: π absorption

Yellow: π CX

Purple: Correlated decay

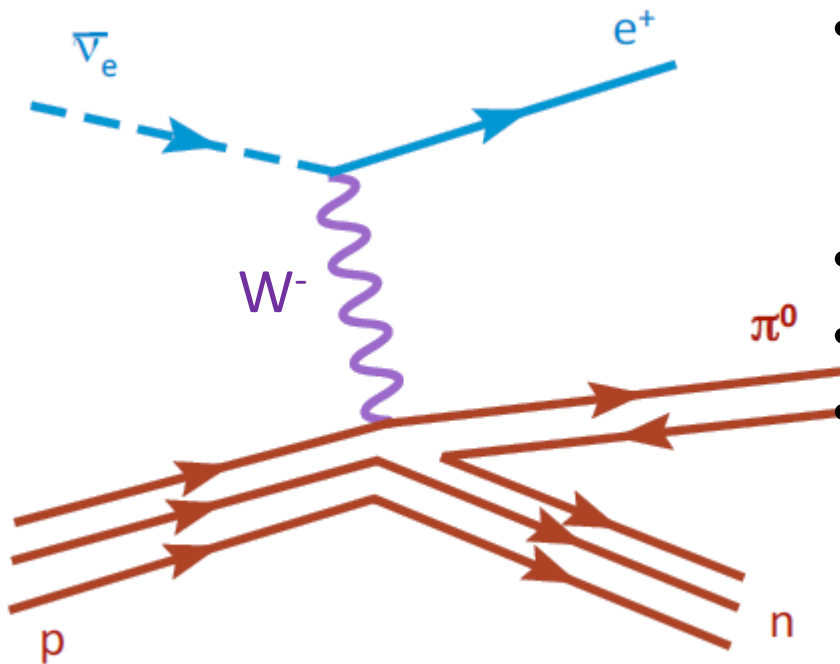
2-2. What's Background events for proton decay searches ?

- **Atmospheric neutrino** is dominant backgrounds for proton decay searches.
 - Visible energy ~ 1 GeV.
 - Solar or SN ν is too low energy.
 - Cosmic ray μ are rejected by outer detector.



Typical background for $p \rightarrow e^+ \pi^0$

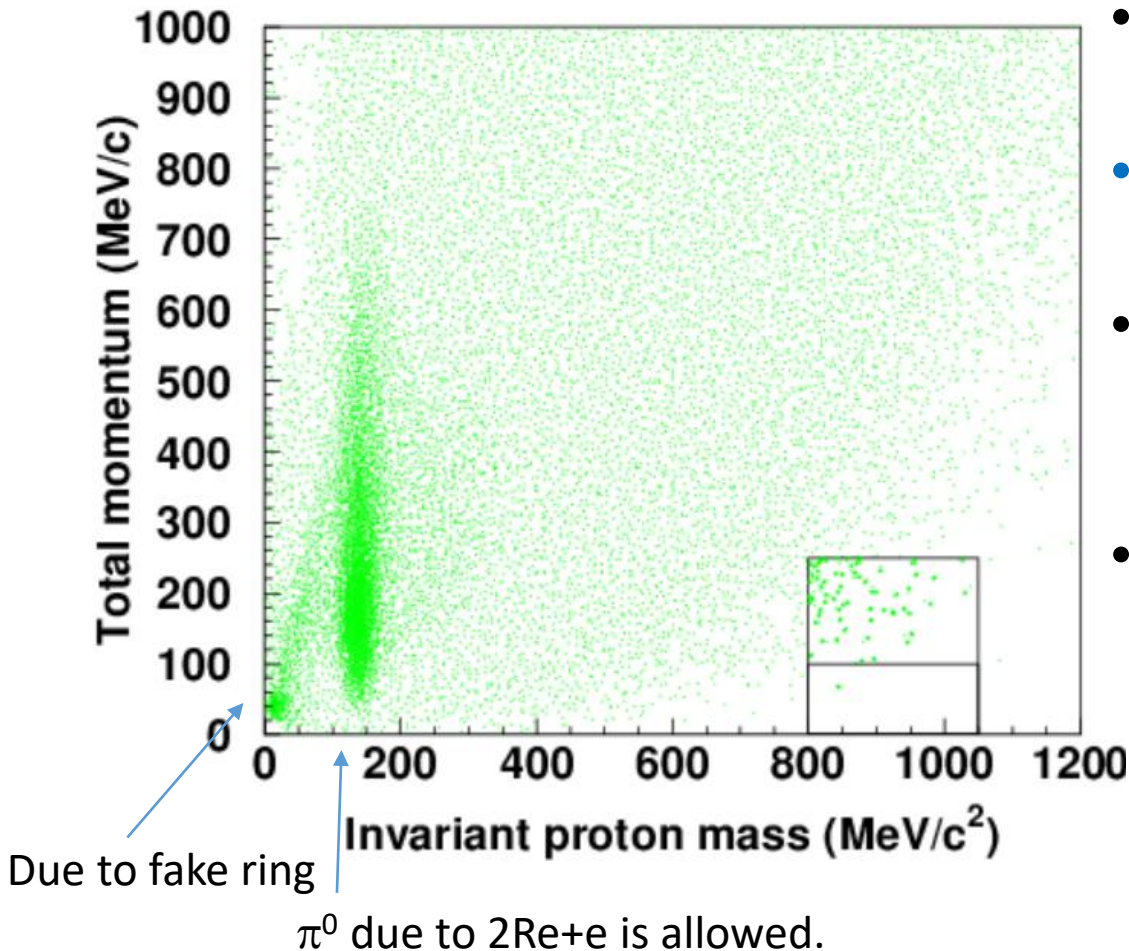
Charged current π^0 production



- Exchange W boson between ν and proton (charged current interaction).
- ν changes to e^+ .
- π^0 and neutron are produced.
- Because neutron doesn't emit Cherenkov light, visible particles after the reaction are same as $p \rightarrow e^+ \pi^0$

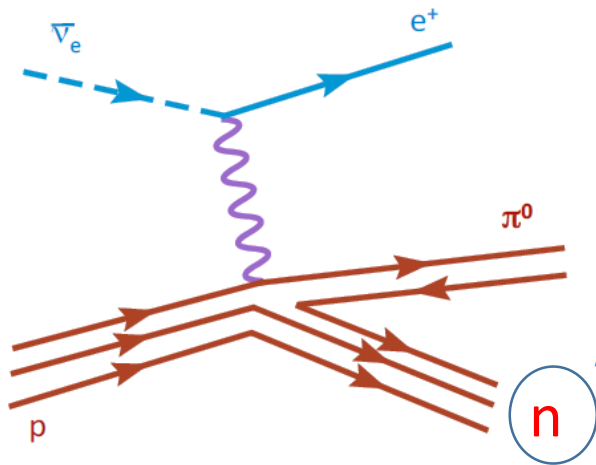
Total mass vs Total momentum for atmospheric ν background MC

(After all cuts except for total mass and momentum)



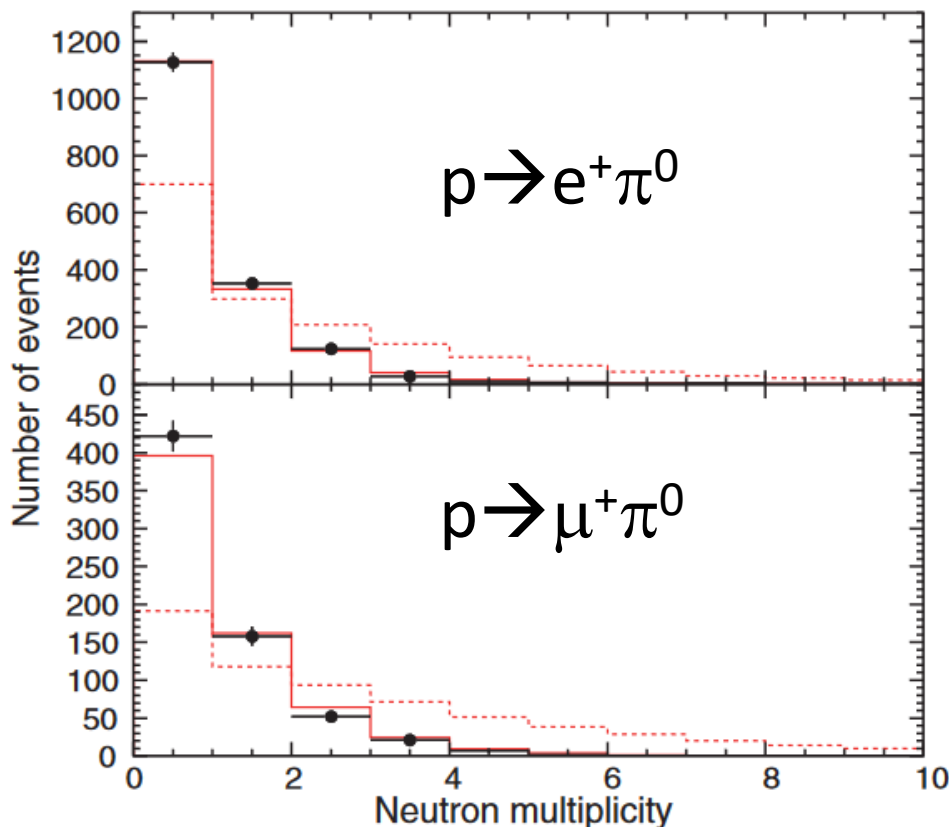
- Generate huge atm.n MC, 2000 year of SK!
- Expected BG:
 $\sim 1.3 \text{ ev/Mton*yr}$
- Neutrino events tend to have higher total momentum.
- Almost background free in lower momentum region ($< 100 \text{ MeV/c}$).
 - The region corresponds to free proton decay.

Further background reduction



- Neutron doesn't emit Cherenkov light.
- However, neutron is thermalized in water and finally captured by hydrogen ($\sim 200 \mu\text{s}$);
$$n + p \rightarrow d + \gamma \text{ (2.2 MeV)}$$
- If we can detect delayed 2.2 MeV γ ray, we can reduce background more.
- Neutron capture is also important for SN Relic ν and separate ν and $\bar{\nu}$ interactions in atmospheric n oscillation analysis.

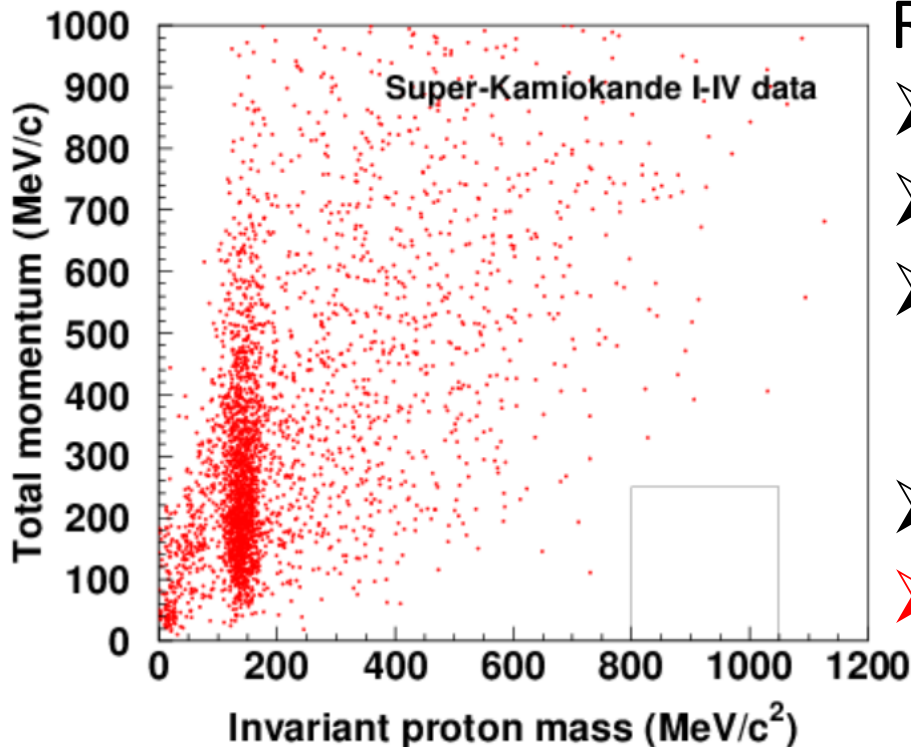
How powerful to reject background



- Sample: out of signal box in M_{tot} vs P_{tot} plot.
 - Dot: data,
 - **Histogram: Atm. ν MC**
(solid: reconstructed, dash: true)
- **$\sim 50\%$ background events are rejected** with neutron=0.
- On the other hand, $\sim 7.5\%$ of $p \rightarrow e^+ \pi^0$ are accompanied with neutron from deexcitation of nucleus. Neutron tagging **reduces a few % in selection efficiency.**

2-3. SK result (so far)

- We have not find any evidences of nucleon decays !



Result of $p \rightarrow e^+ \pi^0$

- Exposure: 450 kton · year
- Efficiency: 38.6 % (SK-IV)
- Expected BG: 0.63 events
 - 0.05 evetns in $P_{\text{tot}} < 100 \text{ MeV/c}$
 - 0.58 events in $100 \sim 250 \text{ MeV/c}$
- Observed: 0 event
- Lower limit of proton life time: $> 2.4 \times 10^{34}$ years

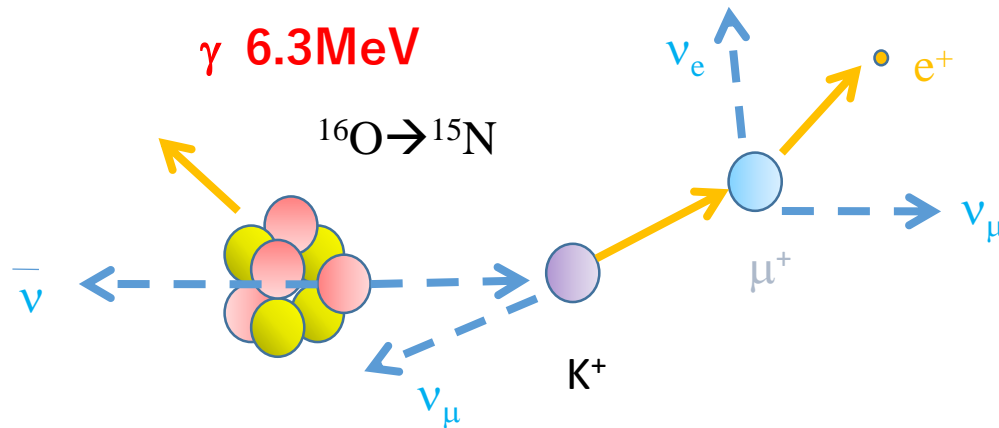
3. SUSY favored decay mode:

$$p \rightarrow \nu K^+$$

3-1 How to find $p \rightarrow \nu K^+$ in Water Cherenkov detector

- K^+ has low momentum, most of them **stop in water** and decay with 12 nsec lifetime.
- Major K^+ decay mode
 - $K^+ \rightarrow \nu \mu^+$: 64 %
 - $K^+ \rightarrow \pi^+ \pi^0$: 21 %
- “Stopping K^+ ” means **two body decay products of K^+ should have monochromatic momentum.**
 - $K^+ \rightarrow \nu \mu^+$: 236 MeV/c
 - $K^+ \rightarrow \pi^+ \pi^0$: 206 MeV/c
- Using this property, Water Cherenkov detector can search for $p \rightarrow \nu K^+$.

3-2. Search for $p \rightarrow \nu K^+$, $K^+ \rightarrow \nu \mu^+$



- Visible particle is only μ^+ with Michel electron.
- Search for data excess around 236 MeV/c of μ comparing with atmospheric ν MC.
- After proton decay, **40 % of remaining nucleus emits 6 MeV γ for deexcitation**. It is useful to reduce background.

