## 

#### present and outlook...

VSON School (Vietnam)

19th July 2022

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D'ORSAY



### disclainner...

This description aims to provide <u>basis of neutrino oscillation with reactor neutrinos</u> based on the **experimental challenges** we faced / are facing now (including an outlook) while highlighting the topics **where reactors have the largest impact**...

Hence, this description will <u>benefit from all other lectures for complementary information</u> not follow a strict chronological account

standard 3 active neutrino

oscillation paradigm

### who I ann?

working on...

• Double Chooz experiment (dismantling now)

### knowns & unknowns...

Weak Flavour Neutrinos (3): **v(e)**, **v(µ)**, **v(t)** — <u>observed</u> 3! (same as quarks)

Mass Neutrinos (3): v(1), v(2), v(3) — <u>assumed</u> ≥3! [cosmology constraints ≤4]

 PMNS matrix (3x3; a la CKM): U, assumed unitarity (→violation?)
 discovery!

 • mixing parameters (3): θ<sub>13</sub>, θ<sub>12</sub>, θ<sub>23</sub> (octant?) — derived J [Jarkslog invariant]
 • CP-violation parameter (1): δ?

Mass Squared Differences (2):  $\delta m^2$  (i.e.  $\Delta m^2_{12}$ )  $\Delta m^2$  (i.e.  $\Delta m^2_{13}$  or  $\Delta m^2_{23}$ )

Mass Ordering (MO):

+ $\delta$ m<sup>2</sup> (solar data — observed!)

 $\pm \Delta m^2 \rightarrow$  which is the lightest neutrino V(1) or V(3)?

unknown [SM]



### reactors' sensitivity...

<u>directly sensitive</u> to:  $\theta$  | 3,  $\theta$  | 2,  $\delta$ m<sup>2</sup>,  $\Delta$ m<sup>2</sup>,  $\pm$ ? $\Delta$ m<sup>2</sup> (mass ordering) via vacuum oscillations

by ≥2030: the world's knowledge of **all those parameters** will be driven by reactors [key **input** for other experiments: accelerators, atmospheric, solar, etc]

indirect (i.e. synergies) boosting the combined sensitivity for  $\delta$  and  $\theta$ 23

**issue!**  $\phi$ (**reactor**) exhibits <u>several inconsistent features</u> ( $\rightarrow$  unsettled debate still)

beyond standard neutrino

### main achievements so far...

#### **Discovery of the neutrino**

#### Tendence mot reactory for solar anomaly a atmospheric anomaly

#### • | 990-2000:

•CHOOZ & Palo Verde support **atmospheric anomaly** Kamiokande's  $v(\mu) \rightarrow v(\tau) [\rightarrow SuperK]$ •(byproduct) direct (stringent) limit on **\thetaI3** (till 2011)

#### •2000-2010:

• KamLAND demonstrates the LMA solution of the "solar anomaly" [key for discovery]

• (byproduct) most spectacular spectral distortion seen so far ( $\theta$  I 2 and  $\delta$ m<sup>2</sup>)

#### •2010-2020:

• Daya Bay $\oplus$  Double Chooz $\oplus$  RENO: first measure the <u>predicted</u>  $\theta$  I 3 — T2K evidence too

•(synergy) critical for first  $\leq 2\sigma$  constraints on  $\delta$  by T2K and NOvA

• (byproduct Double Chooz):  $\phi$ (reactor) issue!  $\rightarrow$  reactor knowledge? vs sterile at  $\sim |eV^2$ ?

experimental setup...



new physics?

# the reactor.

Chooz

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### Nuclear reactors in the world

https://www.carbonbrief.org/mapped-the-worlds-nuclear-power-plants



- 440 operable reactors. 55 reactors under construction (Apr. 2020)
- Reactor's share of total electricity supplies in the world is ~10%



- Nuclear fuel is 3-4% of Uranium-235 (others are Uranium-238) provided by mining, milling, conversion to UF<sub>6</sub> & enrichment
- $\cdot$  Heat from nuclear fission is used to provide steam for a generator
- Control rods are used to control the fission rate
- · As a by-product, reactor core is an electron anti- $\nu$  source

### $\overline{\nu}_{\,e}$ production in reactors



#### **Reactor neutrino flux**

**Energy & neutrinos emitted per fission** 

 $^{235}\text{U} + \text{n} \rightarrow \text{A} + \text{B} + 6.1\text{e}^- + 6.1\,\overline{\nu}_e + 202\,\text{MeV} + 2.4\text{n}$ 

Neutrino flux of commercial reactor (~3 GW on average)

