

Hyper-Kamiokande Status and Prospects with neutrino oscillation measurements

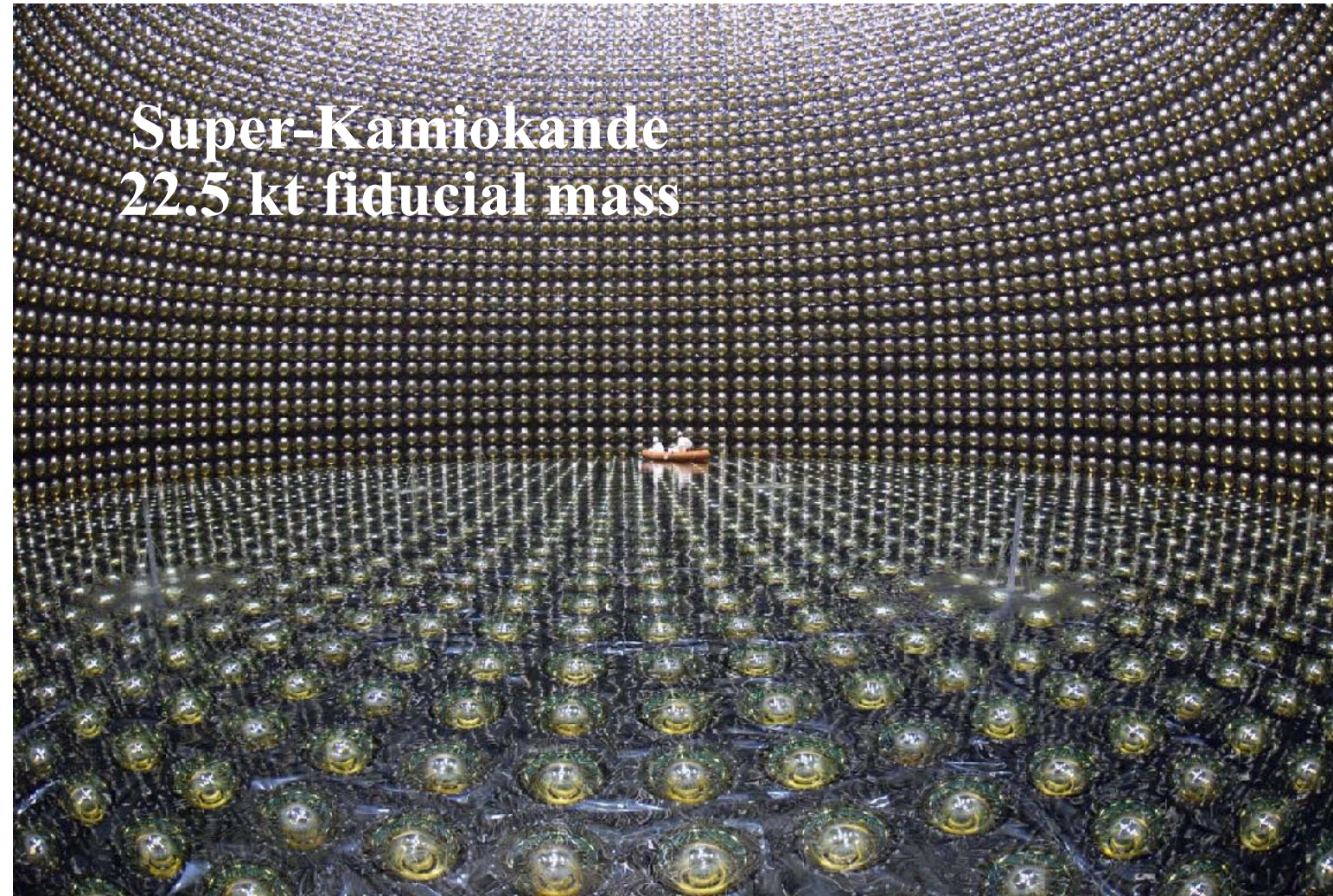


**T. Nakaya (Kyoto U.)
for the Hyper-Kamiokande Collaboration**

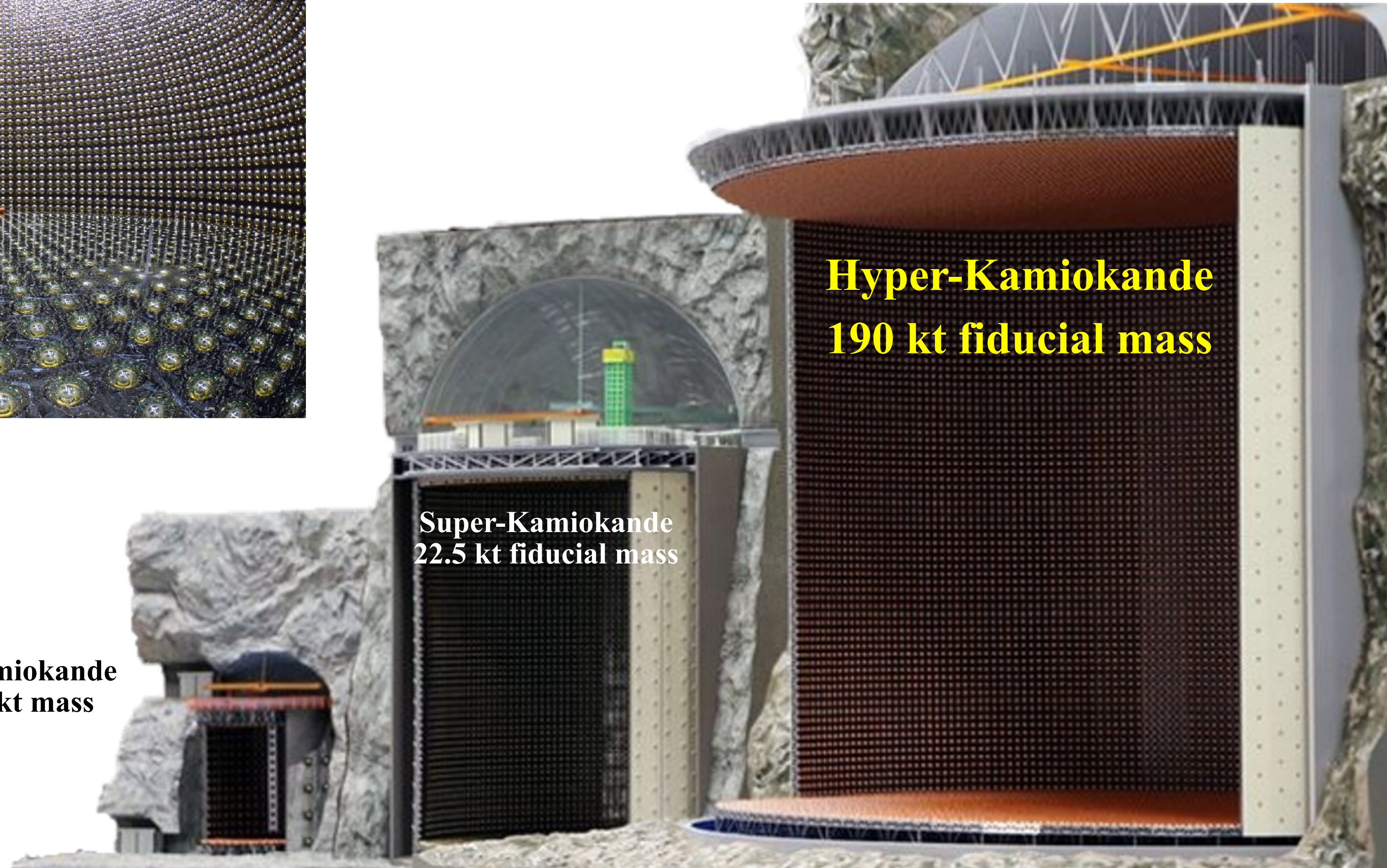


Hyper-Kamiokande

Water Cherenkov detectors in Kamioka