Example of $p \rightarrow \nu K^+$, $K^+ \rightarrow \nu \mu^+$ with γ

Super-Kamiokande IV

Run 999999 Sub 0 Event 69

D_{wall}: 1165.1 cm

Evis: 53.2 MeV

mu-like, $p = 231.0$ MeV/c

Resid(ns)

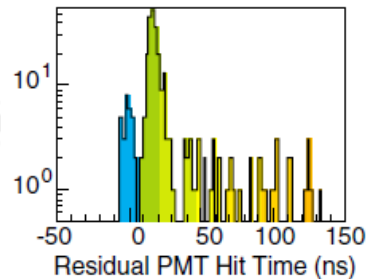
- > 182
- 160- 182
- 137- 160
- 114- 137
- 91- 114
- 68- 91
- 45- 68
- 22- 45
- 0- 22
- -22- 0
- -45- -22
- -68- -45
- -91- -68
- -114- -91
- -137- -114
- <-137

μ

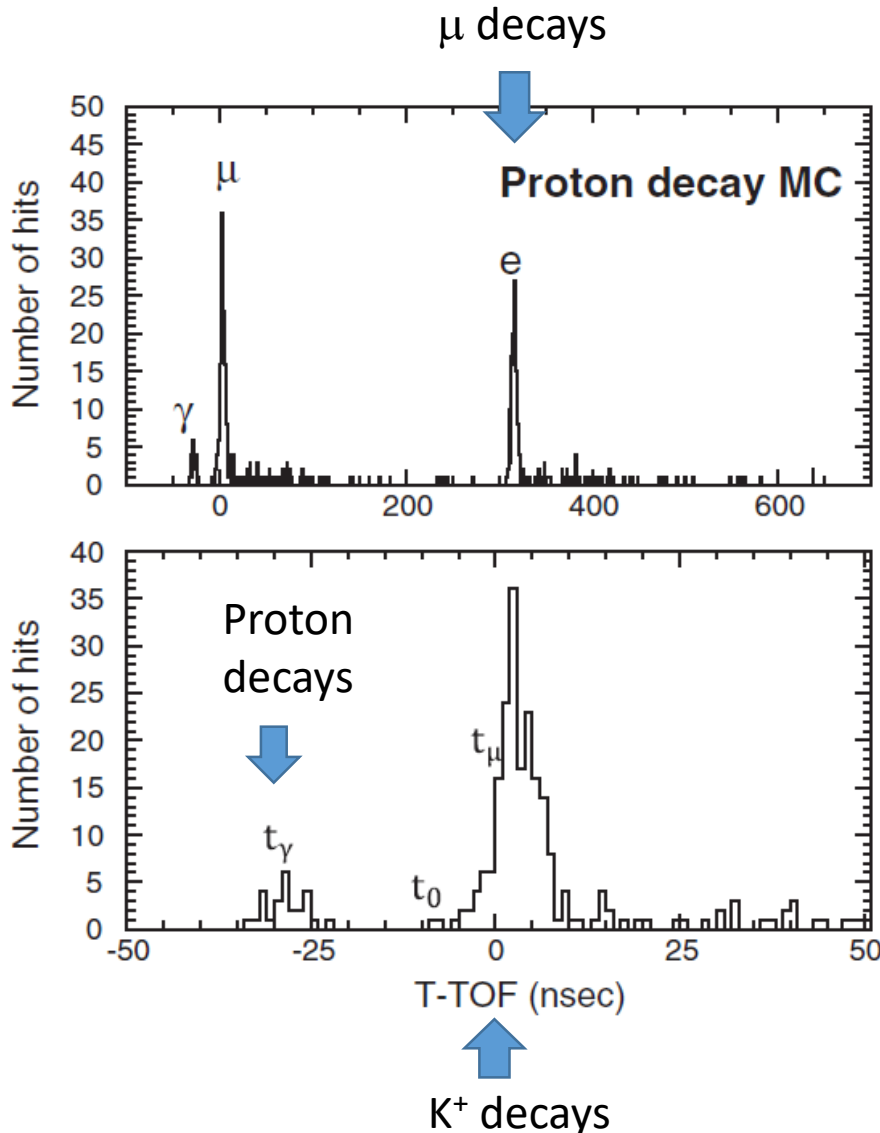
Color of hits corresponds to time.

Cyan corresponds to nuclear γ .

Difficult to identify γ from hit pattern.




Time structure with nuclear γ



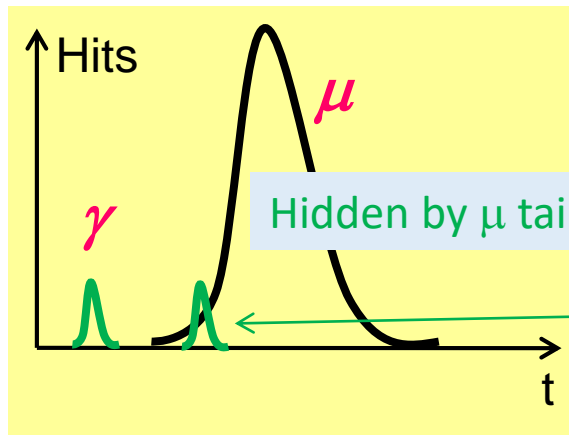
- 3 hit clusters in time should be observed in case of signal.
- The event is triggered by μ hits.
- γ signal is much smaller than μ and easily hidden by tail of μ hits.
- Make 12 nsec time window and slide it toward left from t_0 (end of μ tail) to search for maximum hit cluster.

Selection criteria for $p \rightarrow \nu K^+$, $K^+ \rightarrow \nu \mu^+$

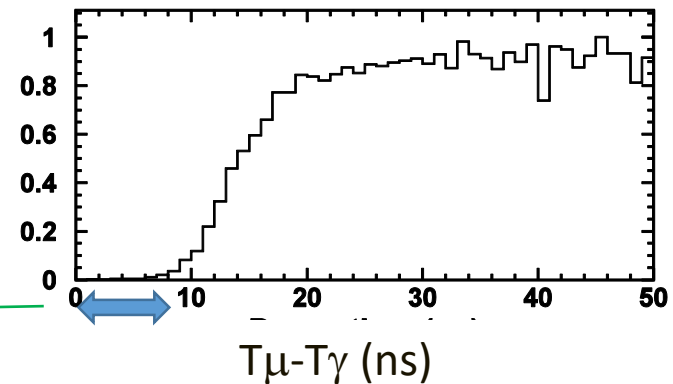
- 1 μ -like ring with Michel electron
- $215 < P_\mu < 260$ MeV/c
- Proton rejection cuts
- Search Max hit cluster  Reduce background by 5×10^{-4} !
by sliding time window (12ns width);
 - $4 < N_\gamma < 30$ hits
 - $T_\mu - T_\gamma < 75$ nsec
- No neutron
- Selection efficiency = (selected events)/(proton decay in fiducial volume):
9 %
 - $\text{Br}(K^+ \rightarrow \nu \mu^+) = 64 \%$, only 40 % emits nuclear $\gamma \rightarrow 26 \%$ even if detector is perfect.

Remark for this analysis

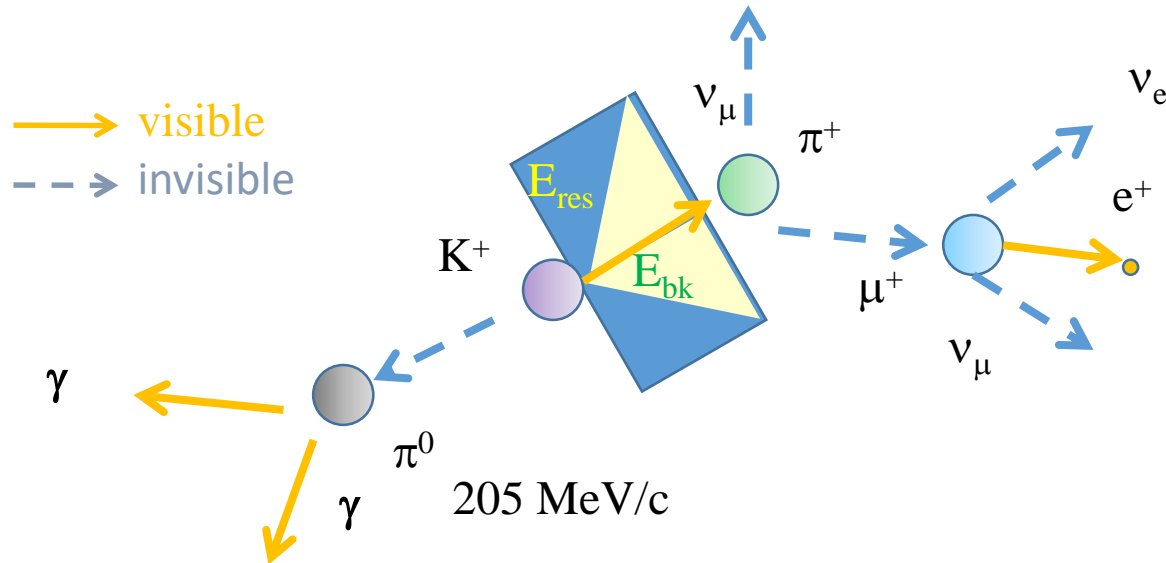
- This analysis is limited by time resolution of PMTs.
 - If γ is close to μ , γ peak is hidden by μ hits.
 - Time resolution of SK PMT is 2.2 nsec at 1 photoelectron.
 - If μ peak becomes sharper, the selection efficiency will be improved.



γ tagging efficiency



3-3. Search for $p \rightarrow \nu K^+$, $K^+ \rightarrow \pi^+ \pi^0$



- Both π^+ and π^0 has **205 MeV/c** in momentum. This is just above Cherenkov threshold for π^+ , thus it is not identified as a ring in most of case.
- π^+ decays into μ (invisible) and ν , μ decays into $e \nu_e \nu_\mu$.
- π^0 decays into 2 γ s.
- Search for 206 MeV/c π^0 with Michel electron.

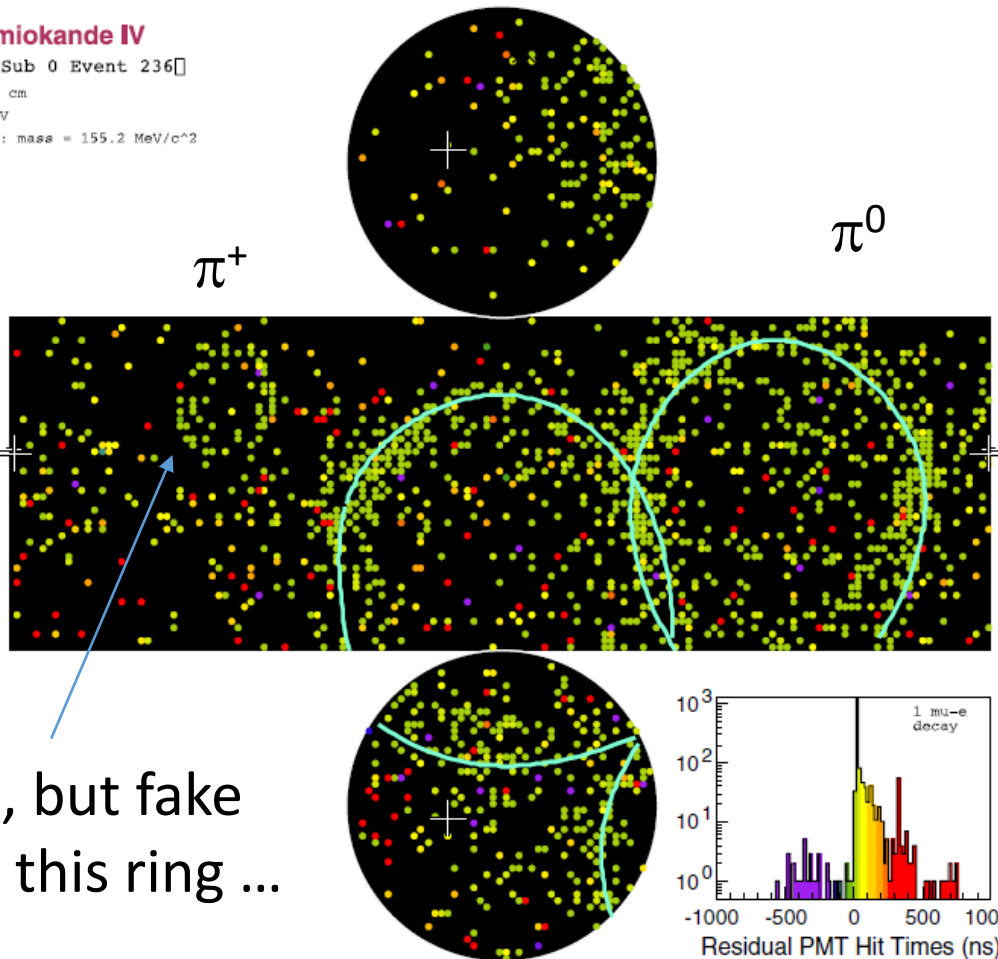
Example of $p \rightarrow \nu K^+$, $K^+ \rightarrow \pi^+ \pi^0$

Super-Kamiokande IV

Run 999999 Sub 0 Event 236
D_{wall}: 1076.4 cm
E_{vis}: 260.4 MeV
2 e-like rings: mass = 155.2 MeV/c²

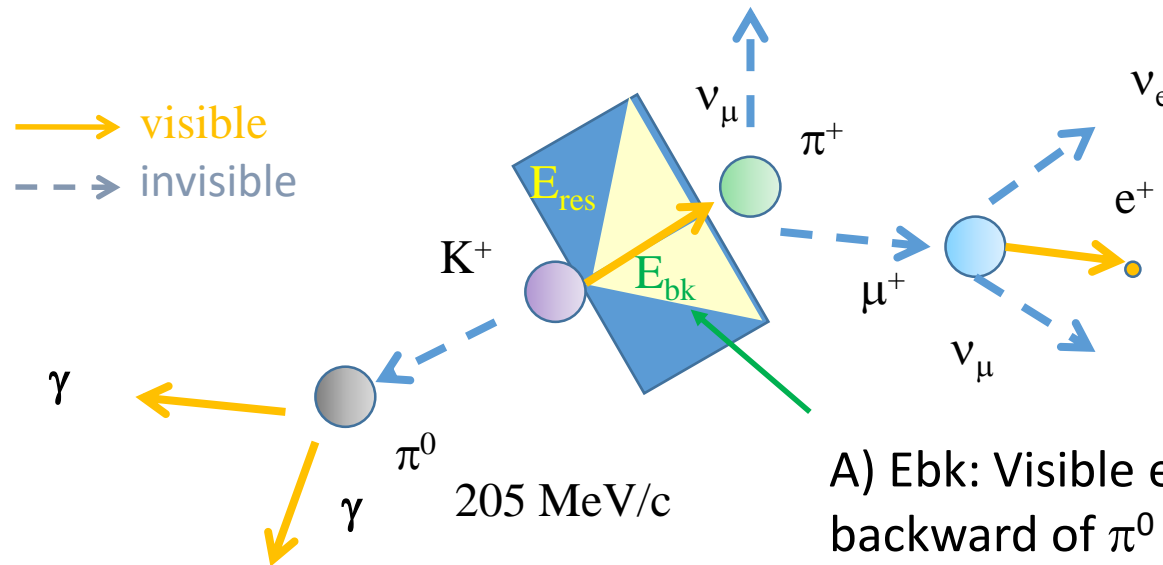
Resid(ns)

- > 251
- 220- 251
- 188- 220
- 157- 188
- 125- 157
- 94- 125
- 62- 94
- 31- 62
- 0- 31
- -31- 0
- -62- -31
- -94- -62
- -125- -94
- -157- -125
- -188- -157
- <-188

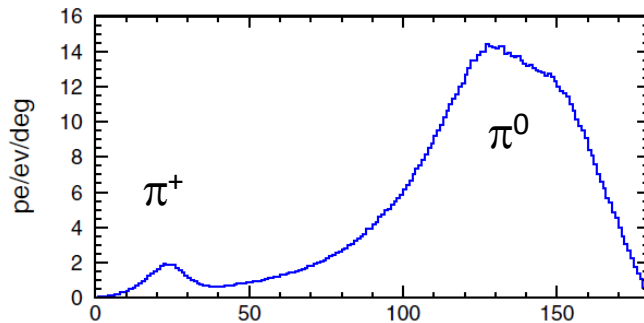


Look like a ring, but fake
ring cut rejects this ring ...

