

# Hyper-Kamiokande and proton decay

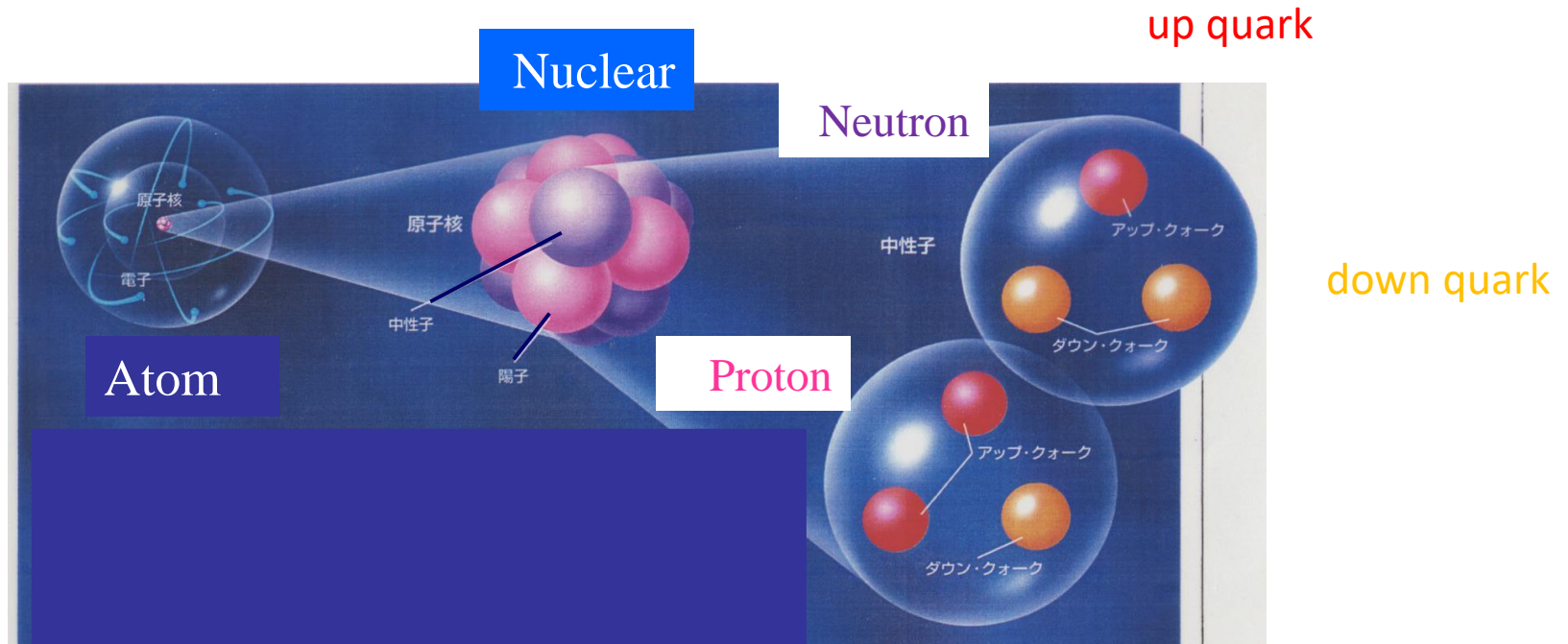
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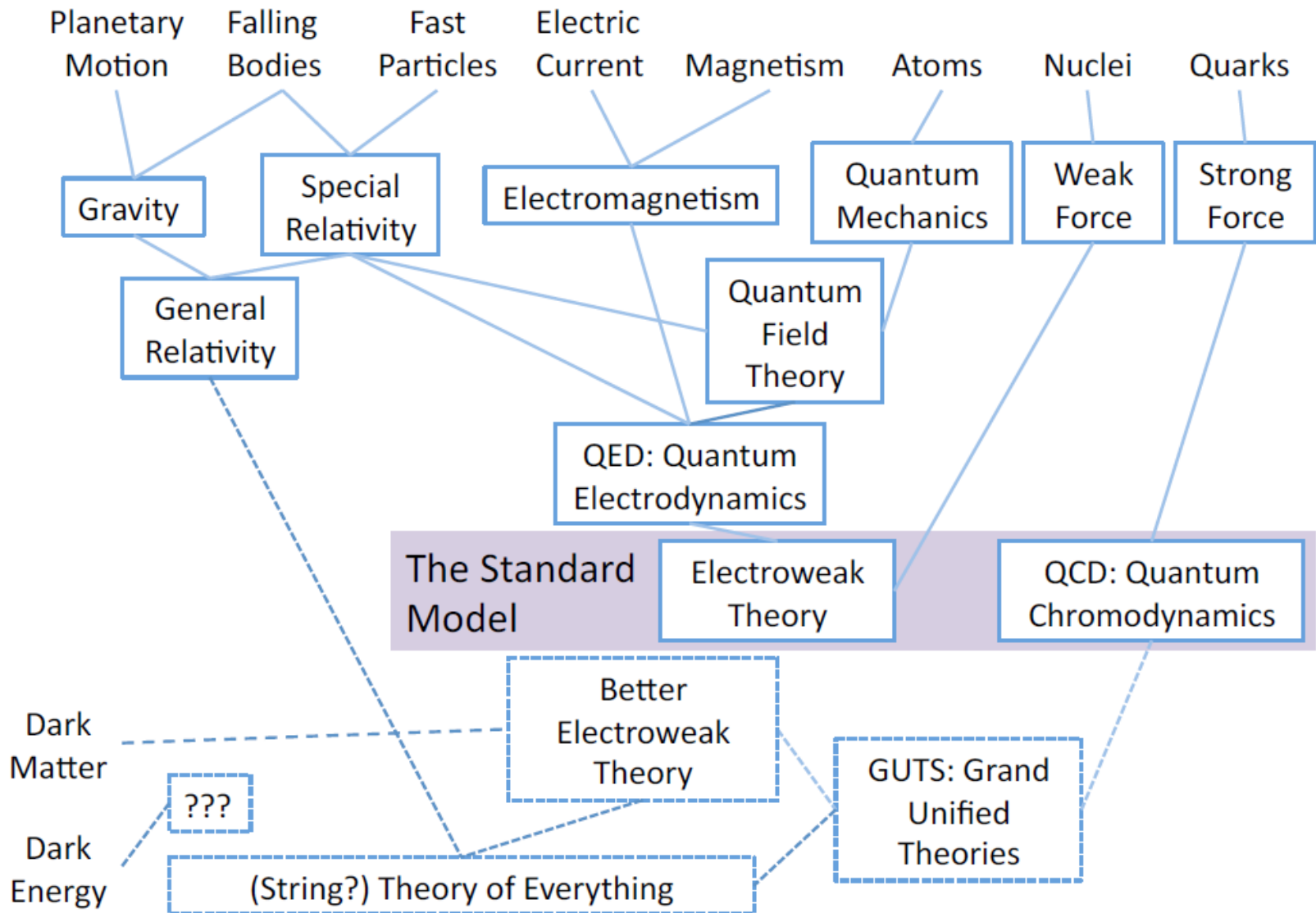
# 1. Does Proton Decay ?

- Nucleus consists of protons and neutrons.

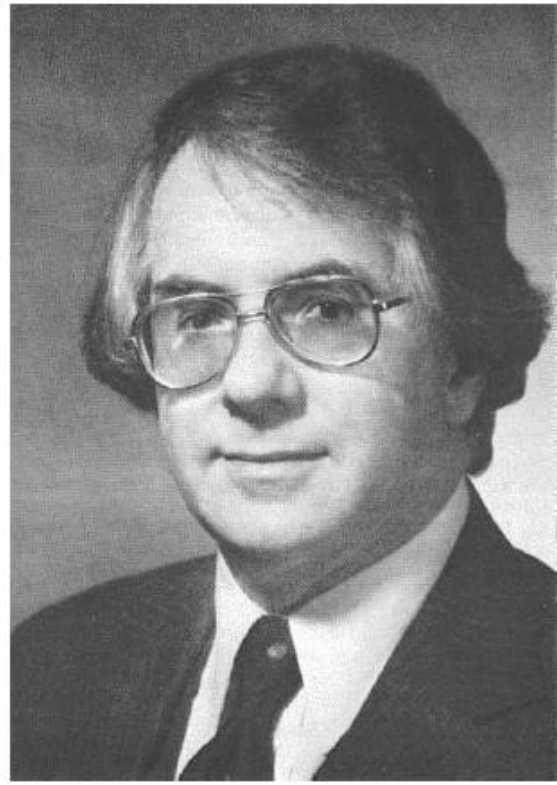


- It is well known that neutron decays spontaneously as  $\beta$ -decay:  $n \rightarrow p + e^- + \bar{\nu}_e$ 
  - Note that  $M_n > M_p$ .
- People thought proton is stable because of baryon number conservation.
  - n,p has baryon number 1.
  - We have never observed phenomena with baryon number violation.
  - Proton is the lightest baryon in the world.
  - 1929: Weyl suggests absolute stability of proton
- But is it really true ?

# Can we explain everything by a single theory ?

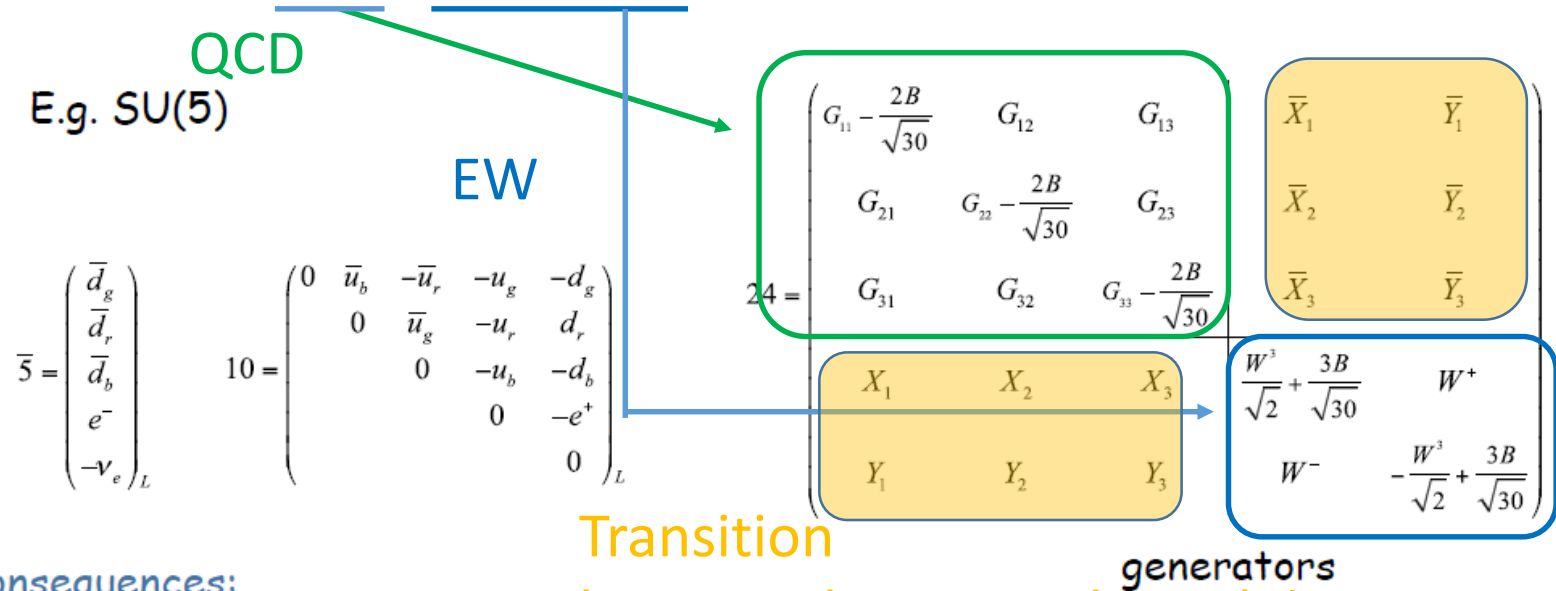


**SU(5) by Georgi and Glashow (1974)**



# Grand Unified Theories

Assume  $SU(3) \otimes SU(2) \otimes U(1)$  is part of a larger symmetry group



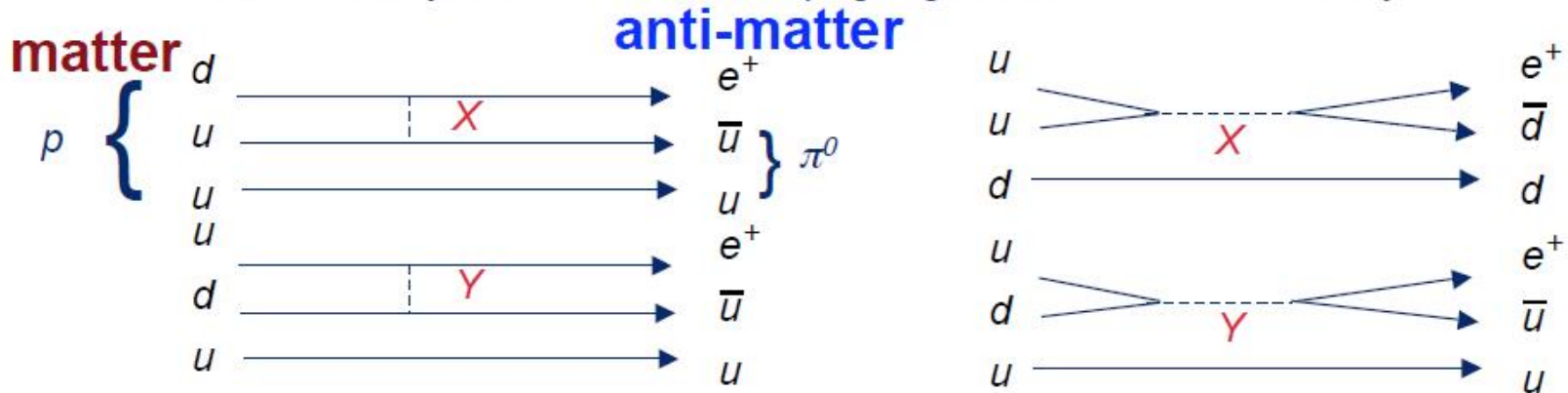
Consequences:

- ◆ Single (unified) coupling
- ◆ Charge quantization:  $Q_d = Q_e/3, Q_u = -2Q_d \Rightarrow Q_p = -Q_e$
- ◆ New gauge interactions ( $X, Y$  bosons)  $\Rightarrow$  proton decay

# Proton Decay in SU(5) By Georgi and Glashow (1974)

- Decay mechanisms

dominated by the dimension=6 op. gauge boson mediated decays



( $X^{\pm 4/3}, Y^{\pm 1/3}$  : new gauge bosons)

- Predictions

$$\tau/B (p \rightarrow e^+ \pi^0) = 4 \times 10^{29 \pm 1.7} \text{ years}, B(p \rightarrow e^+ \pi^0) \approx 40 \sim 60 \%$$

$p \rightarrow e^+ \pi^0$  became the most famous and popular decay mode.

## 2. How to find proton decay

- Watch a proton for very long time ( $> 10^{30}$  years).
  - Age of the universe:  $\sim 10^{10}$  years
  - Obviously impossible.

OR

- Watch many protons for (relatively) short time.
  - Lifetime  $\tau$ :  $N(t) = N(t=0)\exp(-t/\tau)$
  - **Need huge detector !**



# Proton Decay Detectors

IMB

H<sub>2</sub>O

Kamiokande

H<sub>2</sub>O

H<sub>2</sub>O

Super-Kamiokande

Fe

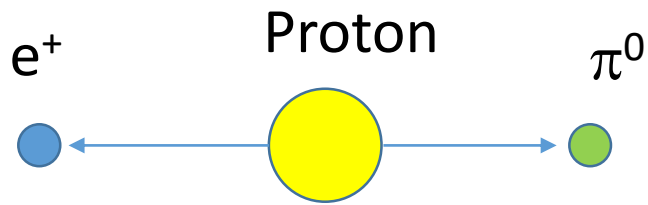
Soudan

Higher sensitivity in Water Cherenkov Detectors

2. Dominant decay mode:

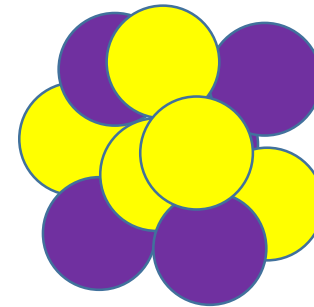
$$\rho \rightarrow e^+ \pi^0$$

# What's important for $p \rightarrow e^+ \pi^0$ ?



In "free" proton case,  $e^+$  and  $\pi^0$  emit in back-to-back. Energy corresponding to proton mass is fully used.

Nucleus

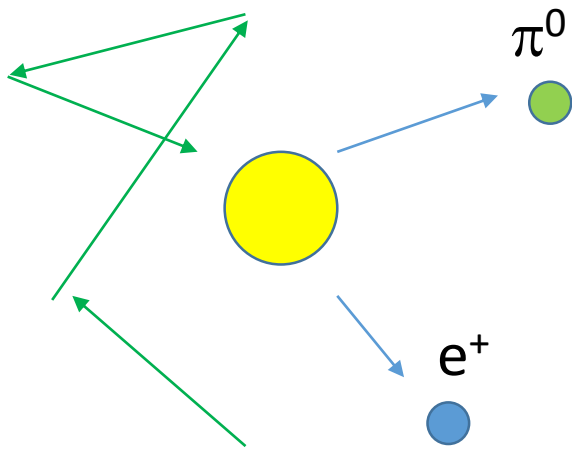


What happens if a bound proton in nucleus decays ?

Inefficiencies and uncertainties of proton decay search come from nuclear effect !

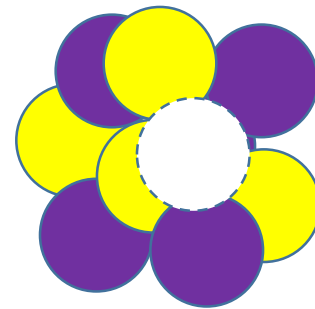
Key 1: Proton never stops in nucleus

- Protons don't exist locally in nucleus. It is always moving in the nuclear potential (Fermi motion,  $p_f \sim 225 \text{ MeV}/c$ ).



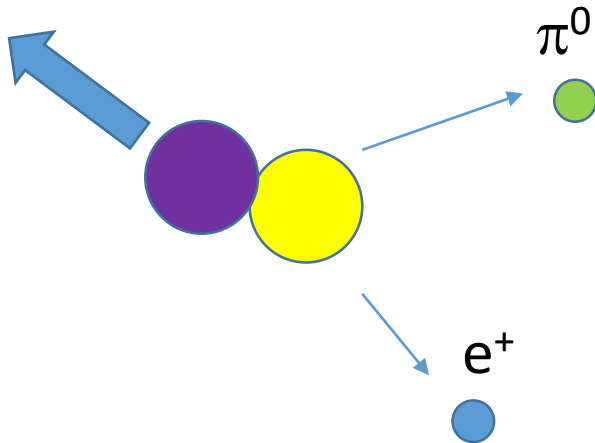
Key 2: Binding energy

- Energy corresponding proton mass should be used for compensating its binding energy (s-state:  $\sim 40 \text{ MeV}$ , p-state:  $15 \text{ MeV}$  in Oxygen).



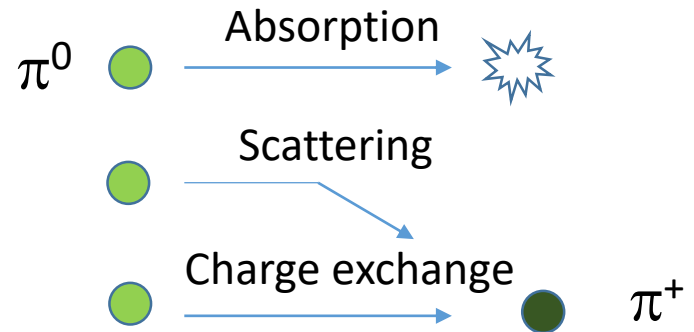
Key 3: Proton strongly binding to other nucleus

- ~ 20% protons are strongly binding to other nucleon which also bring energy when the proton decays (correlated decay)



Key 4:  $\pi$  interacts in nucleus

- Mesons ( $\pi, K, e.t.c.$ ) in decay products are affected in nuclear interactions before exiting nucleus.



