

# Hardware orientation & Training

**Son Cao (IFIRSE)**

6th Vietnam School on Neutrinos, July 12, 2022

# Time allocated for Hardware training



- 45 mins. for hardware orientation
- 4.0 hours x 2 (sec., Wed. And Thu.) for hand-on experience in total
  - But the lab is small and we are lack of human power to have all four groups trained at the same time.
  - Each time, we manage to have 2 groups → 2 hours x 2 (sec.) for each person

# Student grouping & protocol for hardware training

## Same grouping scheme as software training

- Detail of student grouping will be discussed in next section
- Students are divided into 4 groups: Group-A, Group-B, Group-C, and Group-D
- We prepare two independent setups. *They are not identical but the concept is similar*
- Training protocol (for Wed. and Thu. Afternoon section)
  1. Group A + B go for hardware training (13:20 - 15:00) while Group C+D work on mini-projects
  2. 20 minus break (15:00-15:20)
  3. Group C+D go for hardware training (15:20 - 17:00) while Group A+B work on mini-projects

**Hardware training**

**Study place  
for mini-project**

# Goals of hardware training

**Provide some hands-on experience w/ hardwares used in real Neutrino Detector**

**Vietnamese students lack skills with hardware, especially in particle and nuclear physics**

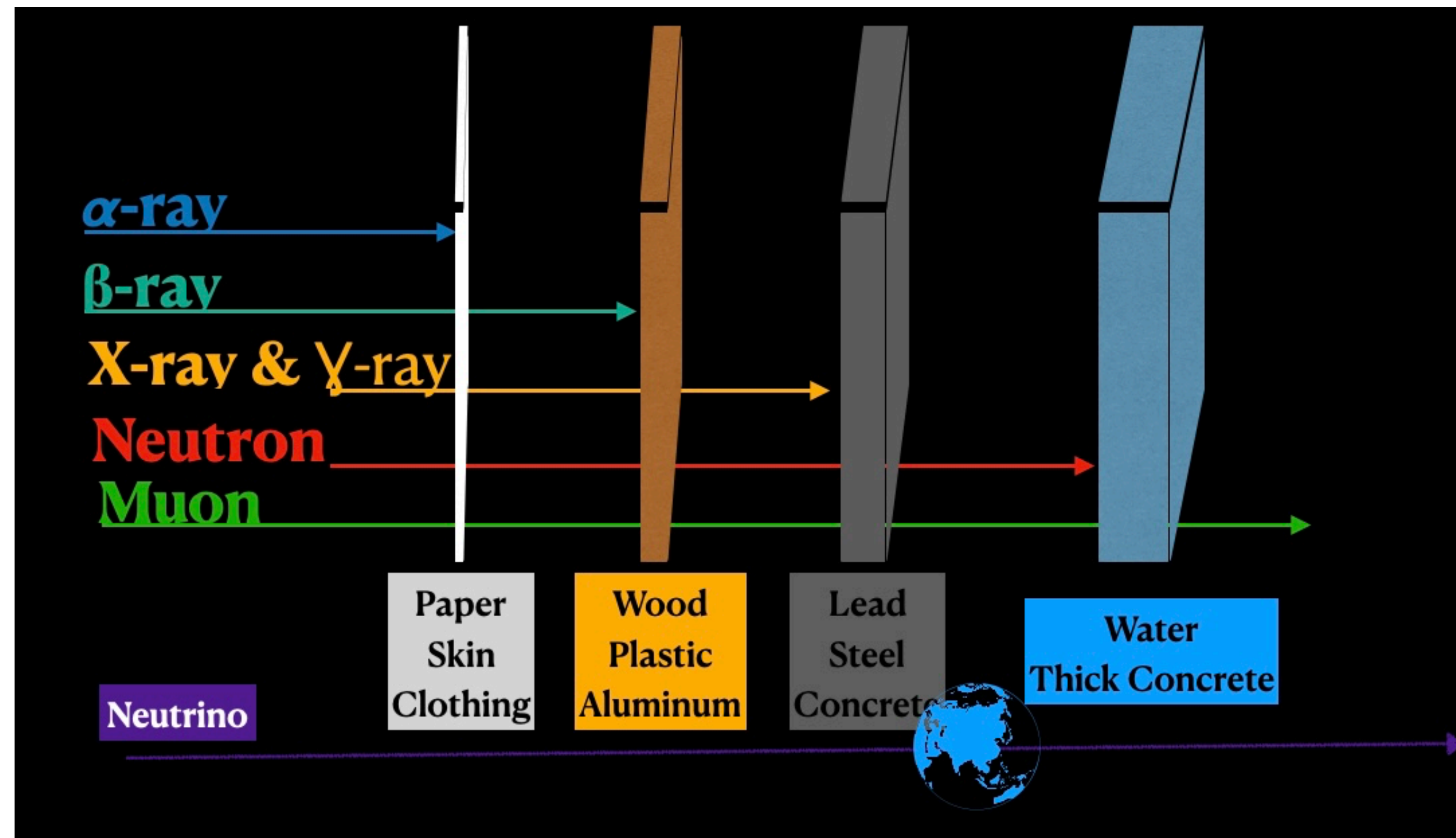
- Did you use multimeter before?
- Did you use oscilloscope before?
- Did you use NIM modules before?
- Did you use photosensor (*not including smartphone's camera*) before?
- ....



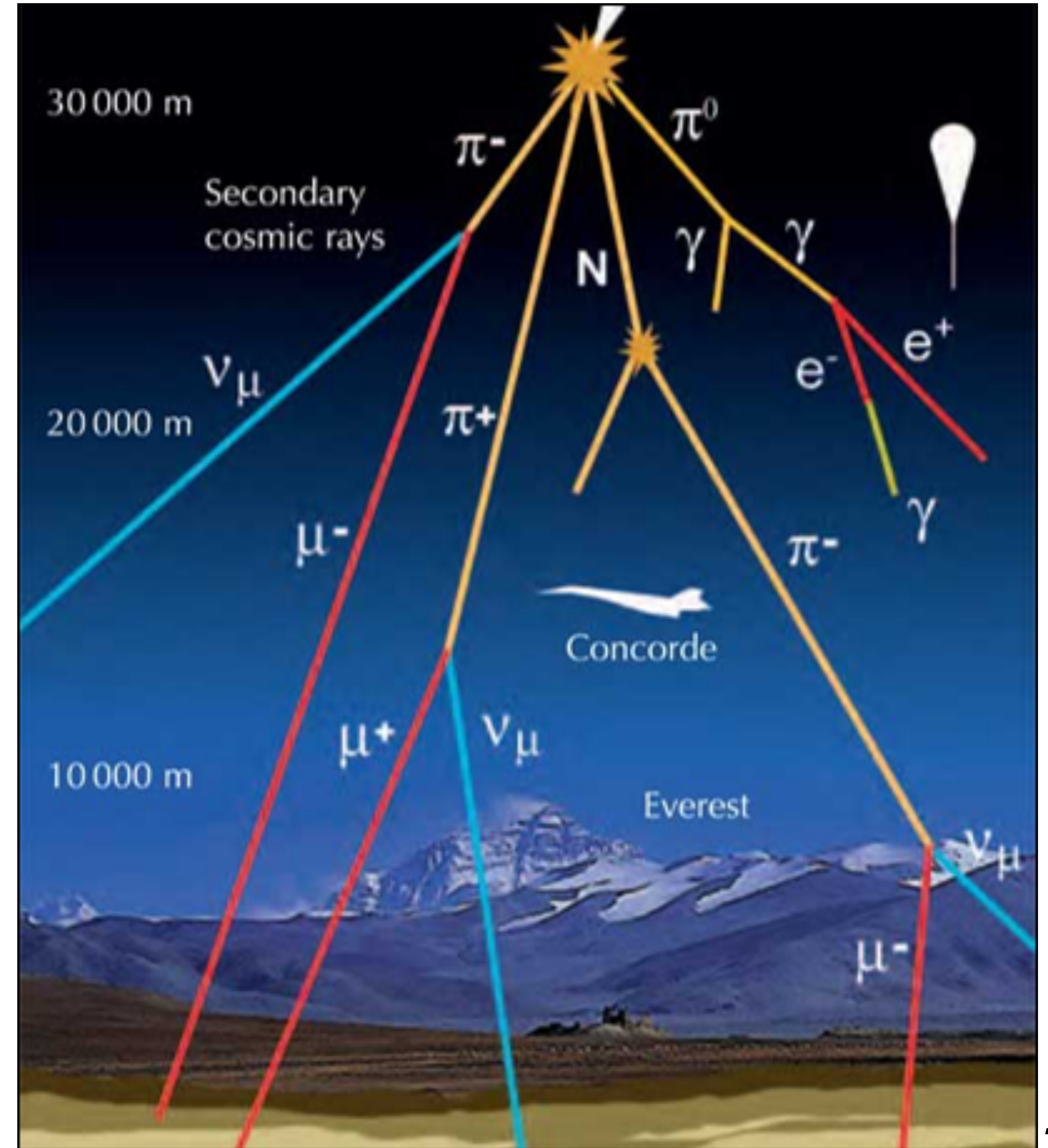
# Goals of hardware training

**Provide some hands-on experience w/ hardwares used in real Neutrino Detector**

- But ...We won't see the neutrino interaction in the lab. *You will need big detector and place near the huge source of neutrino for this. Also real-time identification of neutrino interaction is quite challenging*
- What we can see is with the **cosmic ray muon (big brother/sister of neutrinos)**



**High energy astrophysical particles (eg. hydrogen & helium from the Sun) interact with the Earth's atmosphere**  
**→ produce vast amount of muons**  
**~ 1 muon/cm<sup>2</sup>/s**