Use π^+ information to select events



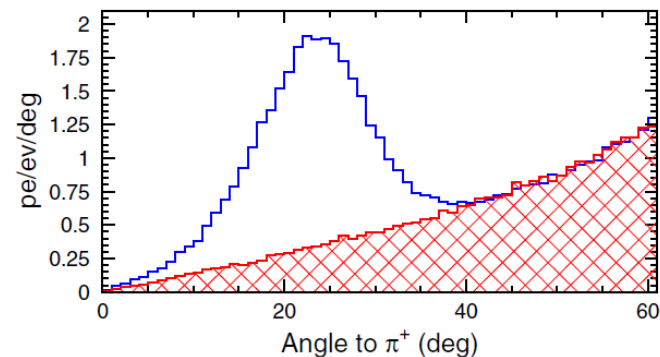
Charge distribution



π^+ direction

B) Make likelihood for hit pattern.

Zoom



Selection criteria for $p \rightarrow \nu K^+, K^+ \rightarrow \pi^+ \pi^0$

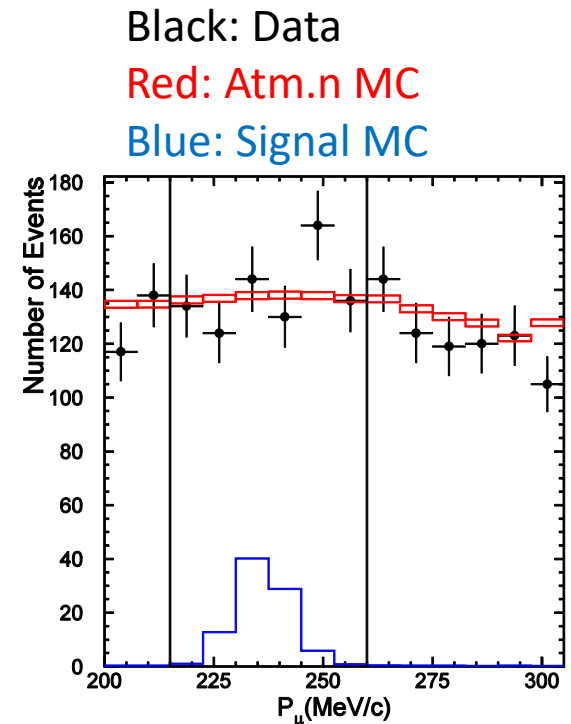
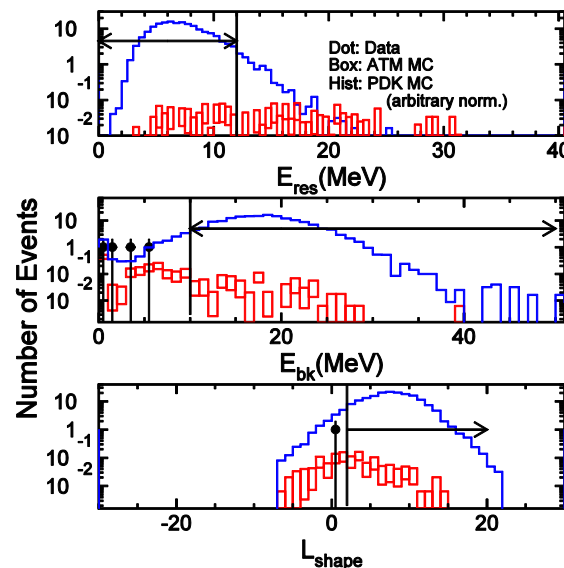
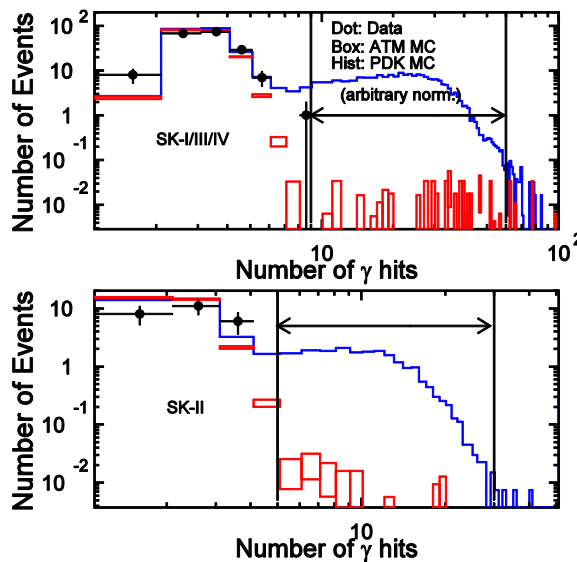
- 1 or 2 e-like rings with decay-e.
- $85 < M\pi^0 < 185$ MeV.
- $175 < P\pi^0 < 250$ MeV/c.
- E_{bk} : visible energy sum in 140-180 deg. of π^0 dir,
 E_{res} : in 90-140 deg,
 L_{shape} : Likelihood based on charge profile
 - $10 < E_{bk} < 50$ MeV
 - $E_{res} < 12$ MeV (20 MeV for 1ring)
 - $L_{shape} > 2.0$ (3.0 for 1ring)
- No neutrons
- Selection efficiency: 10 % ($Br(K^+ \rightarrow \pi^+ \pi^0) = 21$ %)

Background for $p \rightarrow \nu K^+$

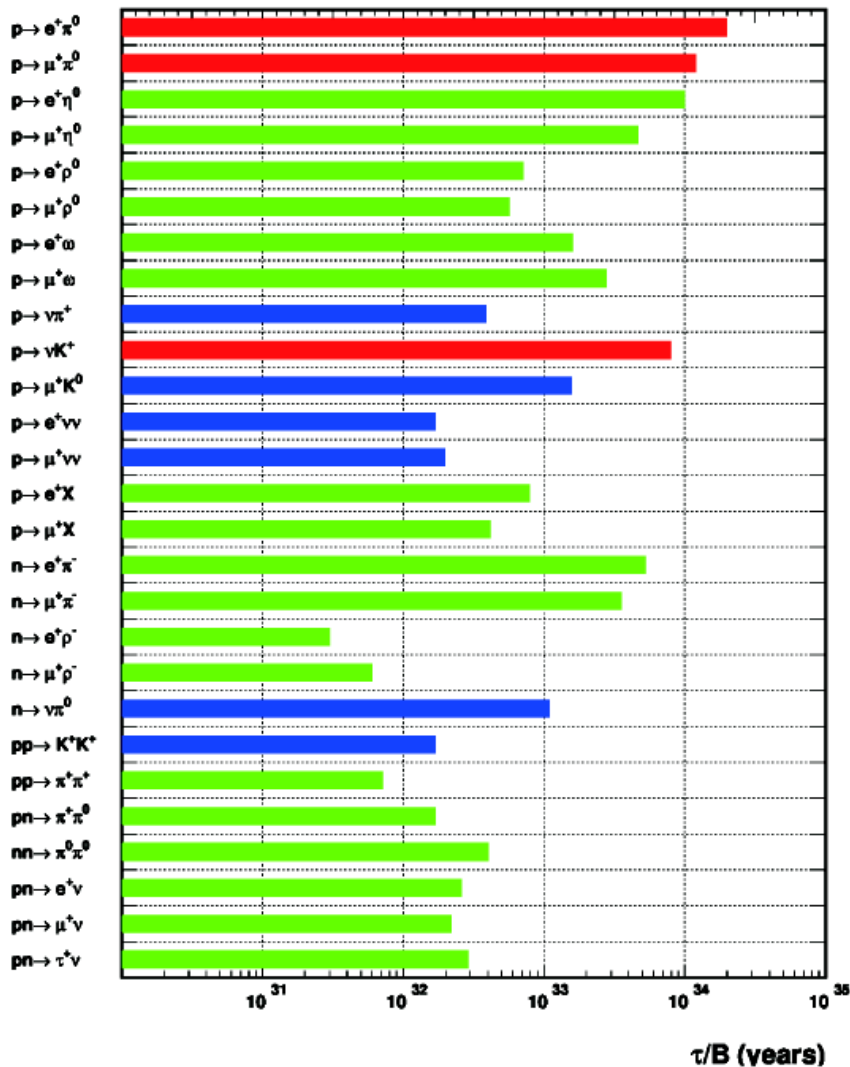
- Dominant background is K^+ production by neutrino interactions.
 - $\nu p \rightarrow \nu \Lambda K^+$, $\Lambda \rightarrow p \pi^-$ (BR:64 %, mostly invisible in WCD)
 - Emit nuclear γ as same as the signal.
- It is also rare interaction and we had poor information from very old bubble chamber. Large uncertainty.
- Recently MINERvA measures K^+ production. It is very useful information for this analysis.

3-4. SK results (So far)

- Exposure: 365 kton · year
- Expected background: 0.3 events for $K^+ \rightarrow \nu \mu$ with nuclear γ , 0.6 events for $K^+ \rightarrow \pi^+ \pi^0$.
- No candidates observed and no excess in momentum distribution.
- Lower lifetime limit: $> 0.8 \times 10^{34}$ year



3. Summary of SK results



- Most of modes have been investigated with > 0.3 Mton \cdot year exposure (red and green in the left figure).
- Super-Kamiokande can cover large number of decay modes.
- Many of them are the most stringent limits on nucleon lifetime.
- We observed some candidates, but still consistent with expected backgrounds and no evidence of nucleon decay has been observed.

Future prospects

- Still no evidence has been found. Major decay modes are explored up to around 10^{34} years.
- Proton lives longer, $\sim 10^{35}$ years ?
 - Run SK 10 times more (~ 200 years)? → Impossible.
- Absolutely, we need larger detector !

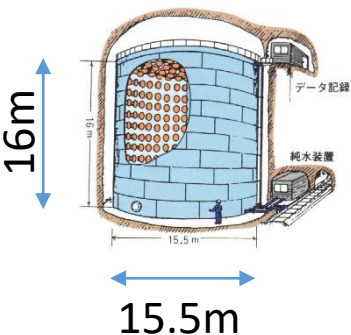
4. Hyper-Kamiokande project



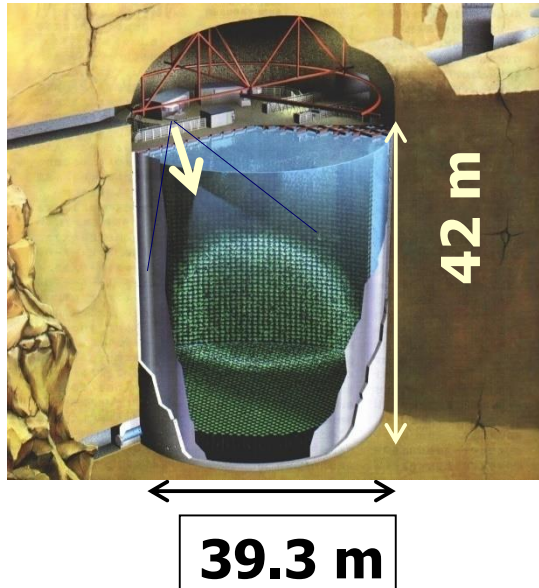
Neutrino oscillation



SN Neutrino



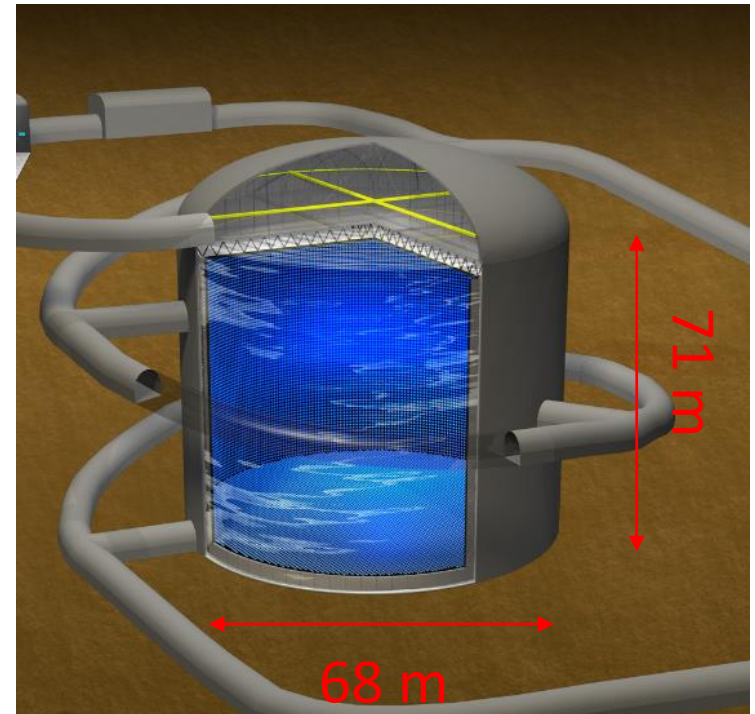
Kamiokande
3kton



Super-Kamiokande
50kton

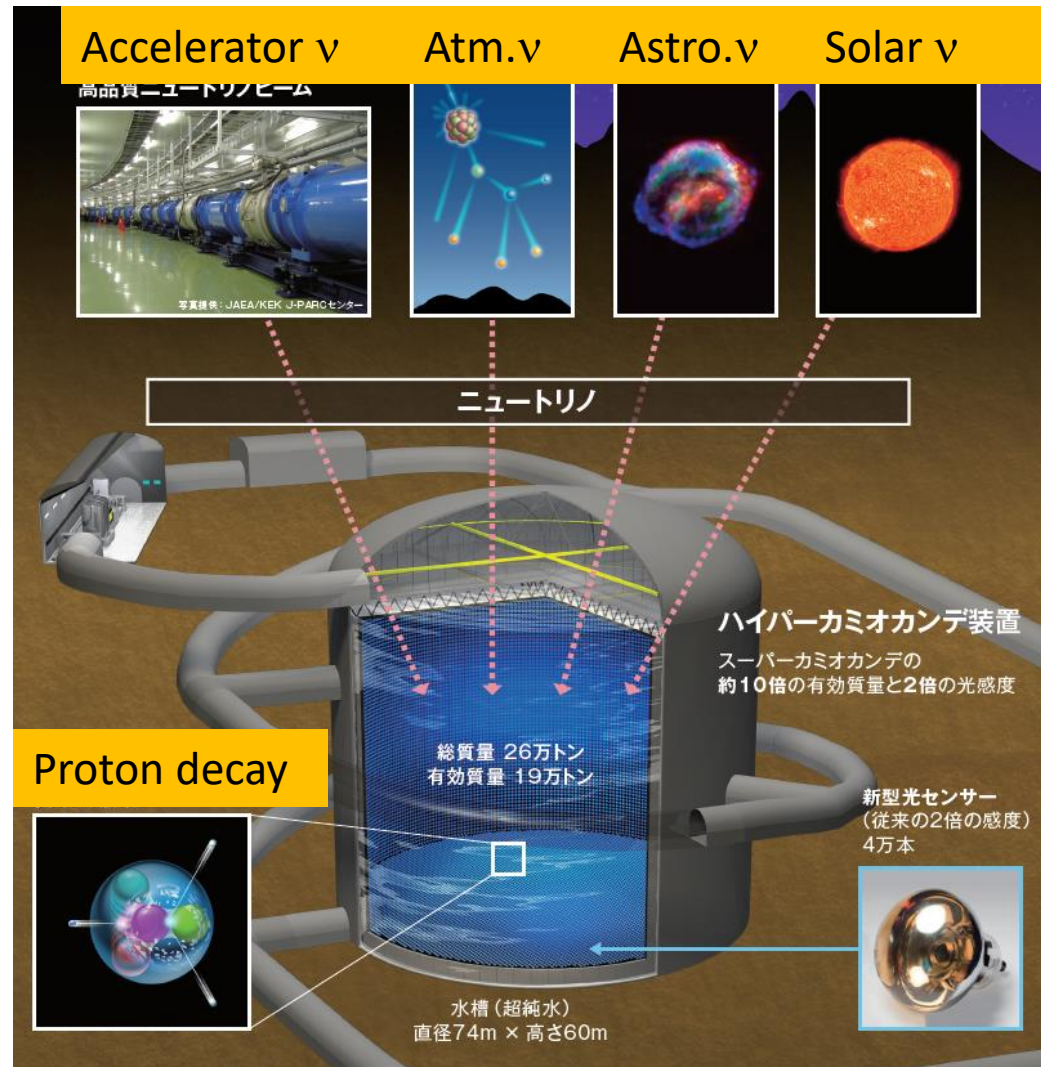


Proton decay ?