 $\cdot$  3 GW<sub>th</sub> = 3 × 10<sup>9</sup> J/s = 1.9 x 10<sup>22</sup> MeV

 $\rightarrow$  Neutrino flux is ~ 6 x 10<sup>20</sup>  $\overline{\nu}_{e}/s$ 

#### Thermal power

- Can be obtained from a power company to estimate  $\nu$  flux prediction
- Thermal power vs time in Chooz reactor for example



#### **Reactor neutrino spectrum**

#### Reference spectra weighted by fuel evolution



# the detector(s).

Double Chooz Near Detector: ~30 ton detection volume

 $d - July 202 \rightarrow dismantling)$ 

### inverse-B decay (IBD) interaction...

### **IBD:** anti- $v_e + p \rightarrow e^+ + n$



#### no e+ PID





### reactor neutrino spectrum...



#### **Double Chooz near-detector spectrum during ON-OFF-ON transition**

# the v discovery (1950's).

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### Reines & Cowan detector (300kg)...



today's inspiration!

### <sup>20</sup> Reines et al **detection strategy legacy**...

**PMT**  $\Leftrightarrow$  transparent medium

v interaction: coincidence and/or tagging

overburden

external shielding

loaded medium (113Cd)

signal modulation

~70years ago → much the same still now!

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### today's version of similar technology...



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### KamLAND/SNO+

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https://liquido.ijclab.in2p3.fr/

under construction



results release with the latest experimental results @ Neutrino 2022 conference (June 2022)

https://media.neutrino2022.org/talk/talk\_session\_apply/108/20220603220651\_33.pdf

LiquidO event-wise imaging...



self-segmentation

## unprecedented PID @ IMeV...

#### potential: redu

#### den / shielding



(reactor) opaque scintillation & (native) self-segmentation

**needless segmentation:** problematic @ IMeV (pollution, cost@complex, etc)

~ I MeV: reactor, geoneutrino, solar,  $\beta\beta$ -decay, etc <sup>25</sup>

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#### status on neutrino oscillation knowledge...

no conclusive sign of any extension so far!!

(inconsistencies vs uncertainties)

#### Δm² & +δm²

must measure all parameters→characterise & test (i.e. over-constrain) Standard Model

≥2030

	best kno	owledge	NuFIT5.0	foreseen	dominant	technique
θ <sub>12</sub>	3,0 %	SK⊕SNO	2,3 %	<1.0%	JUNO	reactor

μΔm²	3,0 %	DYB	1,3 %	≈1.0%	JUNU&DUNE&HK	reactor⊕beam
Mass Ordering	unknown	SK et al	NO @ <b>~3</b> σ	@5σ	JUNO⊕DUNE⊕HK	reactor⊕beam
CPV	unknown	T2K	3/2π @ ≈ <b>2σ</b>	@5σ?	DUNE⊕HK⊕ALL	beam driven
			(now)			(reactor-beam)

JUNO $\oplus$ DUNE $\oplus$ HK will lead precision in the field ( $\rightarrow$  CPV) except  $\theta_{13}$ !

NuFIT 5.0 (2020)

		NuFIT 5.0 (2020)			
2	2.8				_
or 10					
∆m <sup>2</sup>					
			-	$\sin^2 \theta_{12}$	
eV2		<u> </u>	late	$\theta_{12}/^{\circ}$	
0-3	2.2		ic d		
Ξ	E IE	3	heri	$\sin^2 \theta_{23}$	
n <sup>2</sup>			ospl	$\theta_{23}/^{\circ}$	
N -	2.6		tme	• 2.0	
	2°ETTTTTTTTTTTTTTTTTTTTTTTTTTTTTTTTTTTT		Хa	$\sin^2 \theta_{13}$	(
			E S]	$\theta_{13}/^{\circ}$	
			nou	$\delta_{\rm CD}/^{\circ}$	
		F F	vitk	UCP/	
4	TE E		4	$\Delta m_{21}^2$	
				$10^{-5} {\rm eV}^2$	
Sop 1	80			$\Delta m_{3\ell}^2$	
				$10^{-3} \text{ eV}^2$	
				$\sin^2 \theta_{12}$	
			ta	$\theta_{12}/^{\circ}$	
	$0.3  0.4  0.5  0.6  0.7  \Box$		da	127	
	Sill 0 <sub>23</sub>	-	eric	$\sin^2 \theta_{23}$	
			phe	$ heta_{23}/^{\circ}$	
-			nos	· 20	
eVa	75E JE	E	atr	$\sin^{-}\theta_{13}$	
- <sup>-2</sup>			SK	$\theta_{13}/^{\circ}$	
21 []			th	Scp /°	
Δm	ÝE 💙 E		wi	UCP/	
		<b>–</b> 1		$\Delta m^2_{21}$	
		E.LE.		$10^{-5} \text{ eV}^2$	
	0.2 0.25 0.3 0.35 0.4	0.015 0.02 0.025 0.03		$\Delta m^2_{3\ell}$	
	sin <sup>-</sup> θ <sub>12</sub>	sin <sup>-</sup> θ <sub>13</sub>		$10^{-3} \text{ eV}^2$	