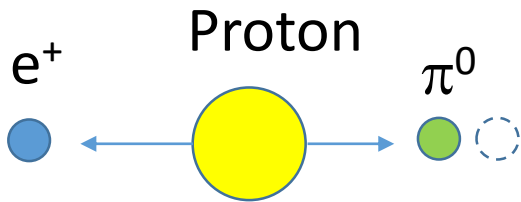
# Why water is used for proton decay search ?

- Easy to construct larger detector.
  - Much cheaper than iron.
  - You can find large water tank everywhere (common technology).
- High efficiency and low uncertainty.
  - H<sub>2</sub>O has two hydrogens which are not affected by nuclear effect . They are regarded as “free” proton.
    - ✓ Bound proton:  $\sim 200 \text{ MeV}/c \sim 0.2c = 6 \times 10^7 \text{ m/s}$
    - ✓ velocity of molecular in liquid  $\sim 10^2 \text{ m/s}$
  - Free protons contribute high selection efficiency and low uncertainty.

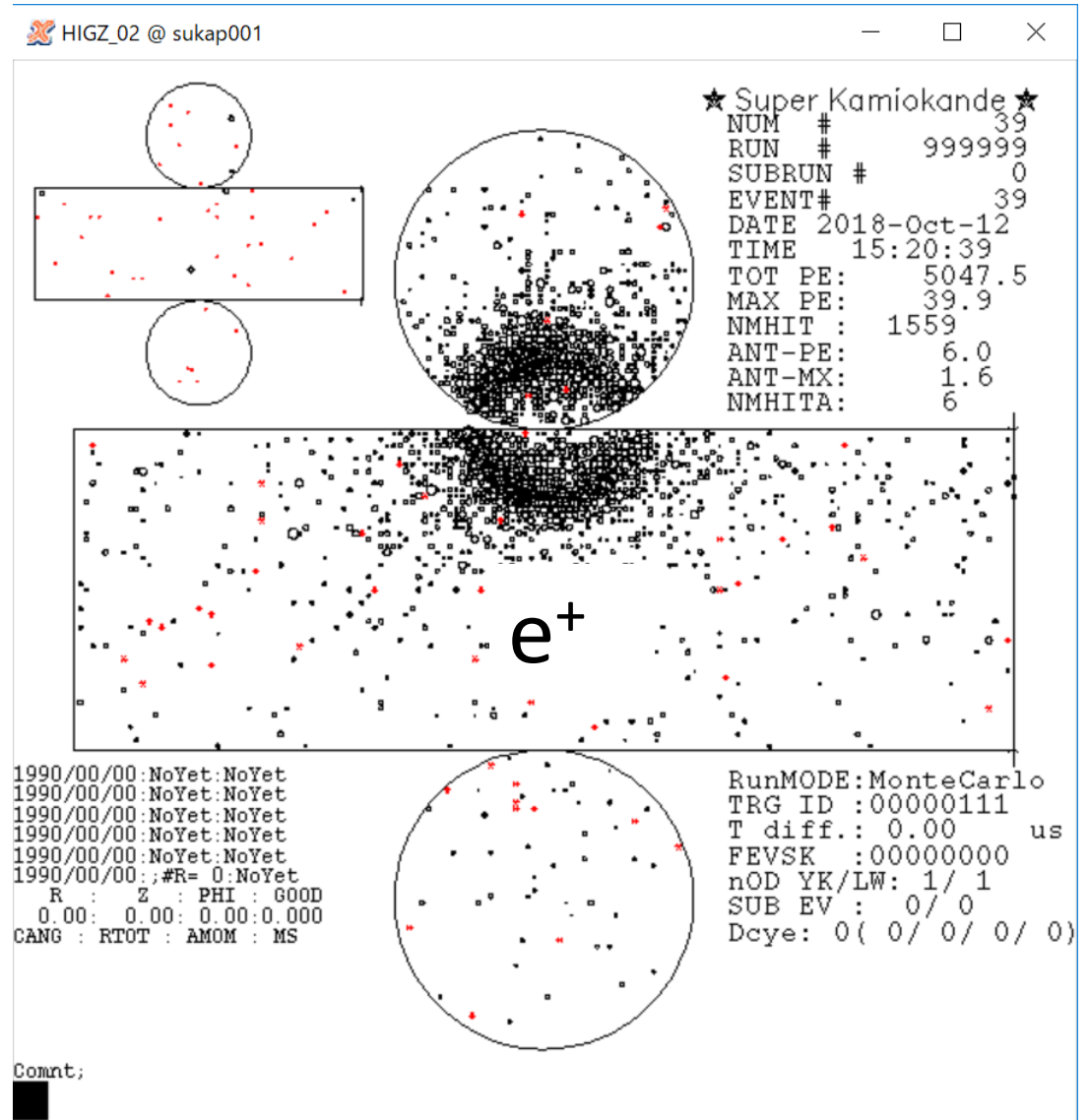




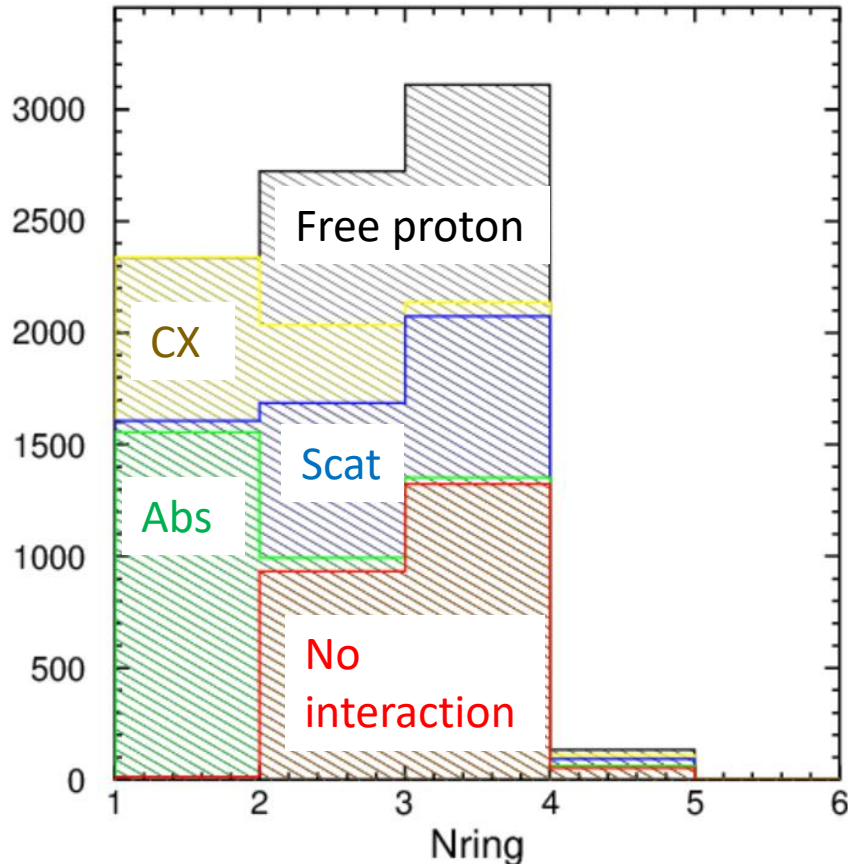




If  $\pi^0$  is absorbed before exiting nucleus, only  $e^+$  is observed (one ring).



# Observed number of ring for $p \rightarrow e^+ \pi^0$



Free proton: H in H<sub>2</sub>O

No interaction in Nucleus

Abs:  $\pi^0$  absorption in Nucleus

Scat: scattered

CX: charge exchange

( $\pi^0 \rightarrow \pi^\pm$ , below threshold)

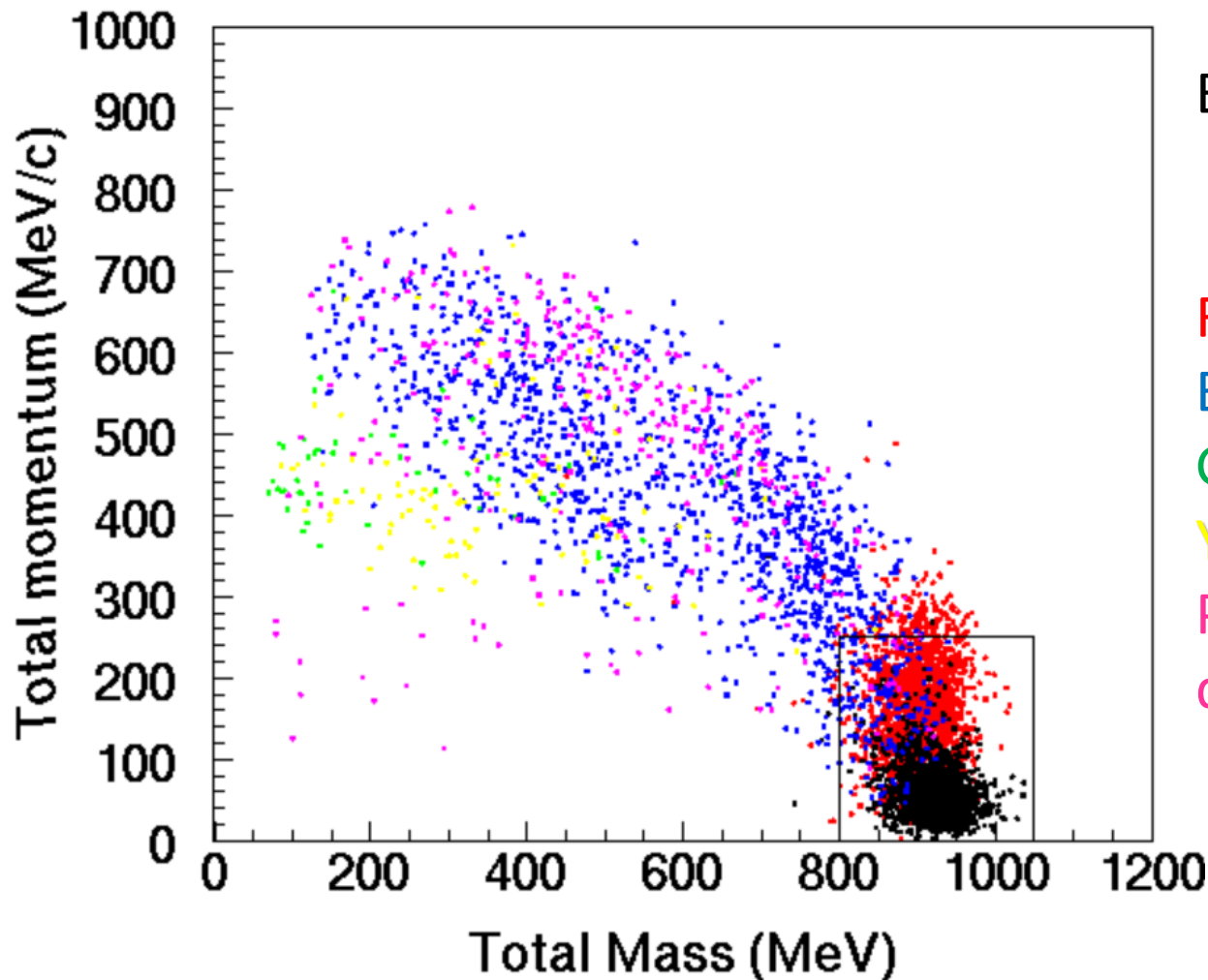
Choose 2 or 3 rings.

# Selection criteria for $p \rightarrow e^+ \pi^0$

1. Event vertex should be located 2 m inward from the tank wall (fiducial volume cut, [22.5kton](#)).
2. 2 or 3 ring event.
3. All ring should be e-like (Particle IDentification).
4. No Michel electrons.
5. Reconstruct  $\pi^0$  mass for 3 ring events. It should be  $85 < M_{\pi^0} < 185 \text{ MeV}/c^2$
6. Reconstruct total mass and momentum should be  $800 < M_{\text{tot}} < 1050 \text{ MeV}/c^2$ ,  $P_{\text{tot}} < 250 \text{ MeV}/c$ .

# Total mass vs Total momentum for $p \rightarrow e^+ \pi^0$

- Selection efficiency  $\sim 40\%$
- Inefficiency is dominated by unavoidable physics processes.



Black: Free proton

Brown: Bound proton

Red: No  $\pi$  interaction

Blue:  $\pi$  scatter

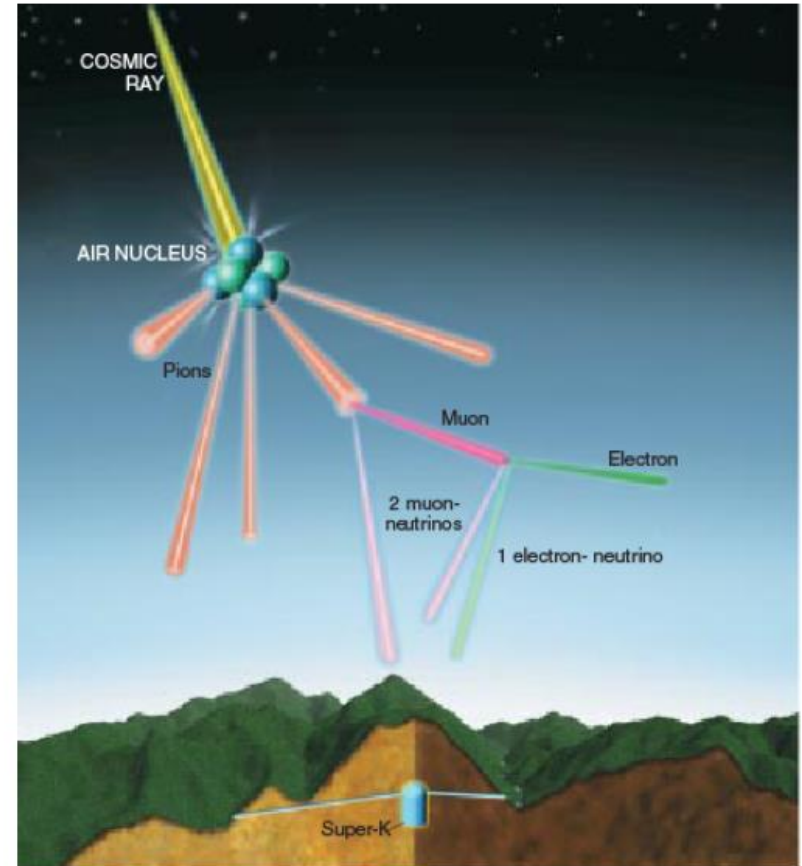
Green:  $\pi$  absorption

Yellow:  $\pi$  CX

Purple: Correlated decay

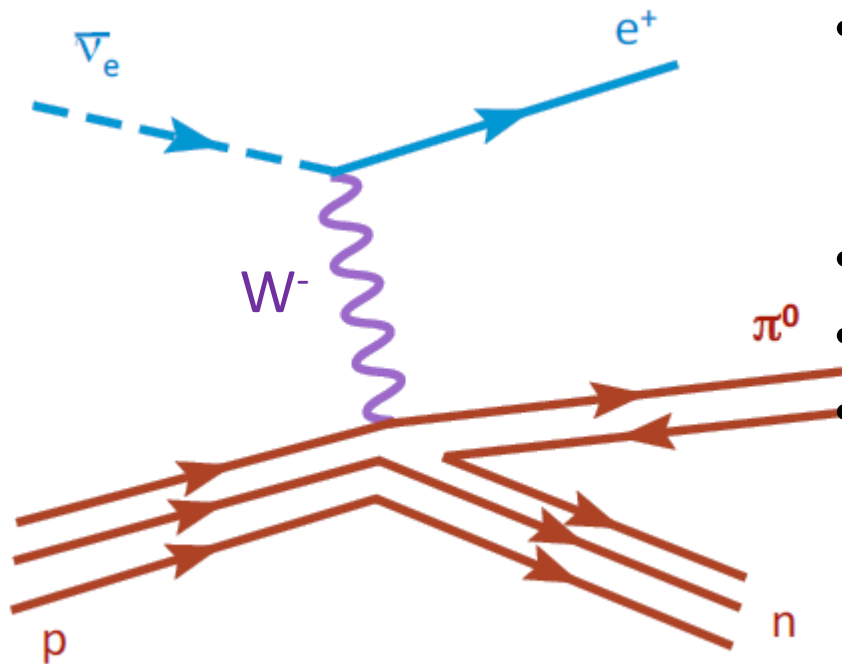
## 2-2. What's Background events for proton decay searches ?

- **Atmospheric neutrino** is dominant backgrounds for proton decay searches.
  - Visible energy  $\sim 1$  GeV.
  - Solar or SN  $\nu$  is too low energy.
  - Cosmic ray  $\mu$  are rejected by outer detector.



# Typical background for $p \rightarrow e^+ \pi^0$

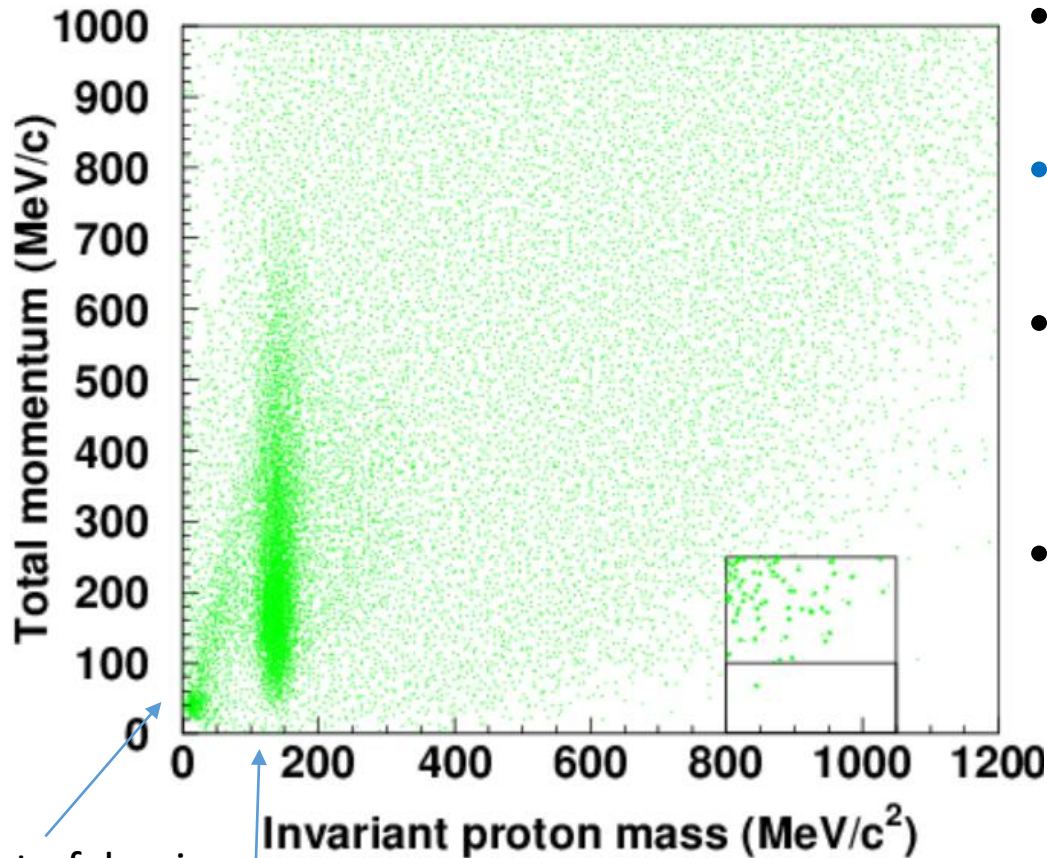
## Charged current $\pi^0$ production



- Exchange  $W$  boson between  $\nu$  and proton (charged current interaction).
- $\nu$  changes to  $e^+$ .
- $\pi^0$  and neutron are produced.
- Because neutron doesn't emit Cherenkov light, visible particles after the reaction are same as  $p \rightarrow e^+ \pi^0$

# Total mass vs Total momentum for atmospheric $\nu$ background MC

(After all cuts except for total mass and momentum)

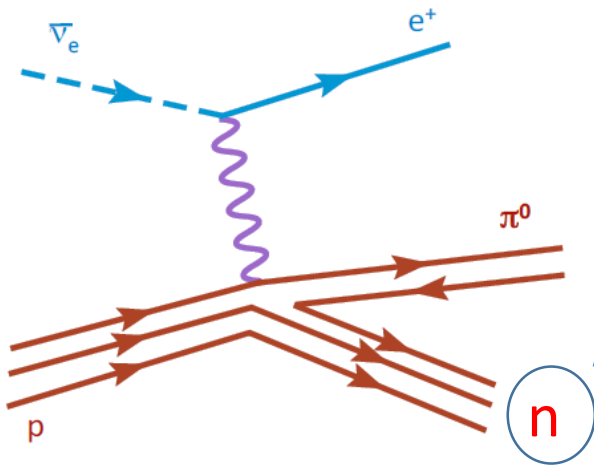


Due to fake ring

$\pi^0$  due to  $2Re+e$  is allowed.