Hyper-Kamiokande
260kton

Hyper-K is multi-purpose detector



HYPER-K COLLABORATION

22 countries, 102 institutes, ~570 people as of July 2023, and growing

Collaborating Institutes



Europe	339 members
Armenia	3
Czech	7
France	33
Germany	1
Greece	4
Italy	59
Poland	45
Russia	23
Spain	48
Sweden	5
Switzerland	15
Ukraine	4
UK	87

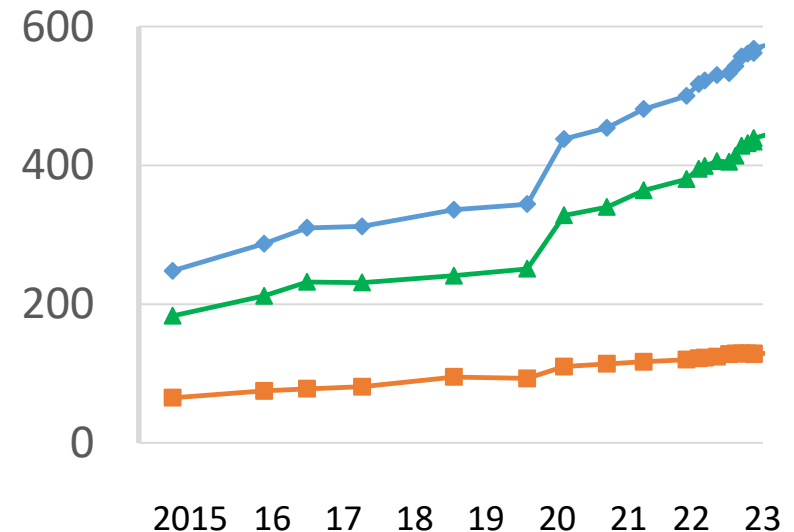
Oceania	160 members
Australia	5
India	10
Korea	16
Japan	129

Americas	62 members
Brazil	3
Canada	42
Mexico	9
USA	8

Africa	12 members
Morocco	12

NUMBER OF COLLABORATORS

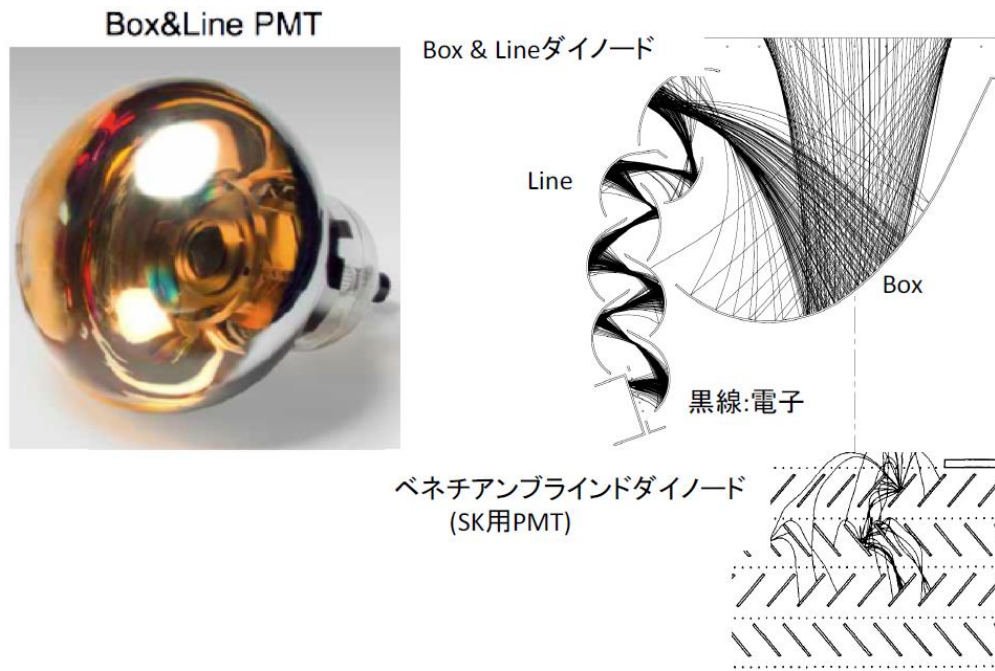
◆ Total ■ Japan ▲ Oversea



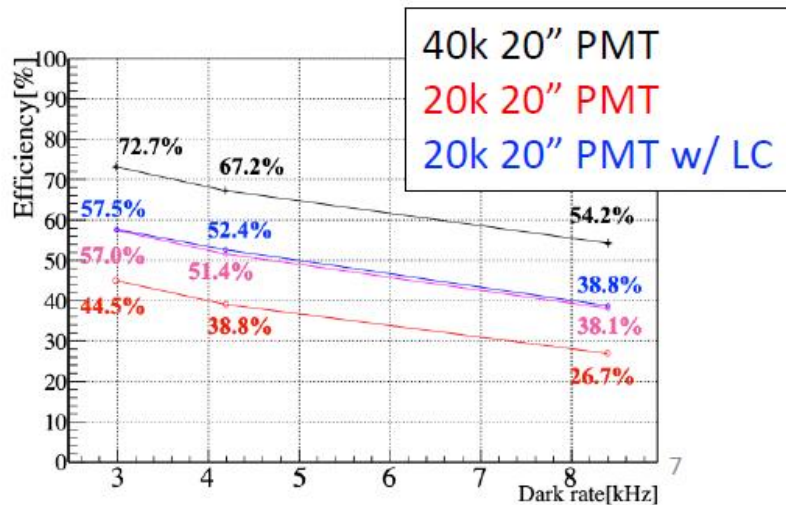
I wish Vietnam joins soon

Enhance proton decay search with HK

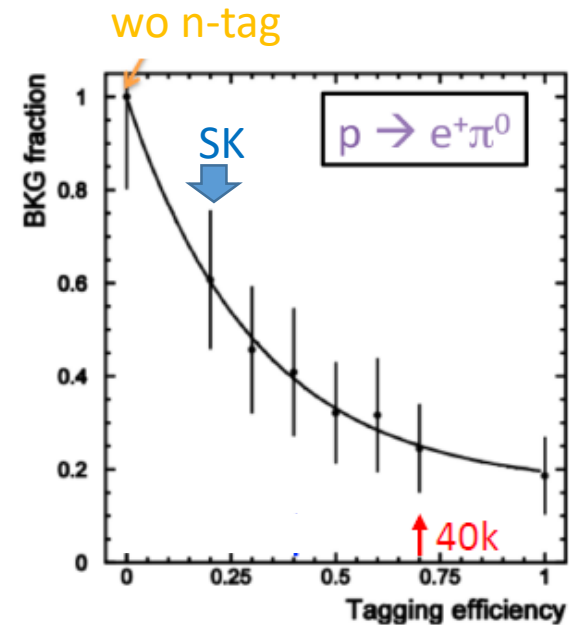
- Fiducial volume: 22.5kton (SK) → 190kton (HK)
- New photo sensor: Box&Line PMT
 - 2 times better photon counting performance
 - a half time resolution



Better photon counting contributes neutron tagging



- **Neutron tagging efficiency** study with several detector set up.
- Efficiency depends on dark rate.
- Achieve **~ 70%** in the current baseline design (black) with **~ 4kHz** dark rate.

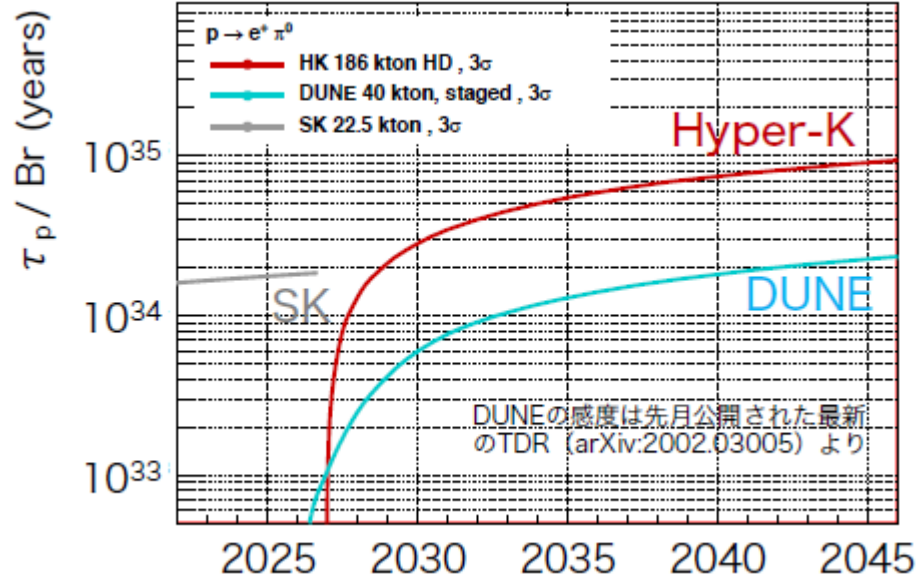
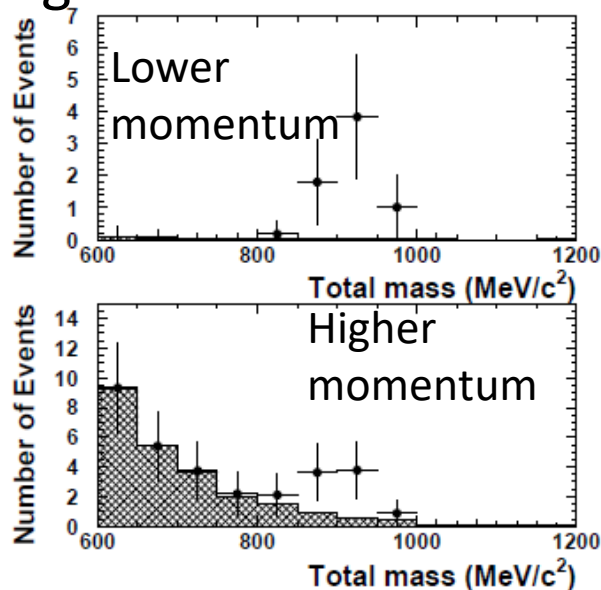


- $p \rightarrow e^+ \pi^0$ background reduction vs. Neutron tagging efficiency
- **Background of HK becomes a half of SK !**

Sensitivity for $p \rightarrow e^+ \pi^0$

Expected signal after 10 years run
assuming the current lifetime limit

3σ discovery potential



$0 < p_{tot} < 100 \text{ MeV}/c$		$100 < p_{tot} < 250 \text{ MeV}/c$	
$\epsilon_{sig} [\%]$	Bkg [/Mton·yr]	$\epsilon_{sig} [\%]$	Bkg [/Mton·yr]
18.7 ± 1.2	0.06 ± 0.02	19.4 ± 2.9	0.62 ± 0.20

(SK: 0.18)

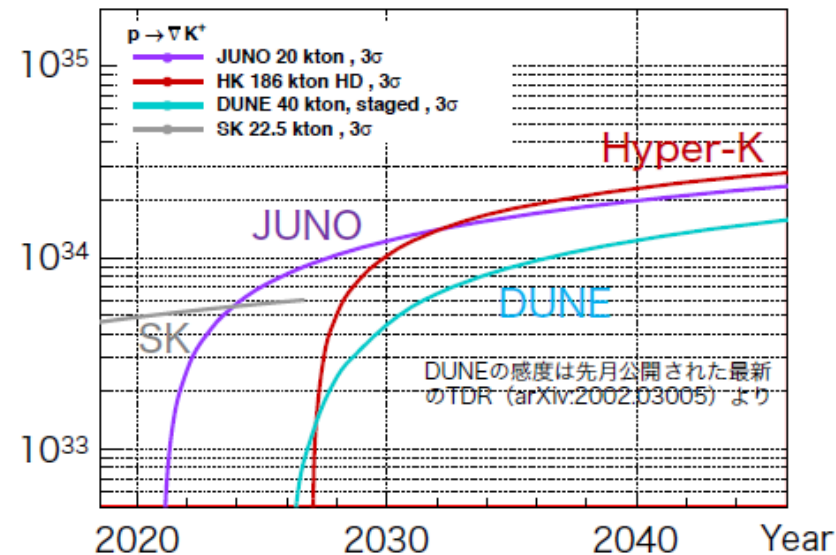
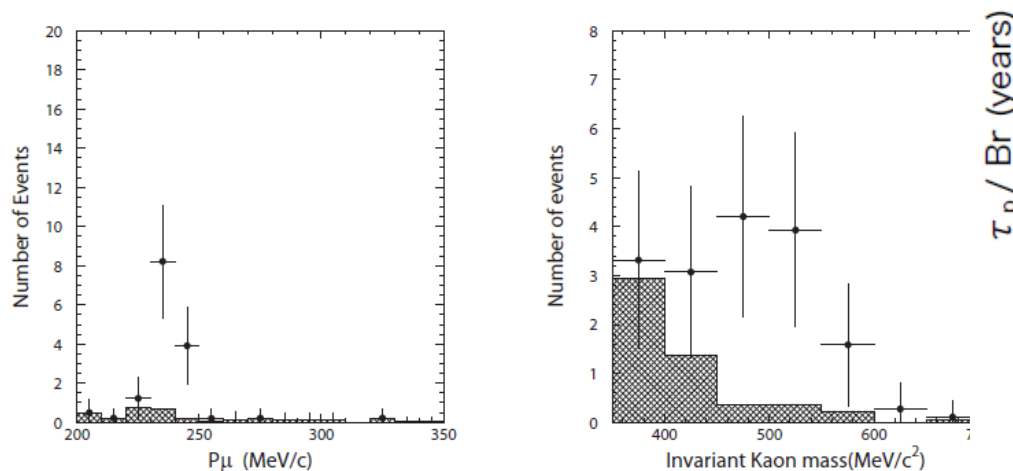
(SK: 1.1)

Reach to 10^{35} years !

Sensitivity for $p \rightarrow \bar{\nu} K^+$

Expected signal after 10 years run
assuming the current lifetime limit

3σ discovery potential



Prompt γ		$\pi^+ \pi^0$		p_μ Spectrum		
ϵ_{sig} [%]	Bkg [/Mton·yr]	ϵ_{sig} [%]	Bkg [/Mton·yr]	ϵ_{sig} [%]	Bkg [/Mton·yr]	σ_{fit} [%]
12.7 ± 2.4	0.9 ± 0.2	10.8 ± 1.1	0.7 ± 0.2	31.0	1916.0	8.0

Better sensitivity than
other detectors !.

What are still unknown in ν oscillation?