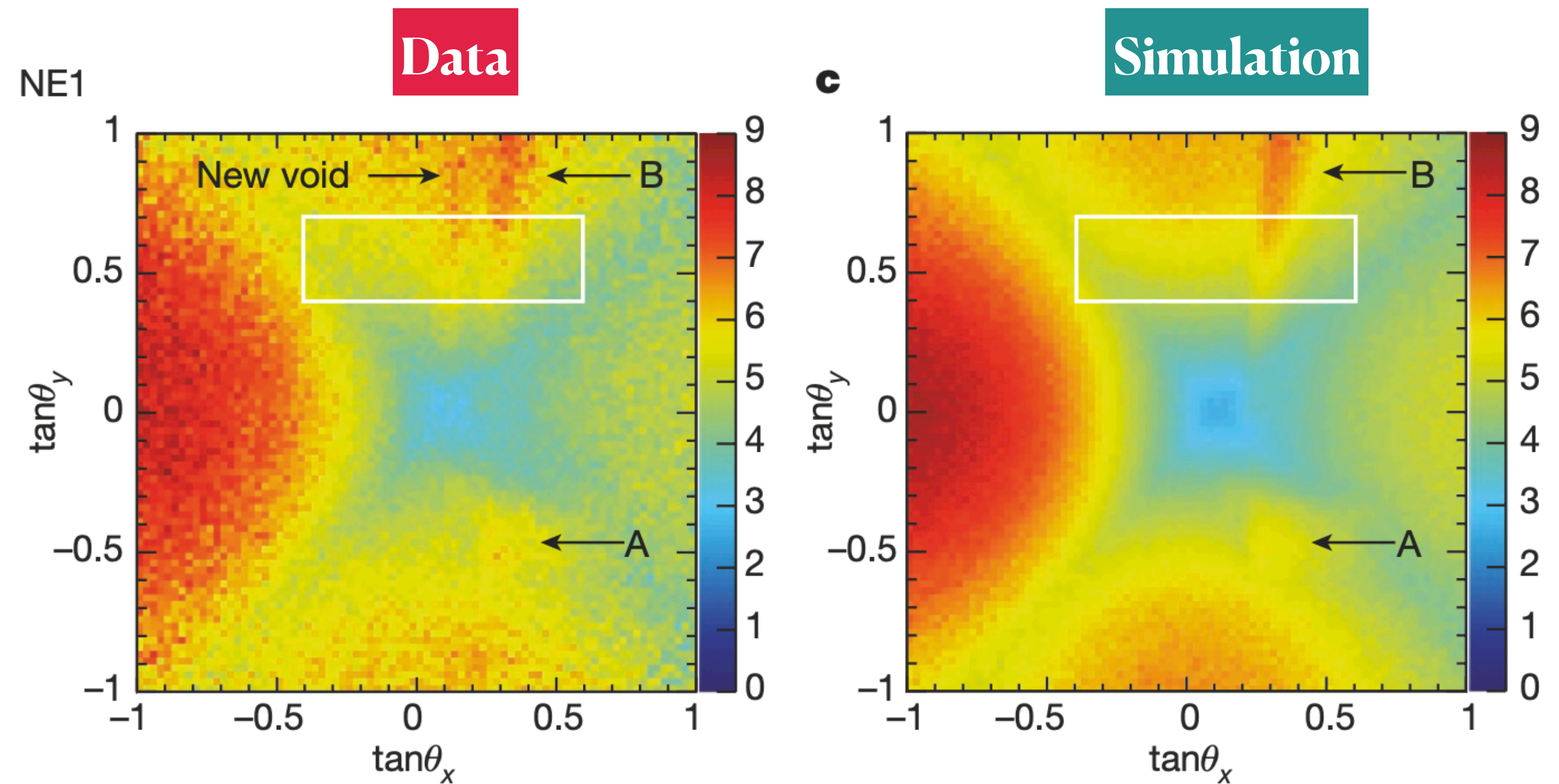


# Muon can be used for a practical application

<https://www.nature.com/articles/nature24647>



called: muon radiography technique

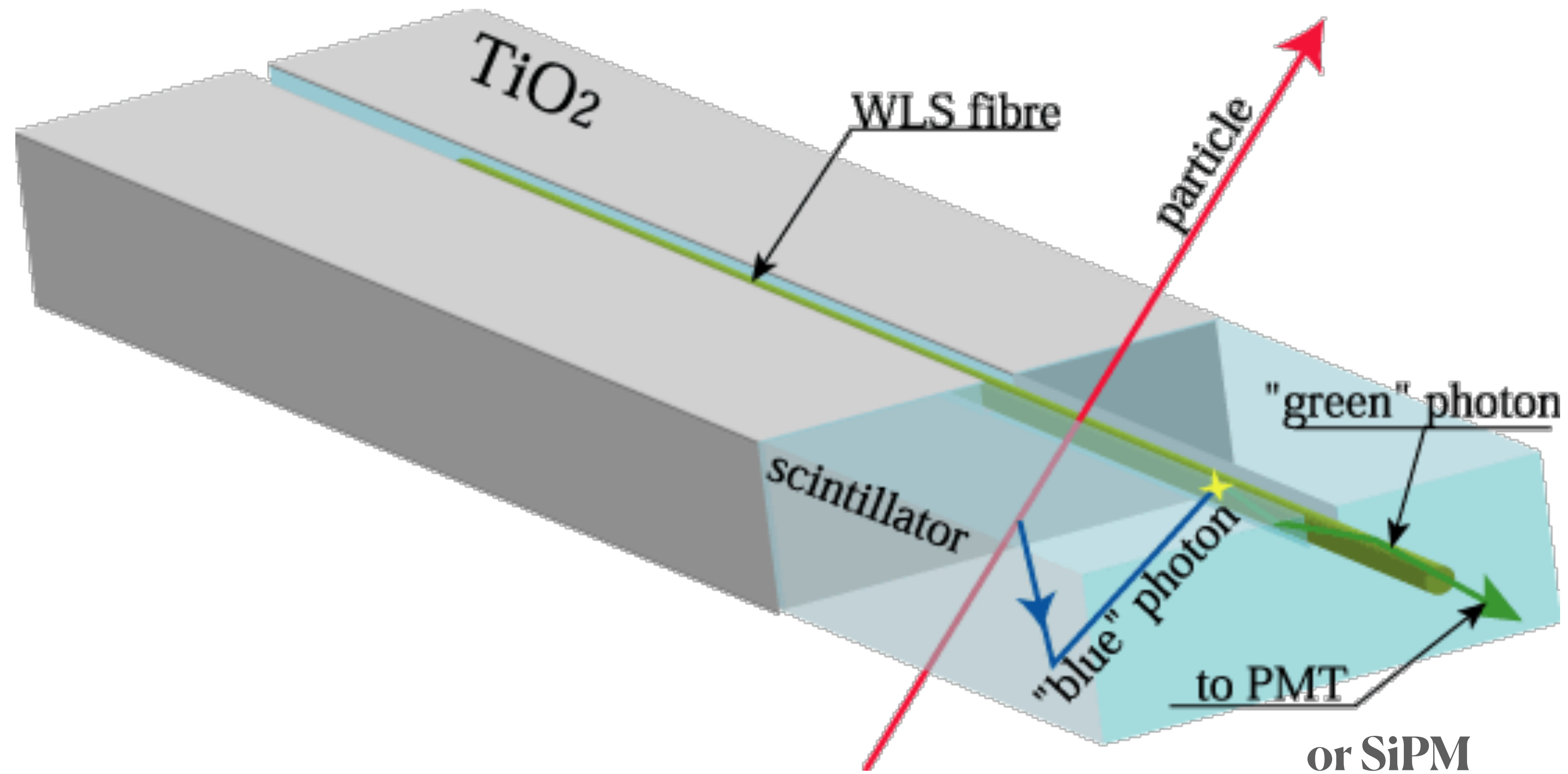


Color is corresponding to intensity of muons.  
Red is with more muons detected



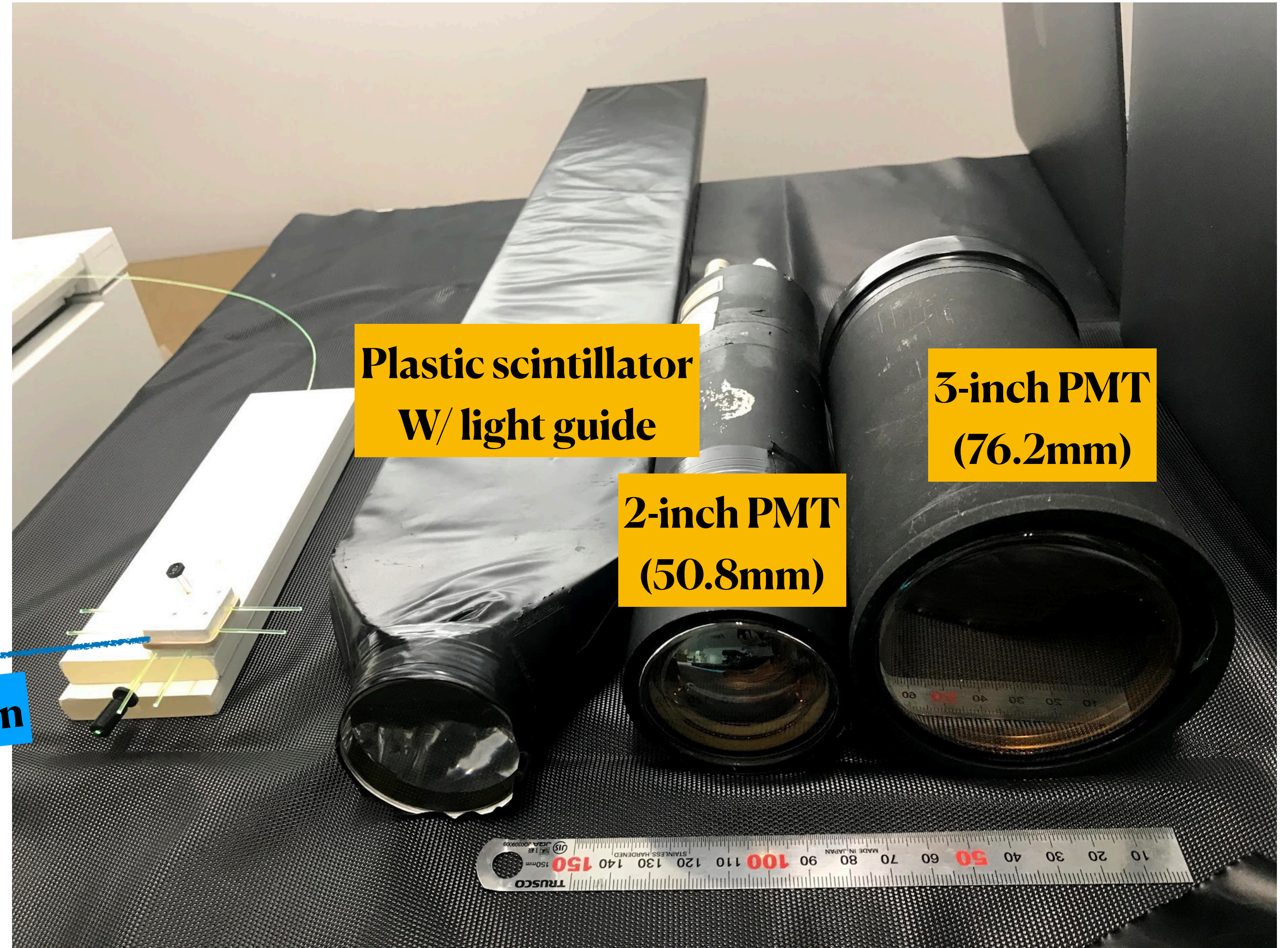
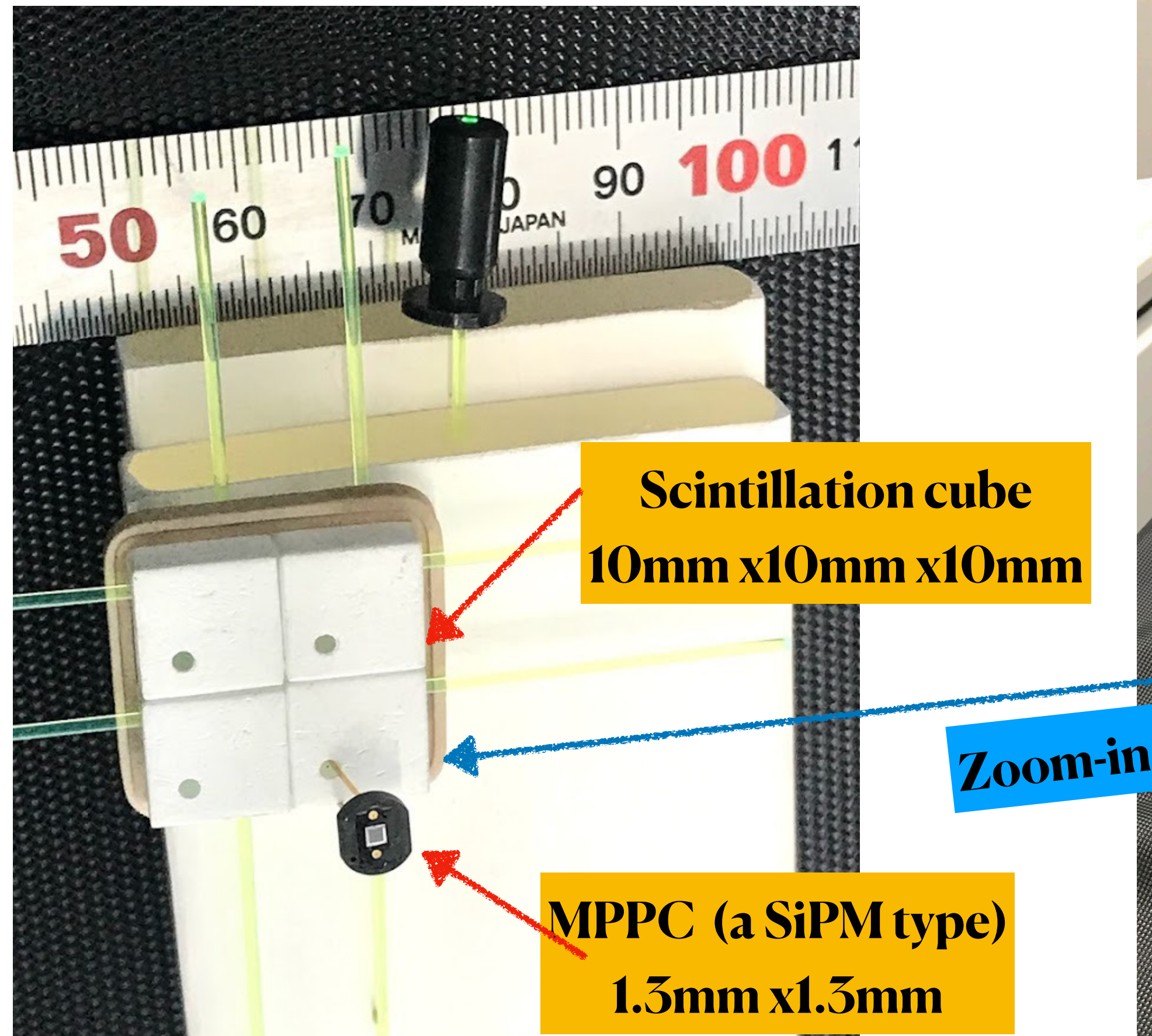
**How can we see muons and  
measure their characteristics  
with *what we have in the lab?***

# Tracking the charged particle w/ scintillator



**When passing through the scintillator, charged particles ( $\mu$ ,  $\pi$ ,  $e$ ,...) deposits energy and excite the scintillation photons, which are collected and guided to the photosensor for converting to the electrical signals (more convenient to manipulate) for data recording.**