- Generate huge atm.n MC, 2000 year of SK!
- Expected BG:  
 $\sim 1.3 \text{ ev/Mton*yr}$
- Neutrino events tend to have higher total momentum.
- Almost background free in lower momentum region ( $< 100 \text{ MeV/c}$ ).
  - The region corresponds to free proton decay.

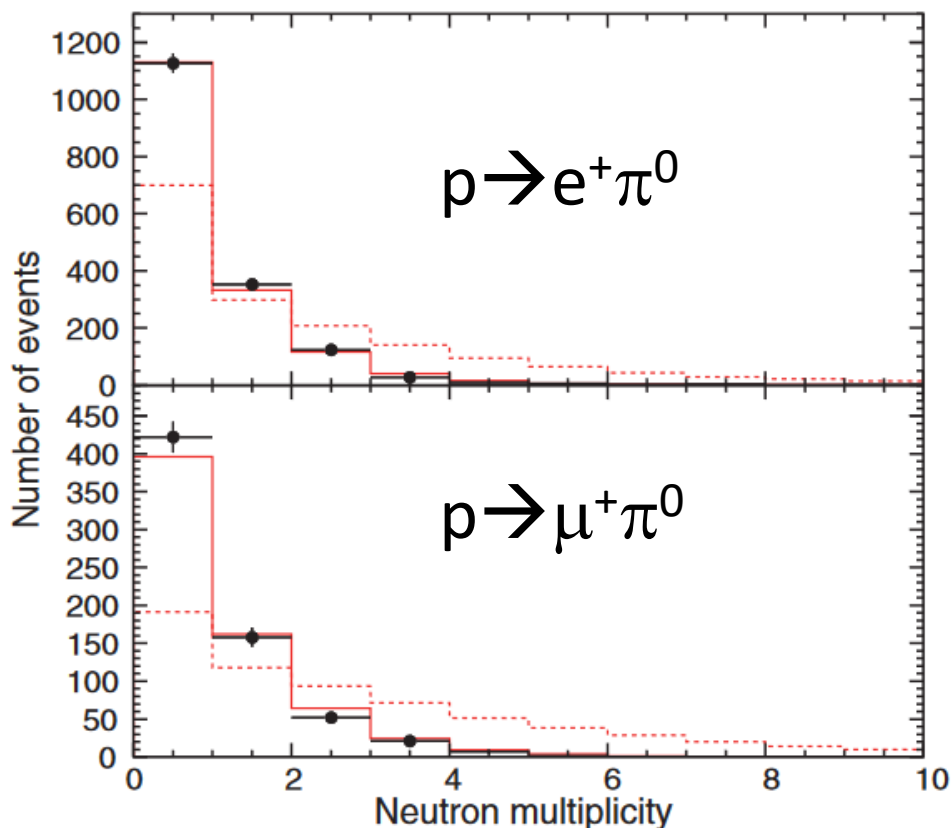
# Further background reduction



- Neutron doesn't emit Cherenkov light.
- However, neutron is thermalized in water and finally captured by hydrogen ( $\sim 200 \mu\text{s}$ );  
$$n + p \rightarrow d + \gamma (2.2 \text{ MeV})$$
- If we can detect delayed 2.2 MeV  $\gamma$  ray, we can reduce background more.
- Neutron capture is also important for SN Relic  $\nu$  and separate  $\nu$  and  $\bar{\nu}$  interactions in atmospheric n oscillation analysis.



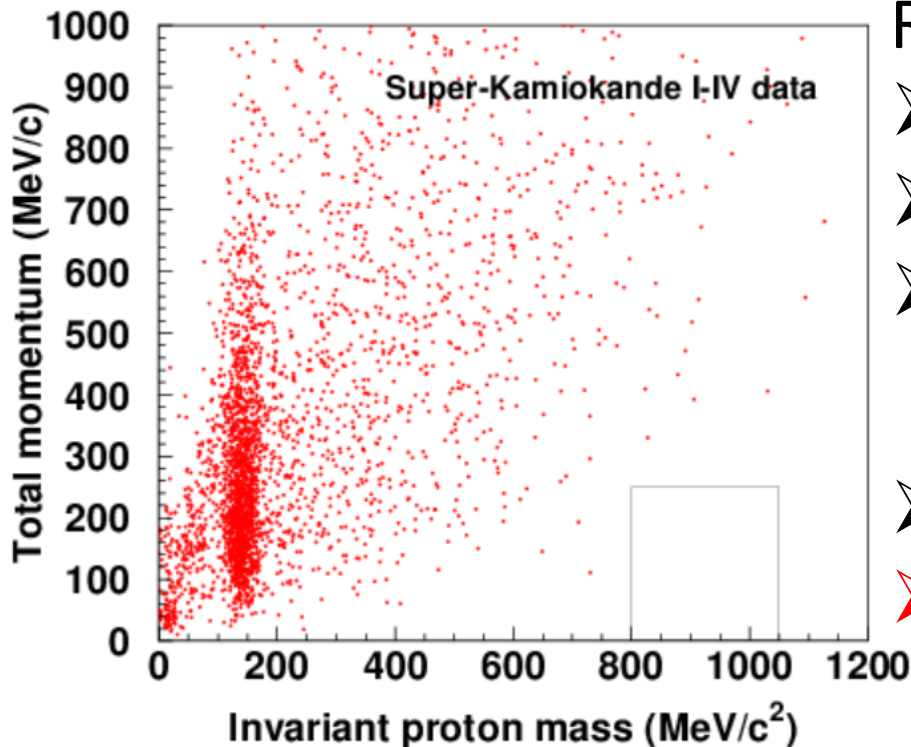
# How powerful to reject background



- Sample: out of signal box in  $M_{\text{tot}}$  vs  $P_{\text{tot}}$  plot.
  - Dot: data,
  - Histogram: Atm.  $\nu$  MC (solid: reconstructed, dash: true)
- $\sim 50\%$  background events are rejected with neutron=0.
- On the other hand,  $\sim 7.5\%$  of  $p \rightarrow e^+ \pi^0$  are accompanied with neutron from deexcitation of nucleus. Neutron tagging reduces a few % in selection efficiency.

## 2-3. SK result (so far)

- We have not find any evidences of nucleon decays !



Result of  $p \rightarrow e^+ \pi^0$

- Exposure: 450 kton · year
- Efficiency: 38.6 % (SK-IV)
- Expected BG: 0.63 events
  - 0.05 evetns in  $P_{\text{tot}} < 100 \text{ MeV/c}$
  - 0.58 events in  $100 \sim 250 \text{ MeV/c}$
- Observed: 0 event
- Lower limit of proton life time:  $> 2.4 \times 10^{34}$  years

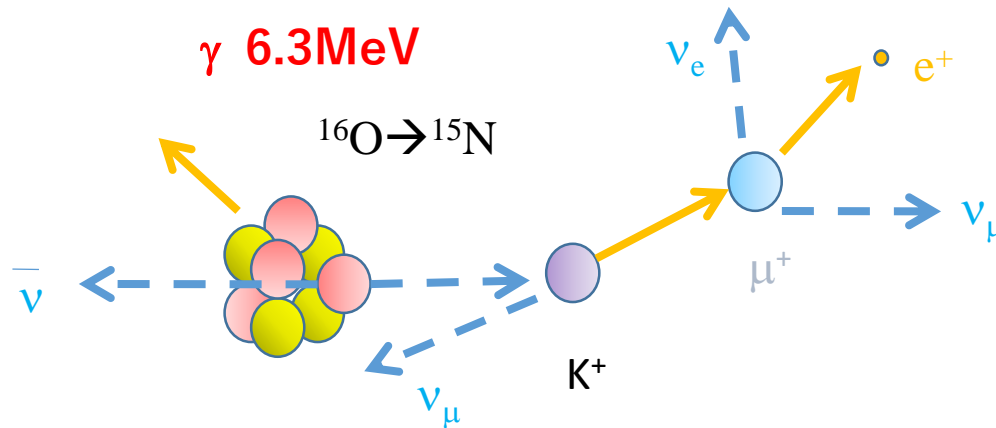
3. SUSY favored decay mode:

$$p \rightarrow \nu K^+$$

# 3-1 How to find $p \rightarrow \nu K^+$ in Water Cherenkov detector

- $K^+$  has low momentum, most of them **stop in water** and decay with 12 nsec lifetime.
- Major  $K^+$  decay mode
  - $K^+ \rightarrow \nu\mu^+$  : 64 %
  - $K^+ \rightarrow \pi^+\pi^0$ : 21 %
- “Stopping  $K^+$ ” means **two body decay products of  $K^+$  should have monochromatic momentum** as seen in the previous question !
  - $K^+ \rightarrow \nu\mu^+$  : 236 MeV/c
  - $K^+ \rightarrow \pi^+\pi^0$ : 206 MeV/c
- Using this property, Water Cherenkov detector can search for  $p \rightarrow \nu K^+$  .

## 3-2. Search for $p \rightarrow \nu K^+$ , $K^+ \rightarrow \nu \mu^+$



- Visible particle is only  $\mu^+$  with Michel electron.
- Search for data excess around 236 MeV/c of  $\mu$  comparing with atmospheric  $\nu$  MC.
- After proton decay, **40 % of remaining nucleus emits 6 MeV  $\gamma$  for deexcitation.** It is useful to reduce background.

# Example of $p \rightarrow \nu K^+$ , $K^+ \rightarrow \nu \mu^+$ with $\gamma$

## Super-Kamiokande IV

Run 999999 Sub 0 Event 69  
D\_wall: 1165.1 cm  
Evis: 53.2 MeV  
mu-like,  $p = 231.0$  MeV/c

### Resid (ns)

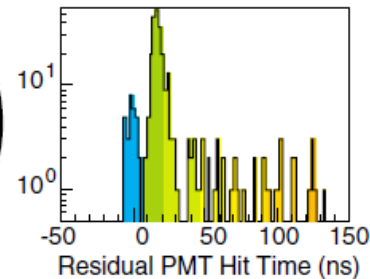
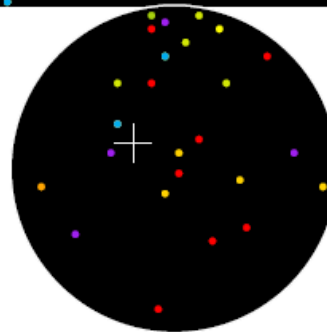
- > 182
- 160- 182
- 137- 160
- 114- 137
- 91- 114
- 68- 91
- 45- 68
- 22- 45
- 0- 22
- -22- 0
- -45- -22
- -68- -45
- -91- -68
- -114- -91
- -137--114
- <-137

$\mu$

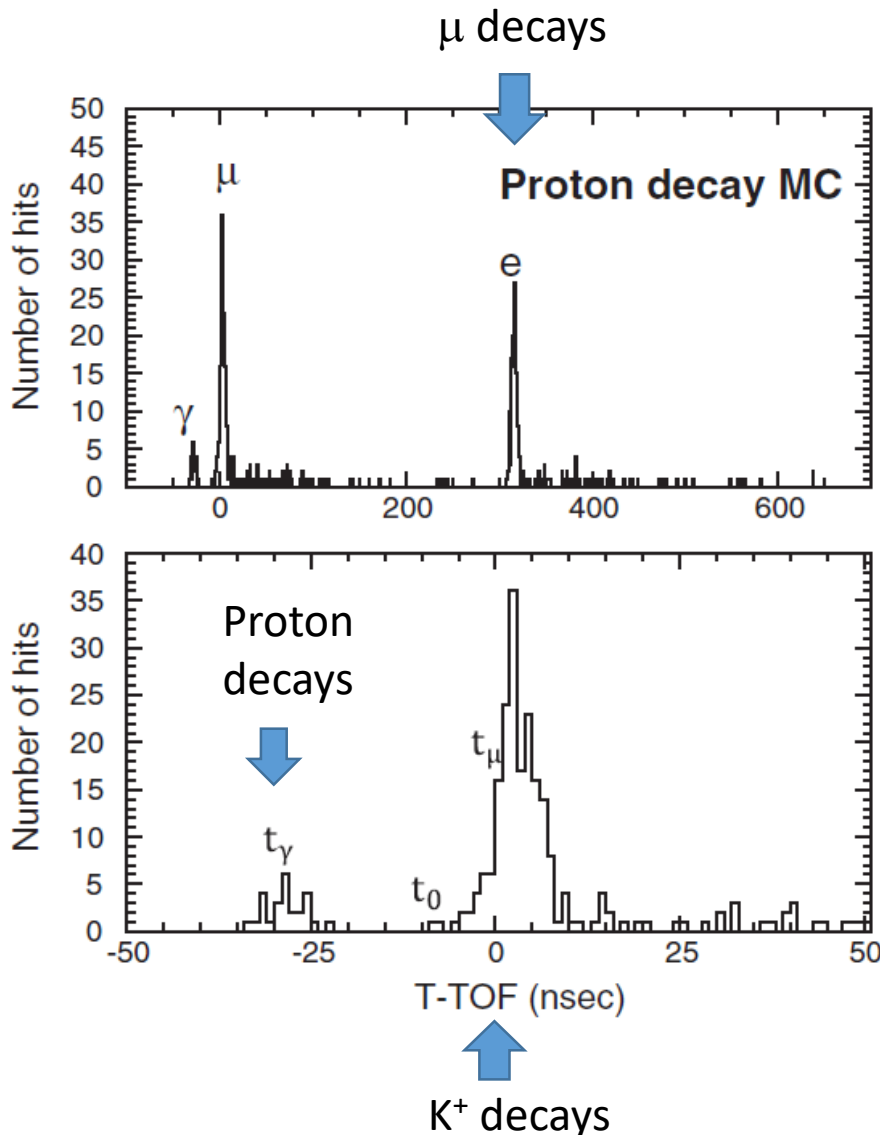
Color of hits corresponds to time.

Cyan corresponds to nuclear  $\gamma$ .

Difficult to identify  $\gamma$  from hit pattern.




# Time structure with nuclear $\gamma$



- 3 hit clusters in time should be observed in case of signal.
- The event is triggered by  $\mu$  hits.
- $\gamma$  signal is much smaller than  $\mu$  and easily hidden by tail of  $\mu$  hits.
- Make 12 nsec time window and slide it toward left from  $t_0$  (end of  $\mu$  tail) to search for maximum hit cluster.

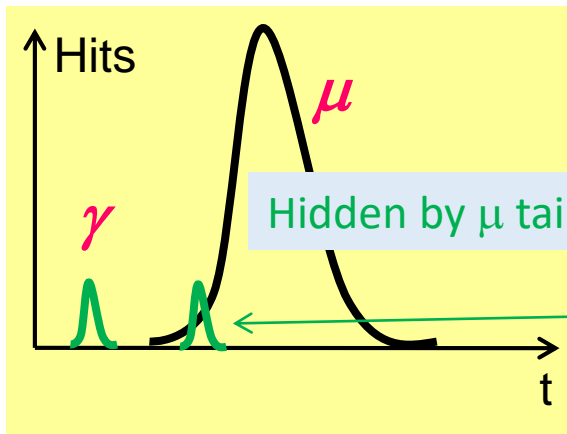
# Selection criteria for $p \rightarrow \nu K^+$ , $K^+ \rightarrow \nu \mu^+$

- 1  $\mu$ -like ring with Michel electron
- $215 < P_\mu < 260$  MeV/c
- Proton rejection cuts
- Search Max hit cluster  Reduce background by  $5 \times 10^{-4}$  !  
by sliding time window (12ns width);
  - $4 < N_\gamma < 30$  hits
  - $T_\mu - T_\gamma < 75$  nsec
- No neutron
- Selection efficiency = (selected events)/(proton decay in fiducial volume):  
9 %
  - $\text{Br}(K^+ \rightarrow \nu \mu^+) = 64$  %, only 40 % emits nuclear  $\gamma \rightarrow 26$  %  
even if detector is perfect.

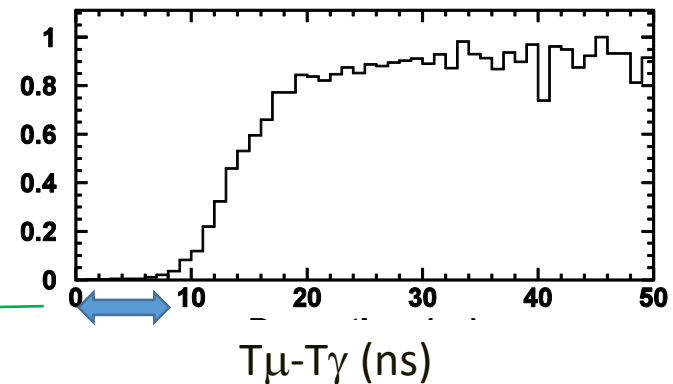


# Remark for this analysis

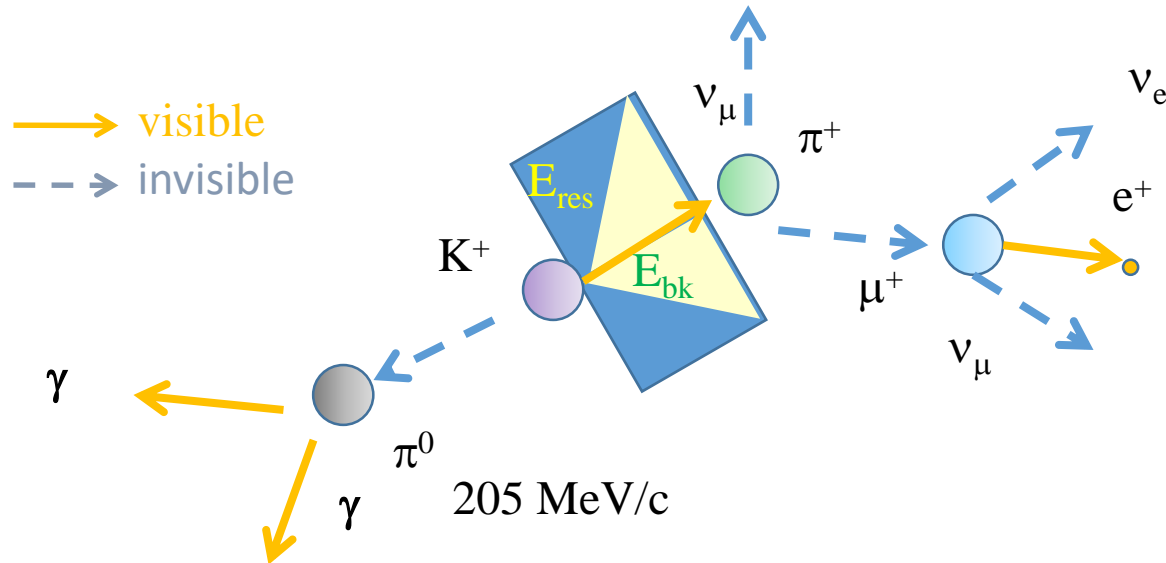
- This analysis is limited by time resolution of PMTs.
  - If  $\gamma$  is close to  $\mu$ ,  $\gamma$  peak is hidden by  $\mu$  hits.
  - Time resolution of SK PMT is 2.2 nsec at 1 photoelectron.
  - If  $\mu$  peak becomes sharper, the selection efficiency will be improved.