		Normal Ordering (best fit)		Inverted Ordering $(\Delta \chi^2 = 2.7)$		
		bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range	
7	$\sin^2 \theta_{12}$	$0.304\substack{+0.013\\-0.012}$	$0.269 \rightarrow 0.343$	$0.304\substack{+0.013\\-0.012}$	$0.269 \rightarrow 0.343$	
uau	$\theta_{12}/^{\circ}$	$33.44_{-0.75}^{+0.78}$	$31.27 \rightarrow 35.86$	$33.45_{-0.75}^{+0.78}$	$31.27 \rightarrow 35.87$	
ITELIC	$\sin^2  heta_{23}$	$0.570\substack{+0.018\\-0.024}$	$0.407 \rightarrow 0.618$	$0.575\substack{+0.017\\-0.021}$	$0.411 \rightarrow 0.621$	
nusp	$ heta_{23}/^{\circ}$	$49.0^{+1.1}_{-1.4}$	$39.6 \rightarrow 51.8$	$49.3^{+1.0}_{-1.2}$	$39.9 \rightarrow 52.0$	
aun	$\sin^2 \theta_{13}$	$0.02221\substack{+0.00068\\-0.00062}$	$0.02034 \rightarrow 0.02430$	$0.02240^{+0.00062}_{-0.00062}$	$0.02053 \rightarrow 0.02436$	
CIC 1	$ heta_{13}/^{\circ}$	$8.57^{+0.13}_{-0.12}$	$8.20 \rightarrow 8.97$	$8.61_{-0.12}^{+0.12}$	$8.24 \rightarrow 8.98$	
ILIIUU	$\delta_{ m CP}/^{\circ}$	$195^{+51}_{-25}$	$107 \rightarrow 403$	$286^{+27}_{-32}$	$192 \rightarrow 360$	
8	$\frac{\Delta m_{21}^2}{10^{-5} \ \mathrm{eV}^2}$	$7.42_{-0.20}^{+0.21}$	$6.82 \rightarrow 8.04$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$	
	$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.514^{+0.028}_{-0.027}$	$+2.431 \rightarrow +2.598$	$-2.497^{+0.028}_{-0.028}$	$-2.583 \rightarrow -2.412$	
		Normal Ore	lering (best fit)	Inverted Orde	ering $(\Delta \chi^2 = 7.1)$	
		Normal Orc bfp $\pm 1\sigma$	lering (best fit) $3\sigma$ range	Inverted Orde bfp $\pm 1\sigma$	$\frac{\text{ering } (\Delta \chi^2 = 7.1)}{3\sigma \text{ range}}$	
	$\sin^2 \theta_{12}$	Normal Ord           bfp $\pm 1\sigma$ 0.304^{+0.012}_{-0.012}	lering (best fit) $3\sigma$ range $0.269 \rightarrow 0.343$	Inverted Orde bfp $\pm 1\sigma$ $0.304^{+0.013}_{-0.012}$	$\frac{\text{ering } (\Delta \chi^2 = 7.1)}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$	
rata	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$	Normal Ord bfp $\pm 1\sigma$ $0.304^{+0.012}_{-0.012}$ $33.44^{+0.77}_{-0.74}$	lering (best fit) $3\sigma$ range $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.86$	Inverted Order bfp $\pm 1\sigma$ $0.304^{+0.013}_{-0.012}$ $33.45^{+0.78}_{-0.75}$	$\frac{\text{ering } (\Delta \chi^2 = 7.1)}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.87$	
sric uata	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$ $\sin^2 \theta_{23}$	Normal Ord bfp $\pm 1\sigma$ $0.304^{+0.012}_{-0.012}$ $33.44^{+0.77}_{-0.74}$ $0.573^{+0.016}_{-0.020}$	$\frac{\text{lering (best fit)}}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.86$ $0.415 \rightarrow 0.616$	Inverted Order bfp $\pm 1\sigma$ $0.304^{+0.013}_{-0.012}$ $33.45^{+0.78}_{-0.75}$ $0.575^{+0.016}_{-0.019}$	$\frac{\text{pring } (\Delta \chi^2 = 7.1)}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.87$ $0.419 \rightarrow 0.617$	
spineric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/\circ}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/\circ}$	Normal Ord bfp $\pm 1\sigma$ $0.304^{+0.012}_{-0.012}$ $33.44^{+0.77}_{-0.74}$ $0.573^{+0.016}_{-0.020}$ $49.2^{+0.9}_{-1.2}$	$\frac{\text{lering (best fit)}}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.86$ $0.415 \rightarrow 0.616$ $40.1 \rightarrow 51.7$	Inverted Order bfp $\pm 1\sigma$ $0.304^{+0.013}_{-0.012}$ $33.45^{+0.78}_{-0.75}$ $0.575^{+0.016}_{-0.019}$ $49.3^{+0.9}_{-1.1}$	$\frac{\text{pring } (\Delta \chi^2 = 7.1)}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.87$ $0.419 \rightarrow 0.617$ $40.3 \rightarrow 51.8$	