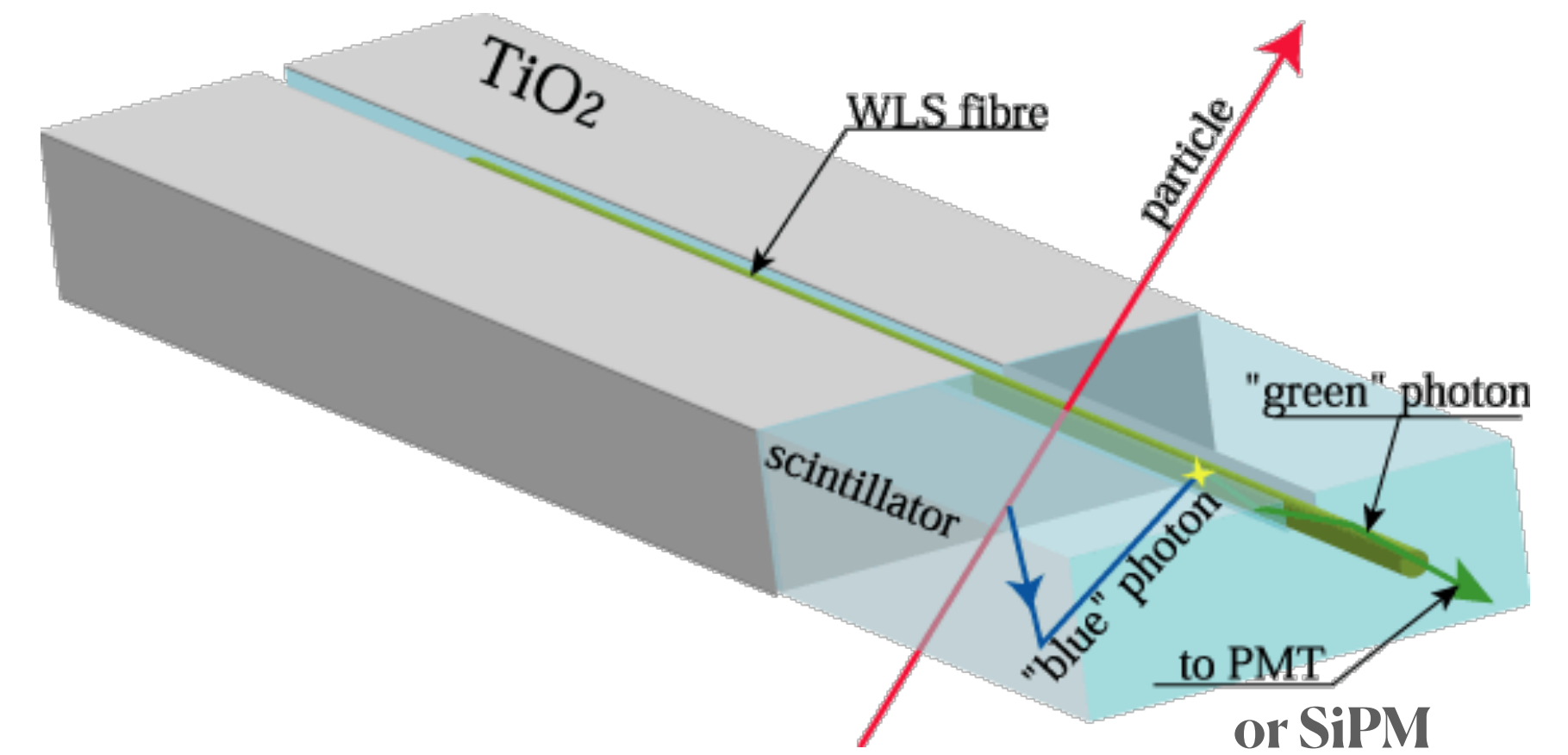
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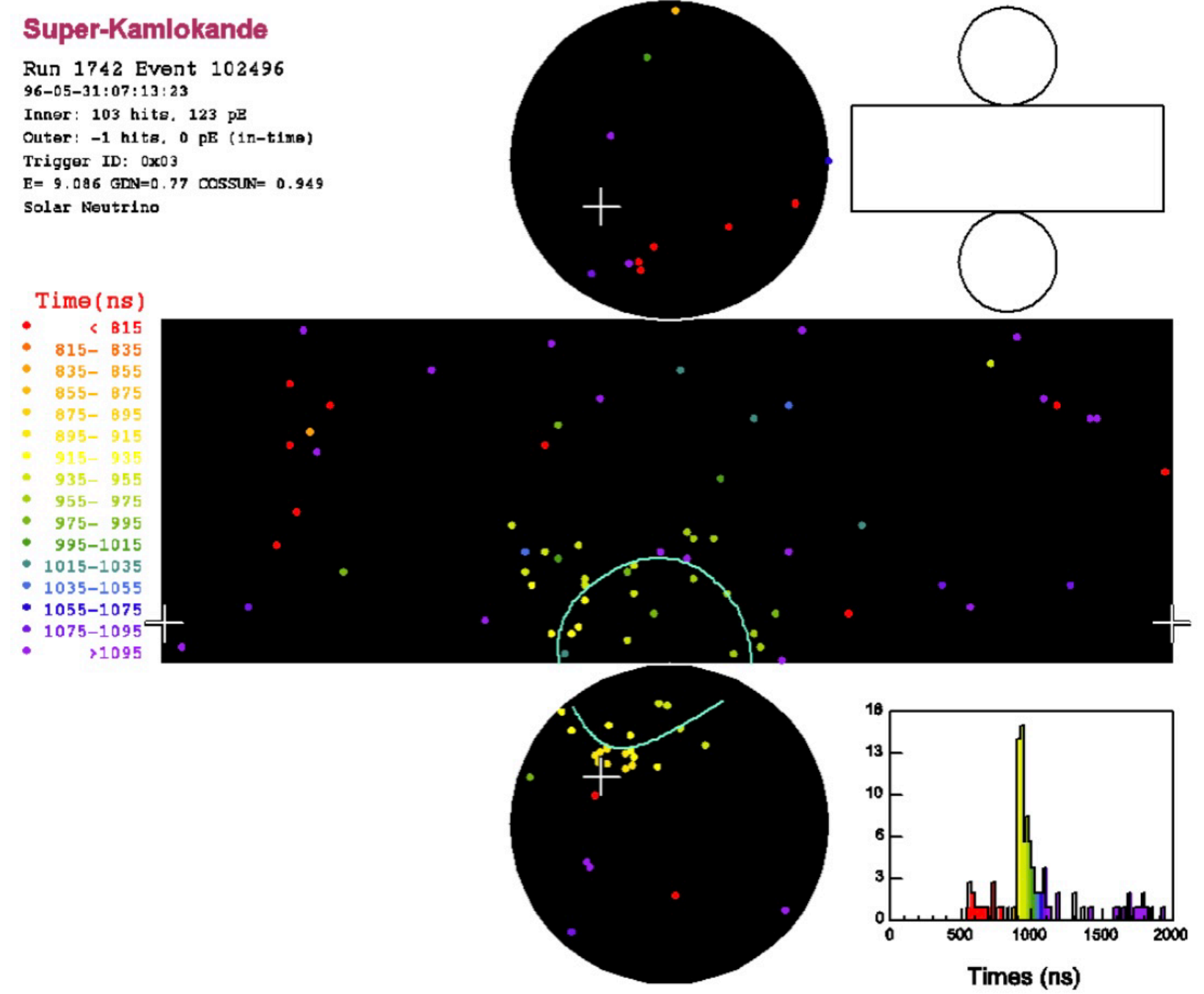
# Tracking the charged particle w/ scintillator

- Muon deposits  $\sim 2\text{MeV}$  per 1cm path in the plastic scintillator
- $2\text{MeV}$  deposit energy will produce  $\sim 10,000$  photons
- Assume the probability for WLS catching the photons is about 1%, then  $\sim 100$  photons are capture and change to green photons
- Detection of photosensor is about 20-40%, so will have about **20-40 photoelectrons observed**
  - Sometime you can get lower due to the aging of scintillator, attenuation in the WLS or light loss from imperfect coupling between the WLS and photosensor

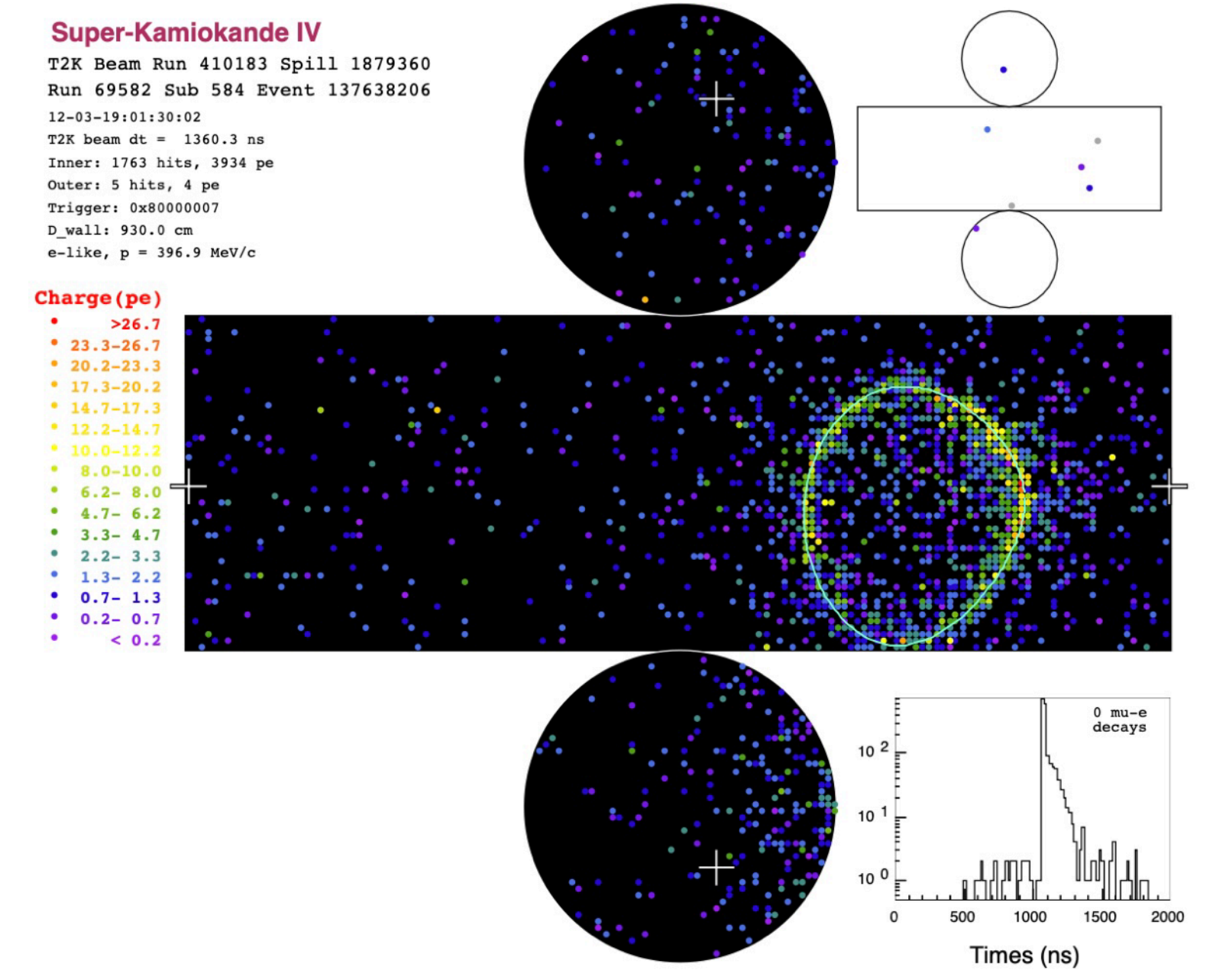


# Trace of neutrinos: (typically) very faint flash of light

“Experimental neutrino experiment in the nutshell”



**A ~ 9MeV solar neutrino candidate**  
**123 p.e. counted in 103 PMT in few 100ns;**  
**~ 1 p.e. per hit PMT**



**A ~400MeV  $\nu_e$  candidate from T2K beam**  
**3934 p.e. counted in 1763 hit PMT in few 100ns**  
**~3-4 p.e. per hit PMT**