$\gamma$  tagging efficiency



### 3-3. Search for $p \rightarrow \nu K^+$ , $K^+ \rightarrow \pi^+ \pi^0$



- Both  $\pi^+$  and  $\pi^0$  has **205 MeV/c** in momentum. This is just above Cherenkov threshold for  $\pi^+$ , thus it is not identified as a ring in most of case.
- $\pi^+$  decays into  $\mu$  (invisible) and  $\nu$ ,  $\mu$  decays into  $e \nu_e \nu_\mu$ .
- $\pi^0$  decays into 2  $\gamma$ s.
- Search for 206 MeV/c  $\pi^0$  with Michel electron.

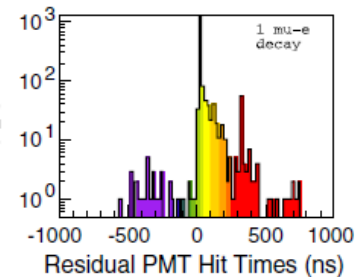
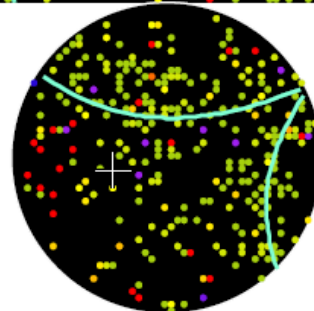
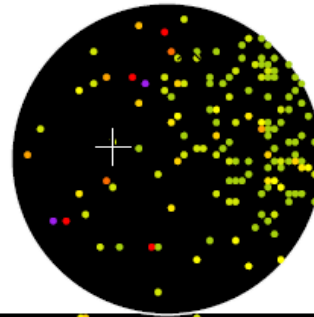
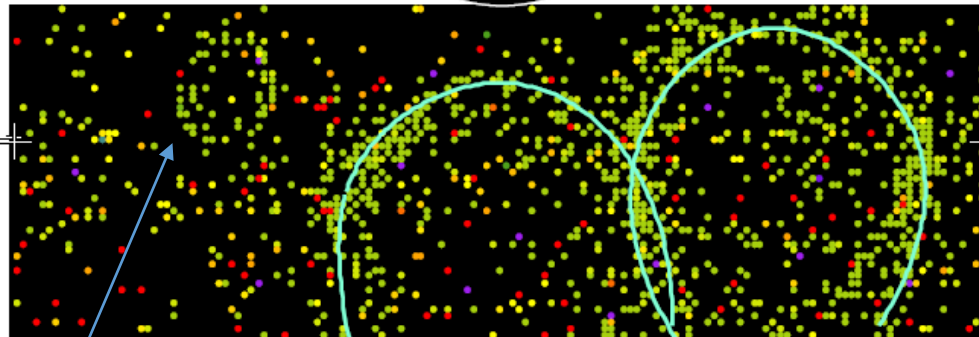
# Example of $p \rightarrow \nu K^+, K^+ \rightarrow \pi^+ \pi^0$

## Super-Kamiokande IV

Run 999999 Sub 0 Event 236  
D<sub>wall</sub>: 1076.4 cm  
E<sub>vis</sub>: 260.4 MeV  
2 e-like rings: mass = 155.2 MeV/c<sup>2</sup>

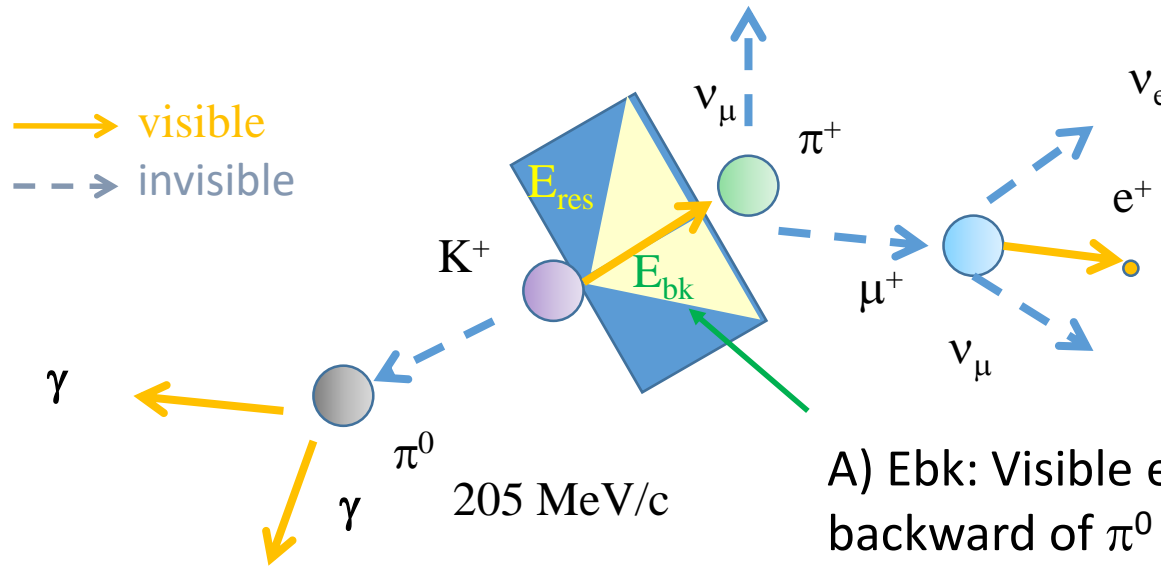
### Resid (ns)

- > 251
- 220- 251
- 188- 220
- 157- 188
- 125- 157
- 94- 125
- 62- 94
- 31- 62
- 0- 31
- -31- 0
- -62- -31
- -94- -62
- -125- -94
- -157- -125
- -188- -157
- < -188

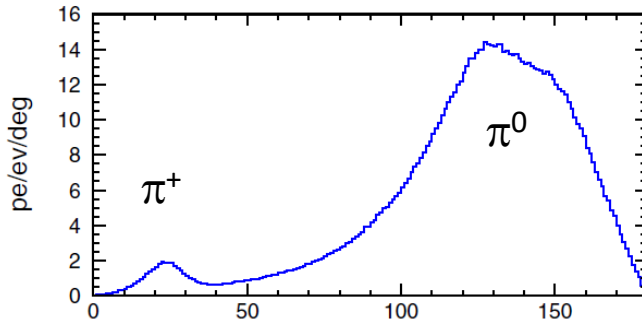


Look like a ring, but fake  
ring cut rejects this ring ...

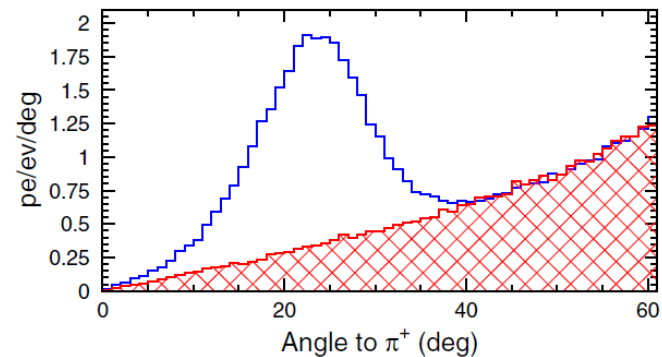
# Use $\pi^+$ information to select events



Charge distribution



Zoom



B) Make likelihood for hit pattern.

↑  
 $\pi^+$  direction

# Selection criteria for $p \rightarrow \nu K^+, K^+ \rightarrow \pi^+ \pi^0$

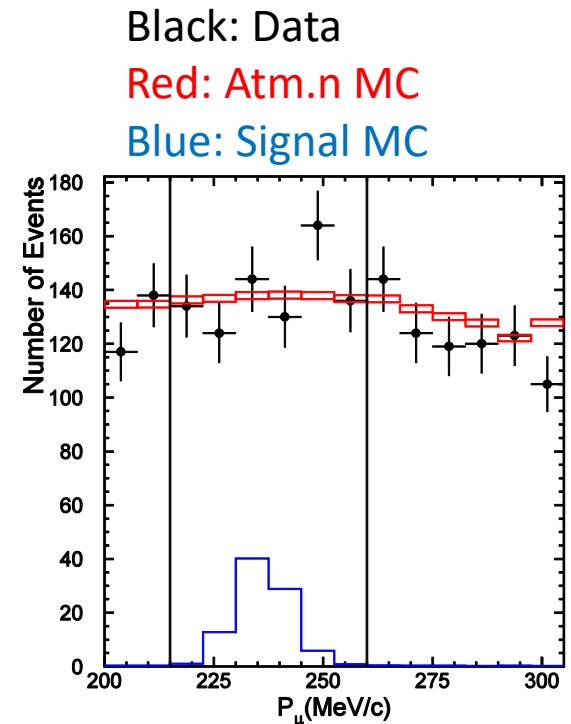
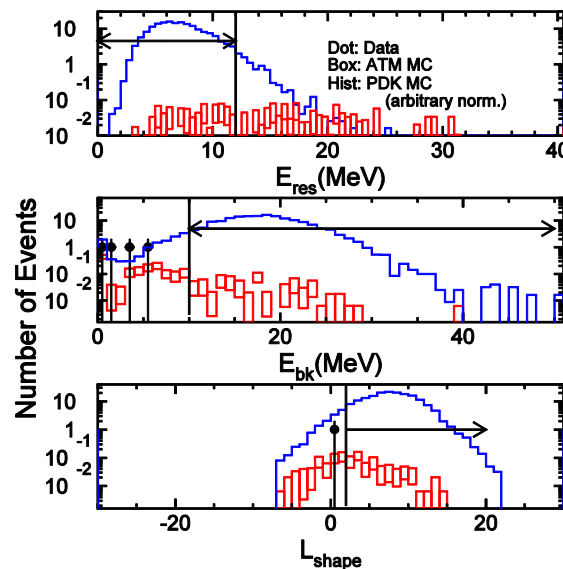
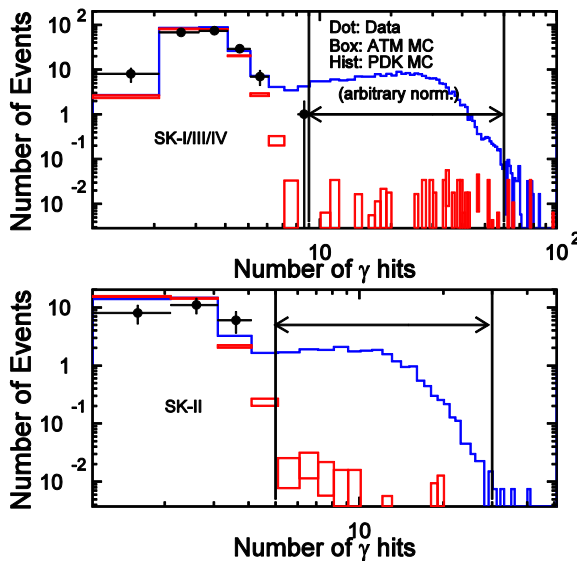
- 1 or 2 e-like rings with decay-e.
- $85 < M_{\pi^0} < 185$  MeV.
- $175 < P_{\pi^0} < 250$  MeV/c.
- $E_{bk}$ : visible energy sum in 140-180 deg. of  $\pi^0$  dir,  
 $E_{res}$ : in 90-140 deg,  
 $L_{shape}$ : Likelihood based on charge profile
  - $10 < E_{bk} < 50$  MeV
  - $E_{res} < 12$  MeV (20 MeV for 1ring)
  - $L_{shape} > 2.0$  (3.0 for 1ring)
- No neutrons
- Selection efficiency: 10 % ( $Br(K^+ \rightarrow \pi^+ \pi^0) = 21$  %)

# Background for $p \rightarrow \nu K^+$

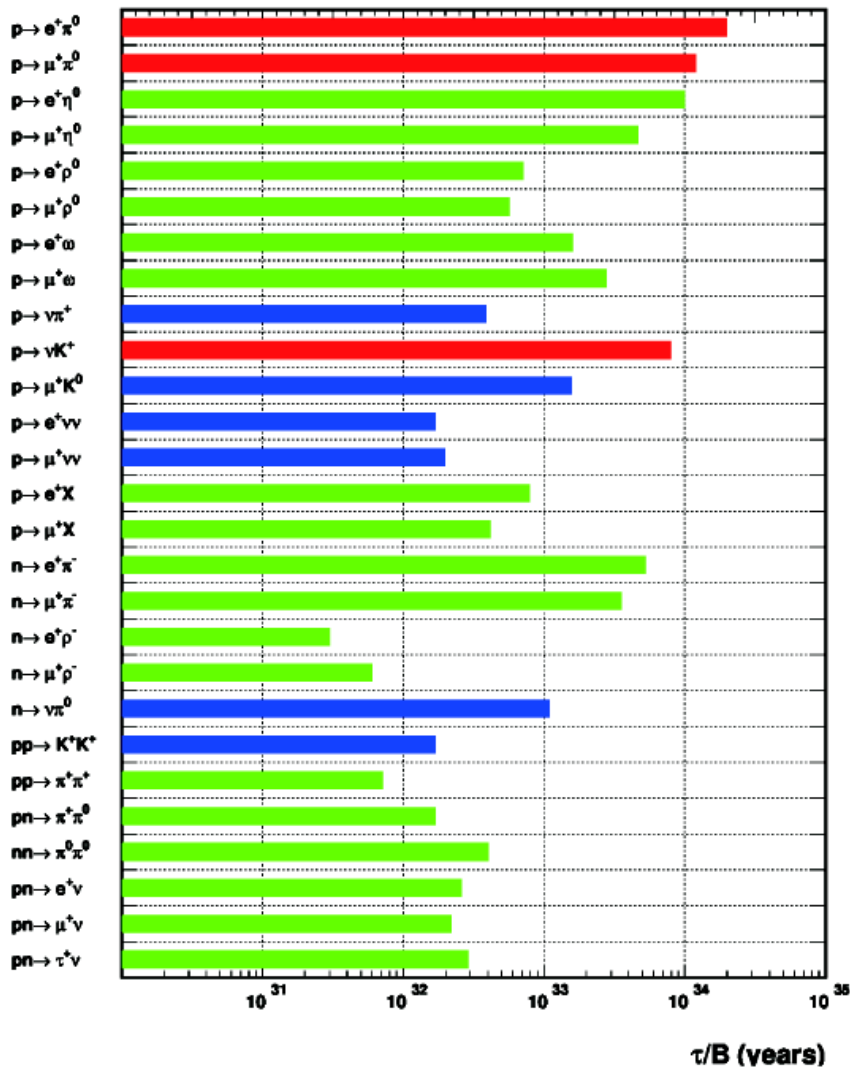
- Dominant background is  $K^+$  production by neutrino interactions.
  - $\nu p \rightarrow \nu \Lambda K^+$  ,  $\Lambda \rightarrow p \pi^-$  (BR:64 %, mostly invisible in WCD)
  - Emit nuclear  $\gamma$  as same as the signal.
- It is also rare interaction and we had poor information from very old bubble chamber. Large uncertainty.
- Recently MINERvA measures  $K^+$  production. It is very useful information for this analysis.

# 3-4. SK results (So far)

- Exposure: 365 kton · year
- Expected background: 0.3 events for  $K^+ \rightarrow \nu\mu$  with nuclear  $\gamma$ , 0.6 events for  $K^+ \rightarrow \pi^+\pi^0$ .
- No candidates observed and no excess in momentum distribution.
- Lower lifetime limit:  $> 0.8 \times 10^{34}$  year



# 3. Summary of SK results



- Most of modes have been investigated with  $> 0.3$  Mton  $\cdot$  year exposure (red and green in the left figure).
- Super-Kamiokande can cover large number of decay modes.
- Many of them are the most stringent limits on nucleon lifetime.
- We observed some candidates, but still consistent with expected backgrounds and **no evidence of nucleon decay has been observed.**



# Future prospects

- Still no evidence has been found. Major decay modes are explored up to around  $10^{34}$  years.
- Proton lives longer,  $\sim 10^{35}$  years ?
  - Run SK 10 times more ( $\sim 200$  years )? → Impossible.
- **Absolutely, we need larger detector !**

# 4. Hyper-Kamiokande project



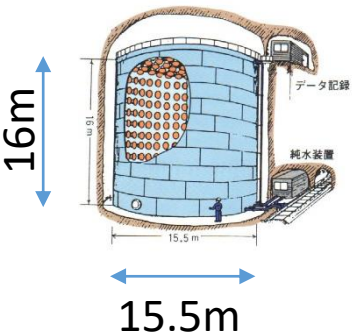
Neutrino oscillation



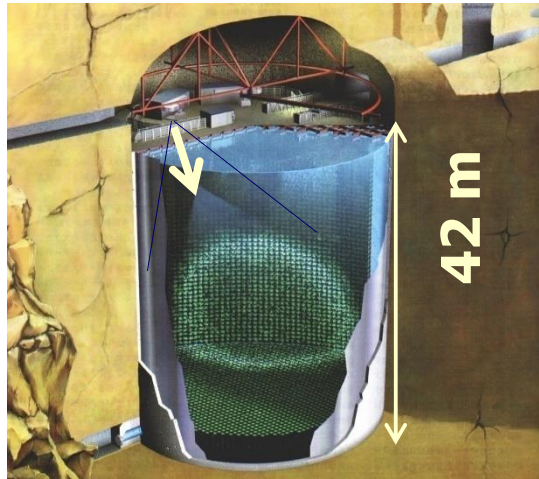
Proton decay ?