aumospineric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/^{\circ}}$ $\sin^2 \theta_{13}$	Normal Ord bfp $\pm 1\sigma$ $0.304^{+0.012}_{-0.012}$ $33.44^{+0.77}_{-0.74}$ $0.573^{+0.016}_{-0.020}$ $49.2^{+0.9}_{-1.2}$ $0.02219^{+0.00062}_{-0.00063}$	$\frac{\text{lering (best fit)}}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.86$ $0.415 \rightarrow 0.616$ $40.1 \rightarrow 51.7$ $0.02032 \rightarrow 0.02410$	Inverted Order bfp $\pm 1\sigma$ $0.304^{+0.013}_{-0.012}$ $33.45^{+0.78}_{-0.75}$ $0.575^{+0.016}_{-0.019}$ $49.3^{+0.9}_{-1.1}$ $0.02238^{+0.00063}_{-0.00062}$	ering (Δ $\chi^2 = 7.1$ ) 3σ range 0.269 → 0.343 31.27 → 35.87 0.419 → 0.617 40.3 → 51.8 0.02052 → 0.02428	
on aunospheric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/\circ}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/\circ}$ $\frac{\sin^2 \theta_{13}}{\theta_{13}/\circ}$	Normal Ord bfp $\pm 1\sigma$ $0.304^{+0.012}_{-0.012}$ $33.44^{+0.77}_{-0.74}$ $0.573^{+0.016}_{-0.020}$ $49.2^{+0.9}_{-1.2}$ $0.02219^{+0.00062}_{-0.00063}$ $8.57^{+0.12}_{-0.12}$	$\frac{\text{lering (best fit)}}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.86$ $0.415 \rightarrow 0.616$ $40.1 \rightarrow 51.7$ $0.02032 \rightarrow 0.02410$ $8.20 \rightarrow 8.93$	Inverted Order bfp $\pm 1\sigma$ $0.304^{+0.013}_{-0.012}$ $33.45^{+0.78}_{-0.75}$ $0.575^{+0.016}_{-0.019}$ $49.3^{+0.9}_{-1.1}$ $0.02238^{+0.00063}_{-0.00062}$ $8.60^{+0.12}_{-0.12}$	ering (Δ $\chi^2 = 7.1$ ) 3σ range 0.269 → 0.343 31.27 → 35.87 0.419 → 0.617 40.3 → 51.8 0.02052 → 0.02428 8.24 → 8.96	
with SIA authospiteric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/\circ}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/\circ}$ $\frac{\sin^2 \theta_{13}}{\theta_{13}/\circ}$ $\delta_{\rm CP}/^\circ$	Normal Ord bfp $\pm 1\sigma$ $0.304^{+0.012}_{-0.012}$ $33.44^{+0.77}_{-0.74}$ $0.573^{+0.016}_{-0.020}$ $49.2^{+0.9}_{-1.2}$ $0.02219^{+0.00062}_{-0.00063}$ $8.57^{+0.12}_{-0.12}$ $197^{+27}_{-24}$	$\frac{\text{lering (best fit)}}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.86$ $0.415 \rightarrow 0.616$ $40.1 \rightarrow 51.7$ $0.02032 \rightarrow 0.02410$ $8.20 \rightarrow 8.93$ $120 \rightarrow 369$	Inverted Order bfp $\pm 1\sigma$ $0.304^{+0.013}_{-0.012}$ $33.45^{+0.78}_{-0.75}$ $0.575^{+0.016}_{-0.019}$ $49.3^{+0.9}_{-1.1}$ $0.02238^{+0.00063}_{-0.00062}$ $8.60^{+0.12}_{-0.12}$ $282^{+26}_{-30}$	$\frac{\text{pring } (\Delta \chi^2 = 7.1)}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.87$ $0.419 \rightarrow 0.617$ $40.3 \rightarrow 51.8$ $0.02052 \rightarrow 0.02428$ $8.24 \rightarrow 8.96$ $193 \rightarrow 352$	
WITH ON AUTIOSPITETIC DATA	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/^{\circ}}$ $\frac{\sin^2 \theta_{13}}{\theta_{13}/^{\circ}}$ $\frac{\delta_{\rm CP}/^{\circ}}{\frac{\Delta m_{21}^2}{10^{-5} \ {\rm eV}^2}}$	Normal Ord bfp $\pm 1\sigma$ $0.304^{+0.012}_{-0.012}$ $33.44^{+0.77}_{-0.74}$ $0.573^{+0.016}_{-0.020}$ $49.2^{+0.9}_{-1.2}$ $0.02219^{+0.00062}_{-0.00063}$ $8.57^{+0.12}_{-0.12}$ $197^{+27}_{-24}$ $7.42^{+0.21}_{-0.20}$	$\frac{\text{lering (best fit)}}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.86$ $0.415 \rightarrow 0.616$ $40.1 \rightarrow 51.7$ $0.02032 \rightarrow 0.02410$ $8.20 \rightarrow 8.93$ $120 \rightarrow 369$ $6.82 \rightarrow 8.04$	Inverted Order bfp $\pm 1\sigma$ $0.304^{+0.013}_{-0.012}$ $33.45^{+0.78}_{-0.75}$ $0.575^{+0.016}_{-0.019}$ $49.3^{+0.9}_{-1.1}$ $0.02238^{+0.00063}_{-0.00062}$ $8.60^{+0.12}_{-0.12}$ $282^{+26}_{-30}$ $7.42^{+0.21}_{-0.20}$	$\frac{\text{pring } (\Delta \chi^2 = 7.1)}{3\sigma \text{ range}}$ $0.269 \rightarrow 0.343$ $31.27 \rightarrow 35.87$ $0.419 \rightarrow 0.617$ $40.3 \rightarrow 51.8$ $0.02052 \rightarrow 0.02428$ $8.24 \rightarrow 8.96$ $193 \rightarrow 352$ $6.82 \rightarrow 8.04$	