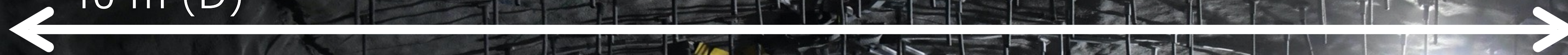
Kamiokande
3 kt mass



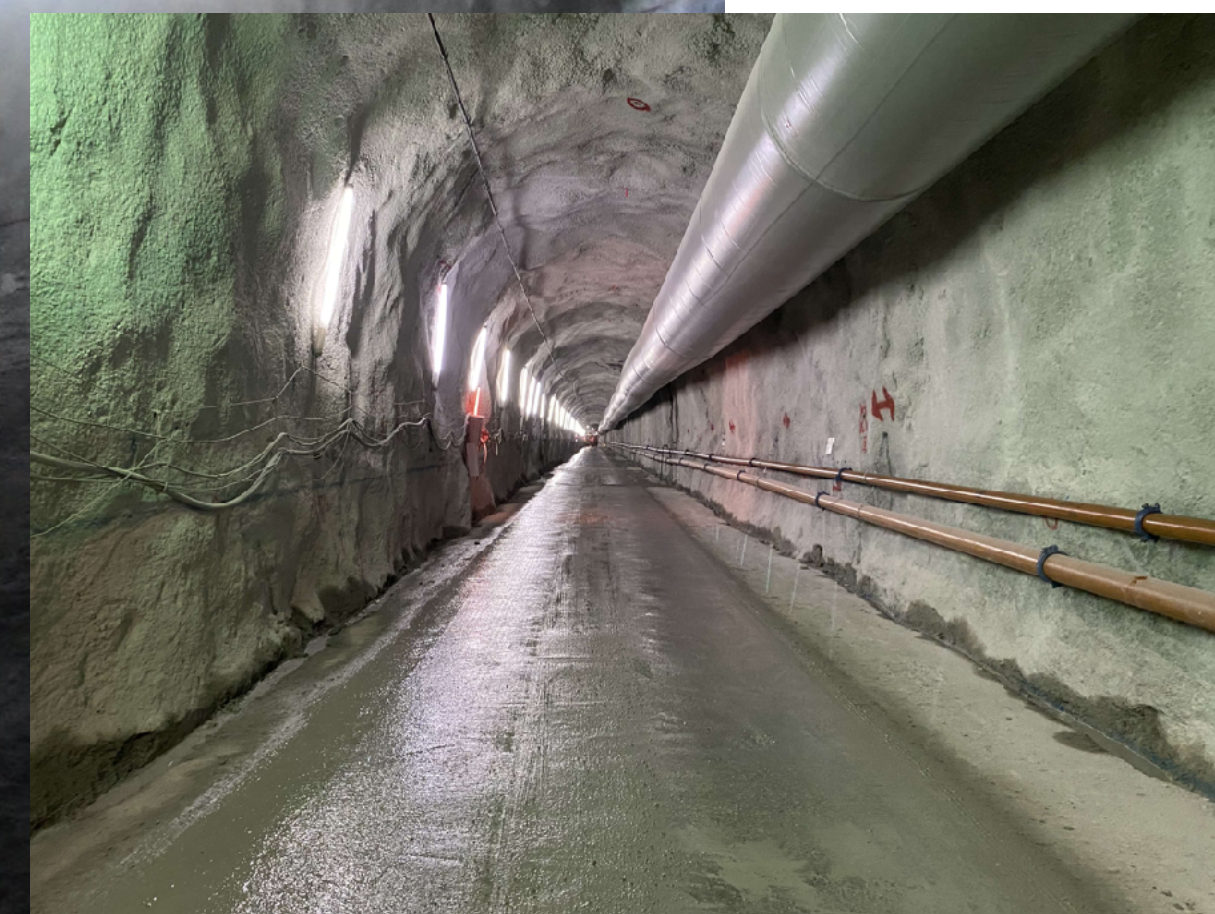
Top of the Detector Cavern

Current Status

40 m (D)

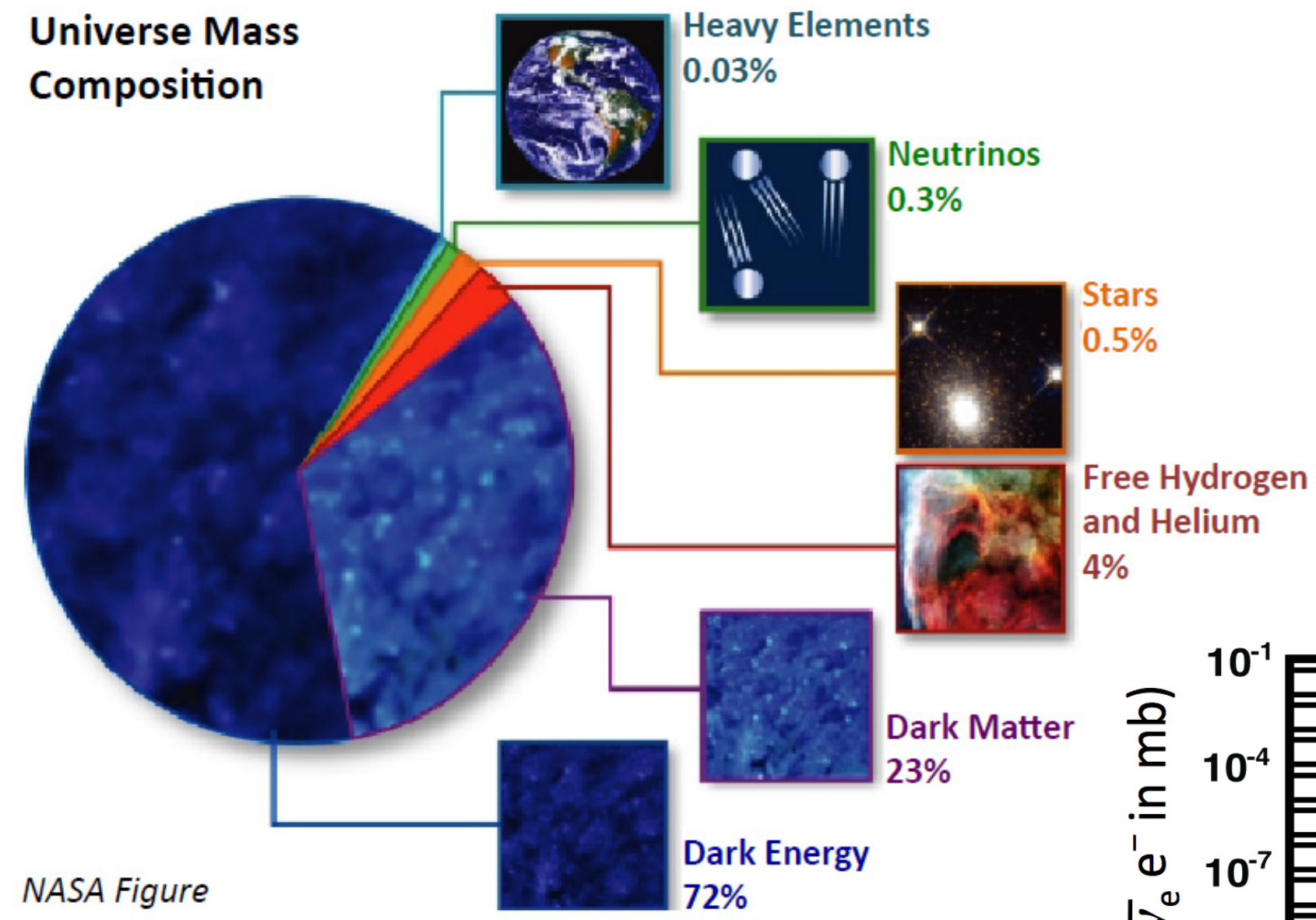


Top of the Detector Cavern
(14th March 2023)

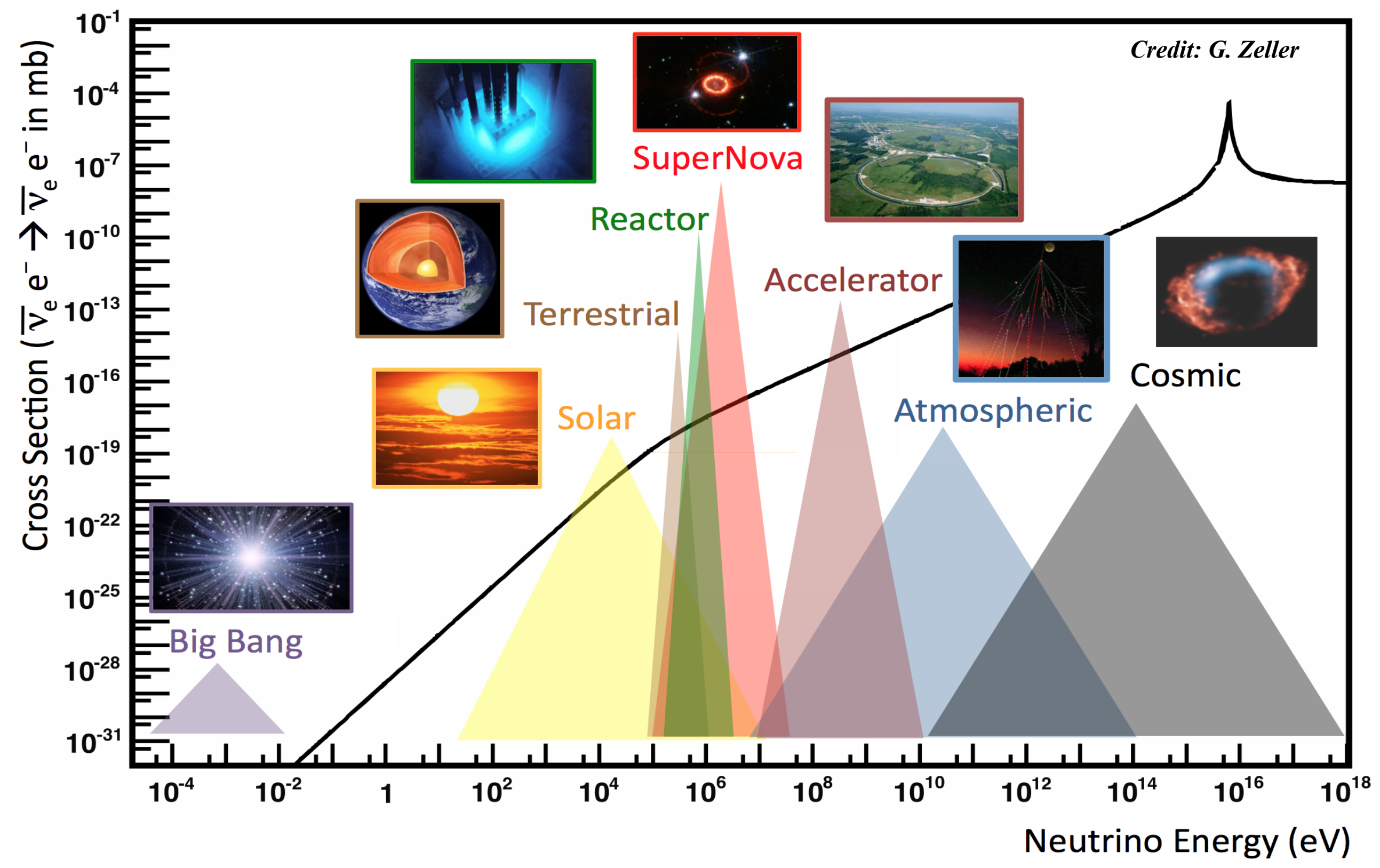


Why Neutrinos?