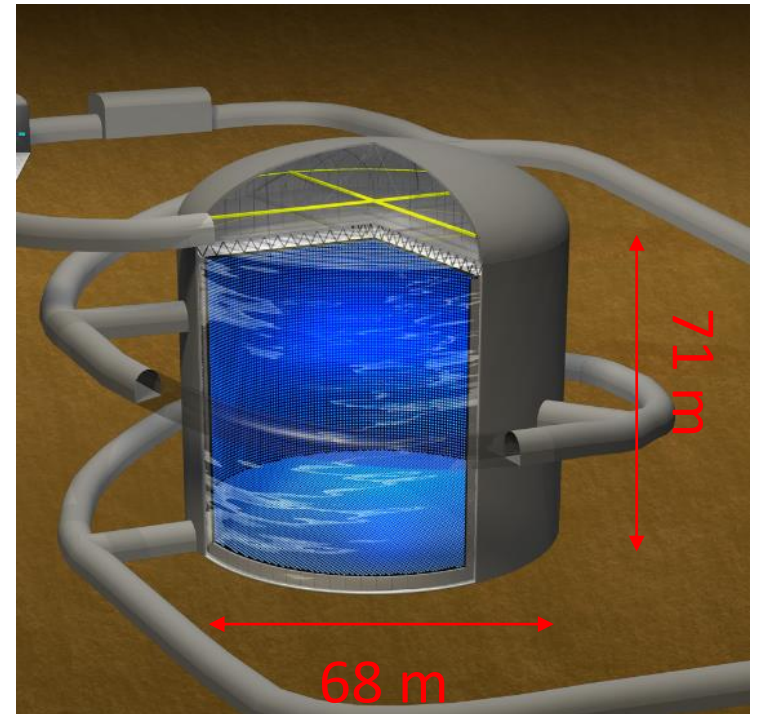
SN Neutrino



Kamiokande  
3kton

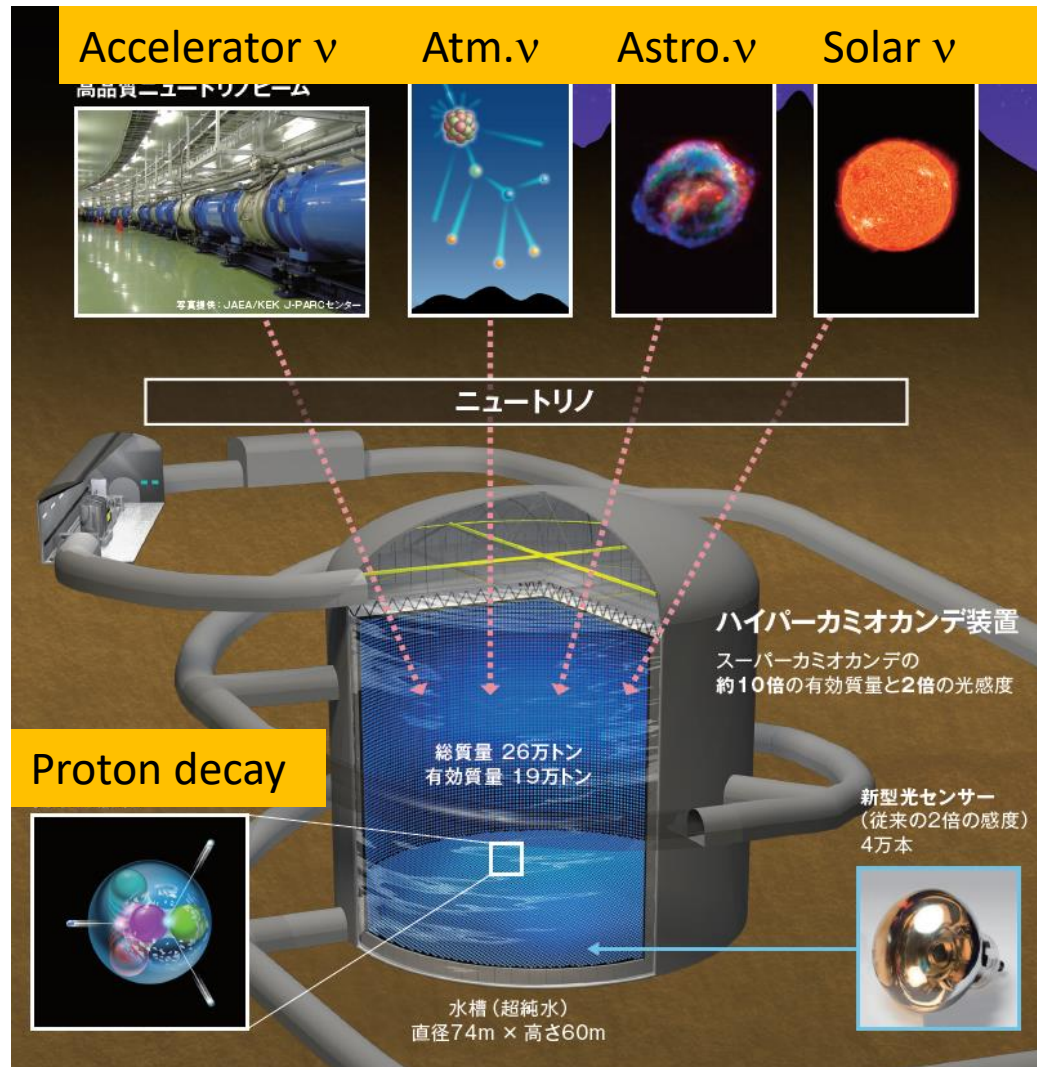


Super-Kamiokande  
50kton



Hyper-Kamiokande  
260kton

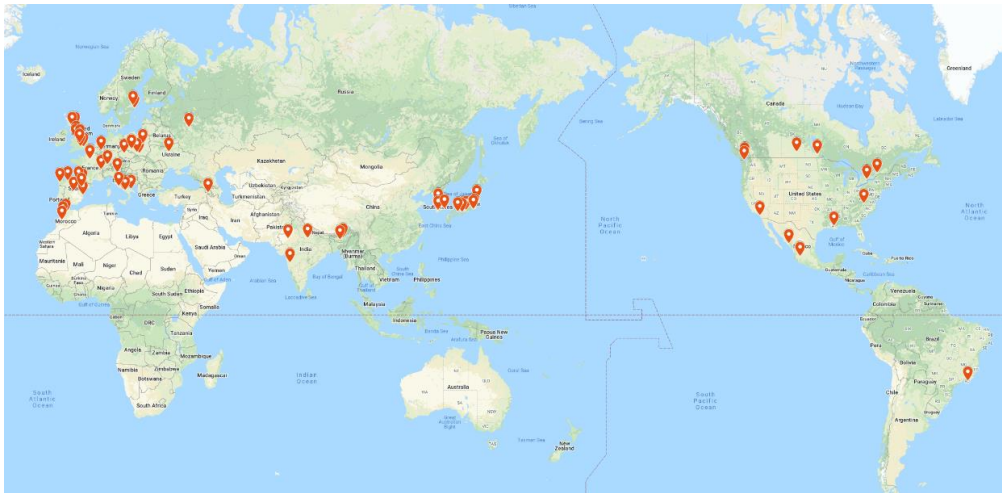
# Hyper-K is multi-purpose detector



# HYPER-K COLLABORATION:

20 countries, 99 institutes, ~500 people as of Jan 2022, and growing

## Collaborating Institutes



Europe	281 members
Armenia	3
Czech	4
France	27
Germany	1
Italy	55
Poland	38
Russia	22
Spain	35
Sweden	5
Switzerland	13
Ukraine	4
UK	74

Asia	149 members
India	8
Korea	18
Japan	123

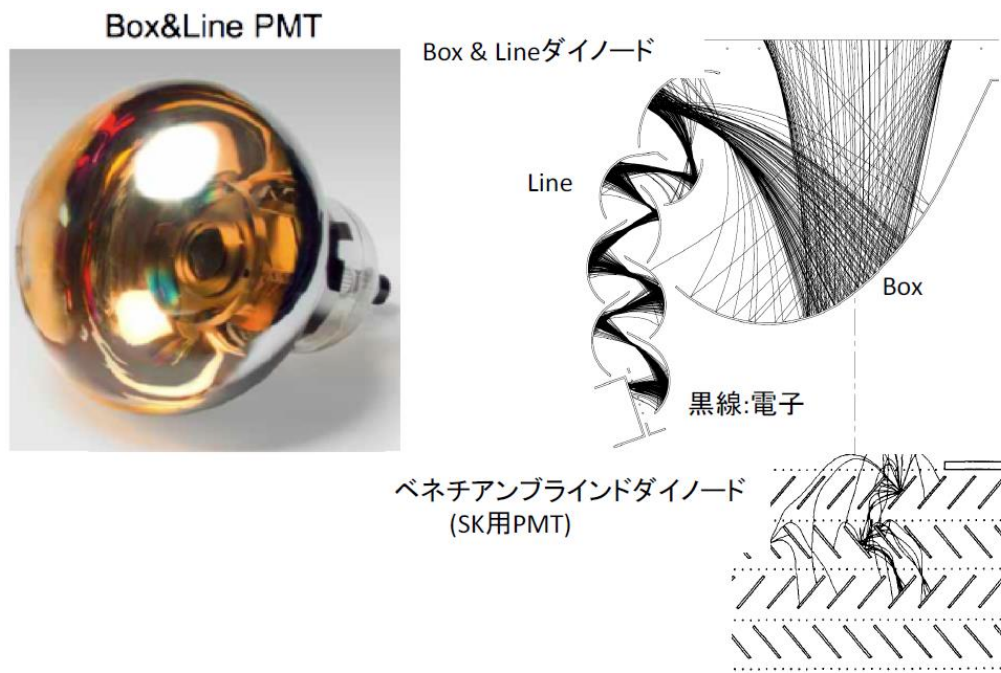
Americas	52 members
Brazil	3
Canada	32
Mexico	8
USA	9

Africa	11 members
Morocco	11

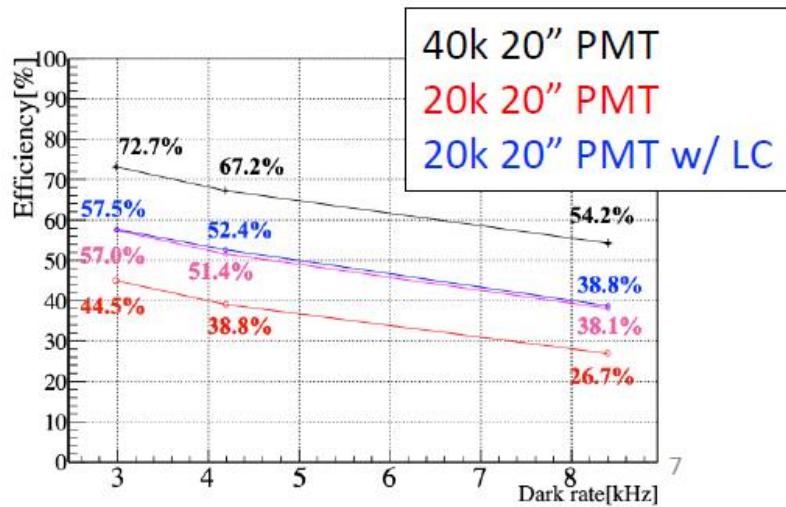
I hope Vietnam joins soon ....

# Enhance proton decay search with HK

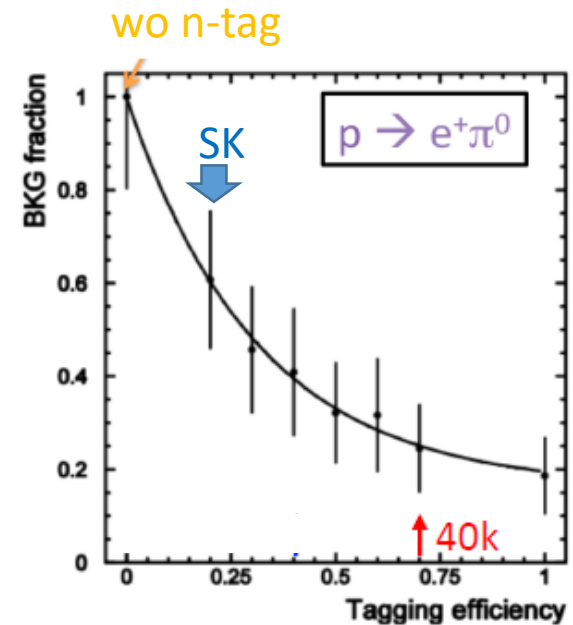
- Fiducial volume: 22.5kton (SK) → 190kton (HK)
- New photo sensor: Box&Line PMT
  - 2 times better photon counting performance
  - a half time resolution



# Better photon counting contributes neutron tagging



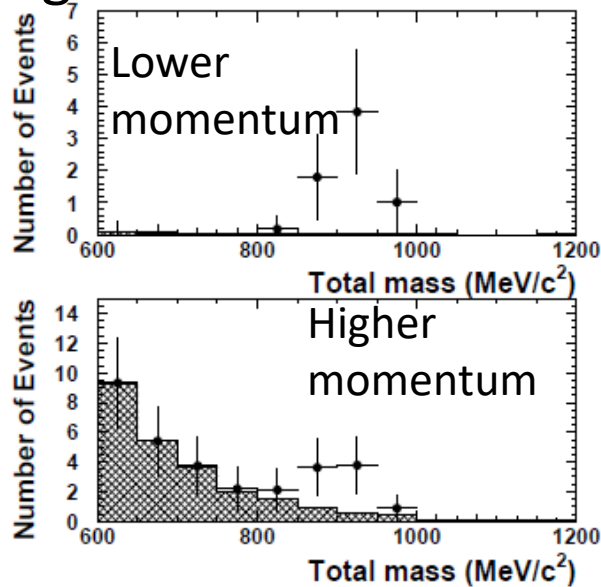
- **Neutron tagging efficiency** study with several detector set up.
- Efficiency depends on dark rate.
- Achieve **~ 70%** in the current baseline design (black) with **~ 4kHz** dark rate.



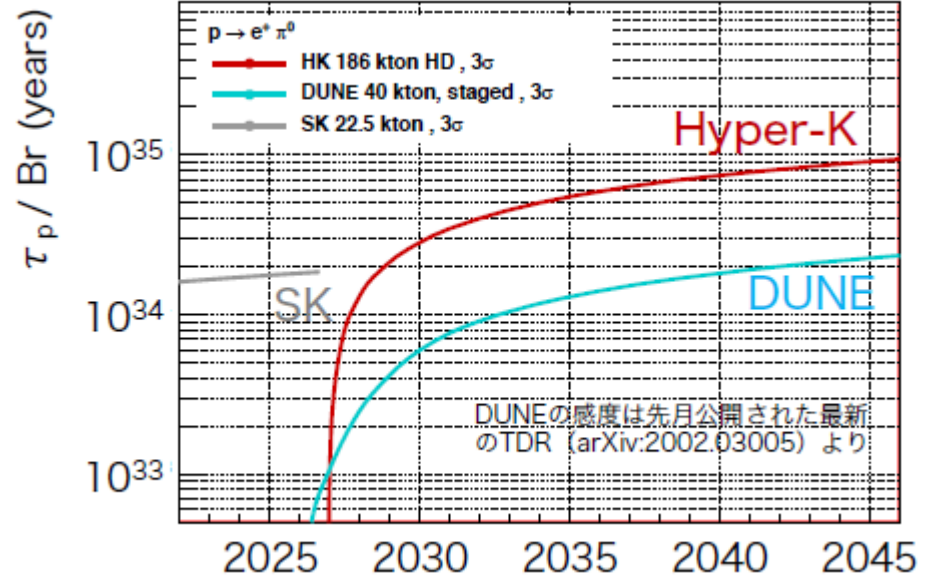
- $p \rightarrow e + \pi^0$  background reduction vs. Neutron tagging efficiency
- **Background of HK becomes a half of SK !**

# Sensitivity for $p \rightarrow e^+ \pi^0$

Expected signal after 10 years run  
assuming the current lifetime limit



3 $\sigma$  discovery potential



$0 < p_{tot} < 100 \text{ MeV}/c$		$100 < p_{tot} < 250 \text{ MeV}/c$	
$\epsilon_{sig}$ [%]	Bkg [/Mton·yr]	$\epsilon_{sig}$ [%]	Bkg [/Mton·yr]
$18.7 \pm 1.2$	$0.06 \pm 0.02$	$19.4 \pm 2.9$	$0.62 \pm 0.20$

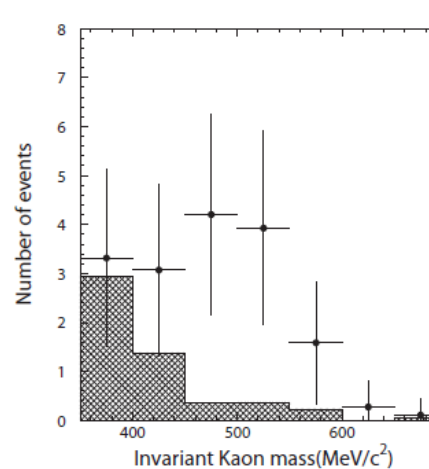
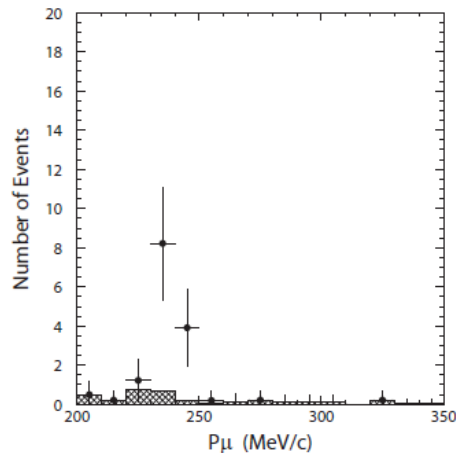
(SK: 0.18)

(SK: 1.1)

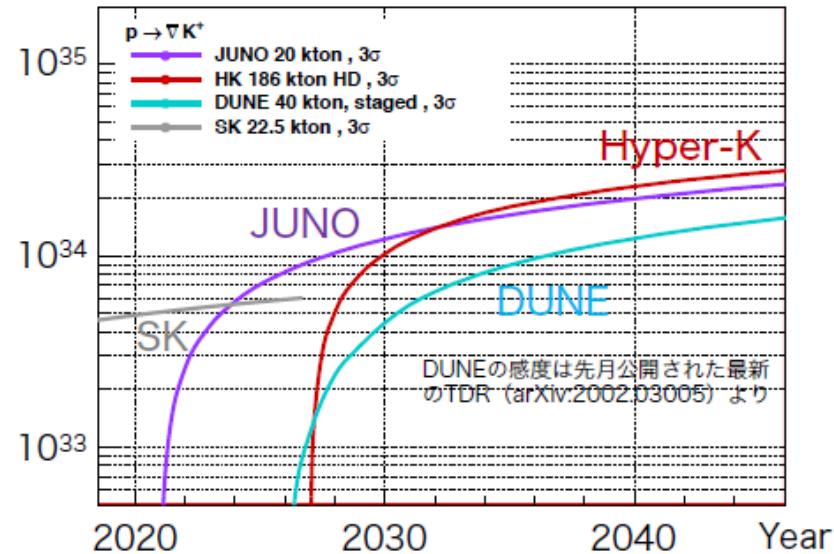
Reach to  $10^{35}$  years !

# Sensitivity for $p \rightarrow \nu K^+$

Expected signal after 10 years run  
assuming the current lifetime limit



$\tau_p / Br$  (years)



$3\sigma$  discovery potential

Prompt $\gamma$		$\pi^+\pi^0$		$p_\mu$ Spectrum		
$\epsilon_{sig}$ [%]	Bkg [/Mton·yr]	$\epsilon_{sig}$ [%]	Bkg [/Mton·yr]	$\epsilon_{sig}$ [%]	Bkg [/Mton·yr]	$\sigma_{fit}$ [%]
$12.7 \pm 2.4$	$0.9 \pm 0.2$	$10.8 \pm 1.1$	$0.7 \pm 0.2$	31.0	1916.0	8.0

Better sensitivity than  
other detectors !.



# What are still unknown in $\nu$ oscillation?

## Unknown parameter

## CP phase

$$c_{ij} = \cos q_{ij}, s_{ij} = \sin q_{ij}$$

$\begin{matrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{matrix}$	$\begin{matrix} n_e \\ n_\mu \\ n_\tau \end{matrix}$	$\begin{matrix} \ddot{\nu} \\ \dot{\nu} \\ \nu \end{matrix}$	$\begin{matrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{matrix}$	<table border="1" style="background-color: #e6f2ff;"> <tr><td>1</td><td>0</td><td>0</td></tr> <tr><td>0</td><td><math>c_{23}</math></td><td><math>s_{23}</math></td></tr> <tr><td>0</td><td><math>-s_{23}</math></td><td><math>c_{23}</math></td></tr> </table>	1	0	0	0	$c_{23}$	$s_{23}$	0	$-s_{23}$	$c_{23}$	$\begin{matrix} \ddot{\nu} \\ \dot{\nu} \\ \nu \end{matrix}$	<table border="1" style="background-color: #ffe6e6;"> <tr><td><math>c_{13}</math></td><td>0</td><td><math>s_{13}e^{-id}</math></td></tr> <tr><td>0</td><td>1</td><td>0</td></tr> <tr><td><math>-s_{13}e^{id}</math></td><td>0</td><td><math>c_{13}</math></td></tr> </table>	$c_{13}$	0	$s_{13}e^{-id}$	0	1	0	$-s_{13}e^{id}$	0	$c_{13}$	$\begin{matrix} \ddot{\nu} \\ \dot{\nu} \\ \nu \end{matrix}$	<table border="1" style="background-color: #e6ffe6;"> <tr><td><math>c_{12}</math></td><td><math>s_{12}</math></td><td>0</td></tr> <tr><td><math>-s_{12}</math></td><td><math>c_{12}</math></td><td>0</td></tr> <tr><td>0</td><td>0</td><td>1</td></tr> </table>	$c_{12}$	$s_{12}$	0	$-s_{12}$	$c_{12}$	0	0	0	1	$\begin{matrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{matrix}$	$\begin{matrix} n_1 \\ n_2 \\ n_3 \end{matrix}$	$\begin{matrix} \ddot{\nu} \\ \dot{\nu} \\ \nu \end{matrix}$	$\begin{matrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{matrix}$	$\begin{matrix} \ddot{\nu} \\ \dot{\nu} \\ \nu \end{matrix}$
1	0	0																																						
0	$c_{23}$	$s_{23}$																																						
0	$-s_{23}$	$c_{23}$																																						
$c_{13}$	0	$s_{13}e^{-id}$																																						
0	1	0																																						
$-s_{13}e^{id}$	0	$c_{13}$																																						
$c_{12}$	$s_{12}$	0																																						
$-s_{12}$	$c_{12}$	0																																						
0	0	1																																						

- $\theta_{12} = 33.6^\circ \pm 1.0^\circ$  Solar  $\nu$ , KamLAND
- $\theta_{23} = 45^\circ \pm 6^\circ$  (90%CL) Atm.  $\nu$ , Acc.  $\nu$

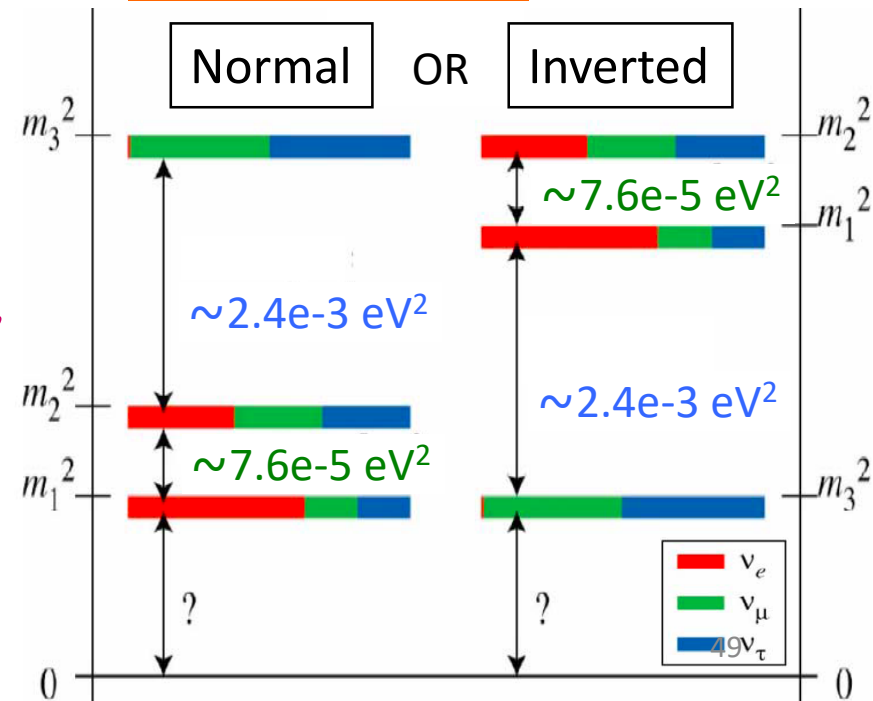
$\theta_{23}$  : How close to  $45^\circ$  ?  
Octant? ( $<45^\circ$ ,  $>45^\circ$  ?)