**In a blinking of LED**



**...~10<sup>15</sup> photons are generated**

**Typically, signature of the comic ray  
muons is also faint**

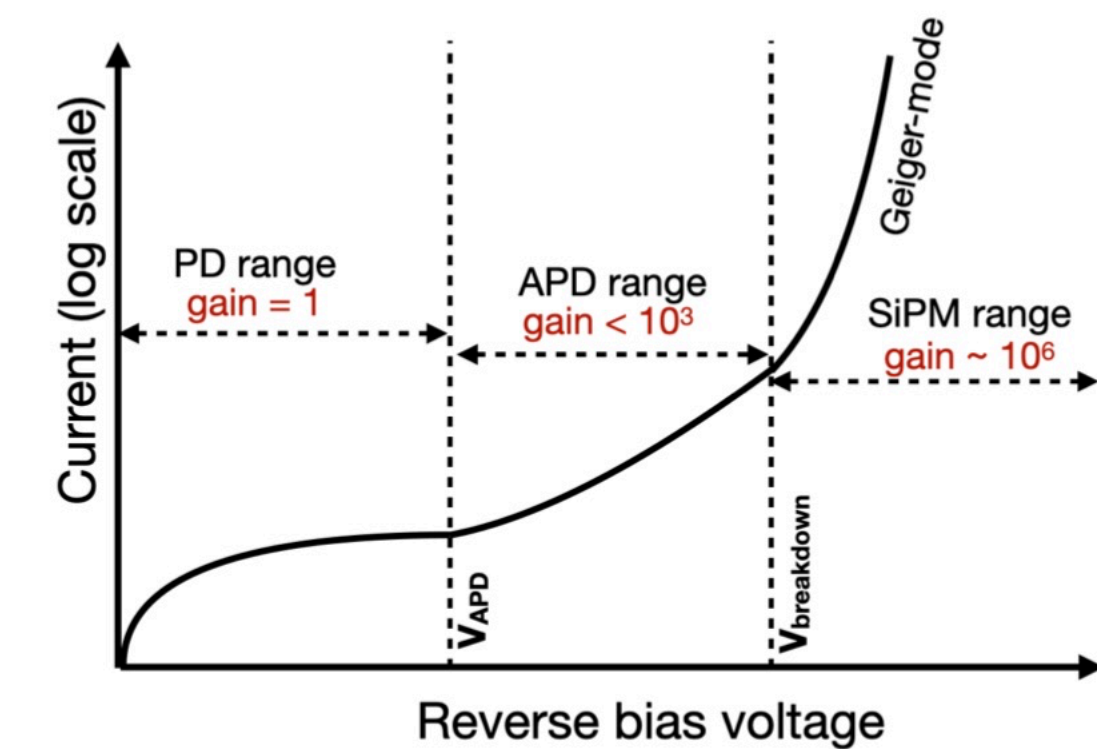
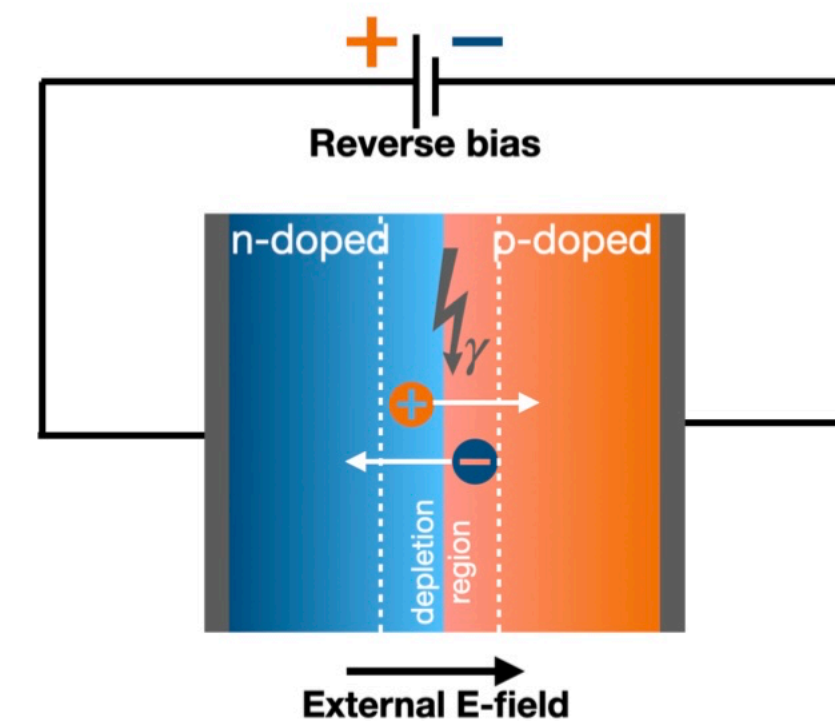
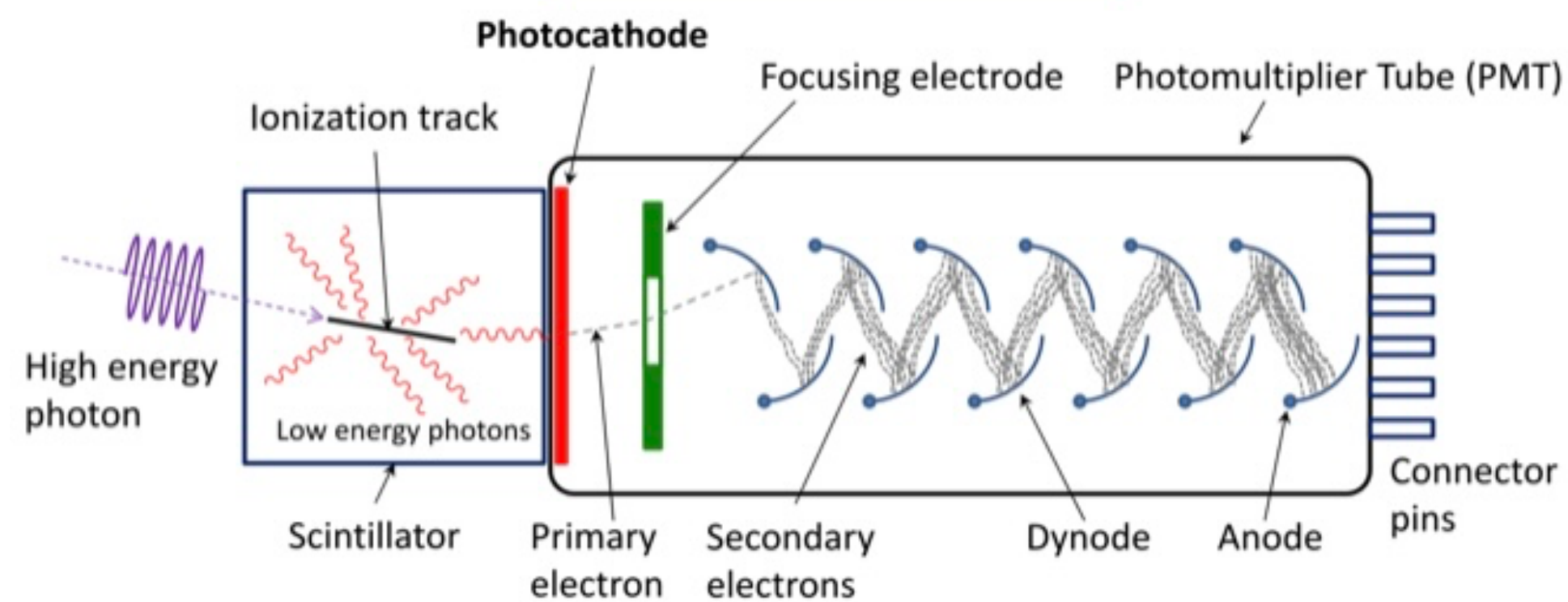


**We need a very good “Eyes”**

# Photosensors: Extremely important to extend particle frontiers (precision, sensitivity, intensity...)

<https://hub.hamamatsu.com/us/en/technical-note/WITS-guide-detector-selection/index.html>

Characteristic	PMT	PD	APD	SiPM
Spectral coverage [nm]	115-1,700	190-13,000	190-1,700	320-900
Peak QE ( $\eta$ ) [%]	< 40	< 90	< 90	< 40 ( <i>PDE</i> )
Active area [mm <sup>2</sup> ]	< 12,000	< 100	< 100	< 10
Gain ( $\mu$ )	$10^5$ - $10^6$	1	< 100	$10^5$ - $10^6$
NEP [W/ $\sqrt{\text{Hz}}$ ]	$> 2 \times 10^{-17}$	$> 6 \times 10^{-16}$	$> 1 \times 10^{-15}$	$> 6 \times 10^{-16}$
Rise time [ns]	$> 0.15$	$> 0.23$	$> 0.35$	$> 1$
Bandwidth [Hz]	$< 2 \times 10^9$	$< 1.5 \times 10^9$	$< 1 \times 10^9$	NA
Time jitter [ns]	$> 0.05$	NA	$> 0.2$	$> 0.2$

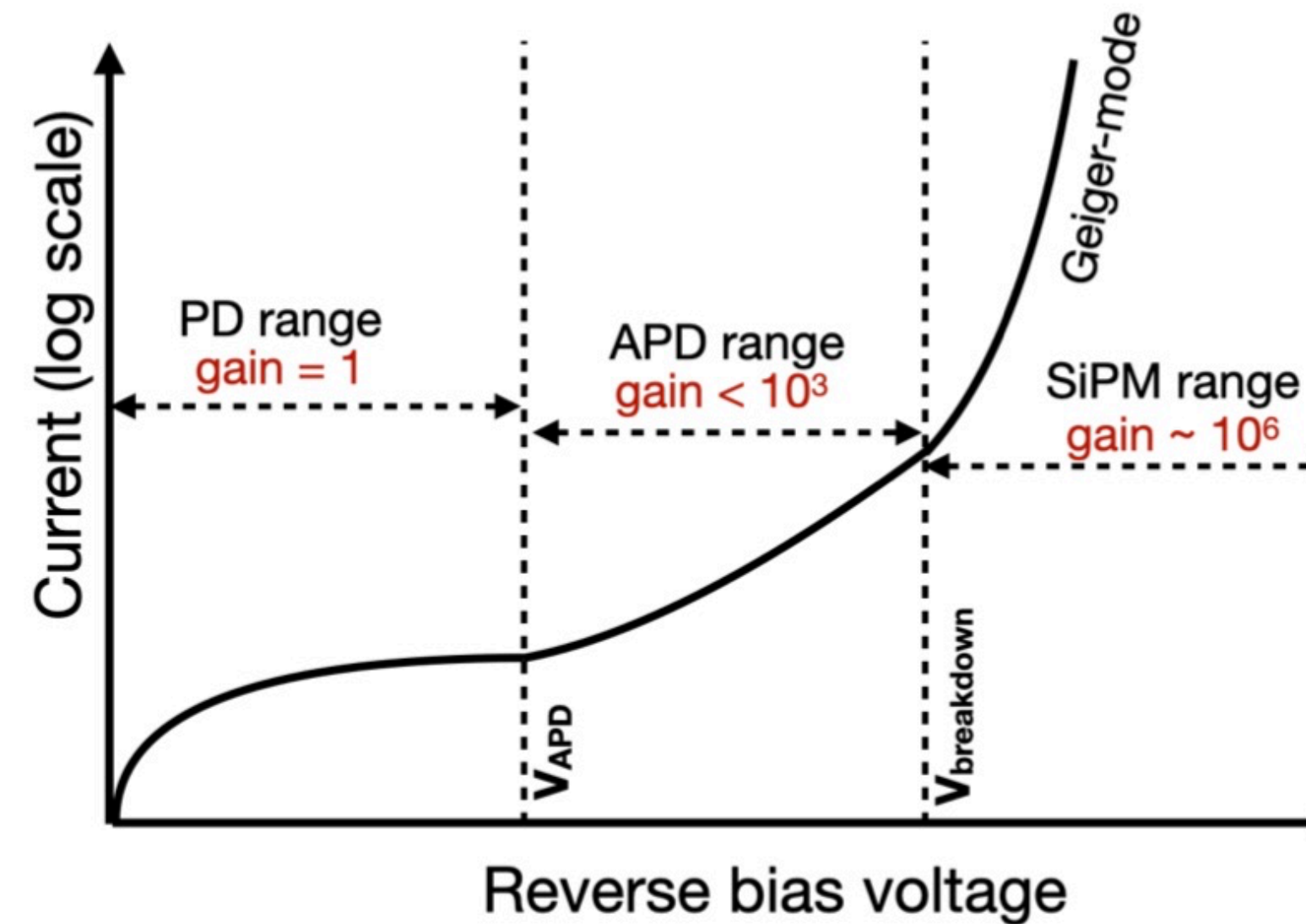
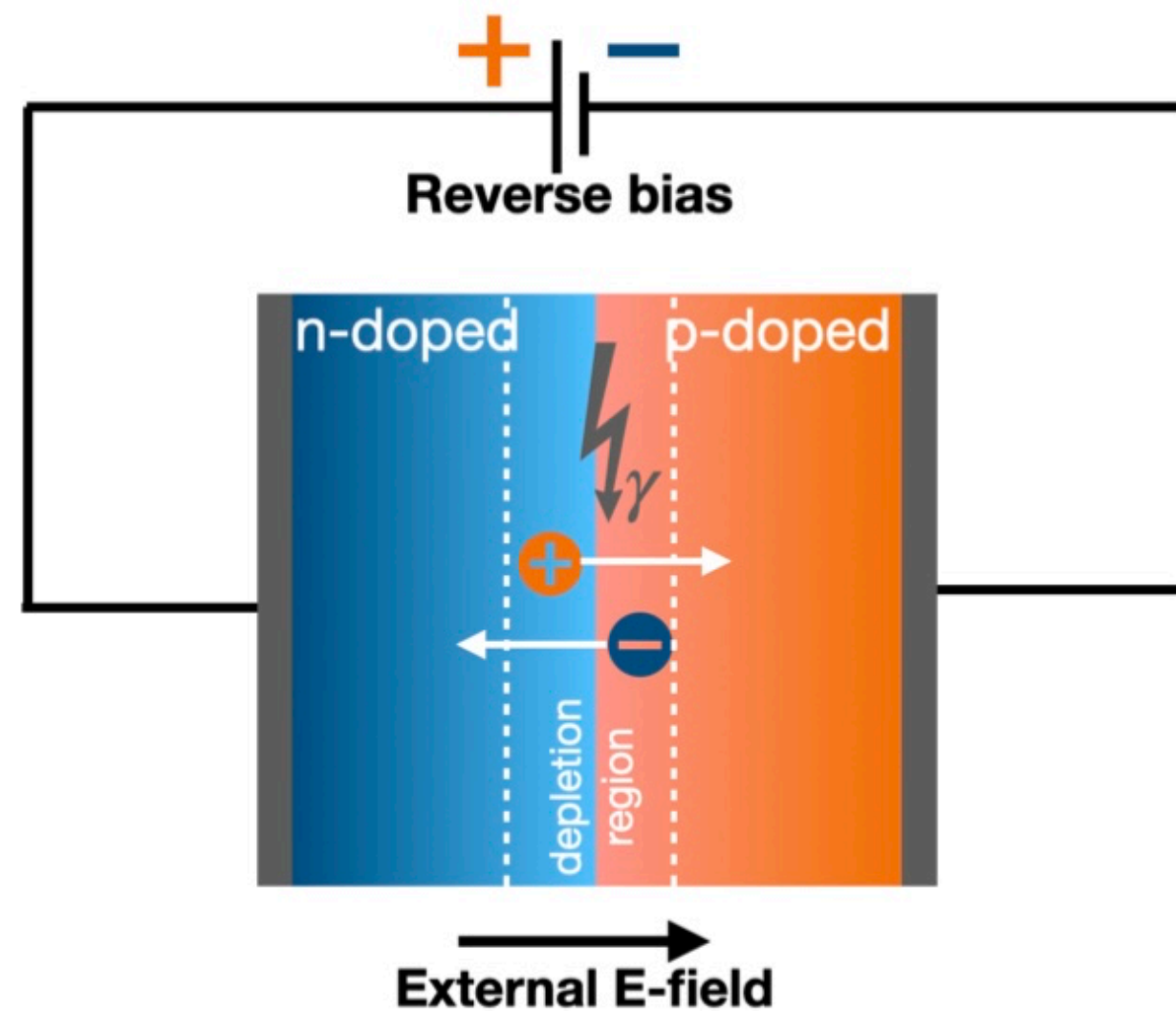