They are always around us and mysterious!



- $T_\nu = 1.95 \text{ K}$
- $\#N_\nu = 112 \times 3 \text{ cm}^{-3}$

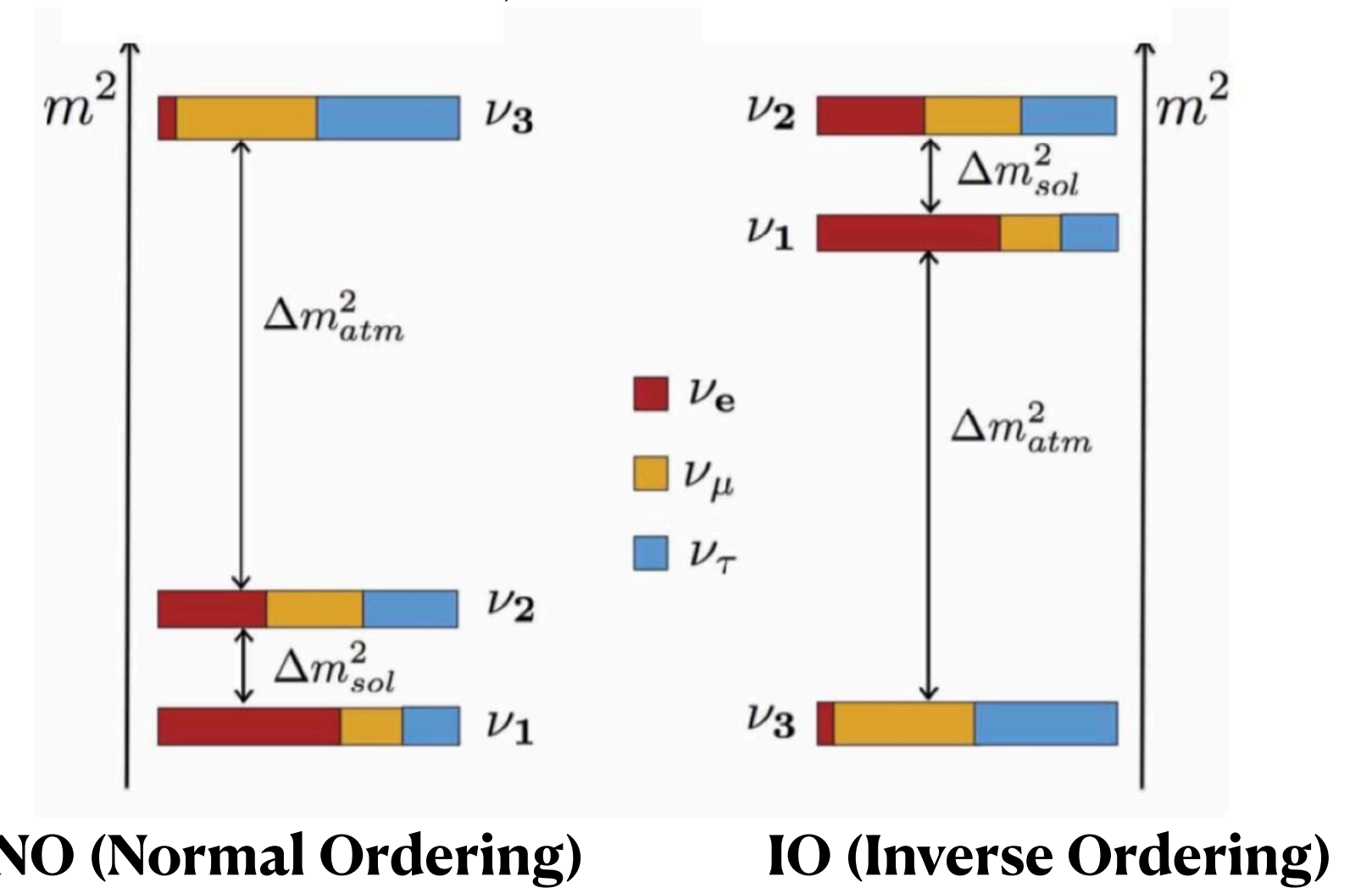
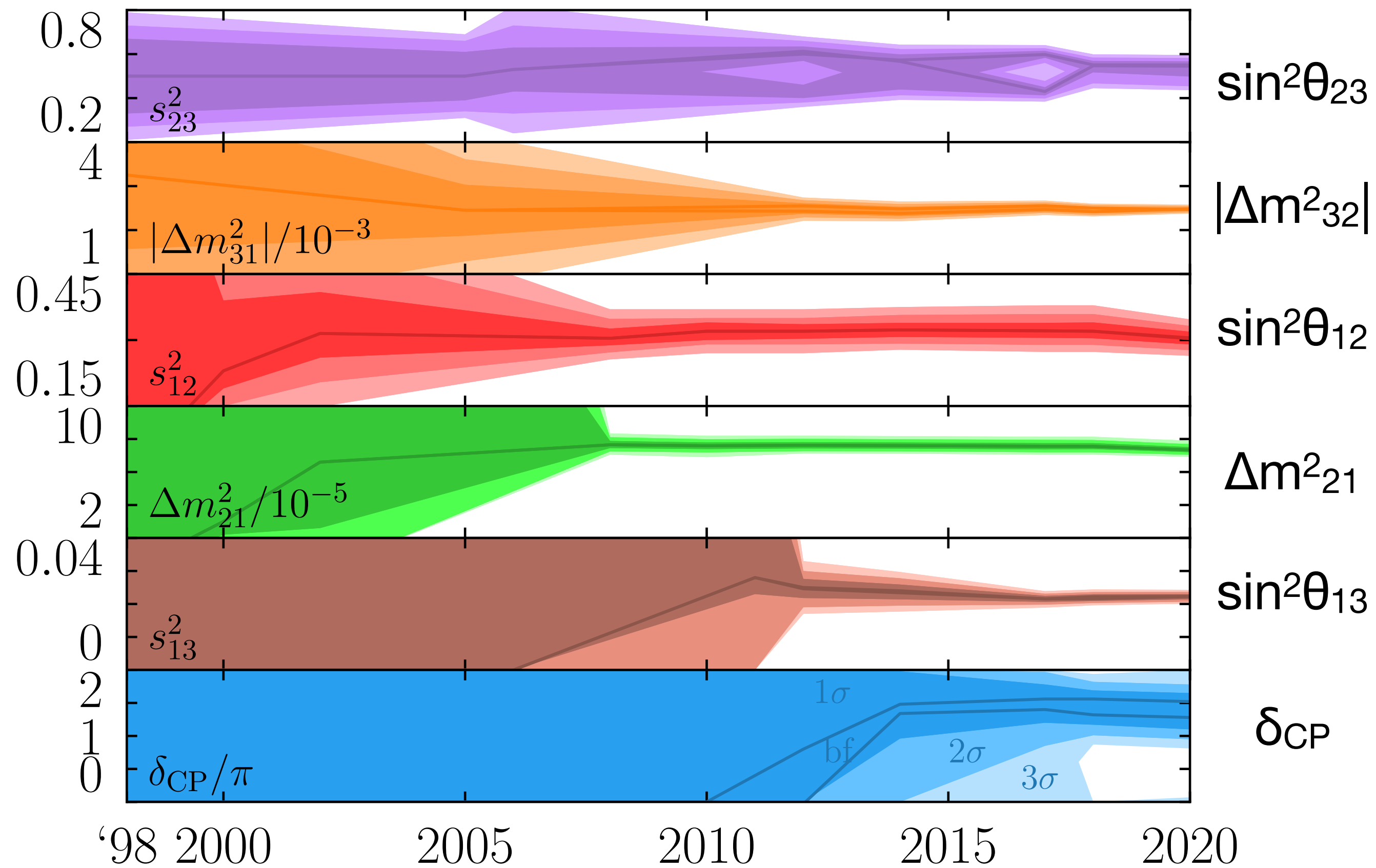


- Particle Physics
- AstroPhysics
- Cosmology
- High energy Astro-particle physics
- Nuclear physics

Neutrino Oscillations

Maki-Nakagawa-Sakata
Mixing matrix

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-\delta_{CP}} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{\delta_{CP}} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