- $\theta_{13} = 9.1^\circ \pm 0.6^\circ$  Reactor, Acc.  $\nu$

Indication of  $\theta_{13} \neq 0$  by T2K  
PRL107, 041801 (2011)

Later precise measurements  
by reactor  $\nu$  experiments

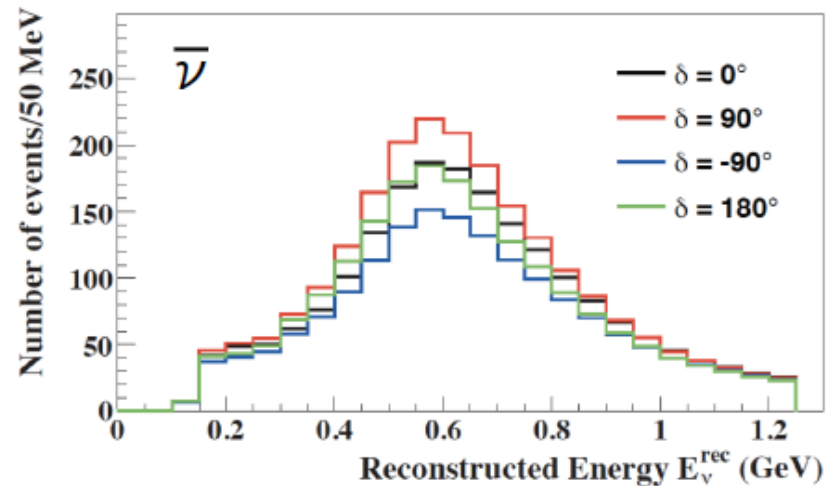
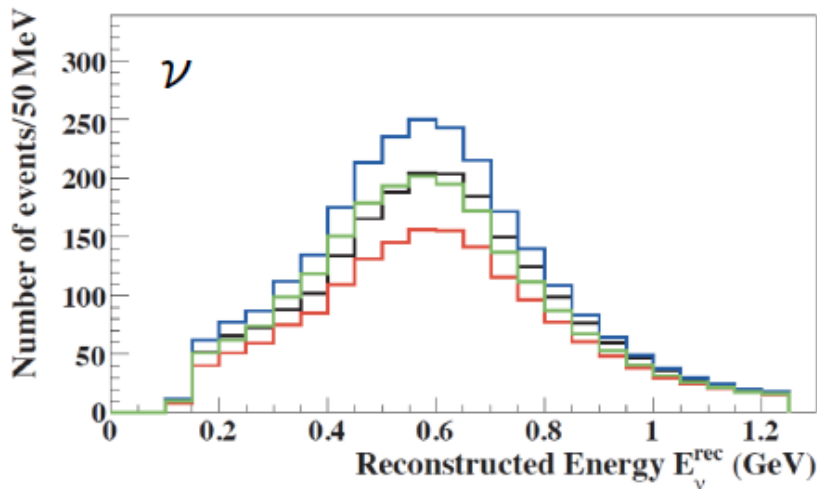
## Mass hierarchy



# Effects of unknown parameters on $\nu$ oscillation are small $\rightarrow$ Need statistics = larger detector !

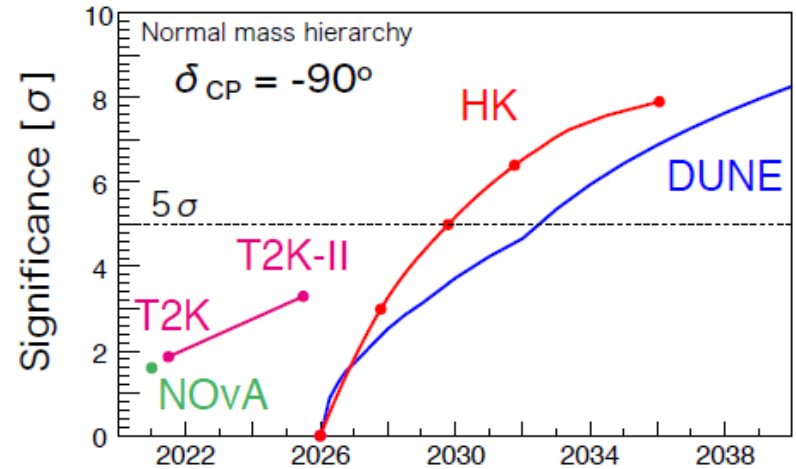
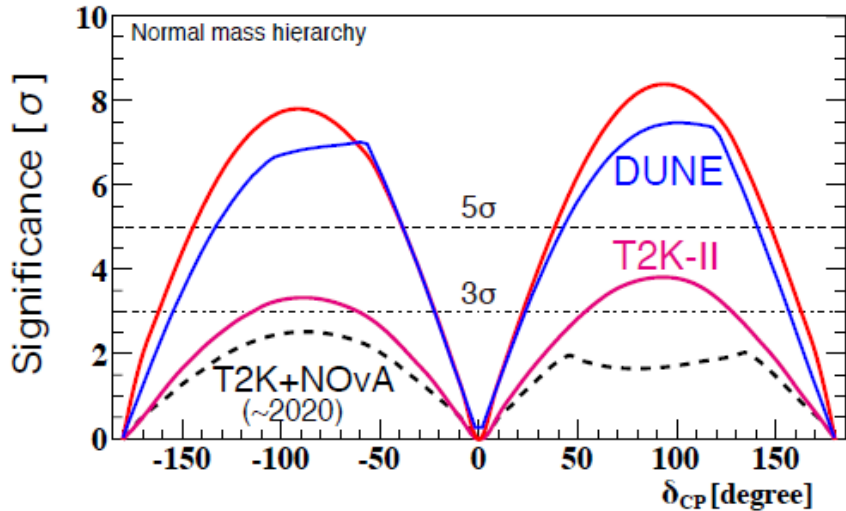
- $\nu$  beam experiments: Can study CP phase by comparing  $\nu$   $\bar{\nu}$  oscillation.

Expected  $\nu_e$  spectrum at HK (assuming 1.3 MW x 10 years)



# Sensitivity of CP violation

HK 10 years



- In 60 % region in  $\delta_{CP}$ , we can discovery  $\delta_{CP}$  with  $5\sigma$ .
- If  $= \delta_{CP} \pm 90$  degree, we can discover it within 5 years.

# Atmospheric $\nu$ : $\nu_\mu \rightarrow \nu_e$ enhancement by matter effect

$\nu_e$  oscillation due to non-zero  $\theta_{13}$  provides atm. nu. observation to investigate mass hierarch effect

$\nu_\mu \rightarrow \nu_e$  osc. probability in matter:

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 \theta_{23} \sin^2 2\theta_{13}^M \sin^2 \left( \frac{1.27 \Delta m_{31,M}^2 L}{E} \right)$$

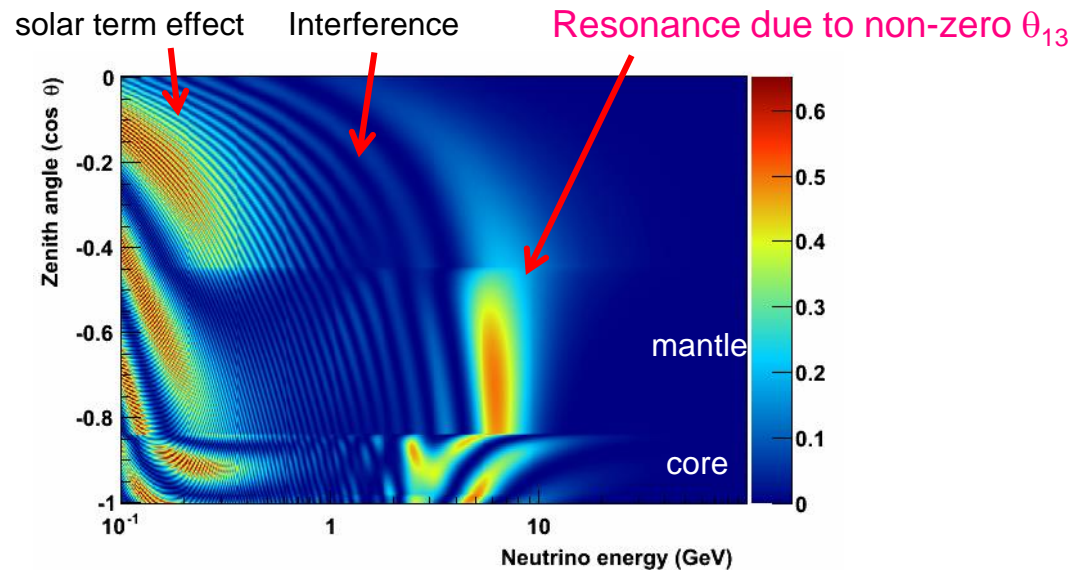
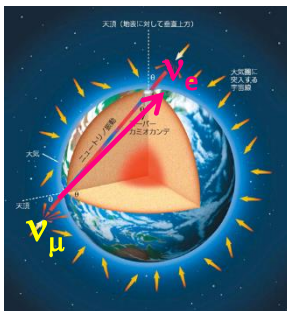
$$\sin^2 2\theta_{13}^M = \frac{\sin^2 2\theta_{13}}{\left( \cos 2\theta_{13} - \frac{A_{CC}}{\Delta m_{31}^2} \right) + \sin^2 2\theta_{13}}$$

$\sin^2 \theta_{13,M}$  has resonance feature when

$$A_{CC} \sim \Delta m_{31}^2 \cos 2\theta_{13} \quad (A_{CC} = 2\sqrt{2}G_F N_e E)$$

$\nu_\mu \rightarrow \nu_e$  resonance in multi-GeV region

$P(\nu_\mu \rightarrow \nu_e)$  of earth-throughgoing  $\nu$  :



# Occurrence of resonance feature depends on neutrino type and mass hierarchy:

$$\sin^2 2\theta_{13}^M = \frac{\sin^2 2\theta_{13}}{\left(\cos 2\theta_{13} - \frac{A_{CC}}{\Delta m_{31}^2}\right) + \sin^2 2\theta_{13}}$$

In case of Inverted hierarchy:

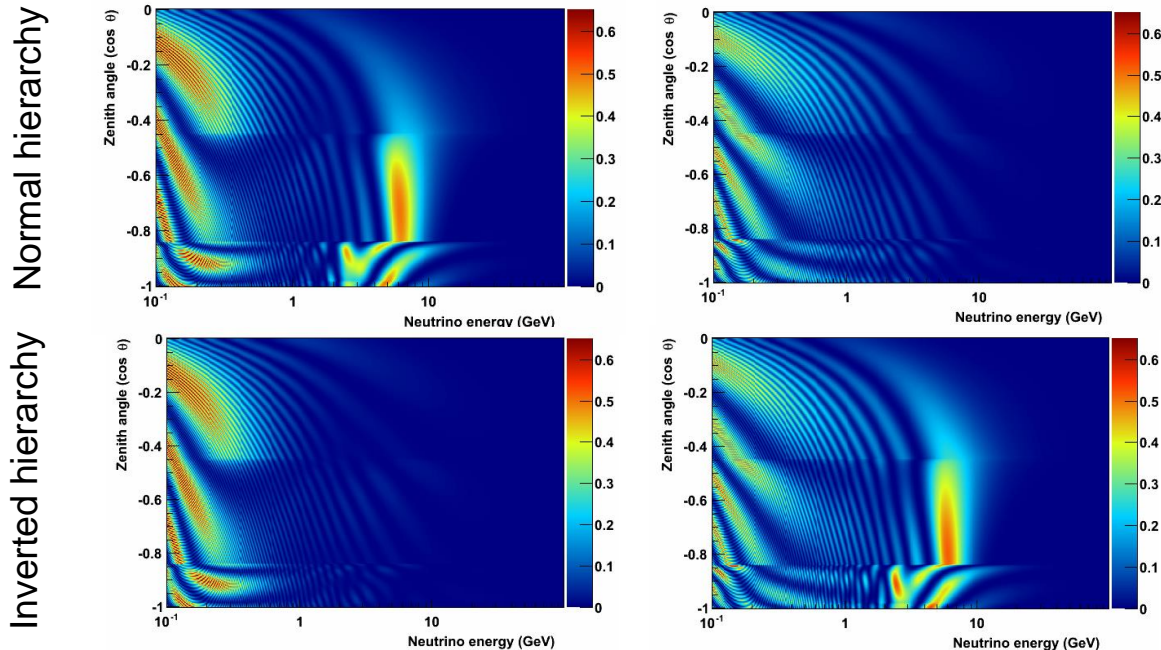
$$\Delta m^2 \rightarrow -\Delta m^2$$

In case of anti neutrino:

$$A_{CC} \rightarrow -A_{CC}$$

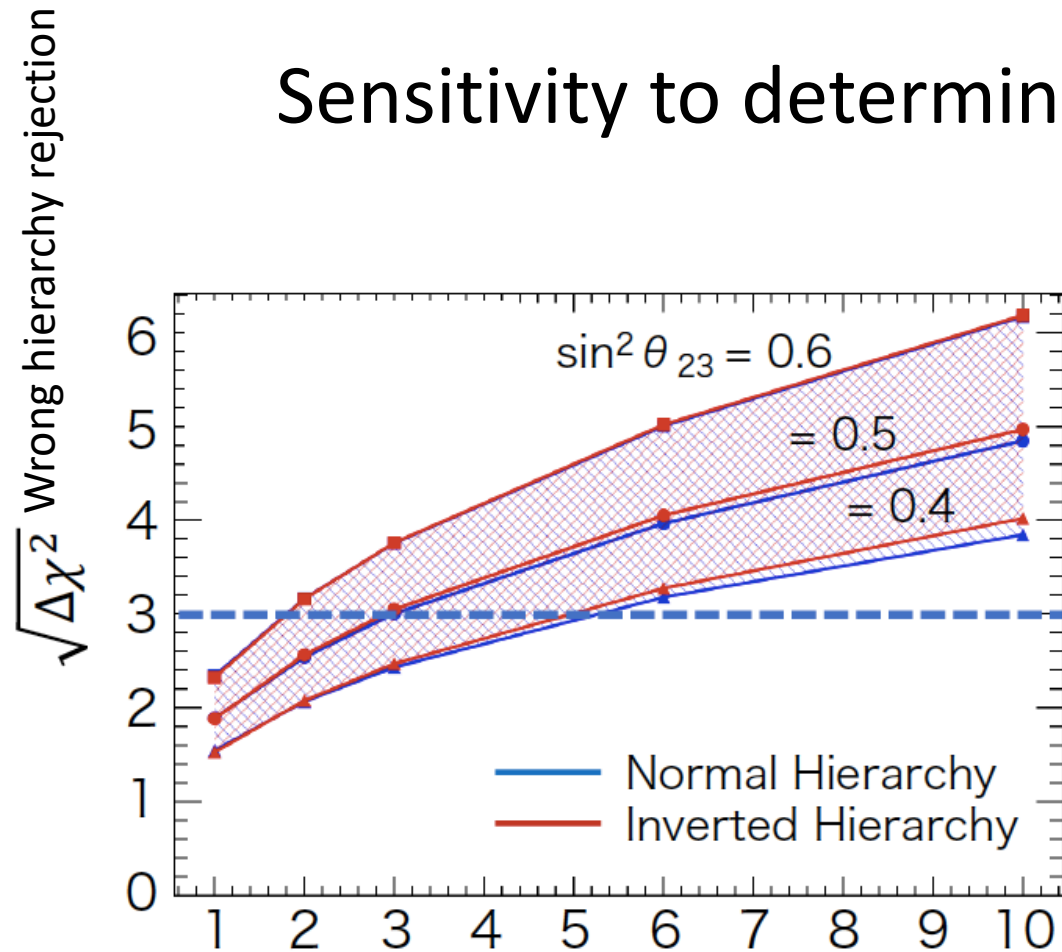
$\nu_\mu \rightarrow \nu_e$

anti- $\nu_\mu \rightarrow$  anti- $\nu_e$



Possible to determine mass hierarchy !

# Sensitivity to determine mass hierarchy



**Beam  $\nu$ :** sensitive to  $\delta\text{CP}$ , but weak in mass hierarchy.

**Atm.  $\nu$ :** large uncertainty from  $\delta\text{CP}$ .

→ Combining both analysis gives **good sensitivity to mass hierarchy.**

**$3\sigma$  determine within 2 ~ 5 years !**

# Construction of HyperK has been started !



Making entrance yard for tunnel.



**Aim to start observation in 2027 !**

Access tunnel has reached to the center of  
HyperK tank in this June !

Detector Cavern is becoming a reality



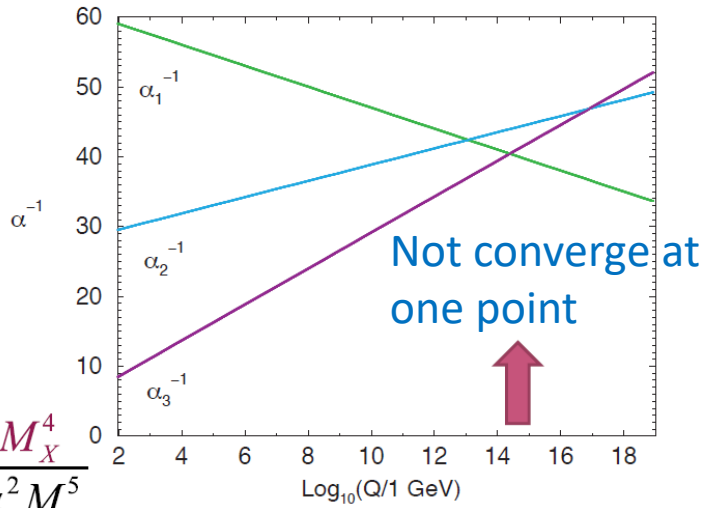


# 4. Summary

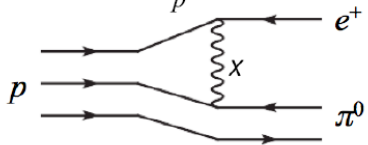
- Proton decay is a key phenomena of Grand Unified Theories beyond the Standard Model.
- Super-Kamiokande is the leading detector to hunt proton decays and have searched for it for more than 20 years.
- However, no evidence has been observed and the current proton lifetime limits are around  $10^{34}$  years.
- It may be around the corner ! Hope three times lucky (3度目の正直 in Japanese) in Hyper-Kamiokande.
- HK also can determine remaining  $\nu$  oscillation parameters .