**“The real voyage of discovery consists, not in seeking new landscapes, but in having new eyes”**

**Marcel Proust**

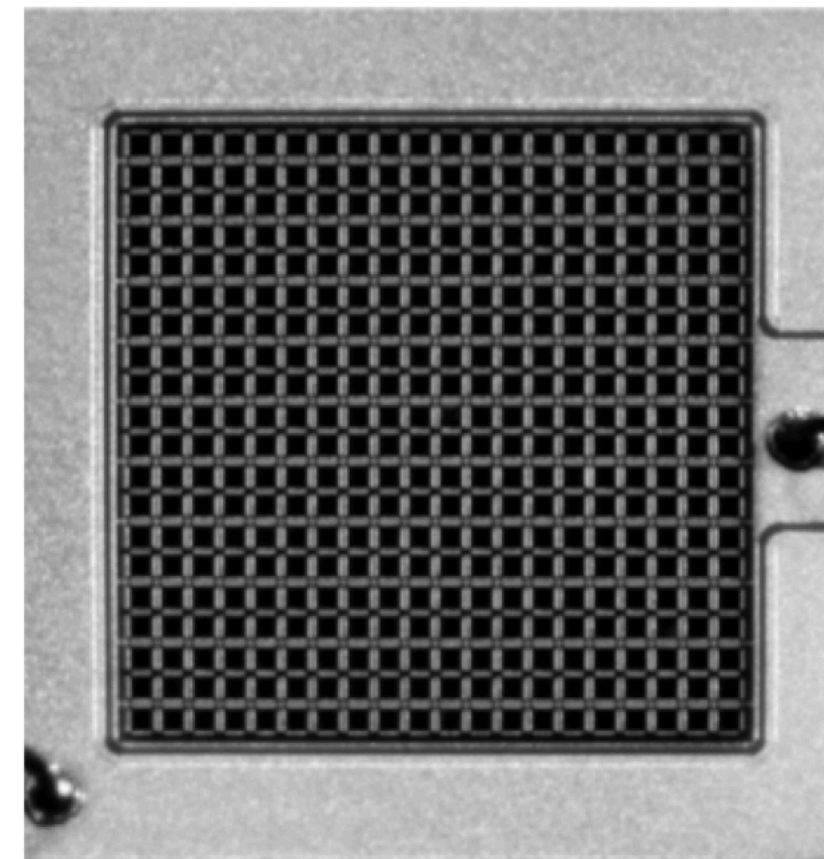
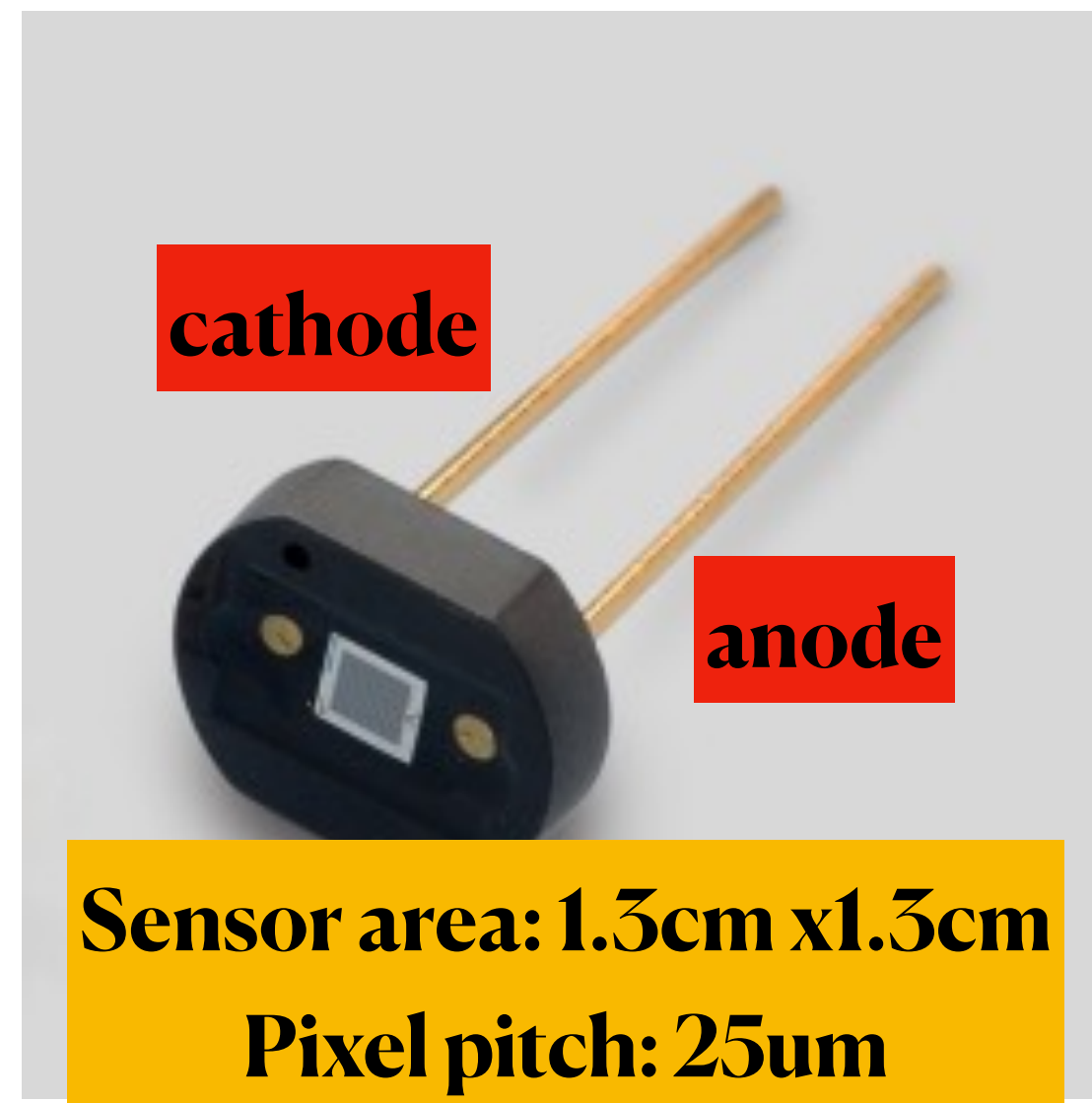


# Photon detection principle w/ Silicon photomultiplier (SiPM)



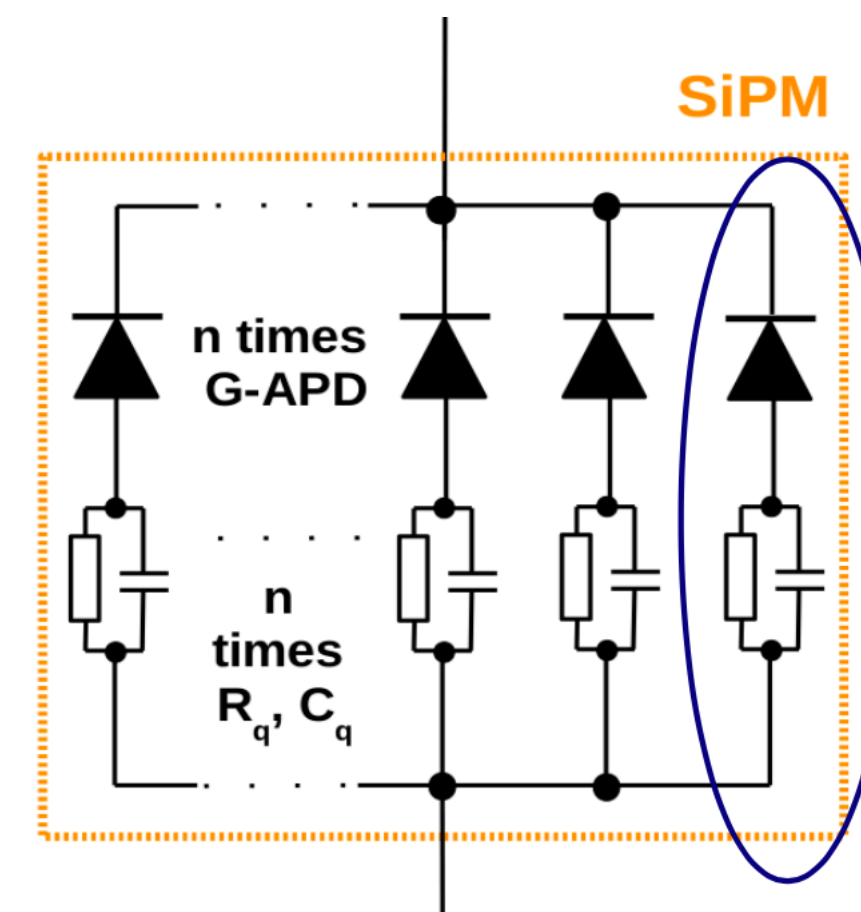
- Based on photoelectric effect: photon strikes and produce a pair of electron/hole
- Various types, selection depending on the measurement
- “Breakdown” here mean both hole and electron play roles in avalanche process

# MPPC: a type of SiPM, developed by Hamamatsu

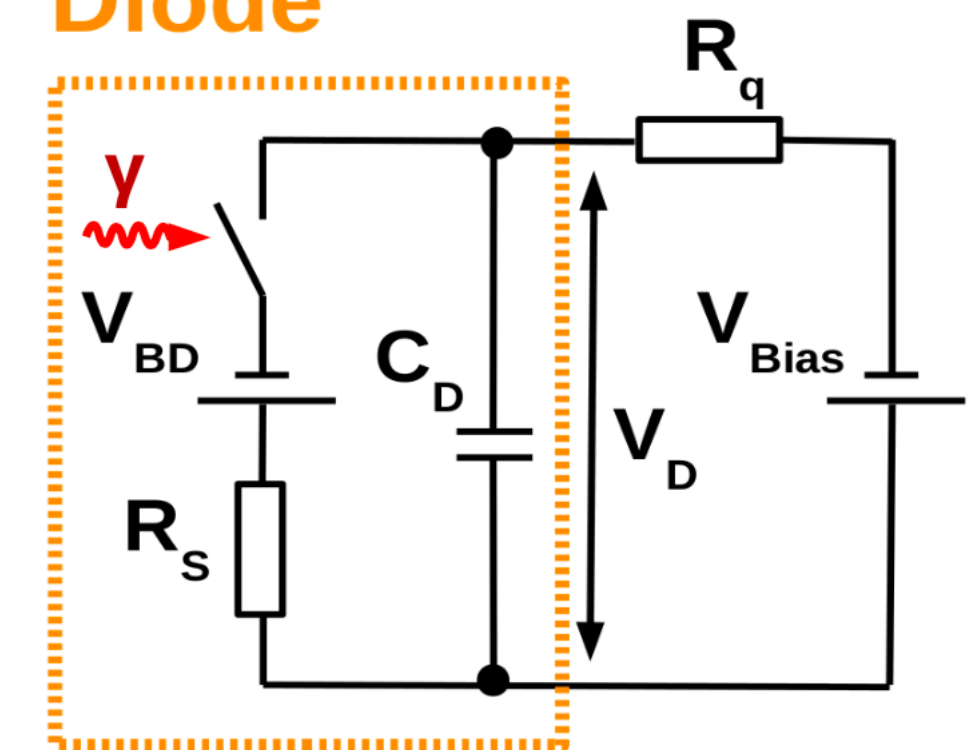


Array of pixels

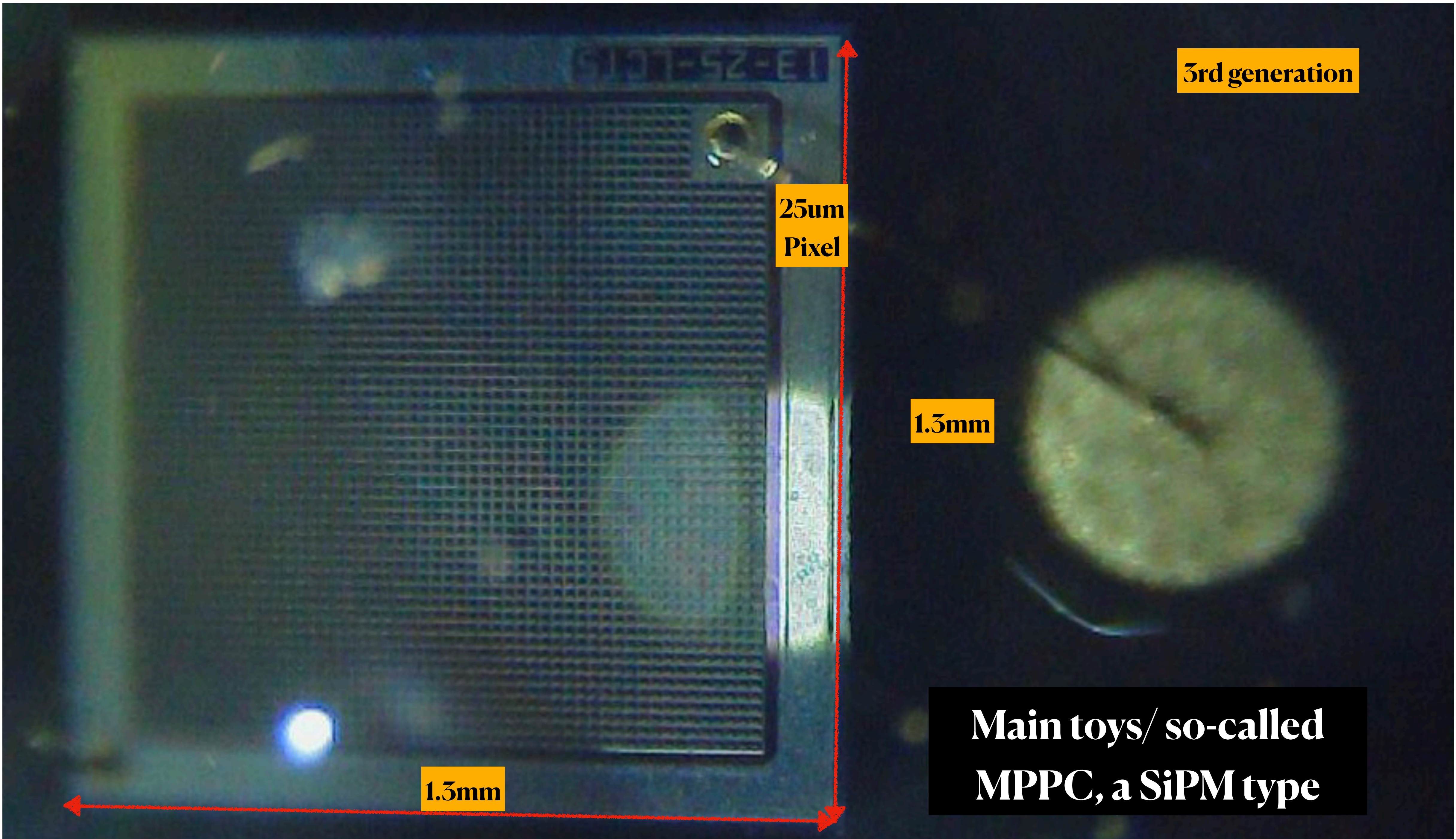
Hamamatsu S13360-1325CS



Diode



- $C_D$ : diode capacitance
- $R_S$ : silicon substrate serial resistor
- $V_{BD}$ : breakdown voltage