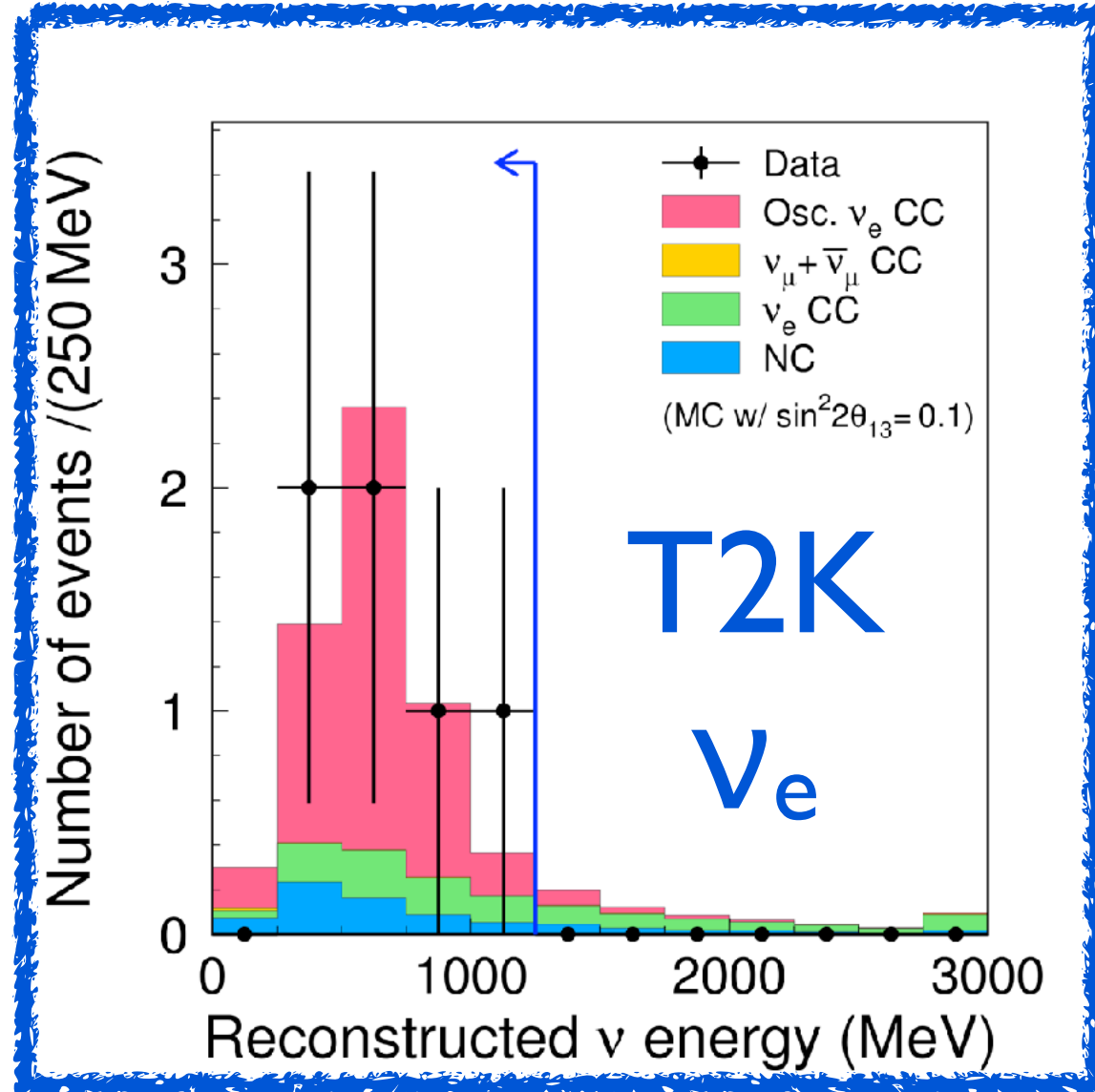


Current major targets

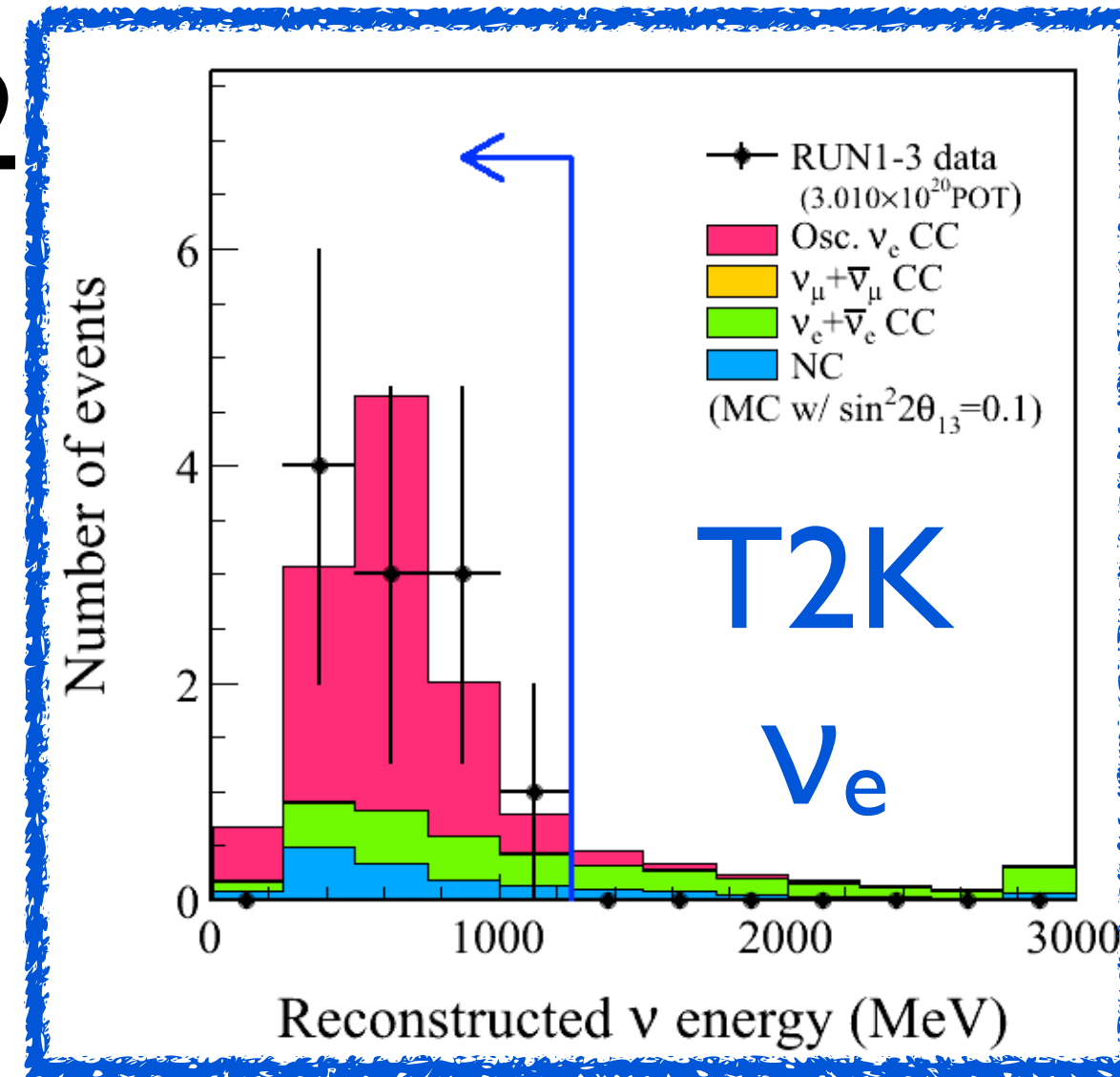
- **CP violation**
- **Mass ordering (NO or IO)**
- **Precision**
 - θ_{23} octant ($\leq 45^\circ$?)