Backup

# Gauge Coupling Unification

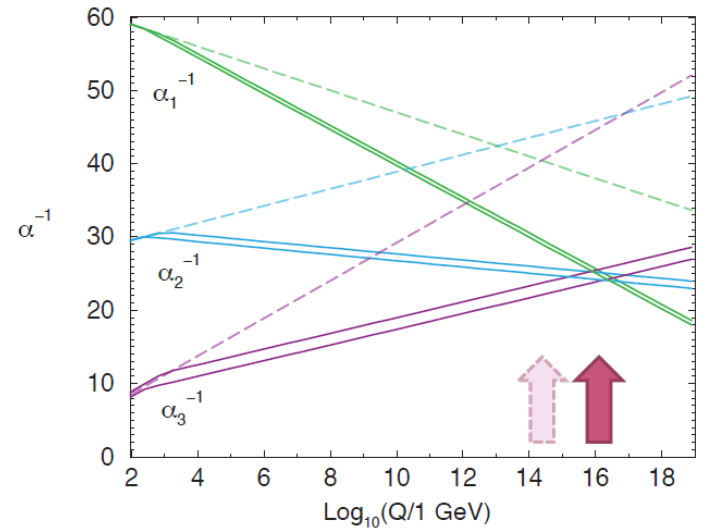


$$\tau \approx \frac{M_X^4}{\alpha^2 M_p^5}$$



$$\tau(e^+ \pi^0) = 4.5 \times 10^{29 \pm 1.7} \text{ years (prediction)}$$

# Problems solved by SUSY ...



Unification scale pushed up...

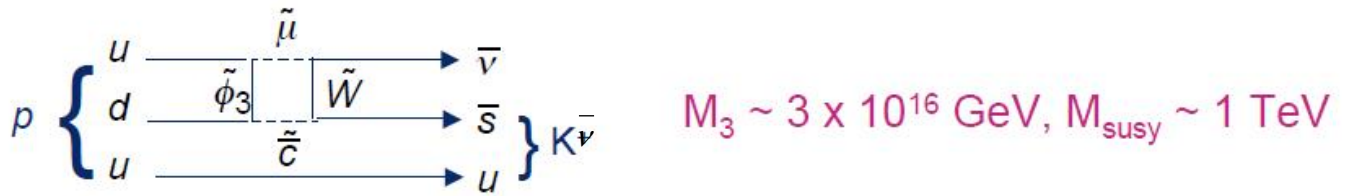
$$\tau(e^+ \pi^0) \approx 10^{35-38} \text{ years}$$

# SUSY GUTs

- Unification scale higher than non-SUSY-GUTs ( $M_x \sim 2 \times 10^{16}$  GeV)  
 suppression of gauge boson mediated decay

$$\tau/B (p \rightarrow e^+ \pi^0) \approx \left( \frac{M_x}{2 \times 10^{16} \text{ GeV}} \right)^4 \times 10^{36 \pm 1} \text{ years}$$

dominated by the D=5 op. (color Higgs triplet,  $q=1/3$ ) mediated decays



$$\tau/B (p \rightarrow K^+ \bar{\nu}) \approx \left( \frac{0.003}{\beta_H} \right)^2 \left( \frac{M_q}{1 \text{ TeV}} \right)^2 \left( \frac{M_3}{10^{16} \text{ GeV}} \right)^2 \left( \frac{2}{\tan \beta} \right)^2 \dots$$

$\times 9 \times 10^{31} \text{ years}$

( $\beta_H$  : hadronic matrix element in  $\text{GeV}^3$ )  
 $\Rightarrow$  highly model dependent

$p \rightarrow \bar{\nu} K^+$  is regarded as dominant mode in SUSY-GUTs.

# Many Other GUTs Beyond This Simple Story

Model	Ref.	Modes	$\tau_N$ (years)
Minimal $SU(5)$	Georgi, Glashow [2]	$p \rightarrow e^+ \pi^0$	$10^{30} - 10^{31}$
Minimal SUSY $SU(5)$	Dimopoulos, Georgi [11], Sakai [12] Lifetime Calculations: Hisano, Murayama, Yanagida [13]	$p \rightarrow \bar{\nu} K^+$ $n \rightarrow \bar{\nu} K^0$	$10^{28} - 10^{32}$
SUGRA $SU(5)$	Nath, Arnowitt [14, 15]	$p \rightarrow \bar{\nu} K^+$	$10^{32} - 10^{34}$
SUSY $SO(10)$ with anomalous flavor $U(1)$	Shafi, Tavartkiladze [16]	$p \rightarrow \bar{\nu} K^+$ $n \rightarrow \bar{\nu} K^0$ $p \rightarrow \mu^+ K^0$	$10^{32} - 10^{35}$
SUSY $SO(10)$ MSSM (std. $d = 5$ )	Lucas, Raby [17], Pati [18]	$p \rightarrow \bar{\nu} K^+$ $n \rightarrow \bar{\nu} K^0$	$10^{33} - 10^{34}$ $10^{32} - 10^{33}$
SUSY $SO(10)$ ESSM (std. $d = 5$ )	Pati [18]	$p \rightarrow \bar{\nu} K^+$	$10^{33} - 10^{34}$ $\lesssim 10^{35}$
SUSY $SO(10)/G(224)$ MSSM or ESSM (new $d = 5$ )	Babu, Pati, Wilczek [19, 20, 21], Pati [18]	$p \rightarrow \bar{\nu} K^+$ $p \rightarrow \mu^+ K^0$	$\lesssim 2 \cdot 10^{34}$ $B \sim (1 - 50)\%$
SUSY $SU(5)$ or $SO(10)$ MSSM ( $d = 6$ )	Pati [18]	$p \rightarrow e^+ \pi^0$	$\sim 10^{34.9 \pm 1}$
Flipped $SU(5)$ in CMSSM	Ellis, Nanopoulos and Wlaker[22]	$p \rightarrow e/\mu^+ \pi^0$	$10^{35} - 10^{36}$
Split $SU(5)$ SUSY	Arkani-Hamed, <i>et. al.</i> [23]	$p \rightarrow e^+ \pi^0$	$10^{35} - 10^{37}$
$SU(5)$ in 5 dimensions	Hebecker, March-Russell[24]	$p \rightarrow \mu^+ K^0$ $p \rightarrow e^+ \pi^0$	$10^{34} - 10^{35}$
$SU(5)$ in 5 dimensions option II	Alciati <i>et.al.</i> [25]	$p \rightarrow \bar{\nu} K^+$	$10^{36} - 10^{39}$
GUT-like models from Type IIA string with D6-branes	Klebanov, Witten[26]	$p \rightarrow e^+ \pi^0$	$\sim 10^{36}$

Uncertainties in the predictions:

Nuclear matrix elements updated w. IQCD, still: x10 uncertainty in lifetime

SUSY masses: ~ x100 uncertainty in lifetime

Proton life time:  
 $10^{30} \sim 10^{35}$  years

TABLE I: Summary of the expected nucleon lifetime in different theoretical models.

## Modes beyond $e^+\pi^0, K^+\nu$ and other antilepton + meson decays

$$p \rightarrow \mu^- \pi^+ K^+$$

$$B + L$$

$$n \rightarrow \bar{n}$$

$$\Delta B = 2, \text{ TeV} < \text{scale} < \text{GUT}$$

$$pp \rightarrow K^+ K^+$$

$$\lambda''_{uds} < 10^{-8}$$

$$p \rightarrow e^- \pi^+ \pi^+ \nu \nu$$

6 dimensions

$$n \rightarrow \nu \nu \nu$$

invisible

$$p \rightarrow e^+ \gamma$$

radiative

there is plenty to keep us busy ...

# How to find 2.2 MeV $\gamma$

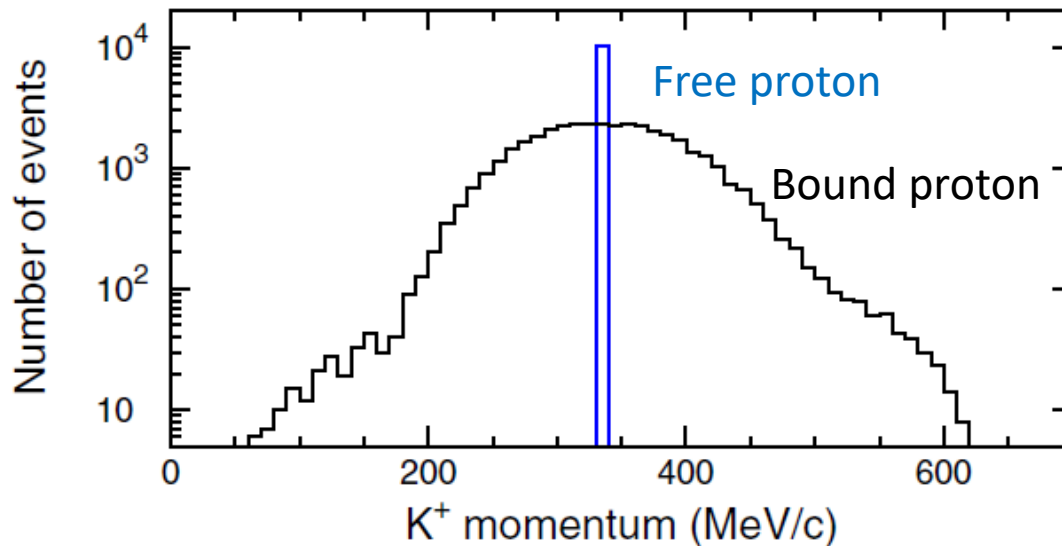


- After Time-of-Flight subtraction, search for 7 hits in 10 nsec time window.  $\rightarrow$  candidates of  $\gamma$ .
- Make 16 variables related to space and time information of each hits (RMS of phi, theta, hit time, e.t.c.)
- Put them into Neural Network to judge  $\gamma$  or not.
- **Neutron tagging efficiency: 21 % (mis-tagging: 1.8 %)**

3.  $p \rightarrow \nu K^+$  search



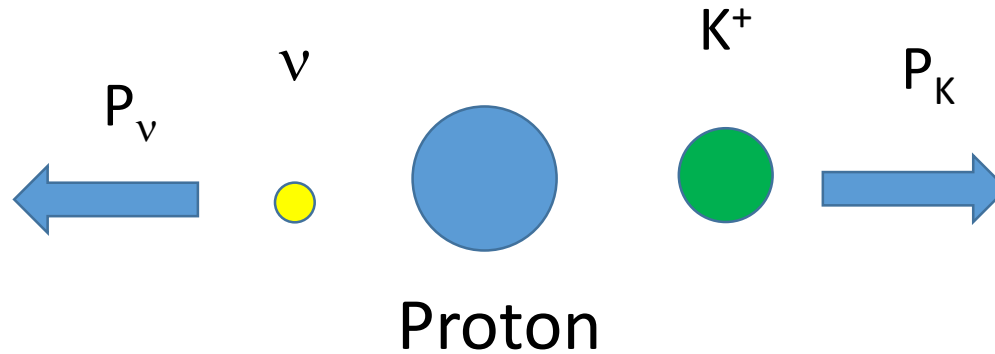
# Difficulty of $p \rightarrow \nu K^+$



- $K^+$  mass: 494 MeV, relatively heavy.
- Cherenkov threshold: 560 MeV/c.
- Most of  $K^+$  can not emit Cherenkov light.

Q. Calculate momentum of  $K^+$  from free proton decay.

Hint: proton mass: 938 MeV, "free" means proton momentum=0.



Conservation of momentum:  $P_\nu = P_K$

Conservation of energy:  $M_P = \sqrt{M_K^2 + P_K^2} + P_\nu$

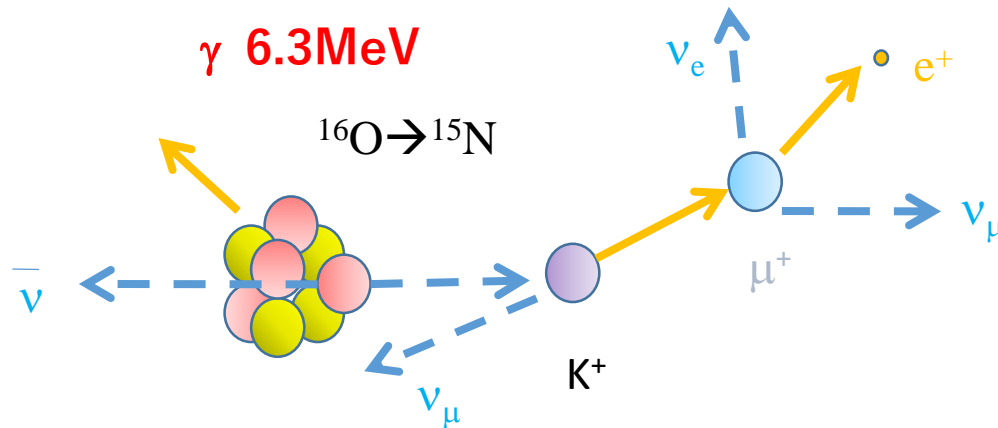
Proton
Electron
Neutrino

$$P_K = \frac{M_P^2 - M_K^2}{2M_P}$$

# 3-1 How to find $p \rightarrow \nu K^+$ in Water Cherenkov detector

- $K^+$  has low momentum, most of them **stop in water** and decay with 12 nsec lifetime.
- Major  $K^+$  decay mode
  - $K^+ \rightarrow \nu\mu^+$  : 64 %
  - $K^+ \rightarrow \pi^+\pi^0$ : 21 %
- “Stopping  $K^+$ ” means **two body decay products of  $K^+$  should have monochromatic momentum** as seen in the previous question !
  - $K^+ \rightarrow \nu\mu^+$  : 236 MeV/c
  - $K^+ \rightarrow \pi^+\pi^0$ : 206 MeV/c
- Using this property, Water Cherenkov detector can search for  $p \rightarrow \nu K^+$  .

## 3-2. Search for $p \rightarrow \nu K^+$ , $K^+ \rightarrow \nu \mu^+$



- Visible particle is only  $\mu^+$  with Michel electron.
- Search for data excess around 236 MeV/c of  $\mu$  comparing with atmospheric  $\nu$  MC.
- After proton decay, **40 % of remaining nucleus emits 6 MeV  $\gamma$  for deexcitation.** It is useful to reduce background.

# Example of $p \rightarrow \nu K^+$ , $K^+ \rightarrow \nu \mu^+$ with $\gamma$

## Super-Kamiokande IV

Run 999999 Sub 0 Event 69  
D\_wall: 1165.1 cm  
Evis: 53.2 MeV  
mu-like,  $p = 231.0$  MeV/c

### Resid (ns)

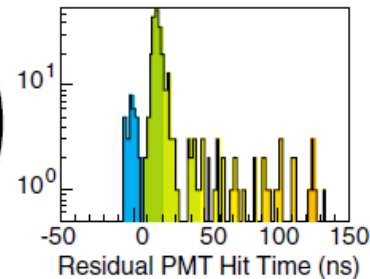
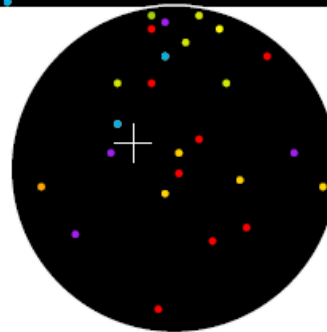
- > 182
- 160- 182
- 137- 160
- 114- 137
- 91- 114
- 68- 91
- 45- 68
- 22- 45
- 0- 22
- -22- 0
- -45- -22
- -68- -45
- -91- -68
- -114- -91
- -137--114
- <-137

$\mu$

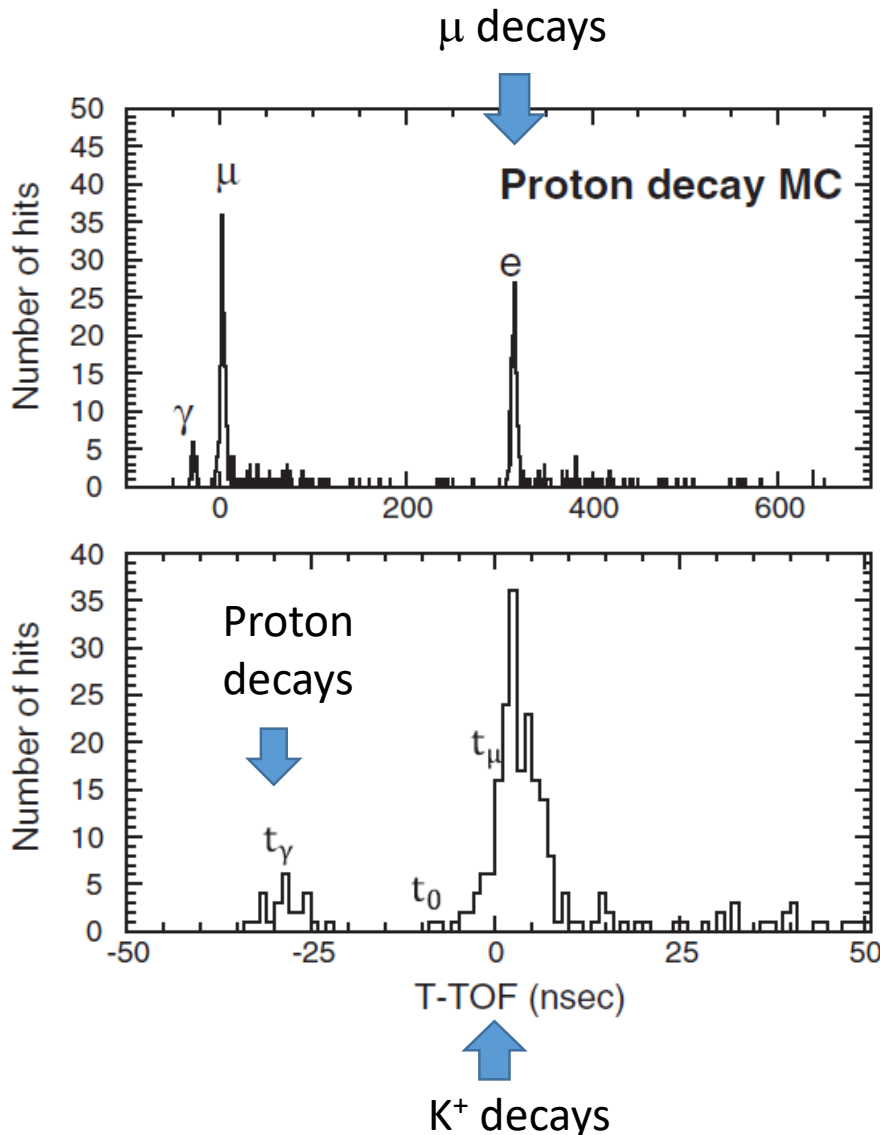
Color of hits corresponds to time.

Cyan corresponds to nuclear  $\gamma$ .

Difficult to identify  $\gamma$  from hit pattern.




# Time structure with nuclear $\gamma$



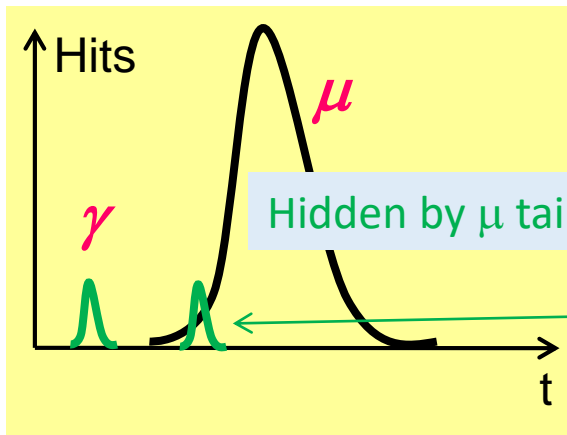
- 3 hit clusters in time should be observed in case of signal.
- The event is triggered by  $\mu$  hits.
- $\gamma$  signal is much smaller than  $\mu$  and easily hidden by tail of  $\mu$  hits.
- Make 12 nsec time window and slide it toward left from  $t_0$  (end of  $\mu$  tail) to search for maximum hit cluster.