SI01-5Z-E1

3rd generation

25um  
Pixel

1.3mm

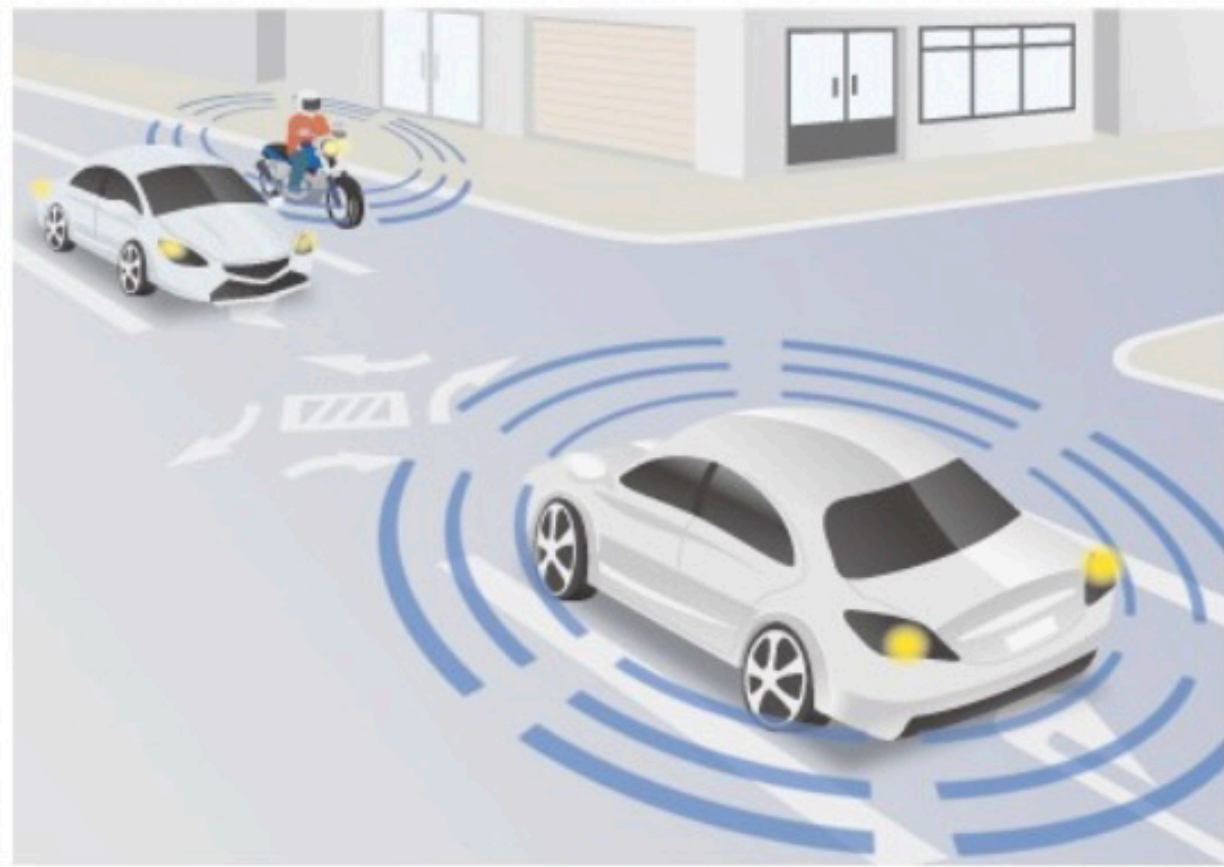
1.3mm

Main toys/ so-called  
MPPC, a SiPM type

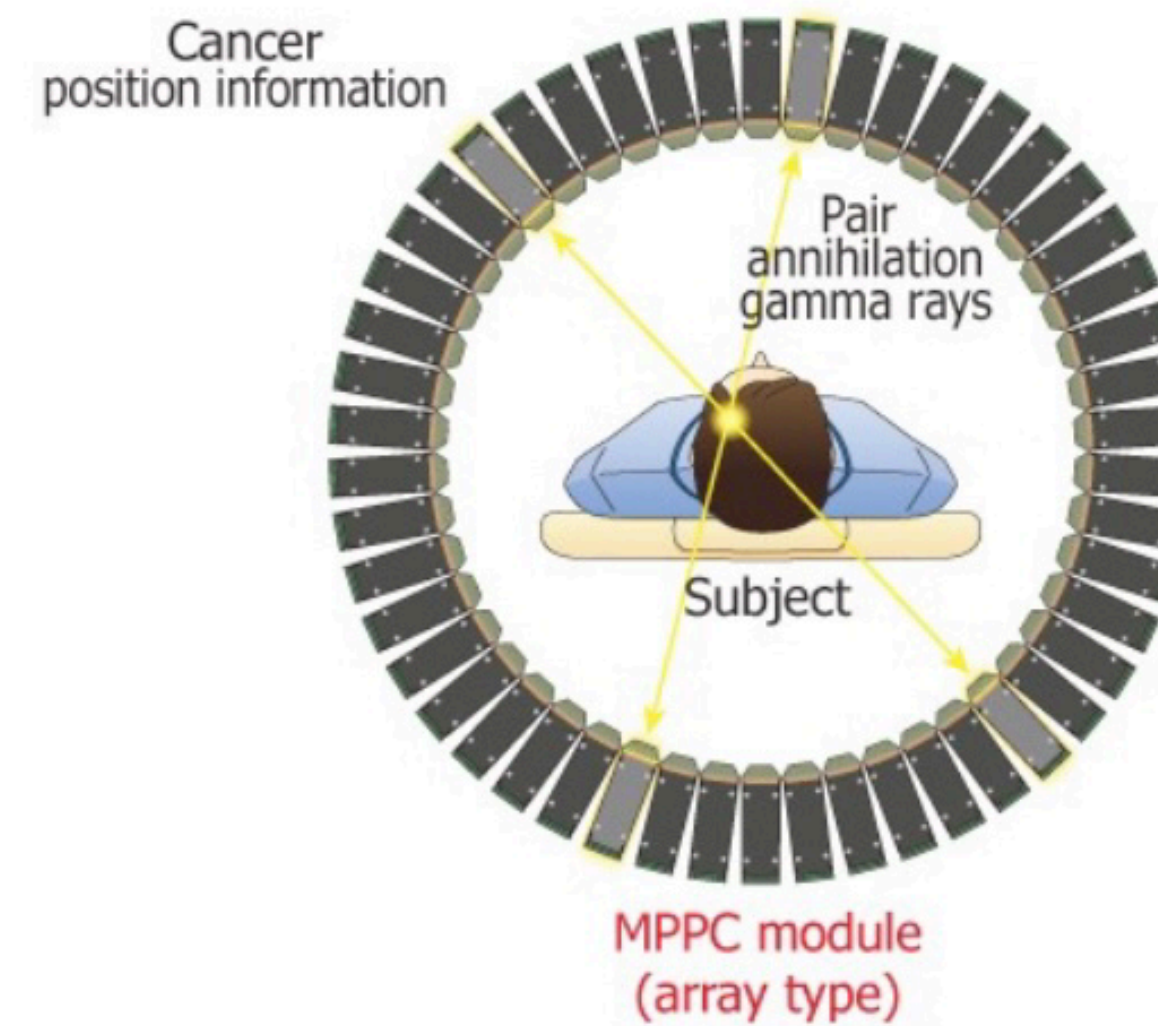
# MPPC applications

<https://www.hamamatsu.com/jp/en/product/optical-sensors/mppc/application.html>

## Distance Measurement (LiDAR)



## PET (Positron Emission Tomography)

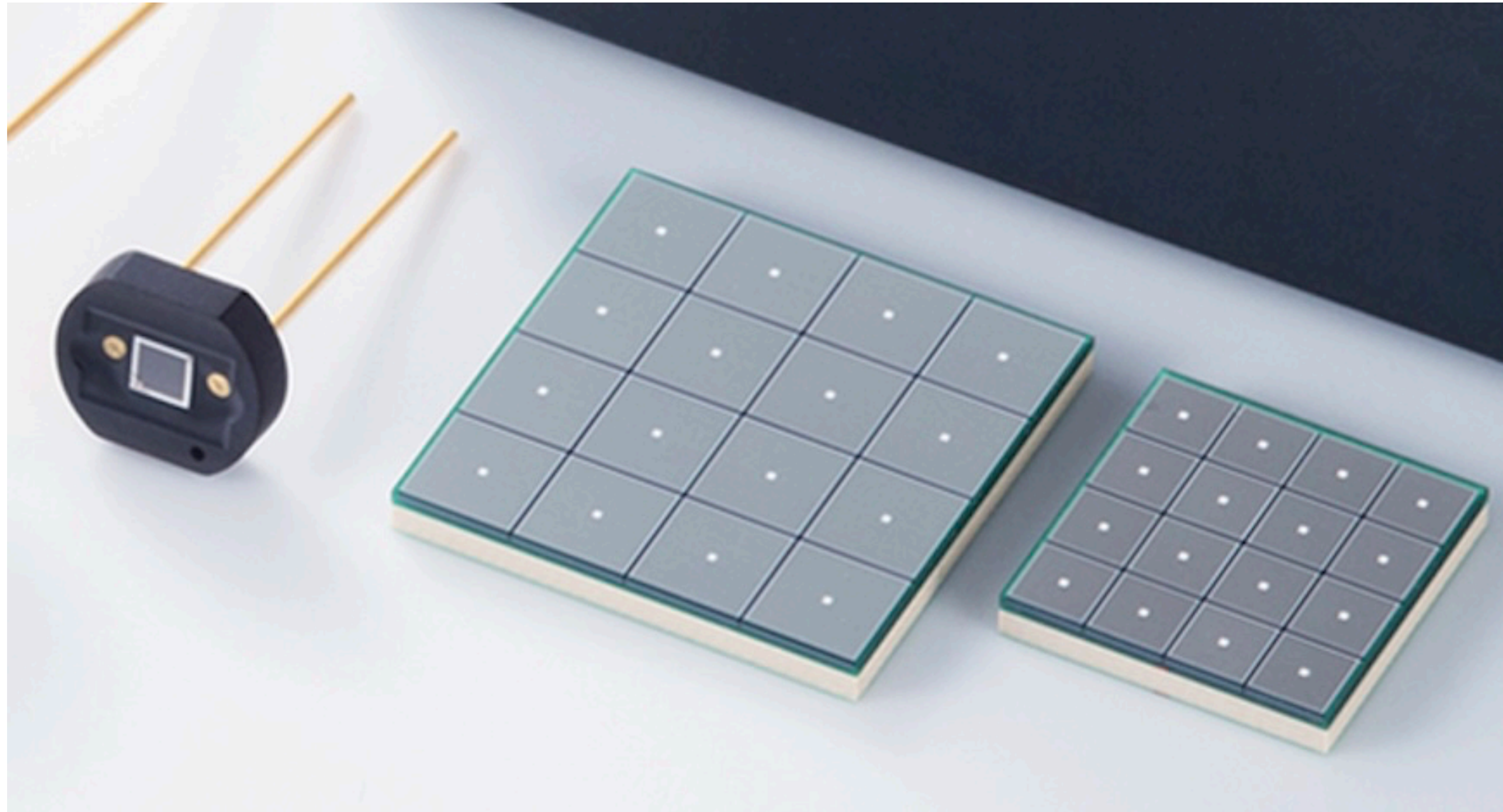