Large θ_{13} opens the window to study CPV

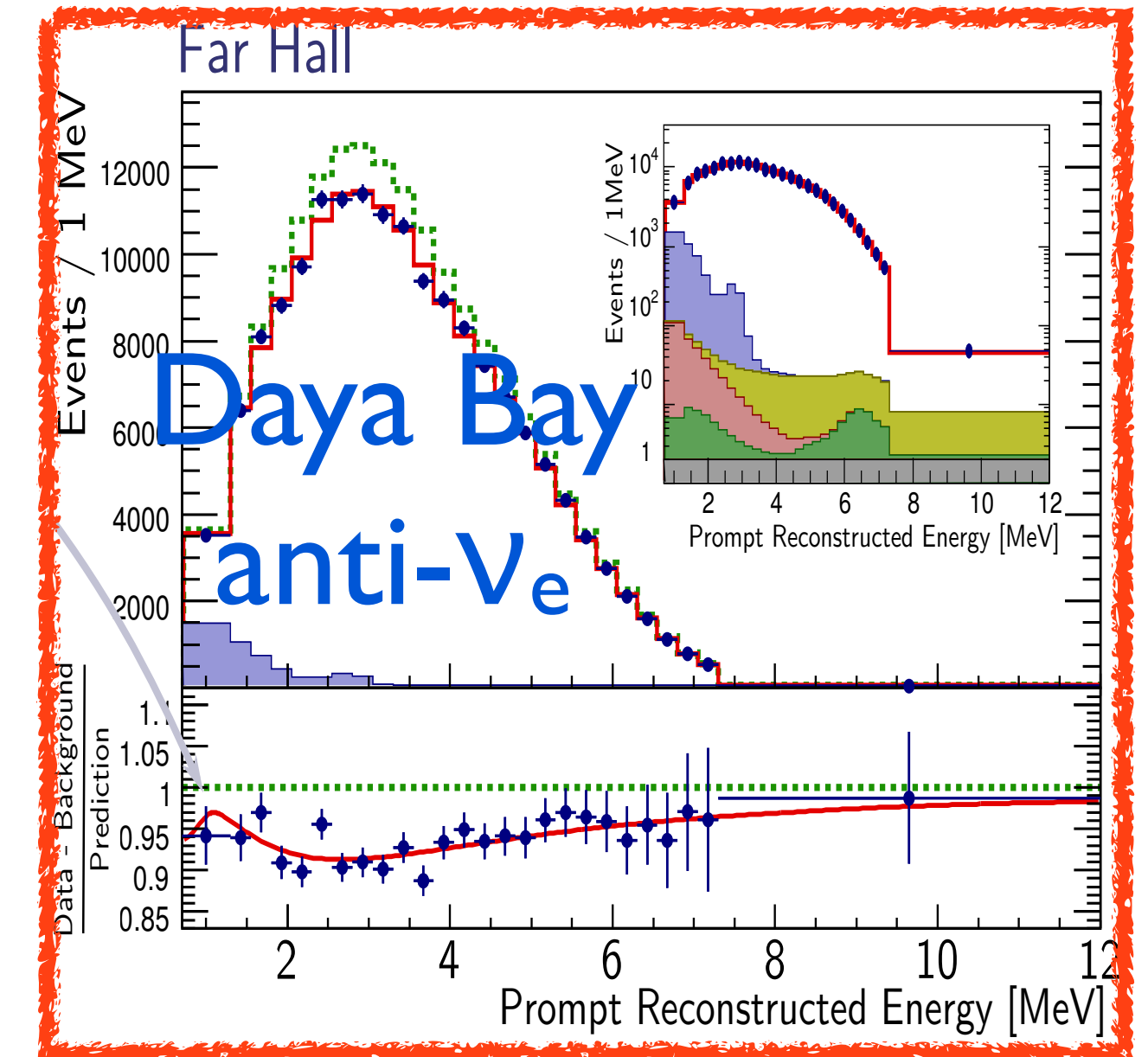
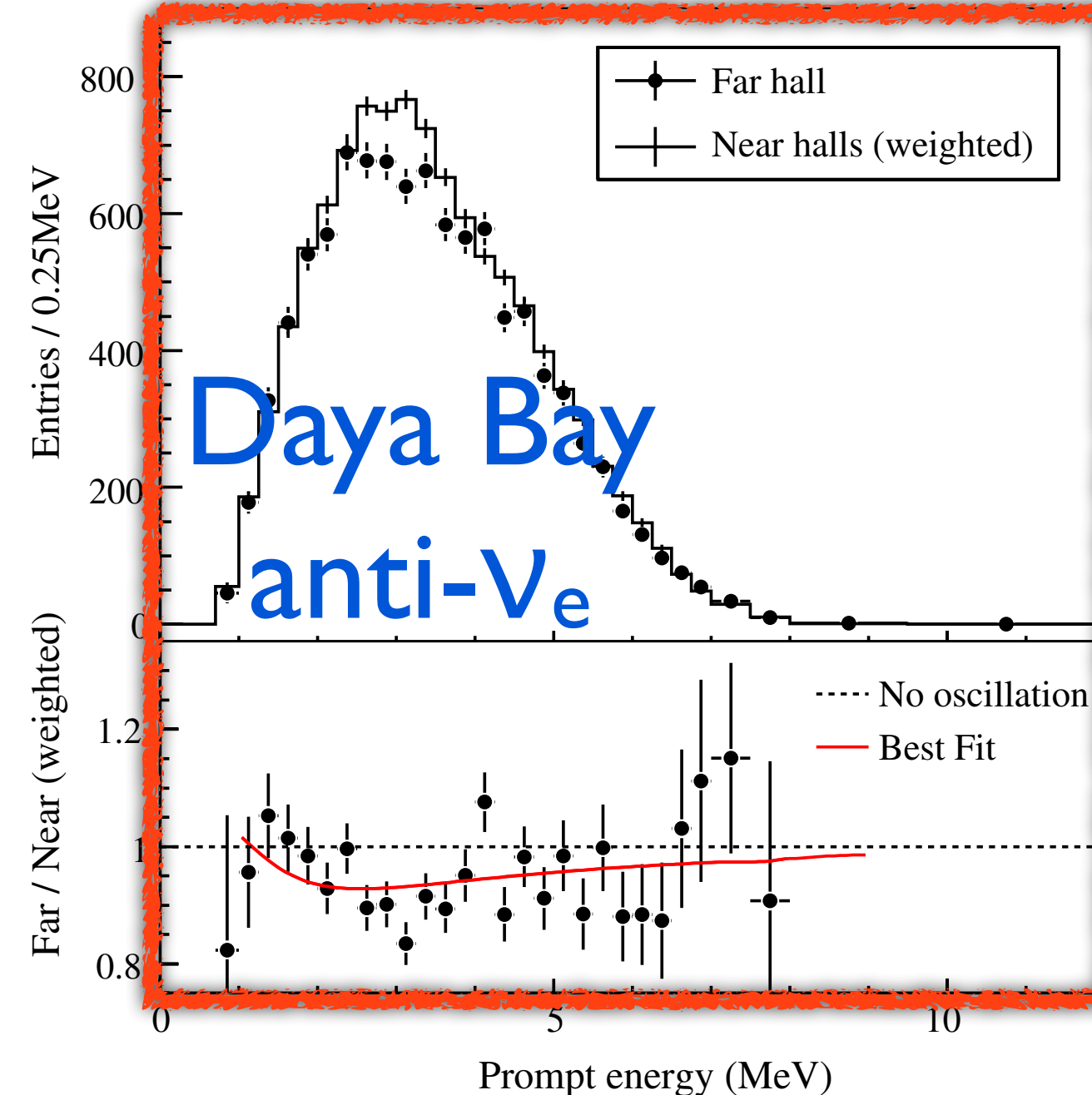
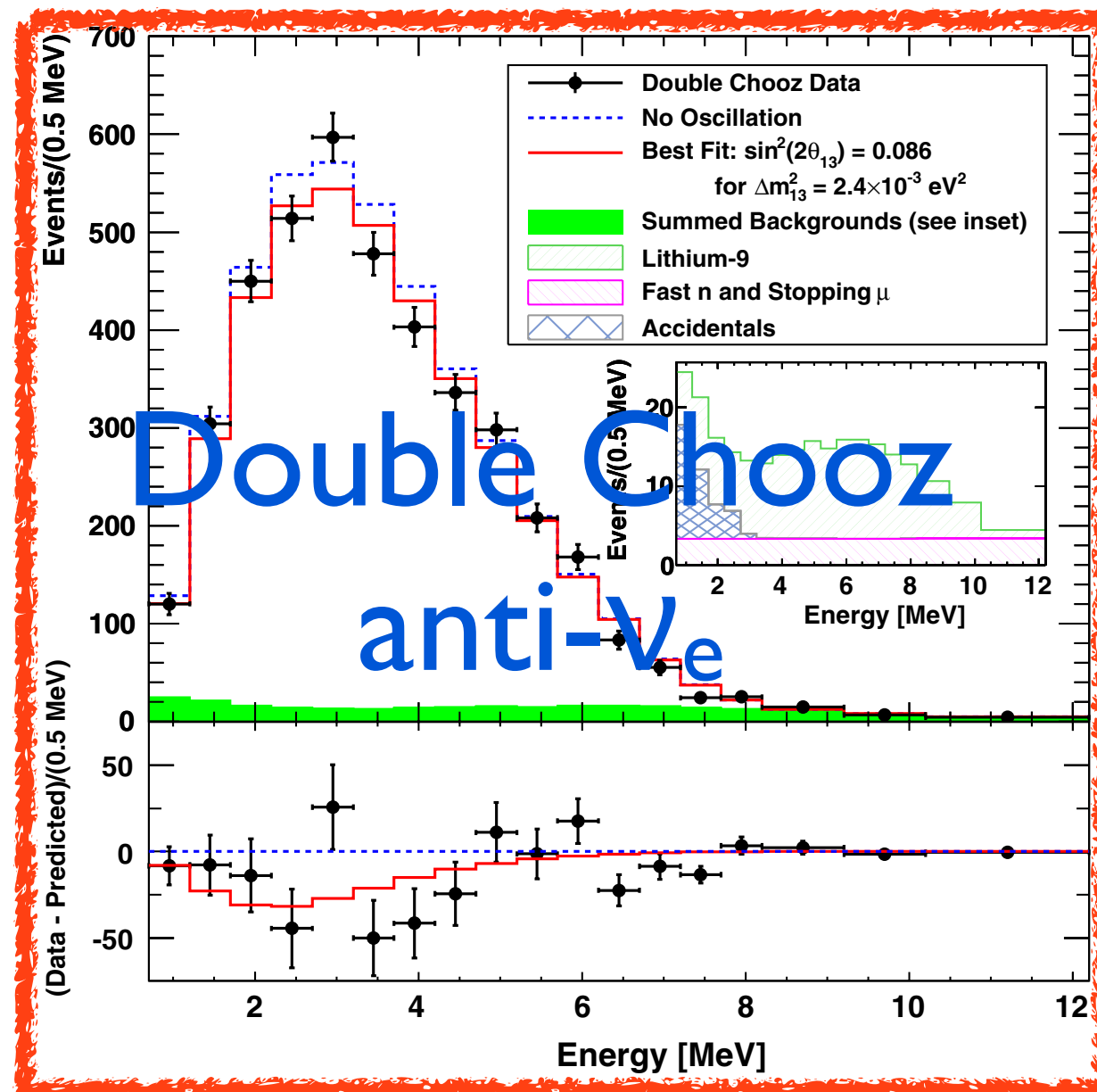
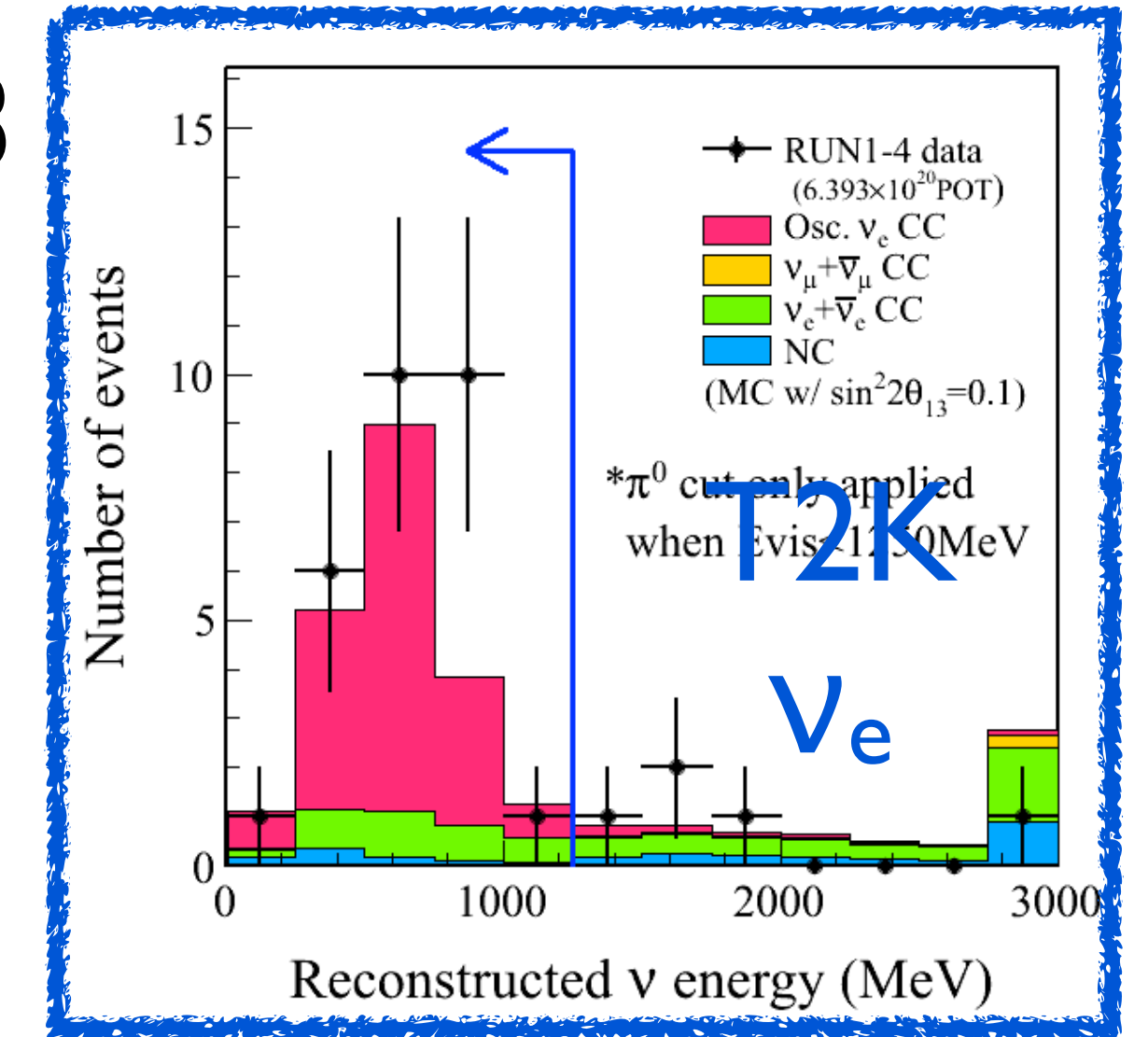
2011