experimental setup...



new physics

# (reactor flux)

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reactor flux discrepancy...



#### now ≤7.0% mismatch between ILL-prediction and data

#### 2019 world reactor flux knowledge...



reactor flux (data) precision <1.0%

#### shape distorsion common across experiments...



only one experiment in tension: **Bugey3** (flat-ish) — spectral reference before **reactor-\thetaI3** 

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reactor flux uncertainty...



nature physics

ARTICLE

First Double Chooz  $\theta_{13}$  Measurement via Total Neutron Capture Detection

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## reactor ultimate flux...

#### Reevaluating reactor antineutrino spectra with new measurements of the ratio between $^{235}$ U and $^{239}$ Pu $\beta$ spectra

V. Kopeikin,<sup>1</sup> M. Skorokhvatov,<sup>1,2</sup> and O. Titov<sup>1,\*</sup>

<sup>1</sup>National Research Centre Kurchatov Institute, 123182, Moscow, Russia <sup>2</sup>National Research Nuclear University MEPhI (Moscow Engineering Physics Institute), 115409, Moscow, Russia (Dated: May 31, 2021)

We report a reanalysis of the reactor antineutrino energy spectra based on the new relative measurements of the ratio  $R = {}^{e}S_{5}/{}^{e}S_{9}$  between cumulative  $\beta$  spectra from  ${}^{235}$ U and  ${}^{239}$ Pu, performed at a research reactor in National Research Centre Kurchatov Institute (KI). A discrepancy with the  $\beta$  spectra measured at Institut Laue-Langevin (ILL) was observed, indicating a steady excess of the ILL ratio by the factor of  $1.054 \pm 0.002$ . We find a value of the ratio between inverse beta decay cross section per fission for  ${}^{235}$ U and  ${}^{239}$ Pu:  $({}^{5}\sigma_{f}/{}^{9}\sigma_{f})_{KI} = 1.45 \pm 0.03$ , and then we reevaluate the converted antineutrino spectra for  ${}^{235}$ U and  ${}^{238}$ U. We conclude that the new predictions are consistent with the results of Daya Bay and STEREO experiments.

arXiv:2103.01684v2 [nucl-ex] 28 May 2021

#### $R=0.925\pm0.010(exp)\pm0.023(model) \rightarrow R\approx1.02\pm0.010(exp)\pm0.02(model)$

experiment flux uncertainty will drive?

solve the reactor "issue" (anomaly):

(discrepancy data and ILL-prediction)

Uncertainty (%)	$\mathbf{ND}$	
Proton Number	0.66	
Thermal Power	0.47	reducible!!
TnC Selection	0.24	
Background	0.18	
Energy per Fission	0.16	
$\theta_{13}$ Correction	0.16	
Statistics	0.22	
Total	0.97	

## 2020 improvements...





#### ≥2020 ab initio estimation





## first LiquidO's experiment...

#### European Innovation Council



#### UK Research and Innovation

Innovation Fundament

#### **CDF**(France) **first time in neutrinos!**

•ČIEMAT (Spain)
 •IJCLab/Université Paris-Saclay (France)
 •J-G Universität Mainz (Germany)
 •Subatech/Nantes Université (France)
 •Sussex University (UK)

Charles University (Czech Republic)
INFN-Padova (Italy)
UC-Irvine (US)
Universidade Estadual de Londrina (Brasil)
PUC-Rio de Janeiro (Brasil)
Queen's University (Canada)
University of Zaragoza (Spain)
Tohoku University / RCNS (Japan)

-OTech'

CLOUD collaboration (EDF I3 institutions over 10 countries)



# θ 3 ⊕ Δm<sup>2</sup>

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#### summary on today's $\Theta$ 3 knowledge/experiments...