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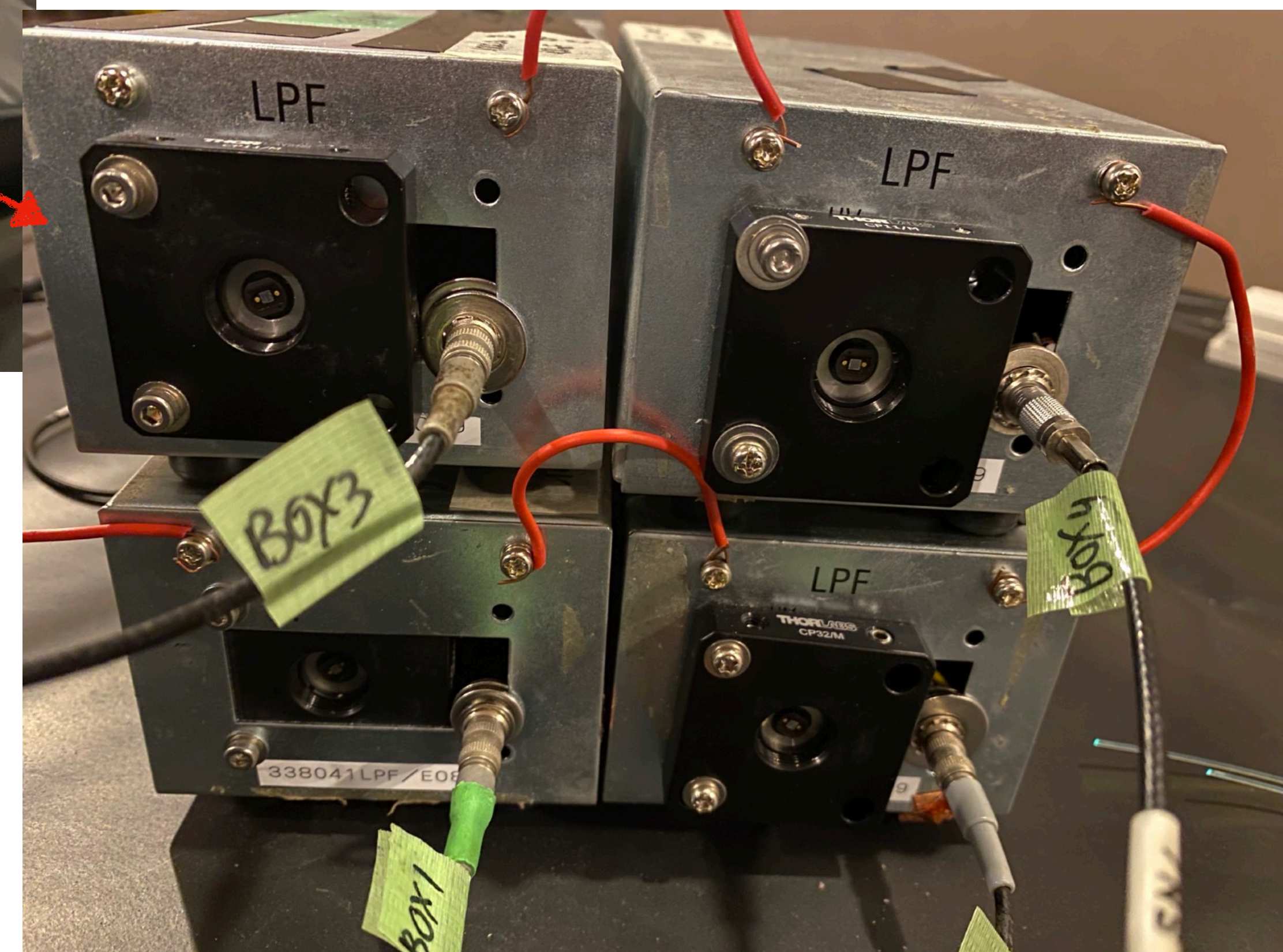
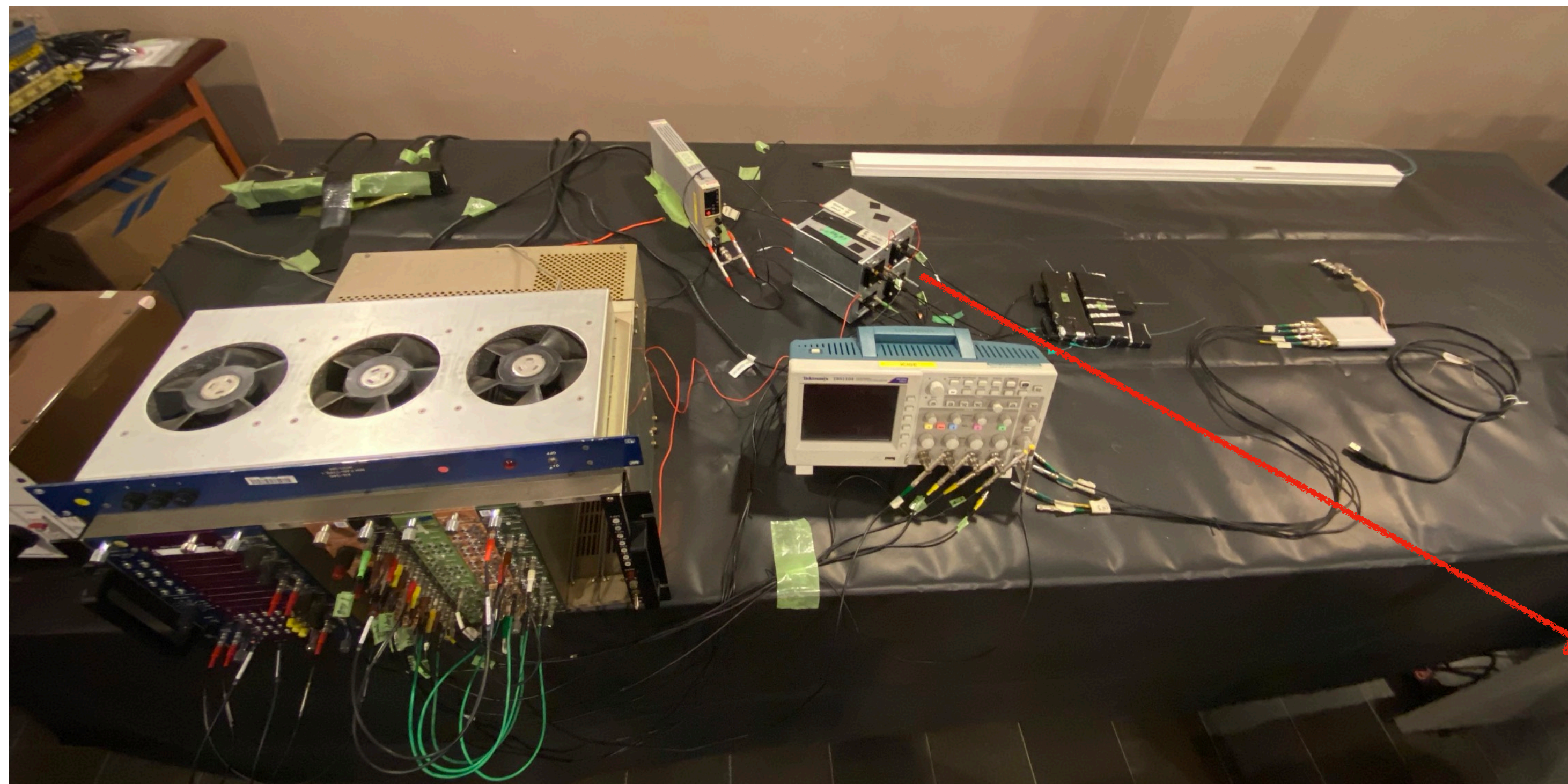
**And many other  
applications**

# We will have two setups

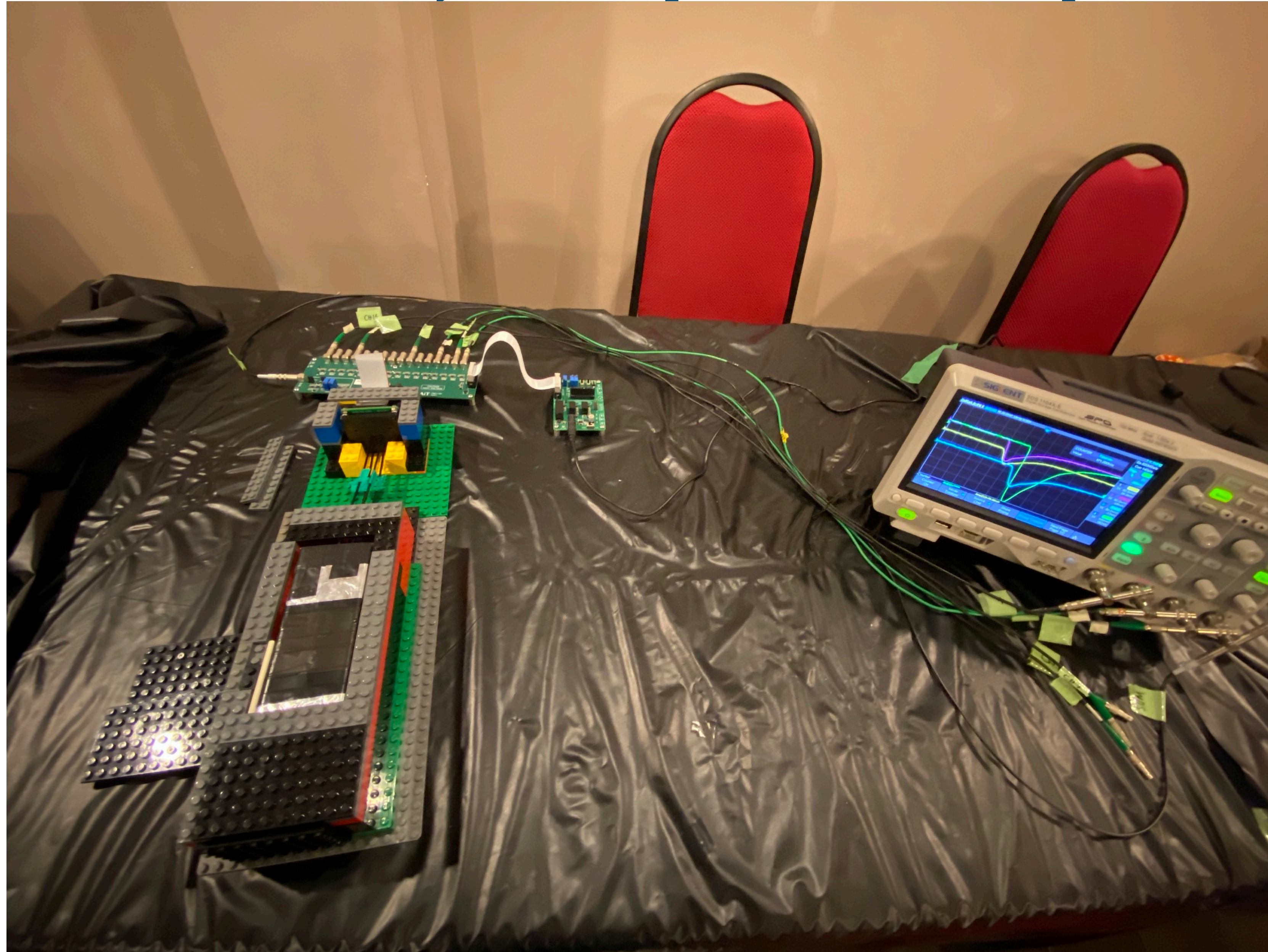
One is with single MPPC and the other is with MPPC arrays



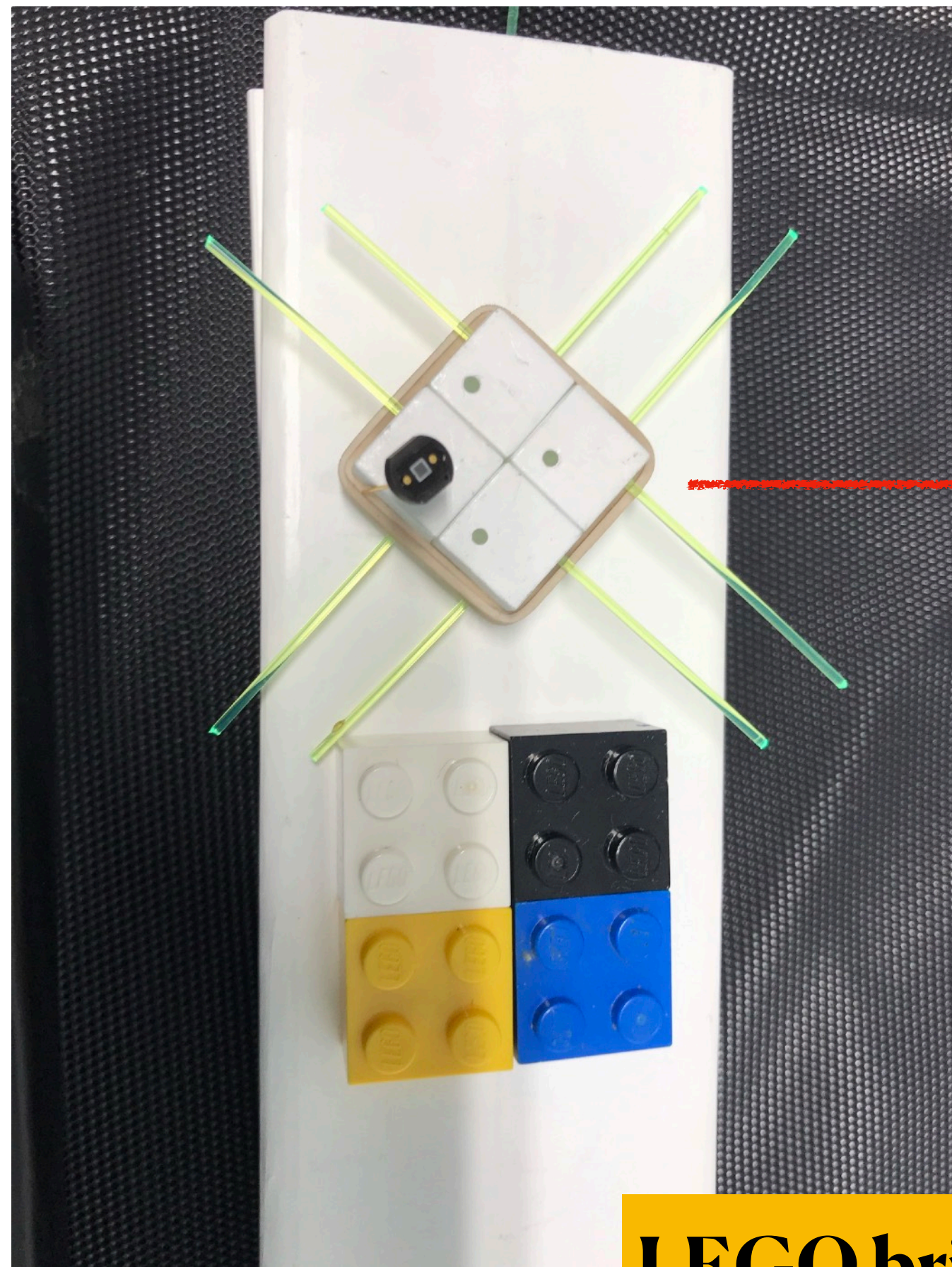
# Setup with single MPPCs (Group A and group C)