2012



2013

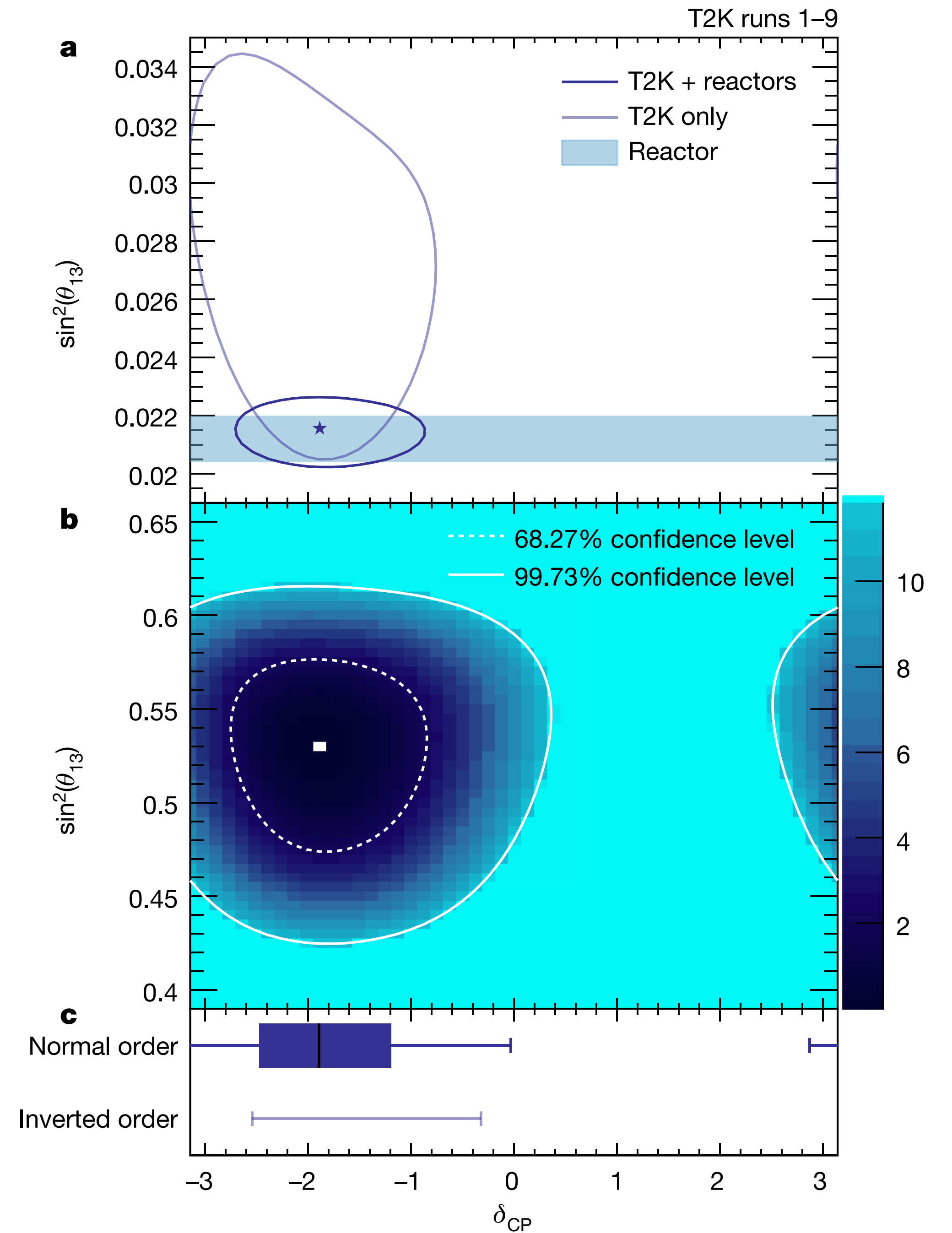


nature

THE MIRROR CRACK'D

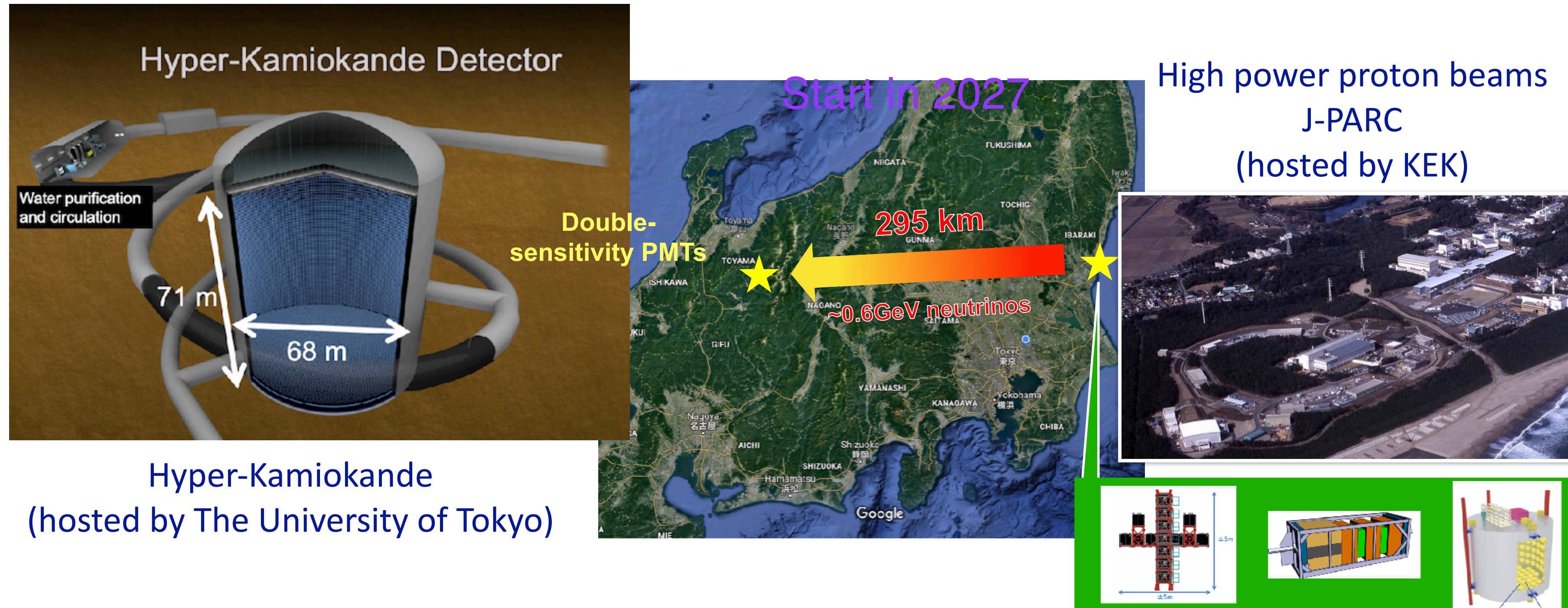
An indication of matter-antimatter
symmetry violation in neutrinos

Vol. 580, No. 7803
0160 2745/nature.com



Hyper-Kamiokande Project

- Long baseline experiment and non-accelerator physics in a single project
- Based on the highly successful tradition of SuperK and T2K (expertise and collaboration)



- **World-largest detector for Nucleon-decay and Neutrino experiment**
 - 8.4 times larger fiducial mass (190 kiloton) than Super-K instrumented with double-sensitivity PMTs
- **World most-intense neutrino beam**
 - J-PARC neutrino beam to be upgraded to 1.3 MW
- **New and upgraded near detectors to control systematic errors**