Unknown parameter

CP phase

$$c_{ij} = \cos \theta_{ij}, s_{ij} = \sin \theta_{ij}$$

$\begin{matrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{matrix}$	$\begin{pmatrix} n_e \\ n_\mu \\ n_\tau \end{pmatrix}$	$\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}$	$\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix}$	$\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$	$\begin{pmatrix} n_1 \\ n_2 \\ n_3 \end{pmatrix}$
------------------------------------------------------------	--------------------------------------------------------	------------------------------------------------------------------------------------------	-----------------------------------------------------------------------------------------------------------------	------------------------------------------------------------------------------------------	---------------------------------------------------

- $\theta_{12} = 33.6^\circ \pm 1.0^\circ$ Solar ν , KamLAND
- $\theta_{23} = 45^\circ \pm 6^\circ$ (90%CL) Atm. ν , Acc. ν

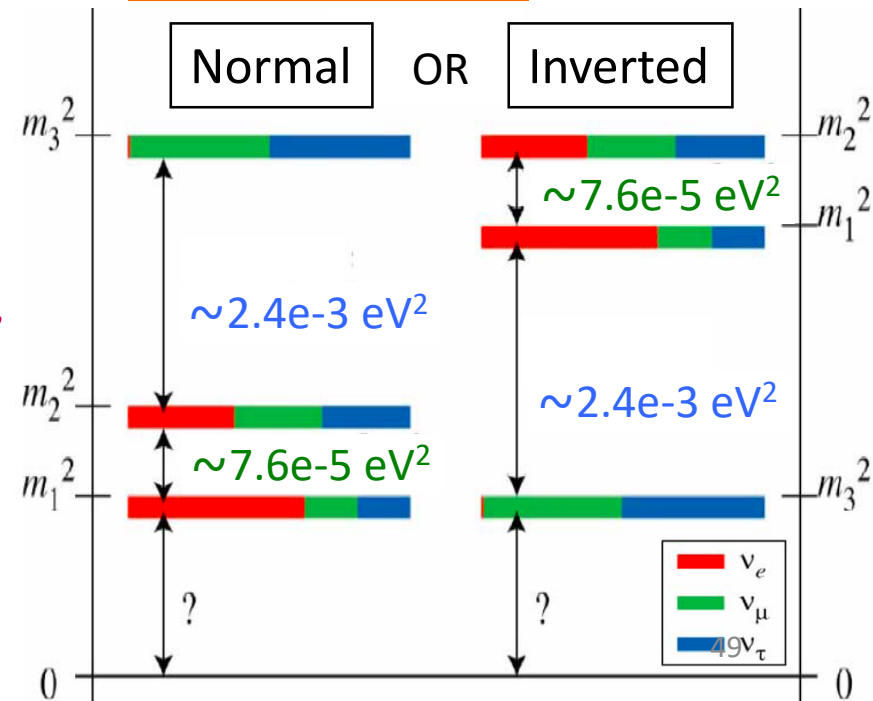
θ_{23} : How close to 45° ?
Octant? ($<45^\circ$, $>45^\circ$?)

- $\theta_{13} = 9.1^\circ \pm 0.6^\circ$ Reactor, Acc. ν

Indication of $\theta_{13} \neq 0$ by T2K
PRL107, 041801 (2011)

Later precise measurements
by reactor ν experiments

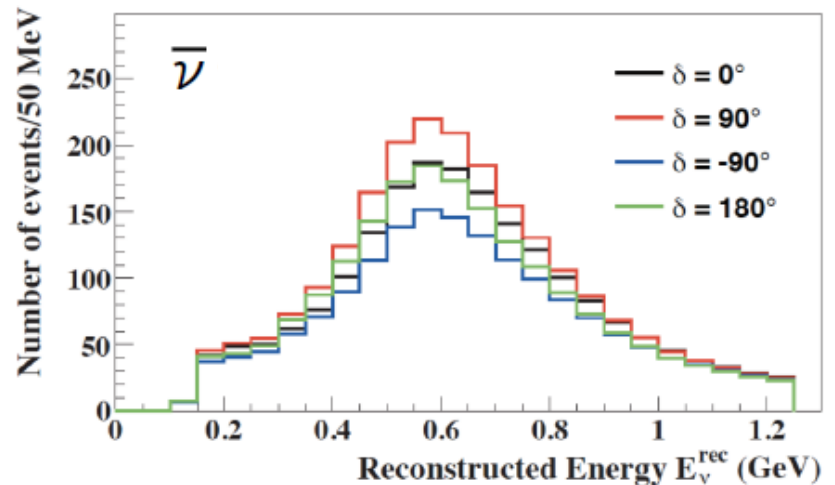
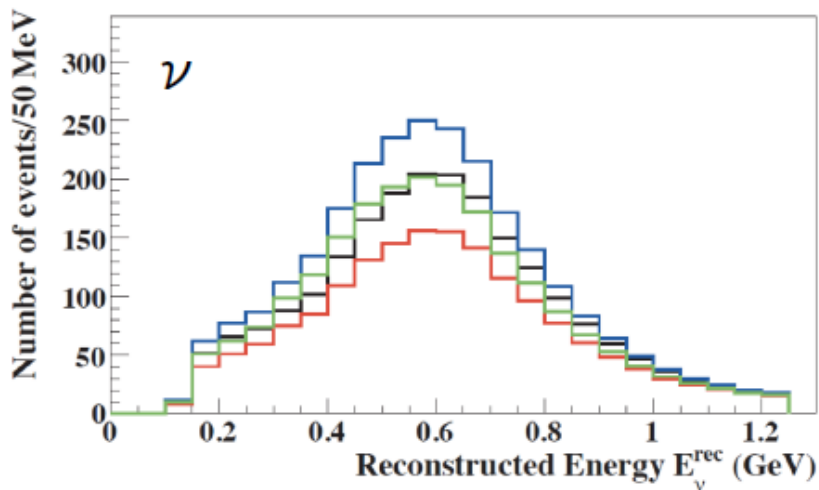
Mass hierarchy



Effects of unknown parameters on ν oscillation are small \rightarrow Need statistics = larger detector !

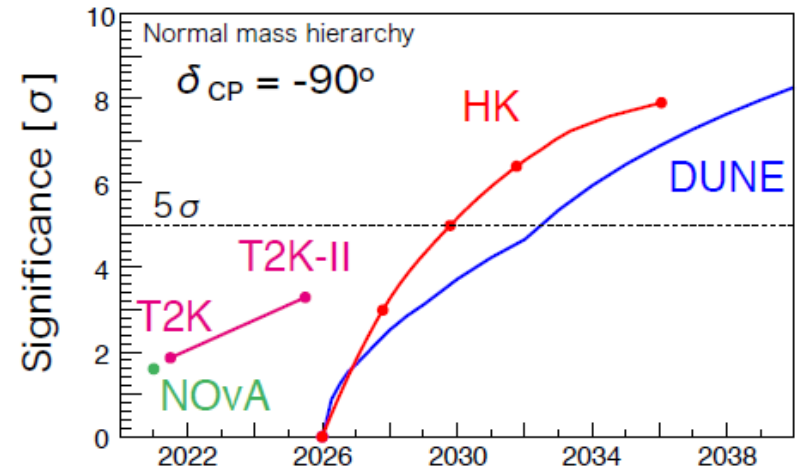
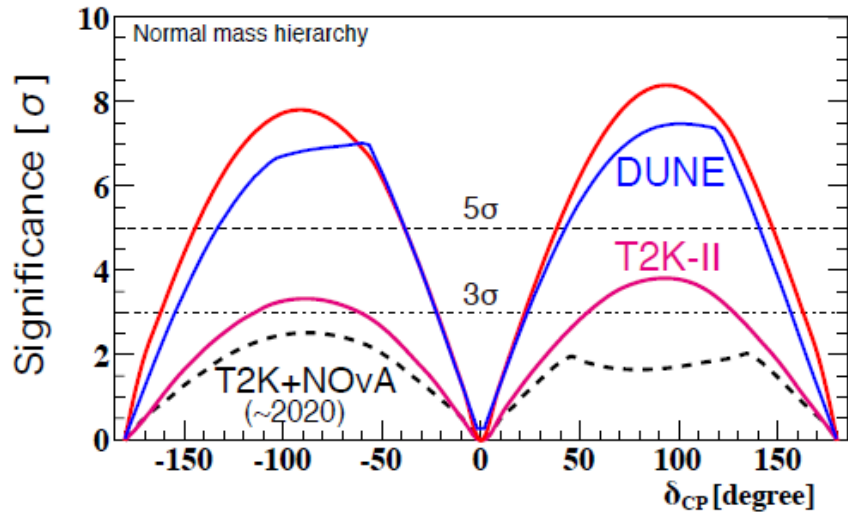
- ν beam experiments: Can study CP phase by comparing ν $\bar{\nu}$ oscillation.

Expected ν_e spectrum at HK (assuming 1.3 MW x 10 years)



Sensitivity of CP violation

HK 10 years



- In 60 % region in δ_{CP} , we can discovery δ_{CP} with 5σ .
- If $\delta_{CP} = \pm 90$ degree, we can discover it within 5 years.

Atmospheric ν : $\nu_\mu \rightarrow \nu_e$ enhancement by matter effect

ν_e oscillation due to non-zero θ_{13} provides atm. nu. observation to investigate mass hierarch effect

$\nu_\mu \rightarrow \nu_e$ osc. probability in matter:

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 \theta_{23} \sin^2 2\theta_{13}^M \sin^2 \left(\frac{1.27 \Delta m_{31,M}^2 L}{E} \right)$$

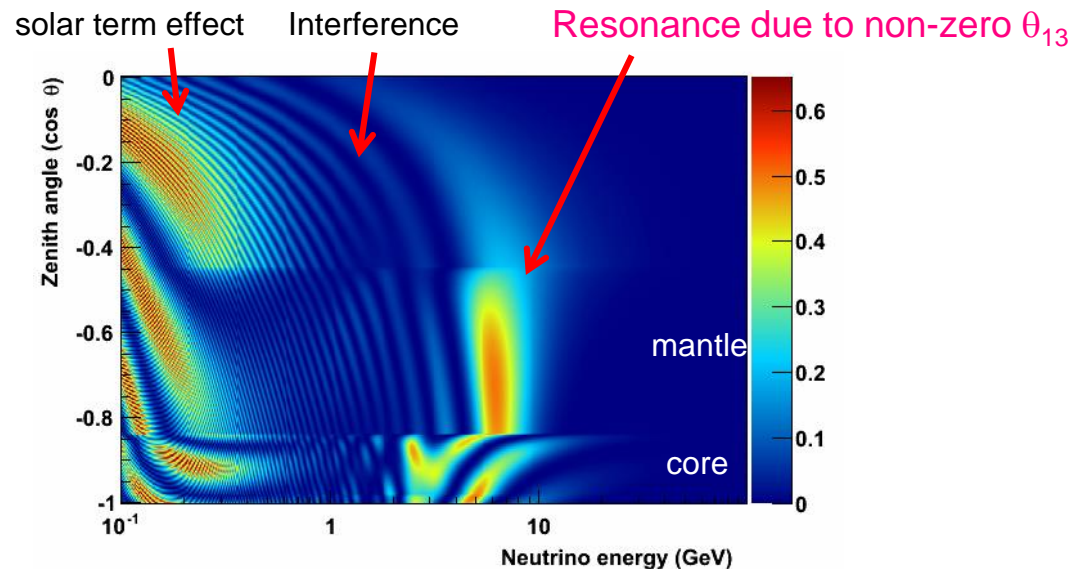
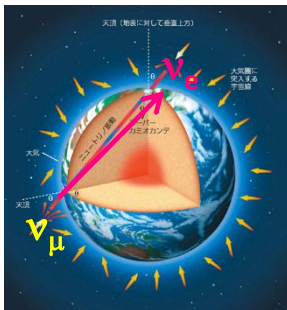
$$\sin^2 2\theta_{13}^M = \frac{\sin^2 2\theta_{13}}{\left(\cos 2\theta_{13} - \frac{A_{CC}}{\Delta m_{31}^2} \right) + \sin^2 2\theta_{13}}$$

$\sin^2 \theta_{13,M}$ has resonance feature when

$$A_{CC} \sim \Delta m_{31}^2 \cos 2\theta_{13} \quad (A_{CC} = 2\sqrt{2}G_F N_e E)$$

$\nu_\mu \rightarrow \nu_e$ resonance in multi-GeV region

$P(\nu_\mu \rightarrow \nu_e)$ of earth-throughgoing ν :



Occurrence of resonance feature depends on neutrino type and mass hierarchy:

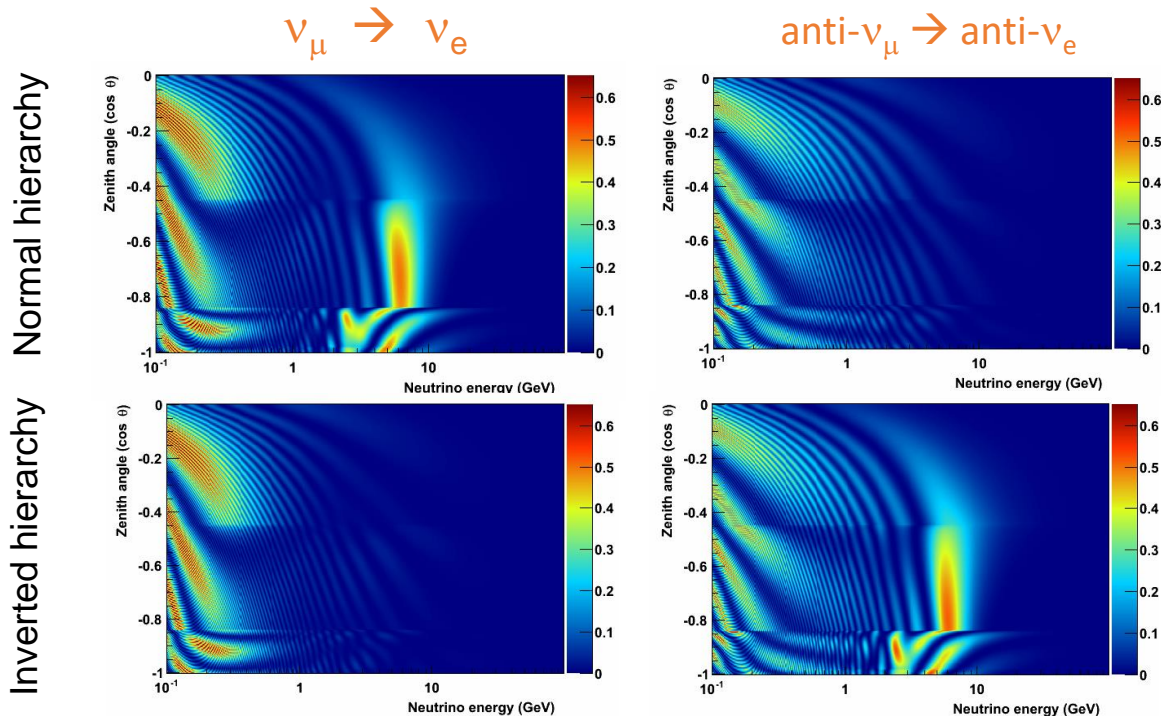
$$\sin^2 2\theta_{13}^M = \frac{\sin^2 2\theta_{13}}{\left(\cos 2\theta_{13} - A_{CC} / \Delta m_{31}^2 \right) + \sin^2 2\theta_{13}}$$