#### reactor-θI3 experiments: DC⊕DYB⊕RENO



"naively extrapolating" from reactor- $\theta$  3 experiments...

~0.01%??!! possible to improve at all?



#### 2020 world status in $\Theta$ 13...

#### θ<sub>13</sub> consistent (≤2σ)



minor tension ( $\leq 2\sigma$ ) & slight increase (2016 $\rightarrow$ 2018)

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T2K@reactor best knowledge CP-Violation...



### θI3 implications

#### CPV phase vs θI3

[constrained by reactor]

#### CPV phase vs θ23

[octant ambiguity]



#### leptonic sector unitarity with LiquidO?

## Iowards Unitarity?

(how far?)

PS-HEP Conference @ Ghent (Belgium

Anatael Cabrera CNRS/IN2P3 LAL@Orsay LNCA@Chooz

Conference @ HEP-European Physics Society (July 2019 @ Ghent Belgium) Web: https://indico.cern.ch/event/577856/contributions/3421609/

## Super Cool

T2K@reactor best knowledge CP-Violation...



## **θI3** implications powerful constraint **CPV** phase vs $\theta$ | 3 [constrained by reactor] CPV phase vs θ23

[octant ambiguity]



# <u>θ\_2⊕δm²</u>

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#### (stunning) experiment rationale...



#### univocal neutrino oscillation pattern...



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#### <sup>s1</sup> KamLAND's spectral distortion (and reactor-off)...

A. Gando et al., Phys. Rev. D 88, 033001 (2013).







#### JUNO location...



 $\theta_{12} \oplus \delta m^2$  (slow)  $\theta_{13} \oplus \Delta m^2$  (fast)

~late 2022

#### rate+shape sensitivity evolution...

#### consider all systematics with state of the art knowledge (KL, DC, DYB)



rate+shape $\rightarrow$  negligible rate uncertainties

## JUNO precision...



	Mass Ordering	$ \Delta m^2_{32} $	$ \Delta m^2_{21} $	$\sin^2 \theta_{12}$	$\sin^2 \theta_{13}$
6 years of data	$3-4\sigma$	$\sim 0.18\%$	$\sim 0.30\%$	$\sim 0.5\%$	$\sim 14\%$
20 years of data	$5\sigma$	$\sim 0.15\%$	$\sim 0.25\%$	$\sim 0.4\%$	$\sim 7\%$



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#### reactor impact in $\theta$ **3**- $\theta$ **12** plane...



## Mass Ordering reactors' role



PREPARED FOR SUBMISSION TO JHEP

IFT-UAM/CSIC-112, YITP-SB-2020-21

#### The fate of hints: updated global analysis of three-flavor neutrino oscillations

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ABSTRACT: Our herein described combined analysis of the latest neutrino oscillation data presented at the Neutrino2020 conference shows that previous hints for the neutrino mass ordering have significantly decreased, and normal ordering (NO) is favored only at the 1.6 $\sigma$ level. Combined with the  $\chi^2$  map provided by Super-Kamiokande for their atmospheric neutrino data analysis the hint for NO is at 2.7 $\sigma$ . The CP conserving value  $\delta_{\rm CP} = 180^{\circ}$ is within 0.6 $\sigma$  of the global best fit point. Only if we restrict to inverted mass ordering, CP violation is favored at the ~ 3 $\sigma$  level. We discuss the origin of these results – which are driven by the new data from the T2K and NOvA long-baseline experiments–, and the relevance of the LBL-reactor oscillation frequency complementarity. The previous 2.2 $\sigma$ tension in  $\Delta m_{21}^2$  preferred by KamLAND and solar experiments is also reduced to the 1.1 $\sigma$ level after the inclusion of the latest Super-Kamiokande solar neutrino results. Finally we present updated allowed ranges for the oscillation parameters and for the leptonic Jarlskog determinant from the global analysis.

KEYWORDS: neutrino oscillations, solar and atmospheric neutrinos



today's world data eads to...

•Super-Kamiokande (most info so far)

2.70 (202

some fragility?

NMO favoured

what are the leading experiments

what's going to happen next?

NuFitv5.0: today's world knowledge — what about tomorrow?

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NuFitv5.0: maginilise today's world knowledge – CPV,  $\theta_{23}$ ,  $\theta_{13}$ , ...