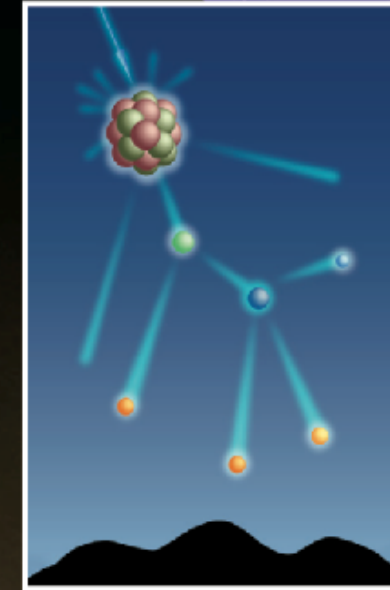
Physics Sensitivity

Hyper-Kamiokande program

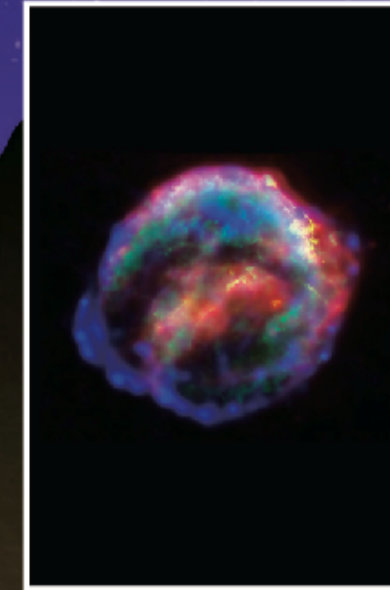
Accelerator Neutrino beam from J-PARC



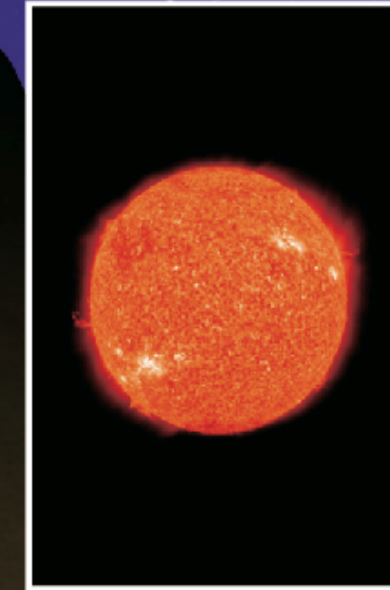
Atmosphere



Supernova



Sun

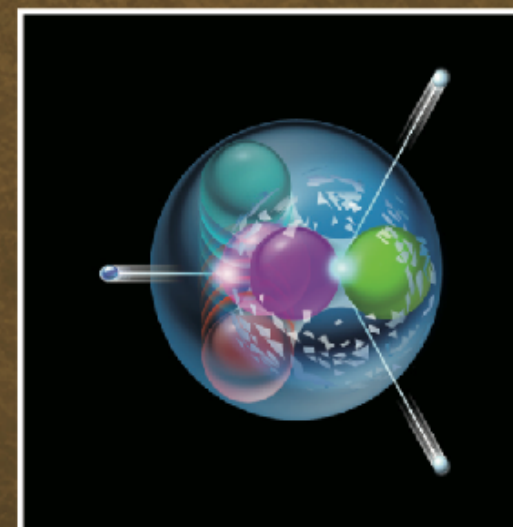


Neutrinos

Hyper-Kamiokande

Total mass 260 kton
Fiducial 190 kton

Proton Decay

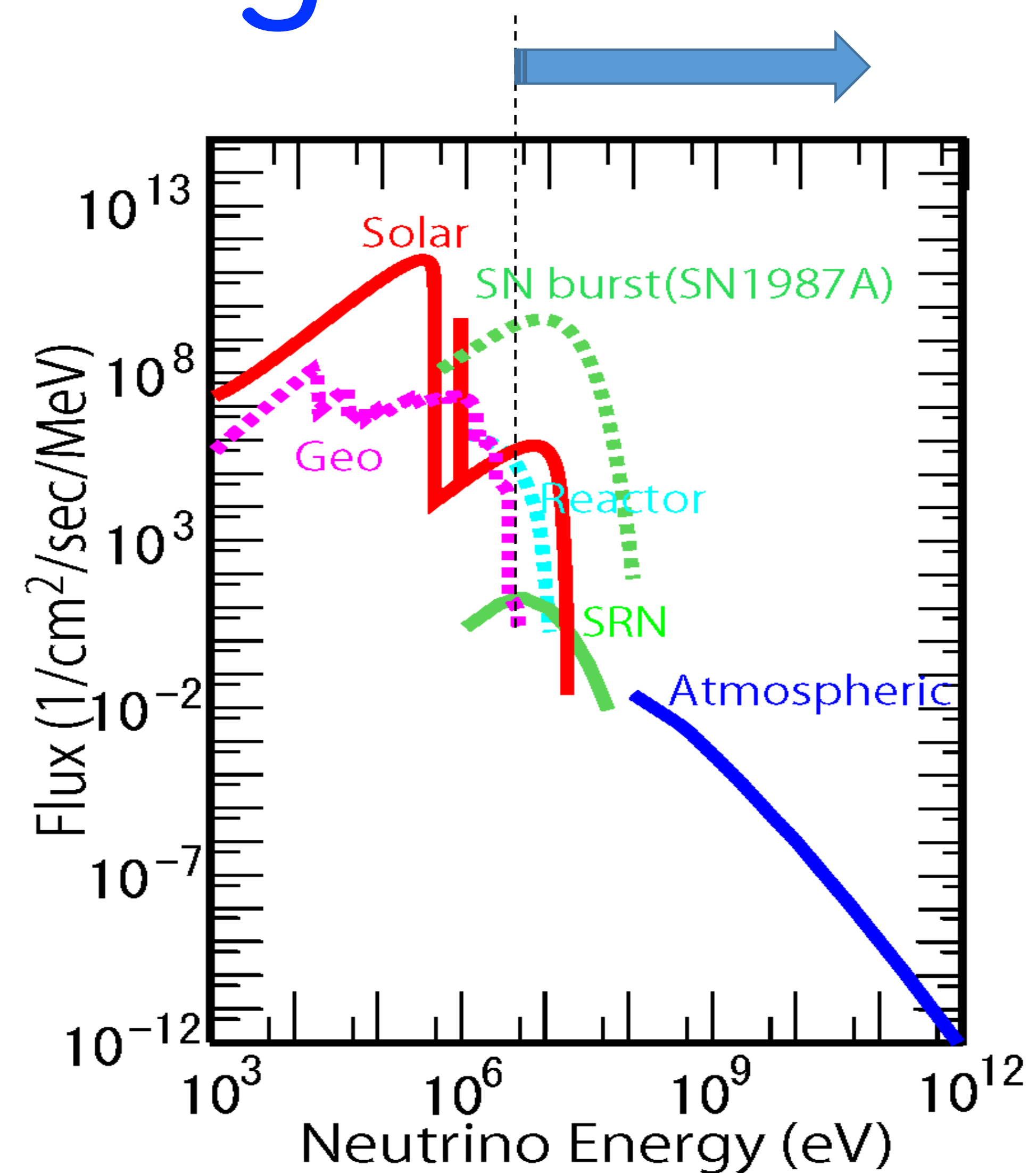
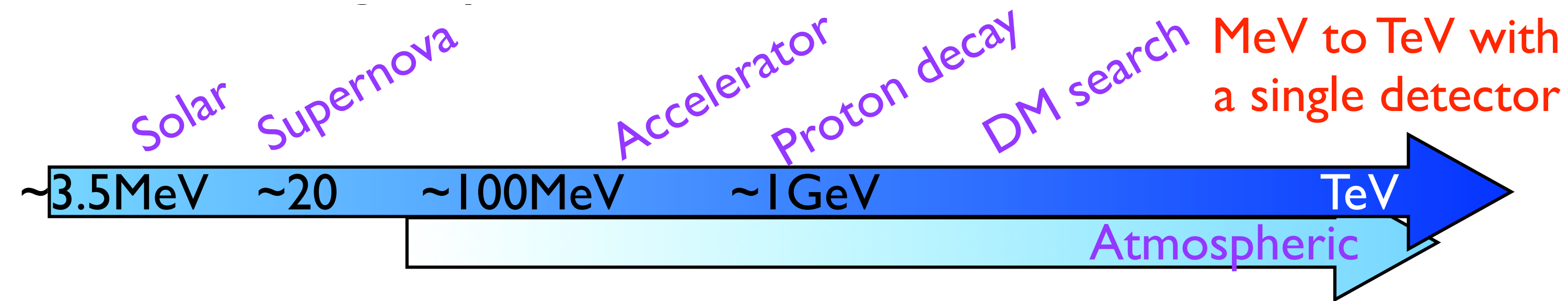