In case of Inverted hierarchy:

$$\Delta m^2 \rightarrow -\Delta m^2$$

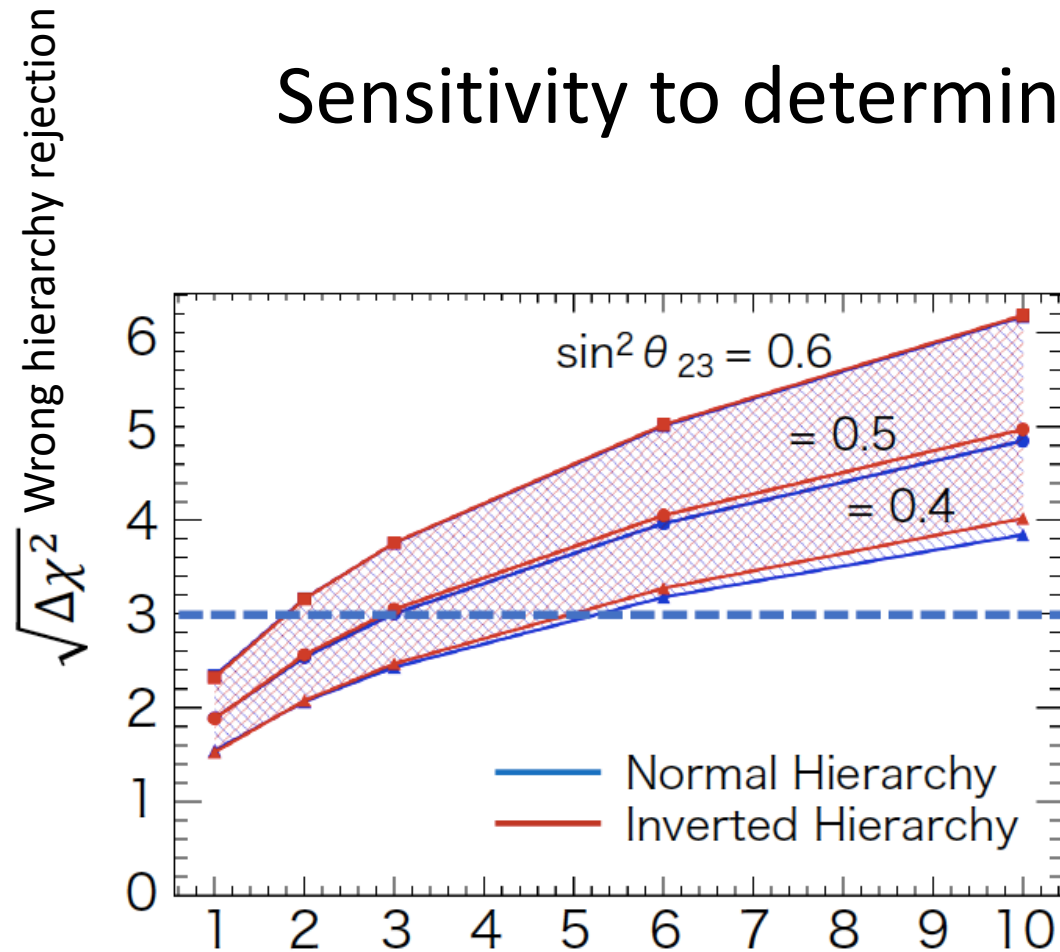
In case of anti neutrino :

$$A_{CC} \rightarrow -A_{CC}$$



Possible to determine mass hierarchy !

Sensitivity to determine mass hierarchy



Beam ν : sensitive to δCP , but weak in mass hierarchy.

Atm. ν : large uncertainty from δCP .

→ Combining both analysis gives good sensitivity to mass hierarchy.

3σ determine within 2 ~ 5 years !

Construction of HyperK has been started !



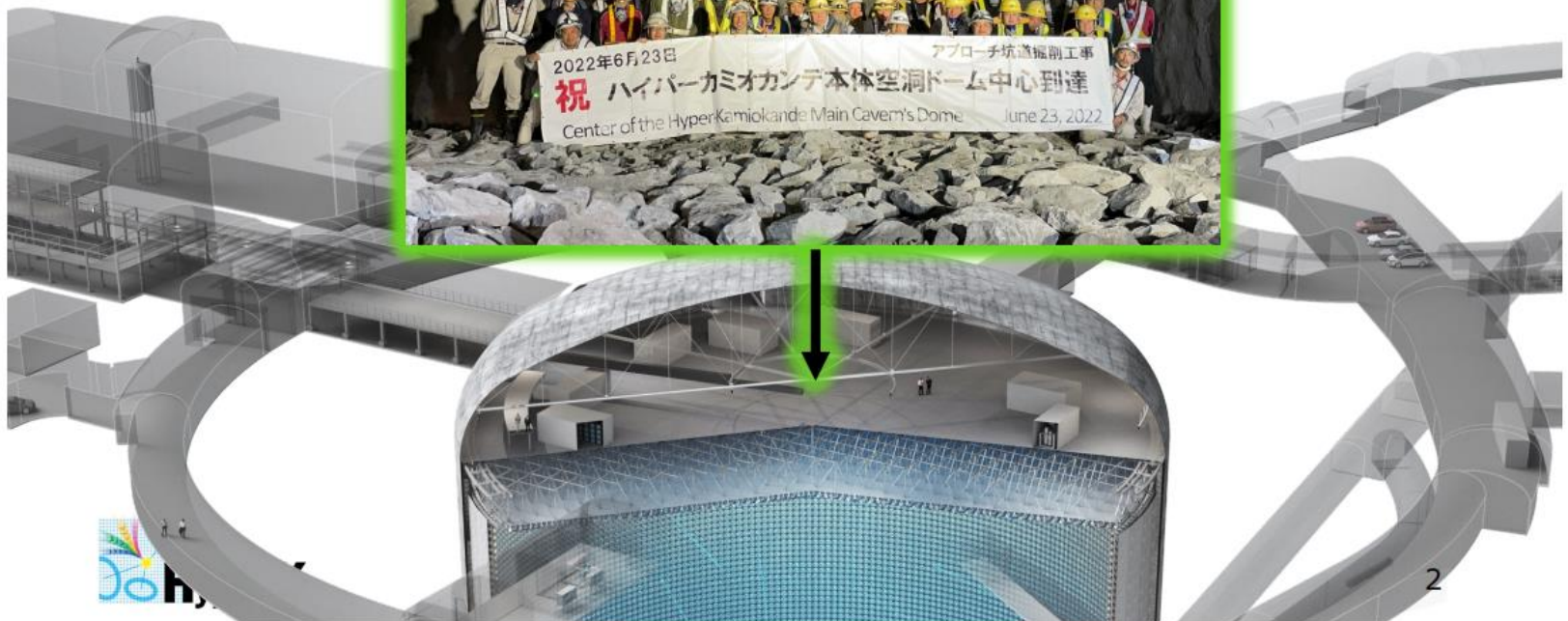
Making entrance yard for tunnel.



Aim to start observation in 2027 !

Access tunnel has reached to the center of
HyperK tank !

Detector Cavern is becoming a reality

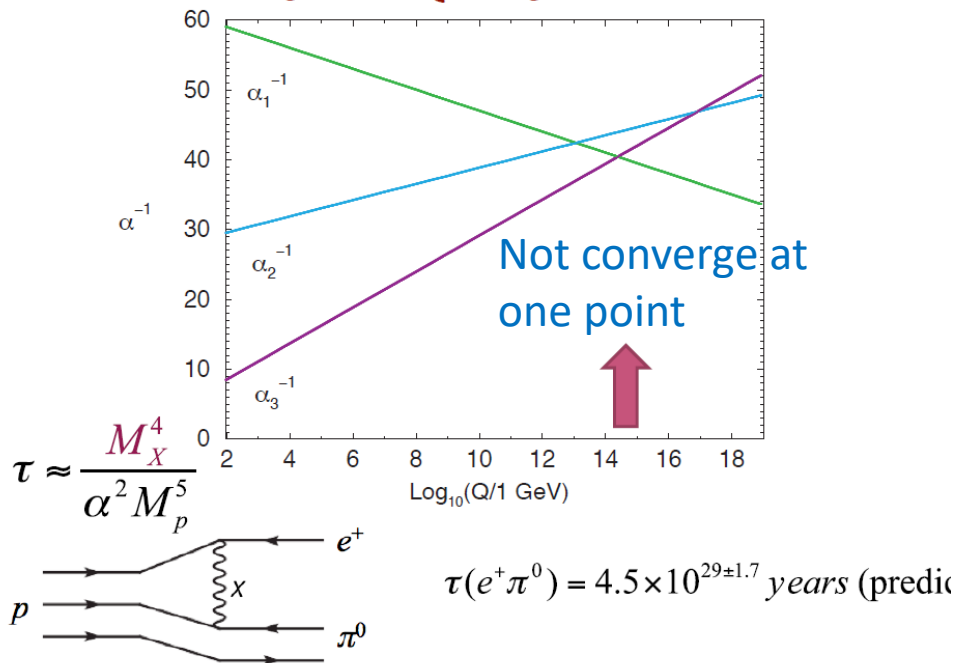


4. Summary

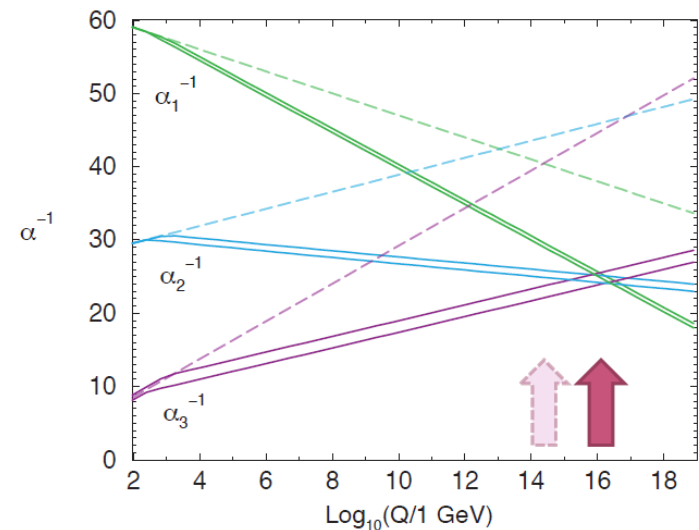
- Proton decay is a key phenomena of Grand Unified Theories beyond the Standard Model.
- Super-Kamiokande is the leading detector to hunt proton decays and have searched for it for more than 20 years.
- However, no evidence has been observed and the current proton lifetime limits are around 10^{34} years.
- It may be around the corner ! Hope three times lucky (3度目の正直 in Japanese) in Hyper-Kamiokande.
- HK also can determine remaining ν oscillation parameters .

Backup

Gauge Coupling Unification



Problems solved by SUSY ...



Unification scale pushed up...

$$\tau(e^+ \pi^0) \approx 10^{35-38} \text{ years}$$

SUSY GUTs

- Unification scale higher than non-SUSY-GUTs ($M_x \sim 2 \times 10^{16}$ GeV)
suppression of gauge boson mediated decay

$$\tau/B (p \rightarrow e^+ \pi^0) \approx \left(\frac{M_x}{2 \times 10^{16} \text{ GeV}} \right)^4 \times 10^{36 \pm 1} \text{ years}$$

dominated by the D=5 op. (color Higgs triplet, $q=1/3$) mediated decays

$$p \left\{ \begin{array}{l} u \xrightarrow{\tilde{\mu}} \bar{\nu} \\ d \xrightarrow{\tilde{\phi}_3} \bar{s} \\ u \xrightarrow{\tilde{c}} u \end{array} \right\} K^+ \quad M_3 \sim 3 \times 10^{16} \text{ GeV}, M_{\text{susy}} \sim 1 \text{ TeV}$$

$$\tau/B (p \rightarrow K^+ \bar{\nu}) \approx \left(\frac{0.003}{\beta_H} \right)^2 \left(\frac{M_{\tilde{q}}}{1 \text{ TeV}} \right)^2 \left(\frac{M_3}{10^{16} \text{ GeV}} \right)^2 \left(\frac{2}{\tan \beta} \right)^2 \dots$$

x 9×10^{31} years

(β_H : hadronic matrix element in GeV^3)

\Rightarrow highly model dependent

$p \rightarrow \bar{\nu} K^+$ is regarded as dominant mode in SUSY-GUTs.

Many Other GUTs Beyond This Simple Story

Model	Ref.	Modes	τ_N (years)
Minimal $SU(5)$	Georgi, Glashow [2]	$p \rightarrow e^+ \pi^0$	$10^{30} - 10^{31}$
Minimal SUSY $SU(5)$	Dimopoulos, Georgi [11], Sakai [12] Lifetime Calculations: Hisano, Murayama, Yanagida [13]	$p \rightarrow \bar{\nu} K^+$ $n \rightarrow \bar{\nu} K^0$	$10^{28} - 10^{32}$
SUGRA $SU(5)$	Nath, Arnowitt [14, 15]	$p \rightarrow \bar{\nu} K^+$	$10^{32} - 10^{34}$
SUSY $SO(10)$ with anomalous flavor $U(1)$	Shafi, Tavartkiladze [16]	$p \rightarrow \bar{\nu} K^+$ $n \rightarrow \bar{\nu} K^0$ $p \rightarrow \mu^+ K^0$	$10^{32} - 10^{35}$
SUSY $SO(10)$ MSSM (std. $d = 5$)	Lucas, Raby [17], Pati [18]	$p \rightarrow \bar{\nu} K^+$ $n \rightarrow \bar{\nu} K^0$	$10^{33} - 10^{34}$ $10^{32} - 10^{33}$
SUSY $SO(10)$ ESSM (std. $d = 5$)	Pati [18]	$p \rightarrow \bar{\nu} K^+$	$10^{33} - 10^{34}$ $\lesssim 10^{35}$
SUSY $SO(10)/G(224)$ MSSM or ESSM (new $d = 5$)	Babu, Pati, Wilczek [19, 20, 21], Pati [18]	$p \rightarrow \bar{\nu} K^+$ $p \rightarrow \mu^+ K^0$	$\lesssim 2 \cdot 10^{34}$ $B \sim (1 - 50)\%$
SUSY $SU(5)$ or $SO(10)$ MSSM ($d = 6$)	Pati [18]	$p \rightarrow e^+ \pi^0$	$\sim 10^{34.9 \pm 1}$
Flipped $SU(5)$ in CMSSM	Ellis, Nanopoulos and Wlaker[22]	$p \rightarrow e/\mu^+ \pi^0$	$10^{35} - 10^{36}$
Split $SU(5)$ SUSY	Arkani-Hamed, <i>et. al.</i> [23]	$p \rightarrow e^+ \pi^0$	$10^{35} - 10^{37}$
$SU(5)$ in 5 dimensions	Hebecker, March-Russell[24]	$p \rightarrow \mu^+ K^0$ $p \rightarrow e^+ \pi^0$	$10^{34} - 10^{35}$
$SU(5)$ in 5 dimensions option II	Alciati <i>et.al.</i> [25]	$p \rightarrow \bar{\nu} K^+$	$10^{36} - 10^{39}$
GUT-like models from Type IIA string with D6-branes	Klebanov, Witten[26]	$p \rightarrow e^+ \pi^0$	$\sim 10^{36}$

Uncertainties in the predictions:

Nuclear matrix elements updated w. IQCD, still: x10 uncertainty in lifetime

SUSY masses: ~ x100 uncertainty in lifetime

Proton life time:
 $10^{30} \sim 10^{35}$ years

TABLE I: Summary of the expected nucleon lifetime in different theoretical models.

Modes beyond $e^+\pi^0, K^+\nu$ and other antilepton + meson decays

$$p \rightarrow \mu^- \pi^+ K^+$$

$$B + L$$

$$n \rightarrow \bar{n}$$

$$\Delta B = 2, \text{TeV} < \text{scale} < \text{GUT}$$

$$pp \rightarrow K^+ K^+$$

$$\lambda''_{uds} < 10^{-8}$$

$$p \rightarrow e^- \pi^+ \pi^+ \nu \nu$$

6 dimensions

$$n \rightarrow \nu \nu \nu$$

invisible

$$p \rightarrow e^+ \gamma$$

radiative

there is plenty to keep us busy ...