# Selection criteria for $p \rightarrow \nu K^+$ , $K^+ \rightarrow \nu \mu^+$

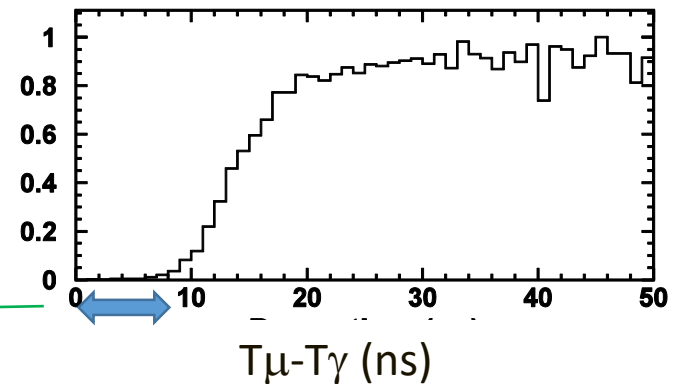
- 1  $\mu$ -like ring with Michel electron
- $215 < P_\mu < 260$  MeV/c
- Proton rejection cuts
- Search Max hit cluster  Reduce background by  $5 \times 10^{-4}$  !  
by sliding time window (12ns width);
  - $4 < N_\gamma < 30$  hits
  - $T_\mu - T_\gamma < 75$  nsec
- No neutron
- Selection efficiency = (selected events)/(proton decay in fiducial volume):  
9 %
  - $\text{Br}(K^+ \rightarrow \nu \mu^+) = 64$  %, only 40 % emits nuclear  $\gamma \rightarrow 26$  %  
even if detector is perfect.

# Remark for this analysis

- This analysis is limited by time resolution of PMTs.
  - If  $\gamma$  is close to  $\mu$ ,  $\gamma$  peak is hidden by  $\mu$  hits.
  - Time resolution of SK PMT is 2.2 nsec at 1 photoelectron.
  - If  $\mu$  peak becomes sharper, the selection efficiency will be improved.

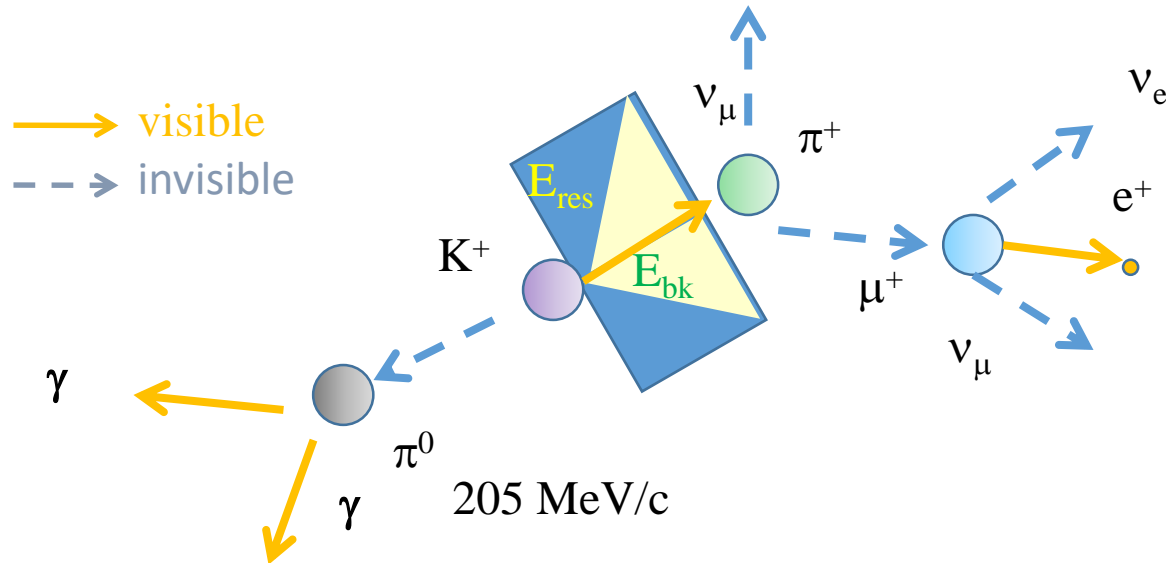


$\gamma$  tagging efficiency





### 3-3. Search for $p \rightarrow \nu K^+$ , $K^+ \rightarrow \pi^+ \pi^0$



- Both  $\pi^+$  and  $\pi^0$  has **205 MeV/c** in momentum. This is just above Cherenkov threshold for  $\pi^+$ , thus it is not identified as a ring in most of case.
- $\pi^+$  decays into  $\mu$  (invisible) and  $\nu$ ,  $\mu$  decays into  $e \nu_e \nu_\mu$ .
- $\pi^0$  decays into 2  $\gamma$ s.
- Search for 206 MeV/c  $\pi^0$  with Michel electron.

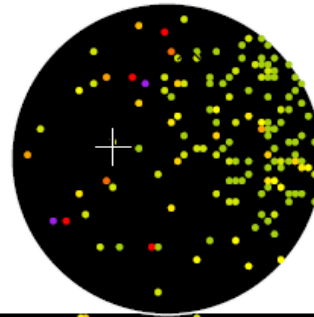
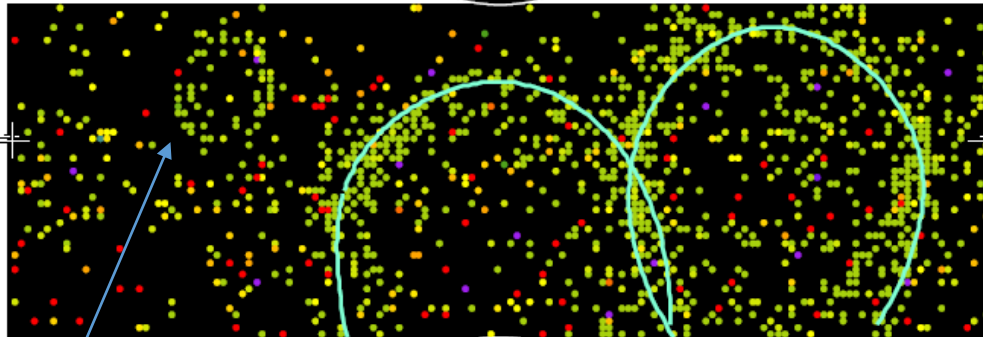
# Example of $p \rightarrow \nu K^+, K^+ \rightarrow \pi^+ \pi^0$

## Super-Kamiokande IV

Run 999999 Sub 0 Event 236  
D<sub>wall</sub>: 1076.4 cm  
E<sub>vis</sub>: 260.4 MeV  
2 e-like rings: mass = 155.2 MeV/c<sup>2</sup>

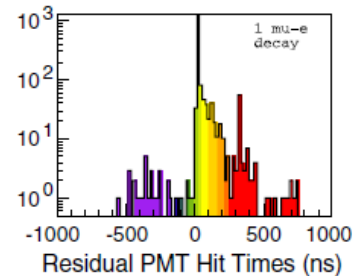
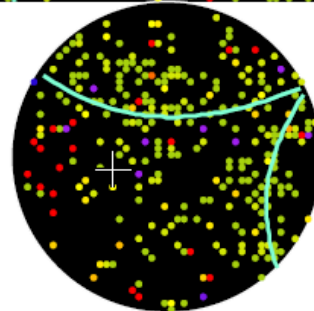
### Resid (ns)

- > 251
- 220- 251
- 188- 220
- 157- 188
- 125- 157
- 94- 125
- 62- 94
- 31- 62
- 0- 31
- -31- 0
- -62- -31
- -94- -62
- -125- -94
- -157- -125
- -188- -157
- < -188



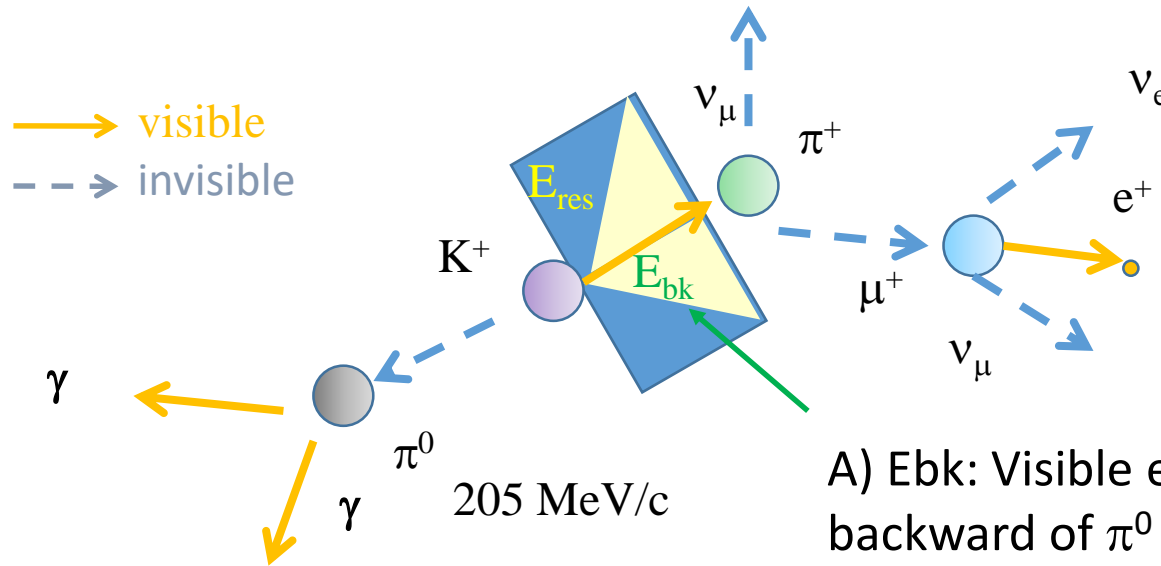
$\pi^+$

$\pi^0$

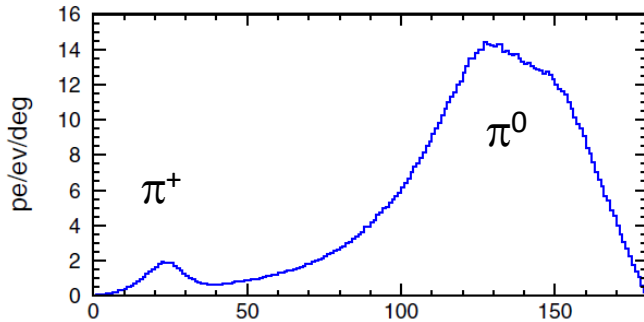


Look like a ring, but fake  
ring cut rejects this ring ...

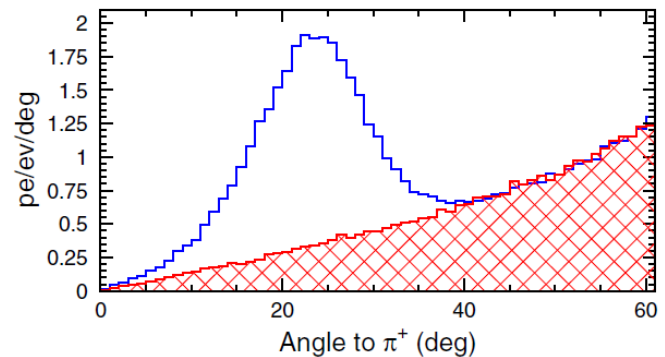
# Use $\pi^+$ information to select events



Charge distribution



Zoom



B) Make likelihood for hit pattern.

$\pi^+$  direction

# Selection criteria for $p \rightarrow \nu K^+, K^+ \rightarrow \pi^+ \pi^0$

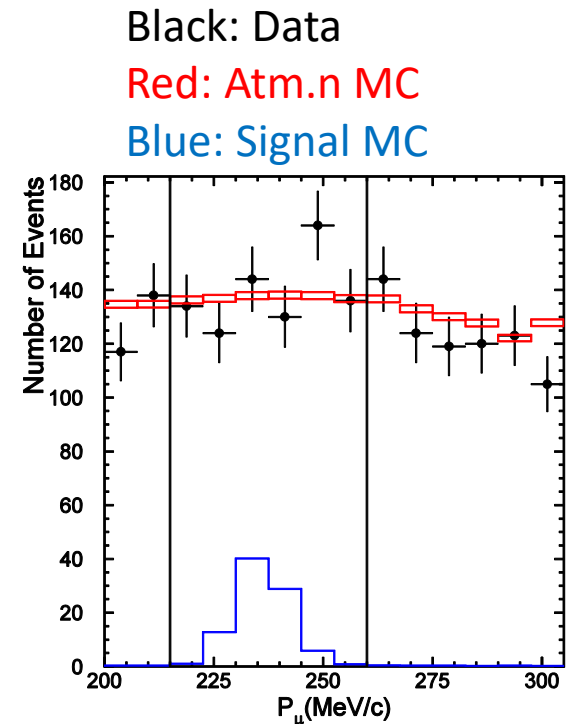
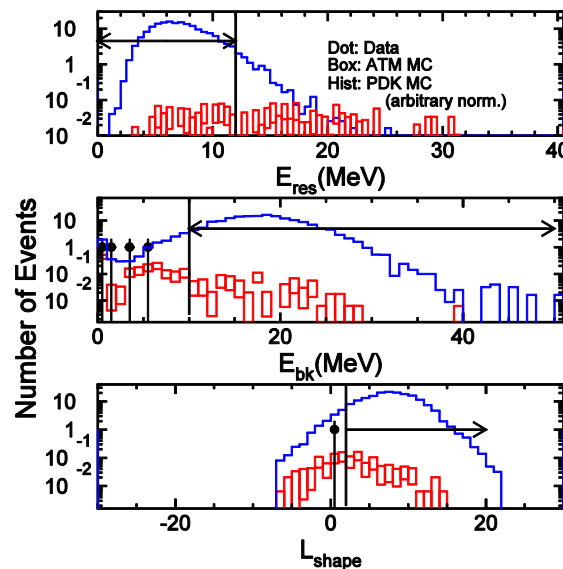
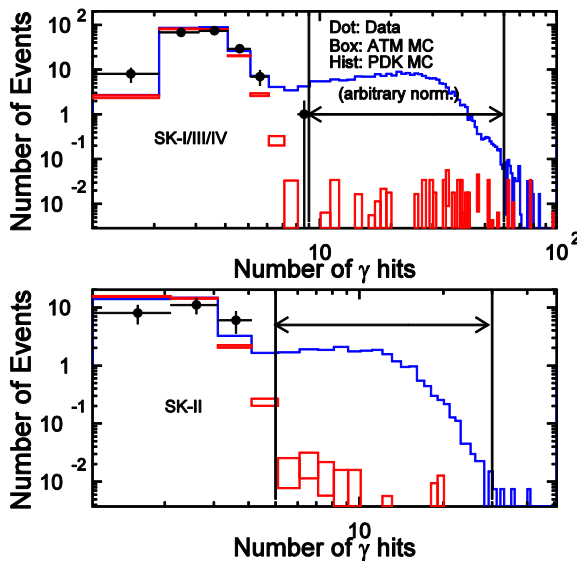
- 1 or 2 e-like rings with decay-e.
- $85 < M_{\pi^0} < 185$  MeV.
- $175 < P_{\pi^0} < 250$  MeV/c.
- $E_{bk}$ : visible energy sum in 140-180 deg. of  $\pi^0$  dir,  
 $E_{res}$ : in 90-140 deg,  
 $L_{shape}$ : Likelihood based on charge profile
  - $10 < E_{bk} < 50$  MeV
  - $E_{res} < 12$  MeV (20 MeV for 1ring)
  - $L_{shape} > 2.0$  (3.0 for 1ring)
- No neutrons
- Selection efficiency: 10 % ( $Br(K^+ \rightarrow \pi^+ \pi^0) = 21$  %)

# Background for $p \rightarrow \nu K^+$

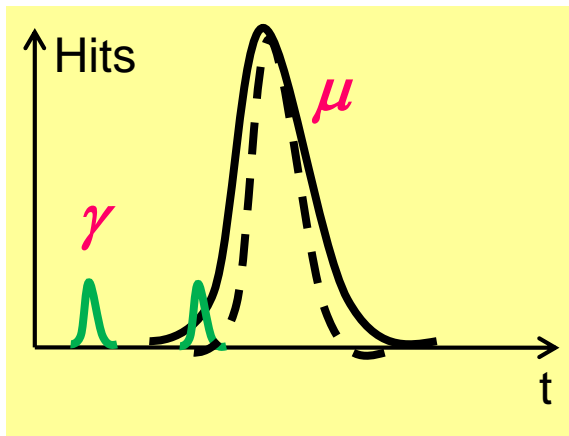
- Dominant background is  $K^+$  production by neutrino interactions.
  - $\nu p \rightarrow \nu \Lambda K^+$  ,  $\Lambda \rightarrow p \pi^-$  (BR:64 %, mostly invisible in WCD)
  - Emit nuclear  $\gamma$  as same as the signal.
- It is also rare interaction and we had poor information from very old bubble chamber. Large uncertainty.
- Recently MINERvA measures  $K^+$  production. It is very useful information for this analysis.

# 3-4. SK results (So far)

- Exposure: 365 kton · year
- Expected background: 0.3 events for  $K^+ \rightarrow \nu\mu$  with nuclear  $\gamma$ , 0.6 events for  $K^+ \rightarrow \pi^+\pi^0$ .
- No candidates observed and no excess in momentum distribution.
- Lower lifetime limit:  $> 0.8 \times 10^{34}$  year

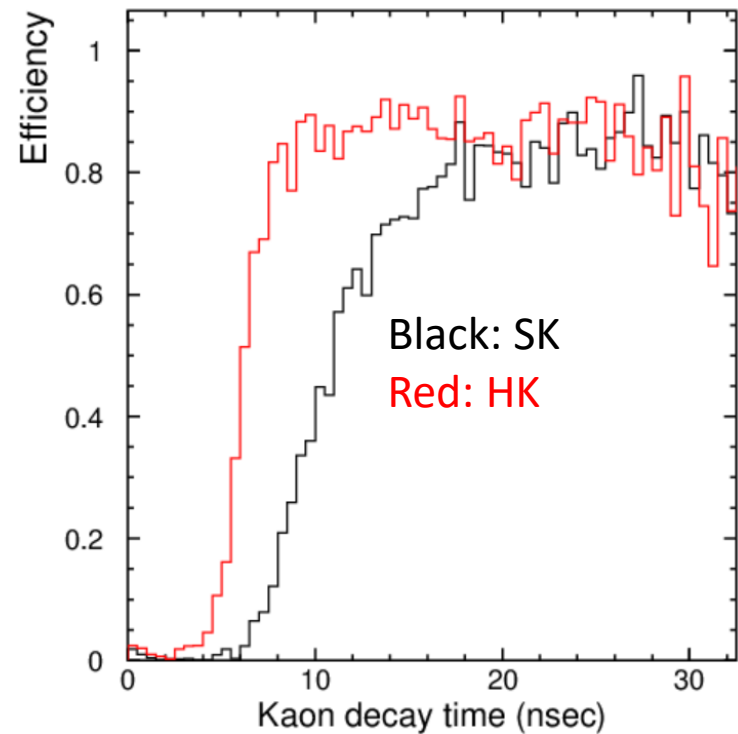


# Faster PMT response improves nuclear $\gamma$ tagging in $p \rightarrow \nu K^+$



- Time resolution: 2.2nsec (SK)  
→ 1.1 nsec (HK).
- Sharper time distribution of  $\mu$   
→  $\gamma$  close to  $\mu$  can be identified !

$\gamma$  tagging efficiency



(Better photon counting also contributes improvement)