## the building blocks...

	direct sensitivity	nuisance	sensitivity	combined sensitivity
JUNO	ultra precision oscillation	θ <sub>13</sub> ?	<b>~</b> 3σ	<mark>δ(</mark> Δm²)≤0.5%
NOvA			<mark>~3-4σ</mark> (~800km baseline)	δ(Δm²)~1.0%
T2K		mainly CPV	<2σ	δ(Δm²)~1.0%
HyperK	fake CPV (due to Earth)	(θ <sub>23</sub> too)	(~250km baseline)	δ(Δm²)~0.5%
DUNE			>5σ! (~1200km baseline)	δ(Δm²)~0.5%
Atmospherics		mainly θ <sub>23</sub> (CPV too)	~3-60 (many baselines)	δ(Δm²) poor

## the building blocks...











#### scientific reports

Explore content $\checkmark$ About the journal $\checkmark$ Publish with us $\checkmark$	Abstract		
	The measurement of neutrino mass ordering (MO) is a fundamental element for the		
nature > scientific reports > articles > article	understanding of leptonic flavour sector of the <i>Standard Model of Particle Physics</i> . Its determination relies on the precise measurement of $\Delta m_{-}^2$ and $\Delta m_{-}^2$ using either neutrino		
<u>Inatore</u> / <u>Scientine reports</u> / <u>articles</u> / article	<i>vacuum oscillations</i> , such as the ones studied by medium baseline reactor experiments, or		
	matter effect modified oscillations such as those manifesting in long-baseline neutrino beams		
Article Open Access Published: 30 March 2022	(LB $ u$ B) or atmospheric neutrino experiments. Despite existing MO indication today, a fully		
	resolved MO measurement ( $\geq 5\sigma$ ) is most likely to await for the next generation of neutrino		
Synergies and prospects for early resolution of the	experiments: JUNO, whose stand-alone sensitivity is $\sim 3\sigma$ , or LBvB experiments (DUNE and		
eynergiee and prospecto for early recondition of the	Hyper-Kamiokande). Upcoming atmospheric neutrino experiments are also expected to		
neutrino mass ordering	provide precious information. In this work, we study the possible context for the earliest full		
i i i i i i i i i i i i i i i i i i i	MO resolution. A firm resolution is possible even before 2028, exploiting mainly vacuum		
	oscillation, upon the combination of JUNO and the current generation of LB $ u$ B experiments		
Anatael Cabrera, Yang Han, Hongzhao Yu + Show authors	(NOvA and T2K). This opportunity is possible thanks to a powerful synergy boosting the		
	overall sensitivity where the sub-percent precision of $\Delta m^2_{32}$ by LBvB experiments is found to		
Scientific Reports 12 Article number: 5393 (2022) Cite this article	be the leading order term for the MO earliest discovery. We also found that the comparison		
	between matter and vacuum driven oscillation results enables unique discovery potential for		
198 Accesses Metrics	physics beyond the Standard Model.		



#### powerful synergy/UNO vs NOvA $\sigma$ T2K: high precision disappearance $\Delta m^{2}_{32...}$



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#### Mass Ordering resolution [now at ~30]...



#### discrepancies may lead to discoveries!

#### JUNO and HK⊕DUNE (Disappearance)

JUNO+NOvA+T2K

## Unitarity

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φ(reactor)/φ(solar) ⊕ θI3(all) ⊕ θI2(all)



## **consider full matrix structure** (not just composition)

### why shape?

•large mixing but a small one!

largest CP-Violation (SM)

•any symmetry behind? [Nature's caprice?]

### U<sub>3x3</sub> unitary?

[so far assumed]

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#### CKM vs PMNS...



d s b



**elegance** (symmetry)



A. De Gouvea, H. Murayama, hep-ph/0301050; PLB, 2015.L. Hall, H. Murayama, N. Weiner, hep-ph/9911341.

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unitarity is behind all our definitions...

#### UNITARITY

**IF 3 neutrino standard states** 

[in agreement with quark's **3 families**]

 $\Rightarrow$  2 mass difference:  $\Delta m^2 \& \delta m^2$ 

 $\Rightarrow$  3 independent mixing angles:  $\theta_{12}, \theta_{23}, \theta_{13}$ 

 $\Rightarrow$  | (Dirac) CP-Violating phase:  $\delta_{CP}$ 

[i.e. a 3x3 unitarity matrix may be complex]

if 4 families

3x3 effective approximation

**UNITARITY** testing for new families + more!!


#### since no CPV (yet) ⇒ test PMNS Unitarity via "each row"

$$U_{l1}|^{2} + |U_{l2}|^{2} + |U_{l3}|^{2} = 1$$

 $|U_{e1}|^2 + |U_{e2}|^2 + |U_{e3}|^2 = 1 \Rightarrow$  explore "electron top-row"

only " $\theta_{12}$ " and " $\theta_{13}$ "

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# today's (**e-row**) **unitarity** knowledge...