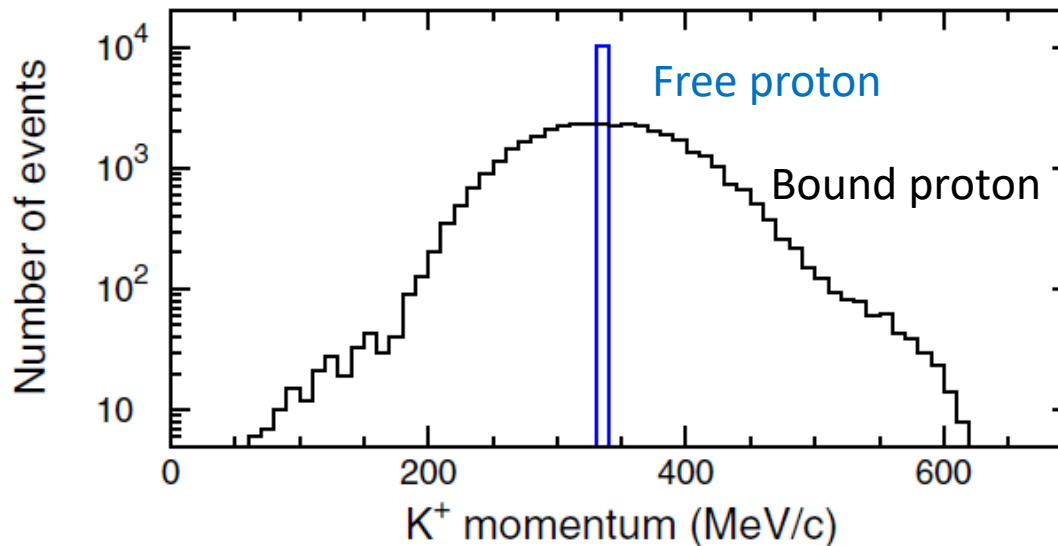
How to find 2.2 MeV γ



- After Time-of-Flight subtraction, search for 7 hits in 10 nsec time window. \rightarrow candidates of γ .
- Make 16 variables related to space and time information of each hits (RMS of phi, theta, hit time, e.t.c.)
- Put them into Neural Network to judge γ or not.
- **Neutron tagging efficiency: 21 % (mis-tagging: 1.8 %)**

3. $p \rightarrow \nu K^+$ search

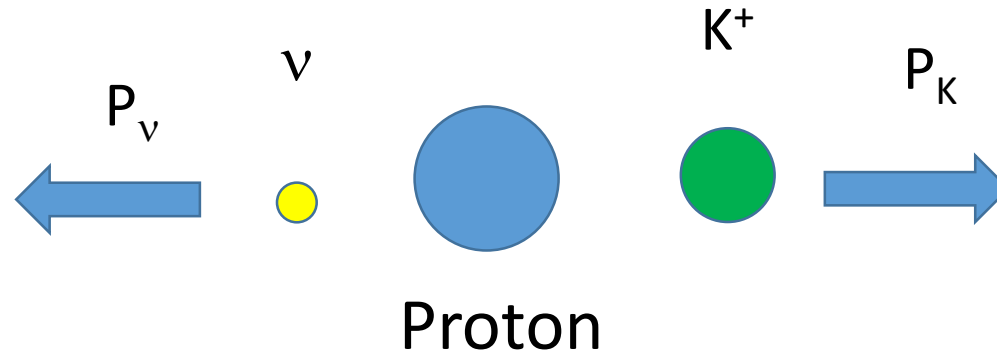
Difficulty of $p \rightarrow \nu K^+$



- K^+ mass: 494 MeV, relatively heavy.
- Cherenkov threshold: 560 MeV/c.
- Most of K^+ can not emit Cherenkov light.

Q. Calculate momentum of K^+ from free proton decay.

Hint: proton mass: 938 MeV, “free” means proton momentum=0.



Conservation of momentum: $P_\nu = P_K$

Conservation of energy: $M_P = \sqrt{M_K^2 + P_K^2} + P_\nu$

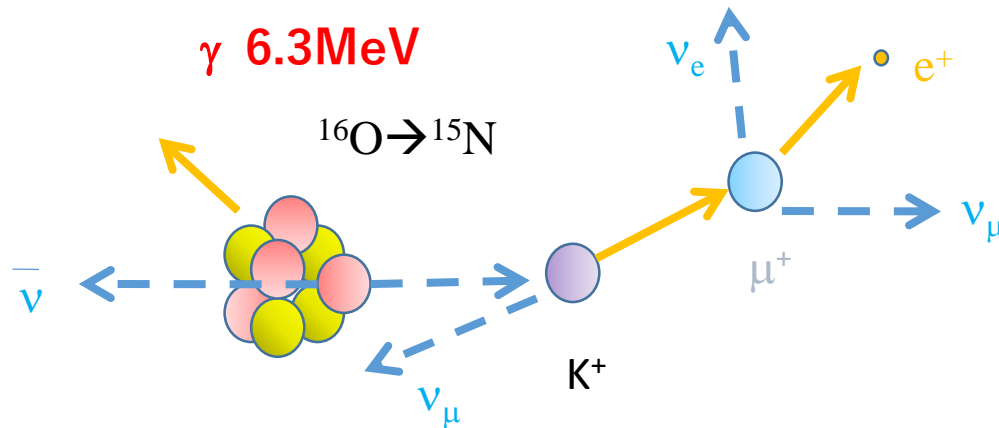
Proton Electron Neutrino

$$P_K = \frac{M_P^2 - M_K^2}{2M_P}$$

3-1 How to find $p \rightarrow \nu K^+$ in Water Cherenkov detector

- K^+ has low momentum, most of them **stop in water** and decay with 12 nsec lifetime.
- Major K^+ decay mode
 - $K^+ \rightarrow \nu \mu^+ : 64 \%$
 - $K^+ \rightarrow \pi^+ \pi^0 : 21 \%$
- “Stopping K^+ ” means **two body decay products of K^+ should have monochromatic momentum** as seen in the previous question !
 - $K^+ \rightarrow \nu \mu^+ : 236 \text{ MeV}/c$
 - $K^+ \rightarrow \pi^+ \pi^0 : 206 \text{ MeV}/c$
- Using this property, Water Cherenkov detector can search for $p \rightarrow \nu K^+$.

3-2. Search for $p \rightarrow \nu K^+$, $K^+ \rightarrow \nu \mu^+$



- Visible particle is only μ^+ with Michel electron.
- Search for data excess around 236 MeV/c of μ comparing with atmospheric ν MC.
- After proton decay, **40 % of remaining nucleus emits 6 MeV γ for deexcitation**. It is useful to reduce background.

Example of $p \rightarrow \nu K^+$, $K^+ \rightarrow \nu \mu^+$ with γ

Super-Kamiokande IV

Run 999999 Sub 0 Event 69

D_{wall}: 1165.1 cm

Evis: 53.2 MeV

mu-like, $p = 231.0$ MeV/c

Resid(ns)

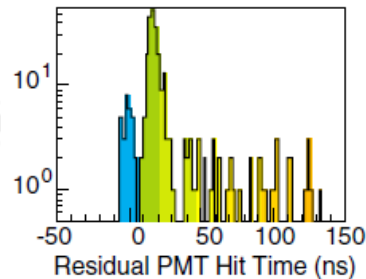
- > 182
- 160- 182
- 137- 160
- 114- 137
- 91- 114
- 68- 91
- 45- 68
- 22- 45
- 0- 22
- -22- 0
- -45- -22
- -68- -45
- -91- -68
- -114- -91
- -137- -114
- <-137

μ

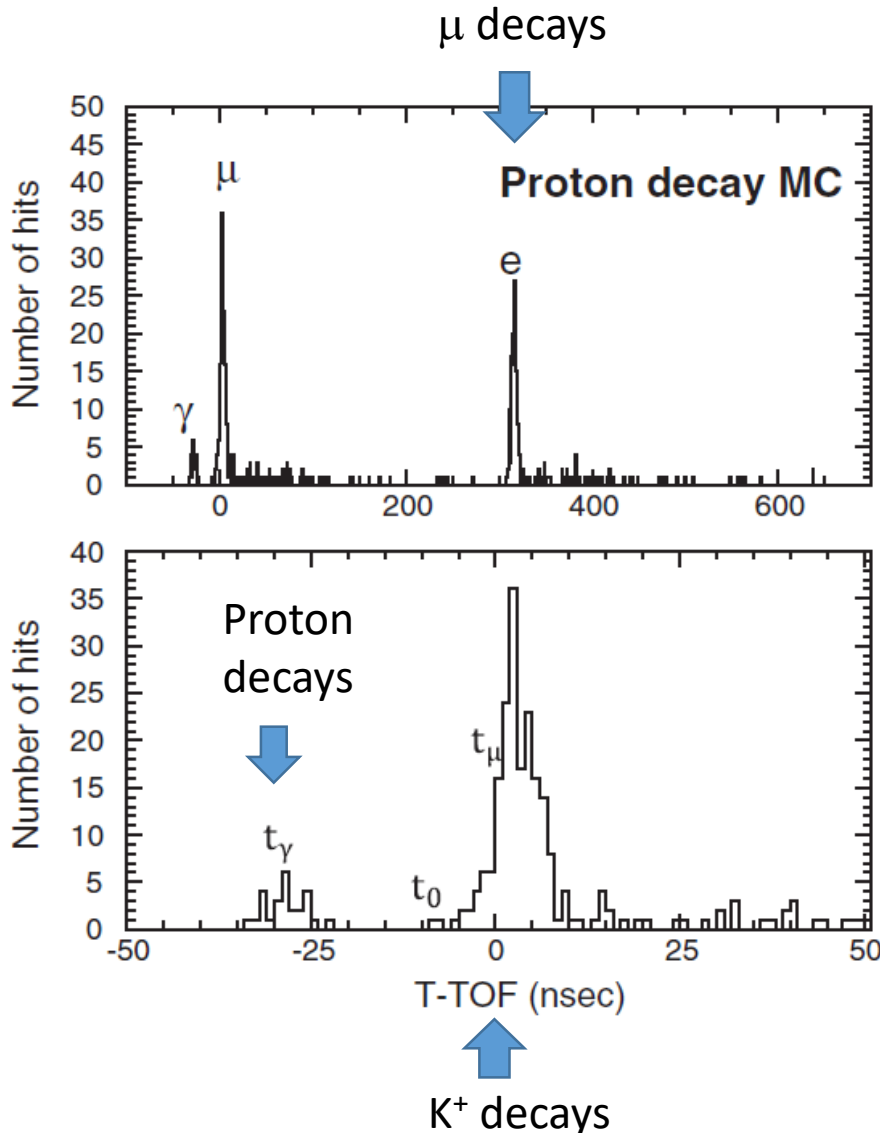
Color of hits corresponds to time.

Cyan corresponds to nuclear γ .

Difficult to identify γ from hit pattern.




Time structure with nuclear γ



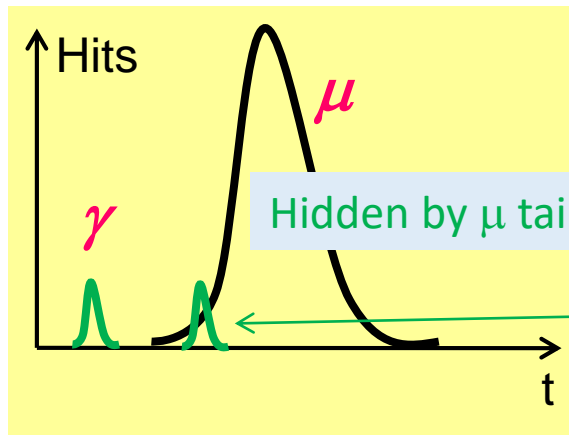
- 3 hit clusters in time should be observed in case of signal.
- The event is triggered by μ hits.
- γ signal is much smaller than μ and easily hidden by tail of μ hits.
- Make 12 nsec time window and slide it toward left from t_0 (end of μ tail) to search for maximum hit cluster.

Selection criteria for $p \rightarrow \nu K^+$, $K^+ \rightarrow \nu \mu^+$

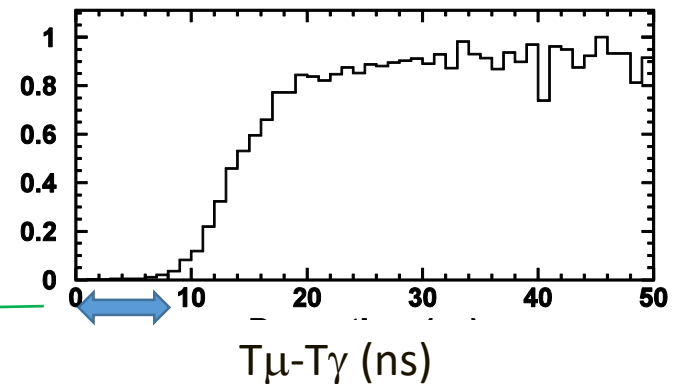
- 1 μ -like ring with Michel electron
- $215 < P_\mu < 260$ MeV/c
- Proton rejection cuts
- Search Max hit cluster  Reduce background by 5×10^{-4} !
by sliding time window (12ns width);
 - $4 < N_\gamma < 30$ hits
 - $T_\mu - T_\gamma < 75$ nsec
- No neutron
- Selection efficiency = (selected events)/(proton decay in fiducial volume):
9 %
 - $\text{Br}(K^+ \rightarrow \nu \mu^+) = 64 \%$, only 40 % emits nuclear $\gamma \rightarrow 26 \%$ even if detector is perfect.

Remark for this analysis

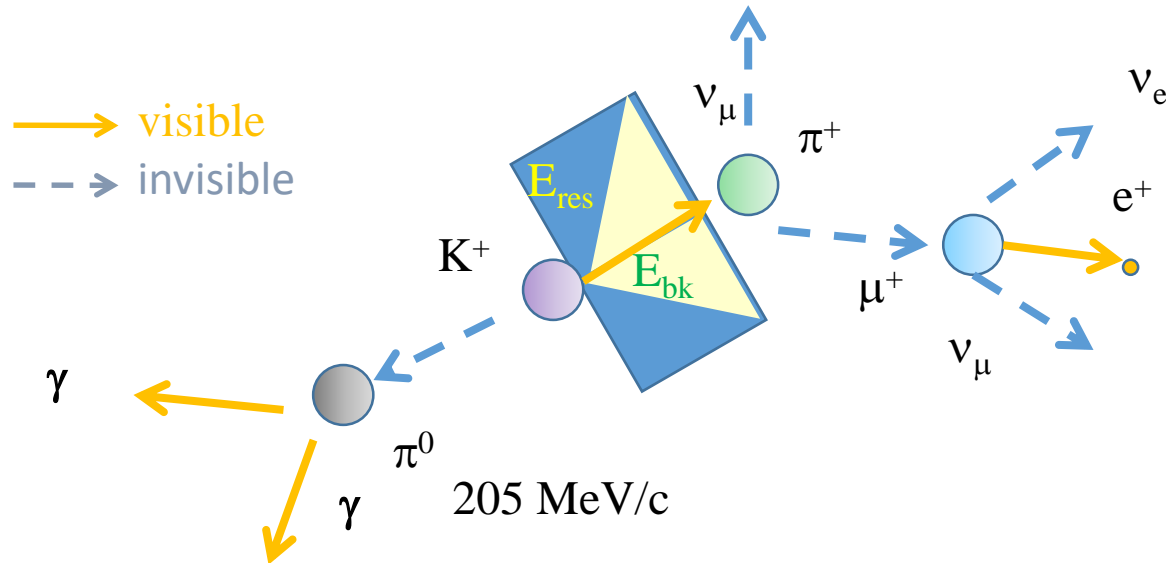
- This analysis is limited by time resolution of PMTs.
 - If γ is close to μ , γ peak is hidden by μ hits.
 - Time resolution of SK PMT is 2.2 nsec at 1 photoelectron.
 - If μ peak becomes sharper, the selection efficiency will be improved.