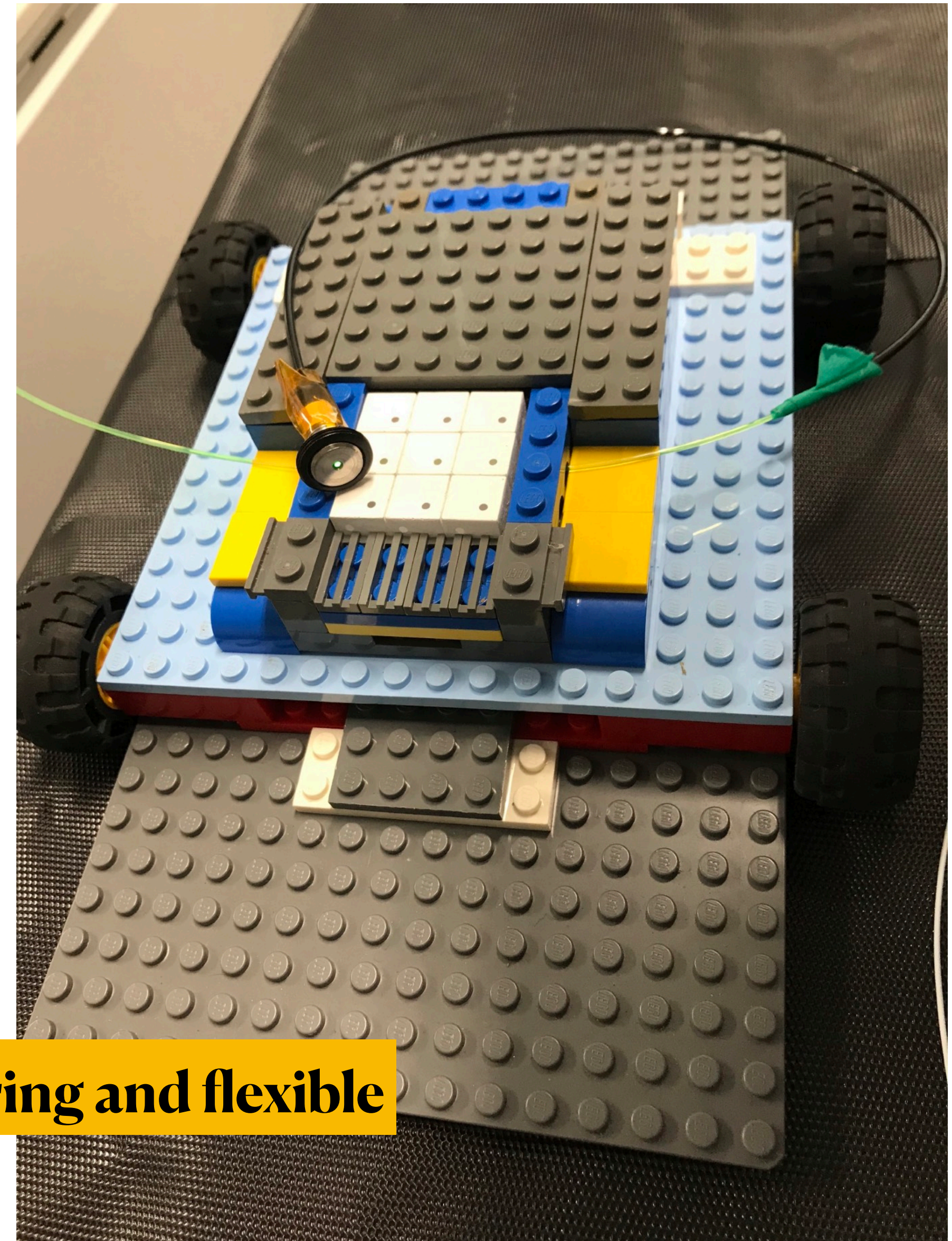
# Setup with MPPC array (Group B and Group D)



# Really LEGO is using...

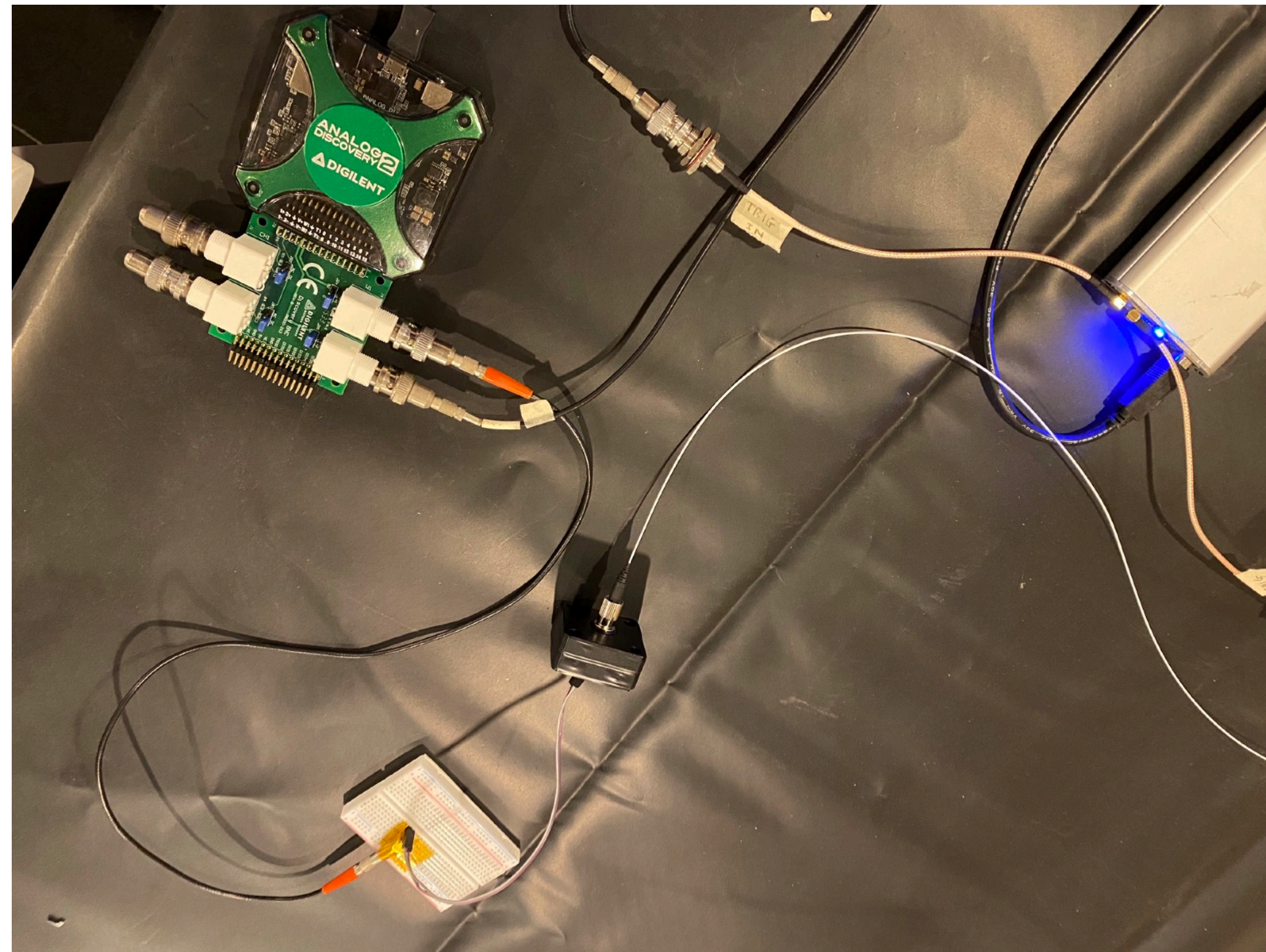


**LEGO bricks: precise manufacturing and flexible**





# Additional tools: Light manipulation



Use for group A and group C



Use for group B and group D

Can reach ~few 10ns optical pulse and level of few photons

# What we will explore (*detail will be explained on whiteboard during the training*)

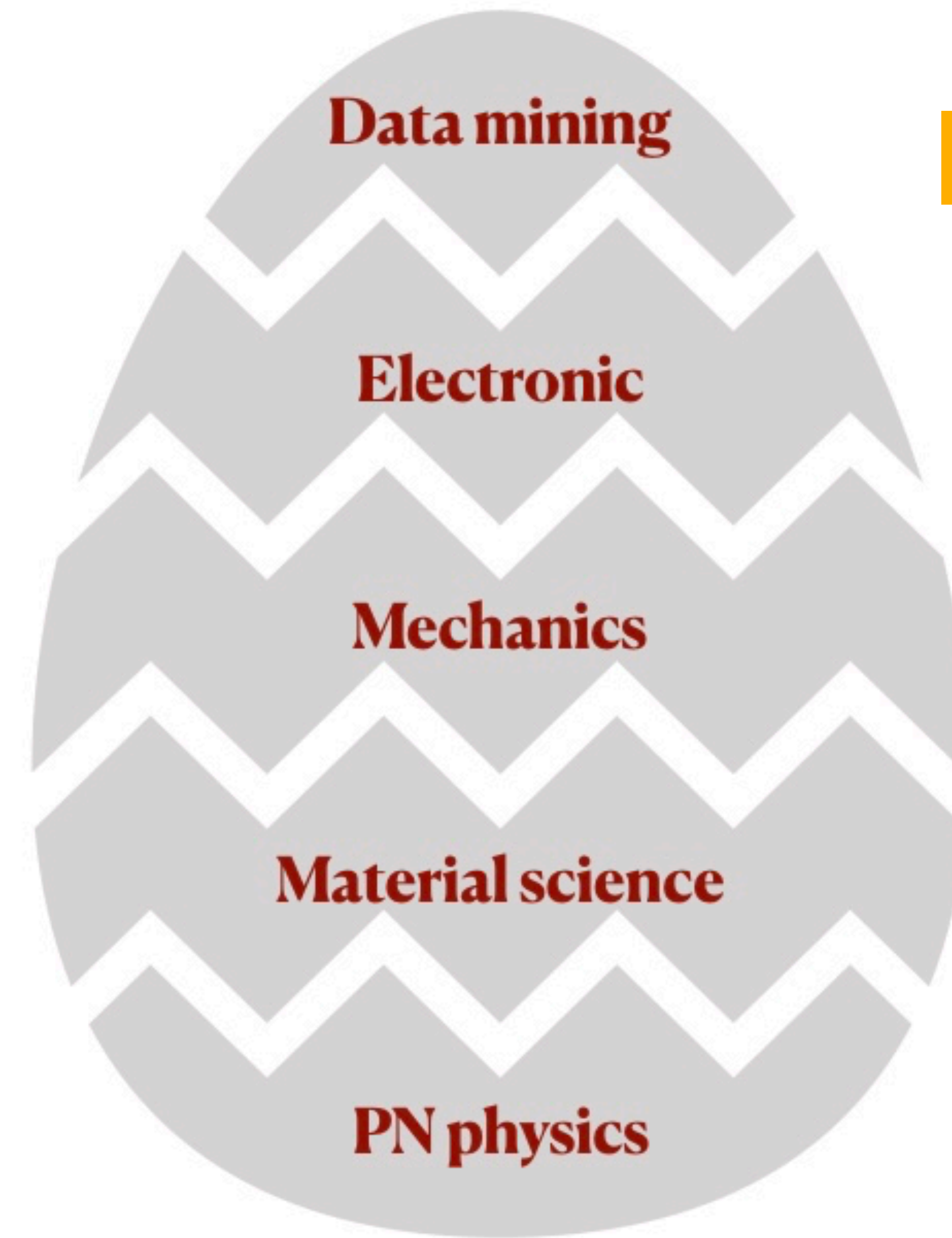
## Day 1, July 13th (Wed.)

- Familiar with hardwares
- Explore MPPC properties
  - Observe single photoelectrons w/ oscilloscope
  - Optical cross talk
  - Charge integration and electric gain calculation
- Signal processing
  - Threshold setting
  - Discriminator
  - Coincidence
  - Counter

## Day 2, July 14th (Thu.)

- Setup with scintillator and wavelength shifting fiber
- Observe cosmic ray muons
- Calculate the rate of muon (*how many triggered muon-like events per cm<sup>2</sup> of scintillator per second*)
- Compute the light yield of muons (*how many photons captured when a muon pass through 1cm thickness of plastic scintillator*)

**We will touch very small part of it.**



**“Experimental neutrino experiment in the nutshell”**

**Neutrino detection is a complicate,  
interdisciplinary field**

# Additional mentors

- Dr. Dung Nguyen
- M.Sc. Quyen Phan, B.Sc. Bao Ton
- Some students may be familiar with the setup

**Time is very limited to play with hardware. You won't satisfy, I'm sure. If you want to play more, please work with us or apply internship**

**<https://ifirse.icise.vn/nugroup/internship/index.html>**

# We thank for your donation



**KEK**



**YOKOHAMA**  
National University

**Without their generosity, this hardware training is impossible.**