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#### unitary explorations limited by absolute flux uncertainty

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reactor flux uncertainty...





ARTICLE

First Double Chooz  $\theta_{13}$  Measurement via Total Neutron Capture Detection

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neutrino last modification of the Standard Model... more discoveries

# Super Cool

rationale...

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# • high precision SM's neutrino oscillation $\implies$ synergise with JUNO & HK $\oplus$ DUNE

# neutrinos as probe BSM→ discoveries? ⇒ beyond today's paradigm?

## status on neutrino oscillation knowledge...

#### **SuperChooz**

#### **SuperChooz**

#### **BSM discovery?**

	3,0 %	SK⊕SNO	2,3 %	≤0.5%	JUNO⊕SC	reactor⊕solar
θ23	5,0 %	NOvA+T2K	2,0 %	≲1.0%?	DUNE⊕HK [SC]	beam (octaint)
			1,5 %	<0.5%	Supervise	
СР	violation?	T2K+NOvA	~3/2π @ ≤2σ	@5σ?	DUNE⊕HK [ <b>SC</b> ]	beam driven
СРТ	violation?			< %?	SuperChooz	reactor⊕solar
Unitarity	violation?			< %?	SuperChooz	reactor⊕solar

#### **reactor** $\oplus$ **beam** & **solar** again via SC — SC's atmospherics under study... nice!!

#### SUPERRCHOODZ pathfinder (i.e. experimental exploration)

#### The SuperChooz Pathfinder Exploration

■ Monday 20 Jun 2022, 11:00 → 12:00 Europe/Paris

https://indico.ijclab.in2p3.fr/event/7663/

• 100/-1-A900 - Auditorium Joliot Curie (IJCLab)

Videoconference

PHE Seminars

▶ Join

#### **11:00** $\rightarrow$ 12:00 The SuperChooz Pathfinder Exploration ¶

A new possible flagship neutrino experiment in Europe opens by exploiting a unique opportunity that has long been hidden in the Chooz nuclear reactor site – Europe's historical and most powerful reactor neutrino science site. The "SuperChooz" project benefits by existence of 2 caverns, formerly hosting the Chooz-A nuclear reactor complex, built in the 60's, that are becoming vacant upon its dismantling completion. The caverns hold a total volume of up to 50 000m^3, thus directly comparable to the size of SuperKamiokande. Its potential use for fundamental science purposes is under active discussion with EDF, thus starting the "pathfinder" exploration era. The SuperChooz caverns combined with the existing ~1km baseline to two of the most powerful N4 Chooz PWR nuclear reactors make this site a unique asset to Europe and specially France. Experimentally, the challenge is the poor overburden (order 100m underground). However, the novel LiquidO technology, born as a byproduct of Double Chooz experiment in the same site, heralds the potential for unprecedented active background rejection of up to two orders of magnitude, thus providing feasibility potential ground for a hypothetical SuperChooz. In this seminar, the rationale of the scientific programme of the hypothetical SuperChooz experiment will be highlighted for the first time. The project is aimed to address some of the most fundamental symmetries (studies under completion) behind the Standard Model, including a design that may open for key synergies that may boost the sensitivities of other neutrino flagship experiments such as DUNE (US), JUNO (China) and HyperKamiokande (Japan).

Speaker: Dr Anatael CABRERA (IJCLab/LNCA - IN2P3/CNRS)

🕲 1h 🖉 🗸

ICLab⊕Subatech teams — Octobre 2020

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EDF CONTS

## CNRS/1712-3 direction EDF CNRS exploring...

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# the remaining challenges... (my view — likely biassed somewhat)

• improve absolute knowledge precision (ex. flux cancels by <u>multi-detector</u>) $\rightarrow$ discoveries?

•intrinsic limitations (no appearance, etc)  $\rightarrow$  empower synergies with accelerators, solar, etc.

• $\theta I 2 \oplus \delta m^2$  precision: is likely hard to improve (few per mille) after JUNO — world best  $\leq I$  year of data

• $\theta I 3 \oplus \Delta m^2$  precision: still improve for  $\theta I 3$  — nobody knows how to!! [ $\rightarrow$ SuperChooz?]

• $\theta$  3 is <u>one of the most intriguing parameter</u> of the PMNS (tiny term among many large terms)  $\Rightarrow$  pointing to a feature(s) or symmetry? certainly this is BSM territory...

•understand meaning of structure of the PMNS (i.e. large mixing) — very different from CKM

challenge remains on BACKGROUND control

? JUNO

goal:

S/BG ≥100 with minimal overburden

LiquidO?