γ tagging efficiency



3-3. Search for $p \rightarrow \nu K^+$, $K^+ \rightarrow \pi^+ \pi^0$



- Both π^+ and π^0 has **205 MeV/c** in momentum. This is just above Cherenkov threshold for π^+ , thus it is not identified as a ring in most of case.
- π^+ decays into μ (invisible) and ν , μ decays into $e \nu_e \nu_\mu$.
- π^0 decays into 2 γ s.
- Search for 206 MeV/c π^0 with Michel electron.

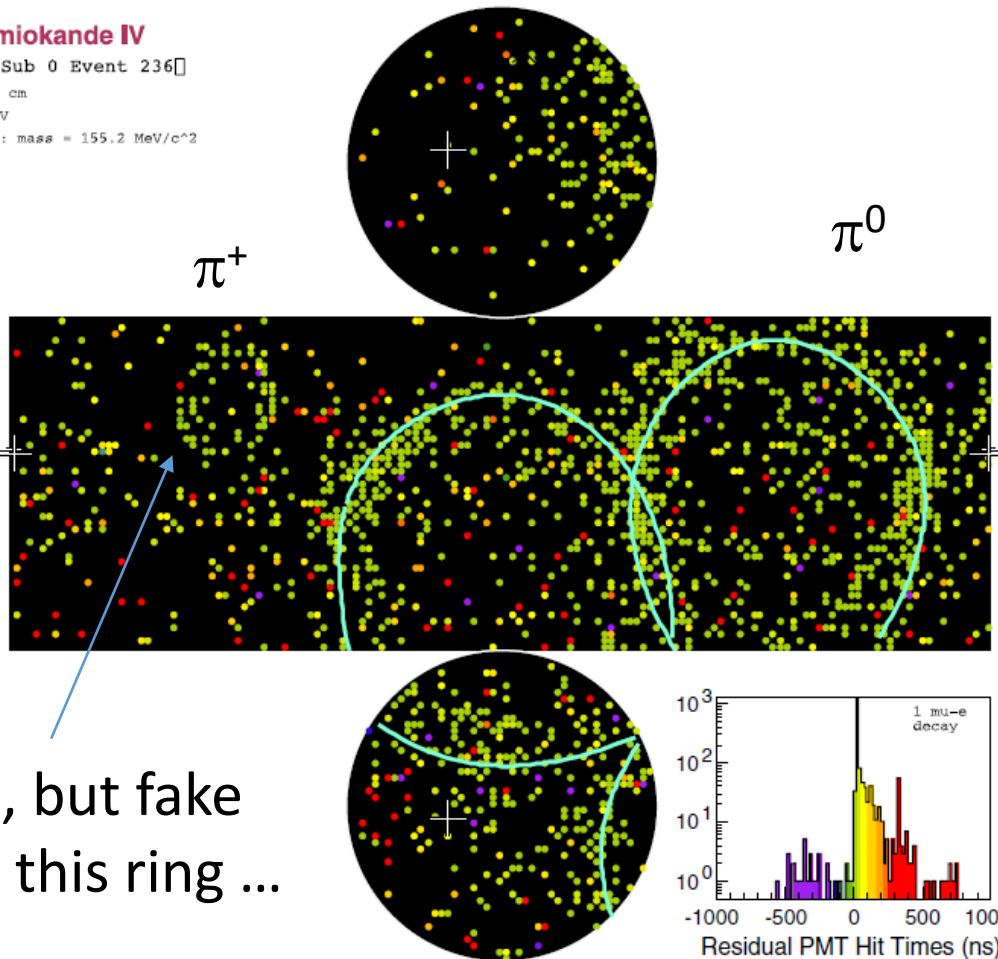
Example of $p \rightarrow \nu K^+$, $K^+ \rightarrow \pi^+ \pi^0$

Super-Kamiokande IV

Run 999999 Sub 0 Event 236
D_{wall}: 1076.4 cm
E_{vis}: 260.4 MeV
2 e-like rings: mass = 155.2 MeV/c²

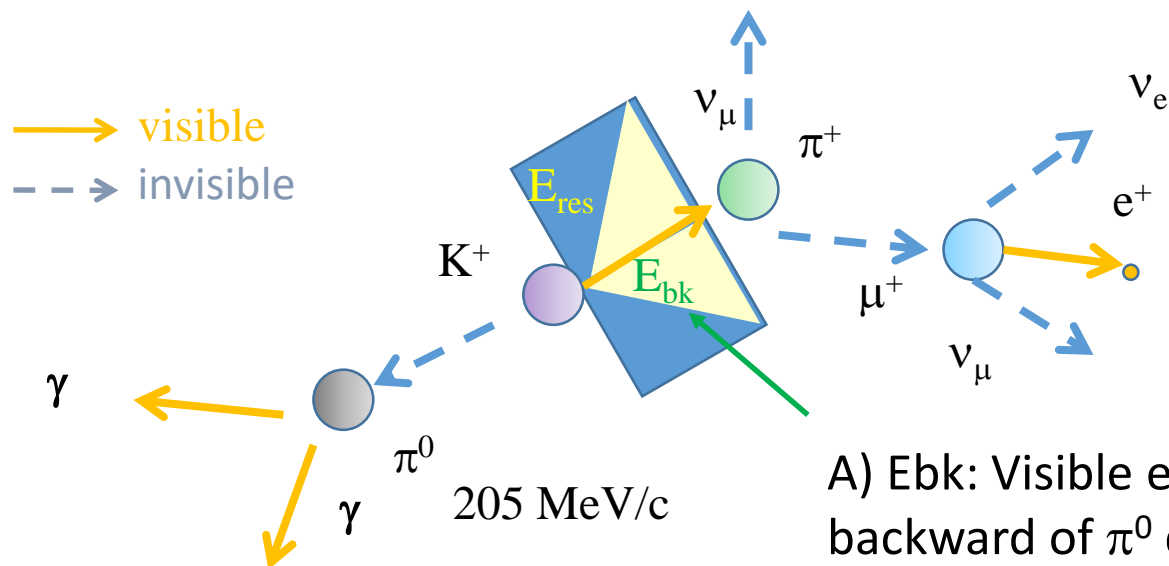
Resid(ns)

- > 251
- 220- 251
- 188- 220
- 157- 188
- 125- 157
- 94- 125
- 62- 94
- 31- 62
- 0- 31
- -31- 0
- -62- -31
- -94- -62
- -125- -94
- -157- -125
- -188- -157
- <-188

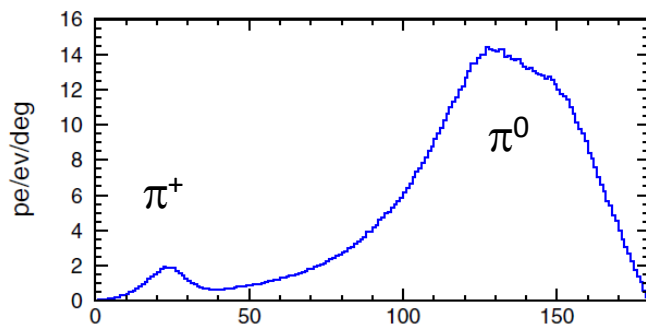


Look like a ring, but fake
ring cut rejects this ring ...

Use π^+ information to select events

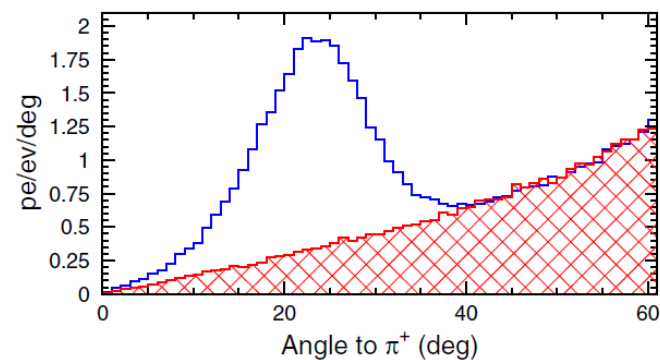


Charge distribution



↑
 π^+ direction

Zoom



B) Make likelihood for hit pattern.

Selection criteria for $p \rightarrow \nu K^+, K^+ \rightarrow \pi^+ \pi^0$

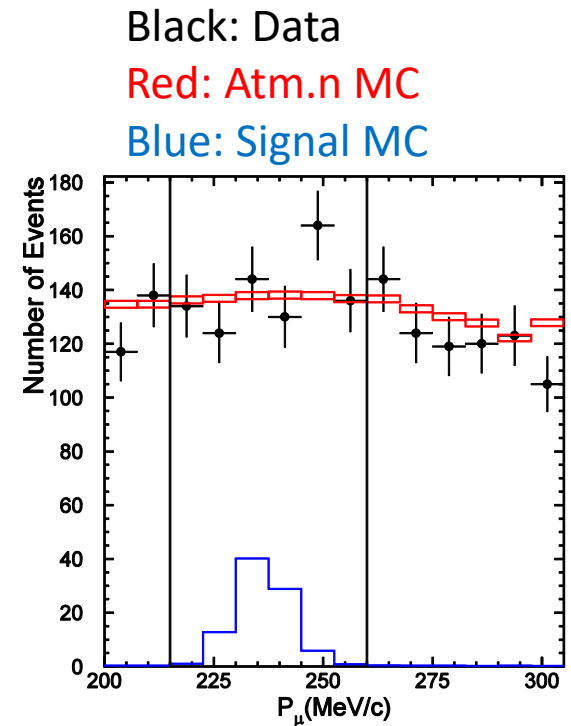
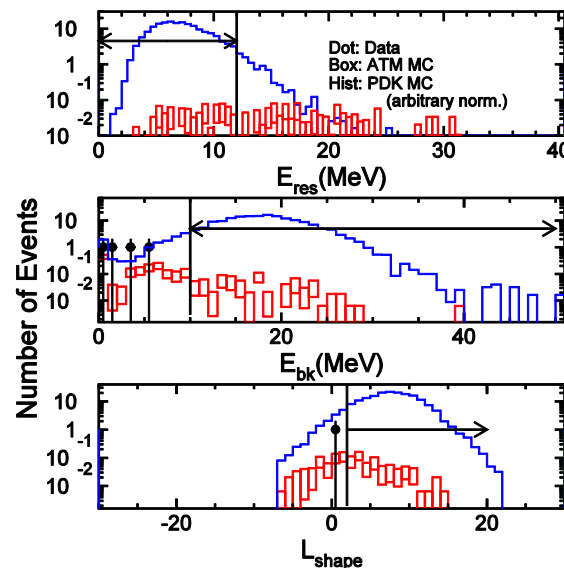
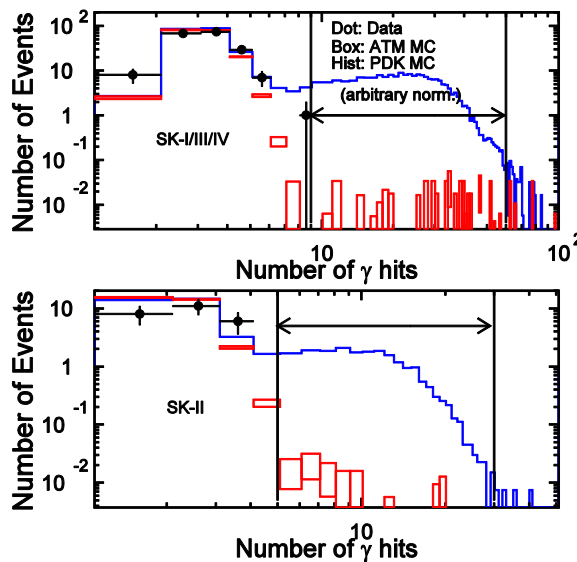
- 1 or 2 e-like rings with decay-e.
- $85 < M\pi^0 < 185$ MeV.
- $175 < P\pi^0 < 250$ MeV/c.
- E_{bk} : visible energy sum in 140-180 deg. of π^0 dir,
 E_{res} : in 90-140 deg,
 L_{shape} : Likelihood based on charge profile
 - $10 < E_{bk} < 50$ MeV
 - $E_{res} < 12$ MeV (20 MeV for 1ring)
 - $L_{shape} > 2.0$ (3.0 for 1ring)
- No neutrons
- Selection efficiency: 10 % ($Br(K^+ \rightarrow \pi^+ \pi^0) = 21$ %)

Background for $p \rightarrow \nu K^+$

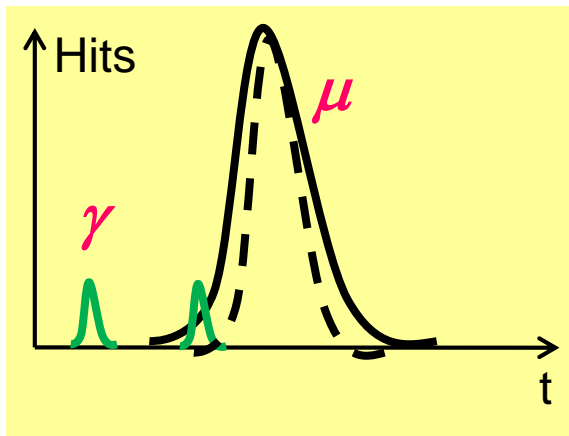
- Dominant background is K^+ production by neutrino interactions.
 - $\nu p \rightarrow \nu \Lambda K^+$, $\Lambda \rightarrow p \pi^-$ (BR:64 %, mostly invisible in WCD)
 - Emit nuclear γ as same as the signal.
- It is also rare interaction and we had poor information from very old bubble chamber. Large uncertainty.
- Recently MINERvA measures K^+ production. It is very useful information for this analysis.

3-4. SK results (So far)

- Exposure: 365 kton · year
- Expected background: 0.3 events for $K^+ \rightarrow \nu \mu$ with nuclear γ , 0.6 events for $K^+ \rightarrow \pi^+ \pi^0$.
- No candidates observed and no excess in momentum distribution.
- Lower lifetime limit: $> 0.8 \times 10^{34}$ year

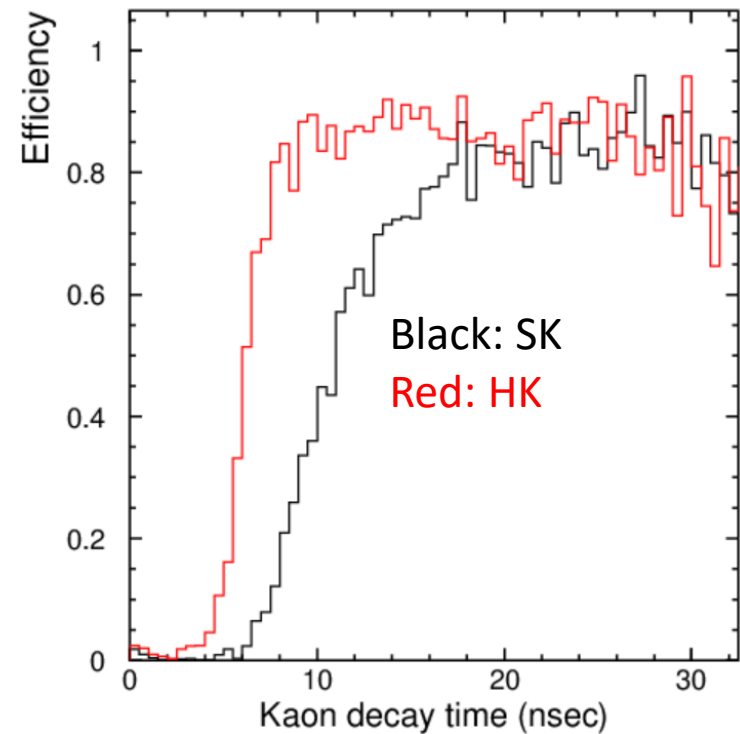


Faster PMT response improves nuclear γ tagging in $p \rightarrow \nu K^+$



- Time resolution: 2.2nsec (SK)
→ 1.1 nsec (HK).
- Sharper time distribution of μ
→ γ close to μ can be identified !

γ tagging efficiency



(Better photon counting also
contributes improvement)