New photo-sensors



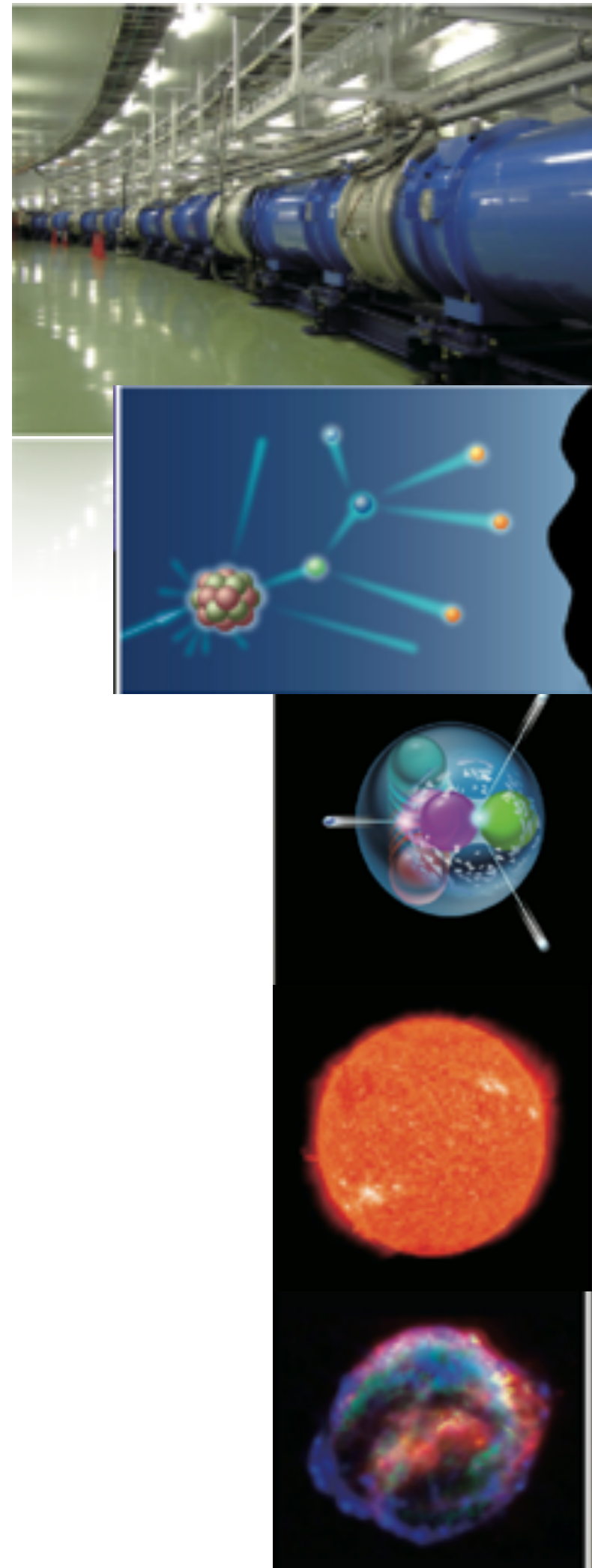
Tank filled with pure water 68m (D) x 71m (H)

Many Physics targets



- The **mass** of the detector with the **wide energy coverage** is the Key to probe new physics.
- It is an only unique choice to search for the proton decay up to 10^{35} years.

Target sensitivity

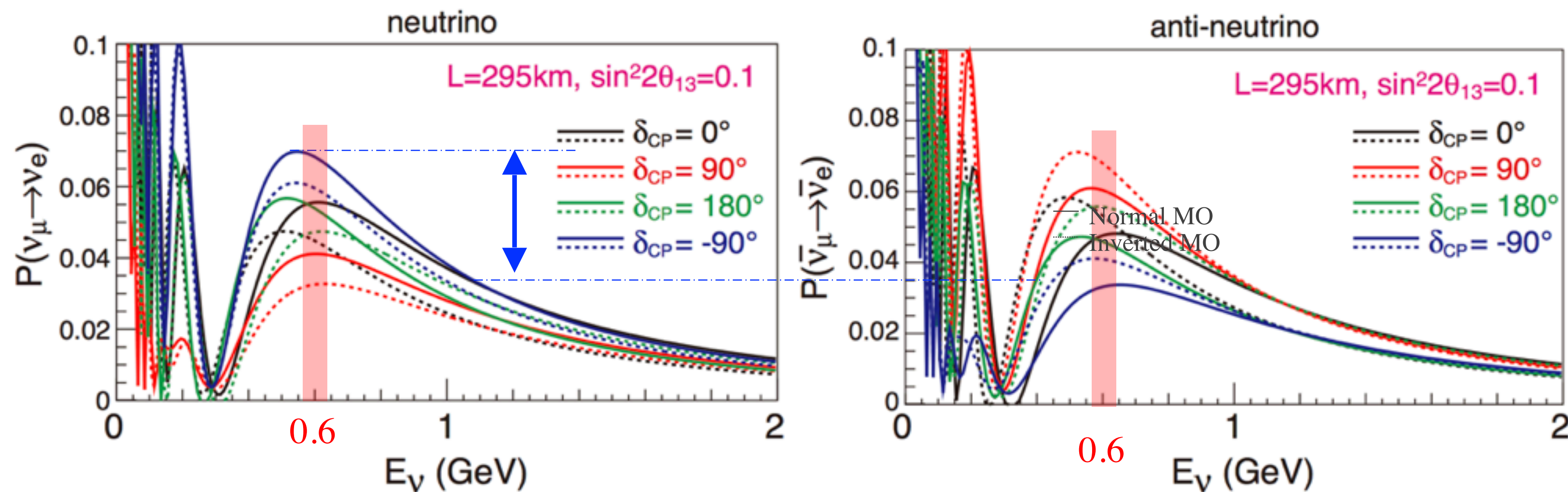


Physics category	Parameters	Sensitivity
LBL (1.3MW×10years)	δ precision	7°-20°
	CPV coverage (3/5 σ)	76%/58%
	$\sin^2\theta_{23}$ error (for 0.5)	± 0.017
ATM+LBL (10 years)	MO determination	$>3.8\sigma$
	Octant determination (3 σ)	$ \theta_{23}-45^\circ >2^\circ$
Proton Decay (20 years)	τ for $e^+\pi^0$ (3 σ)	1×10^{35}
	τ for $\bar{\nu}K$ (3 σ)	3×10^{34}
Solar (10 years)	Day/Night (from 0/from KL)	8 σ /4 σ
	Upturn	$>3\sigma$
Supernova	Burst (10kpc)	54k-90k
	Relic	70 ν 's / 10 years

Long-baseline program with the J-PARC neutrino beam

Experimental setup

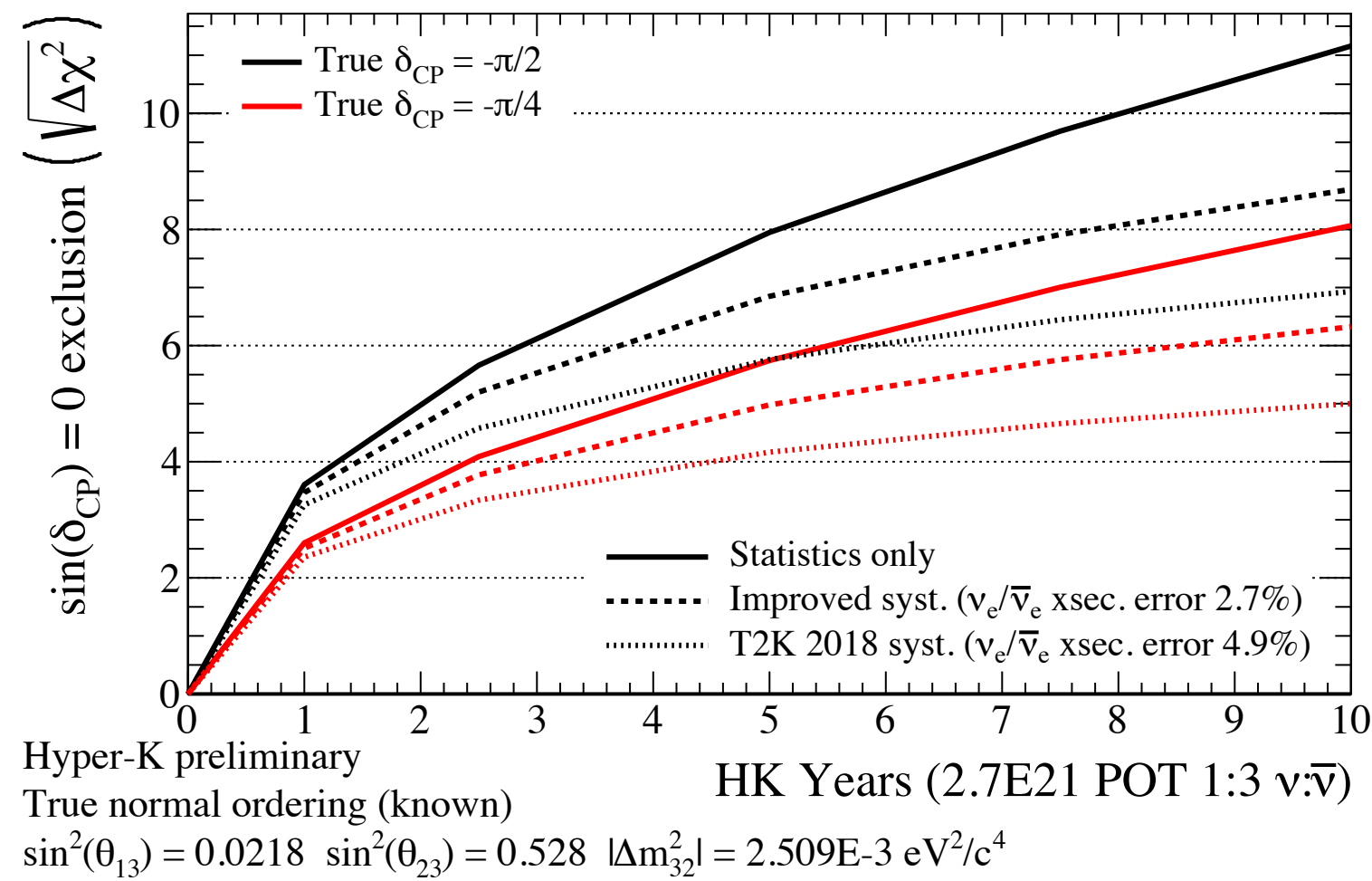
- 2.5° off-axis ν_μ and $\bar{\nu}_\mu$ beam peaked at 0.6 GeV (oscillation maximum at 295km)
 - Major component is QE: E_ν determined from (p, θ) of charged lepton
- Measures CP violation in neutrinos by comparing $P(\nu_\mu \rightarrow \nu_e)$ and $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$



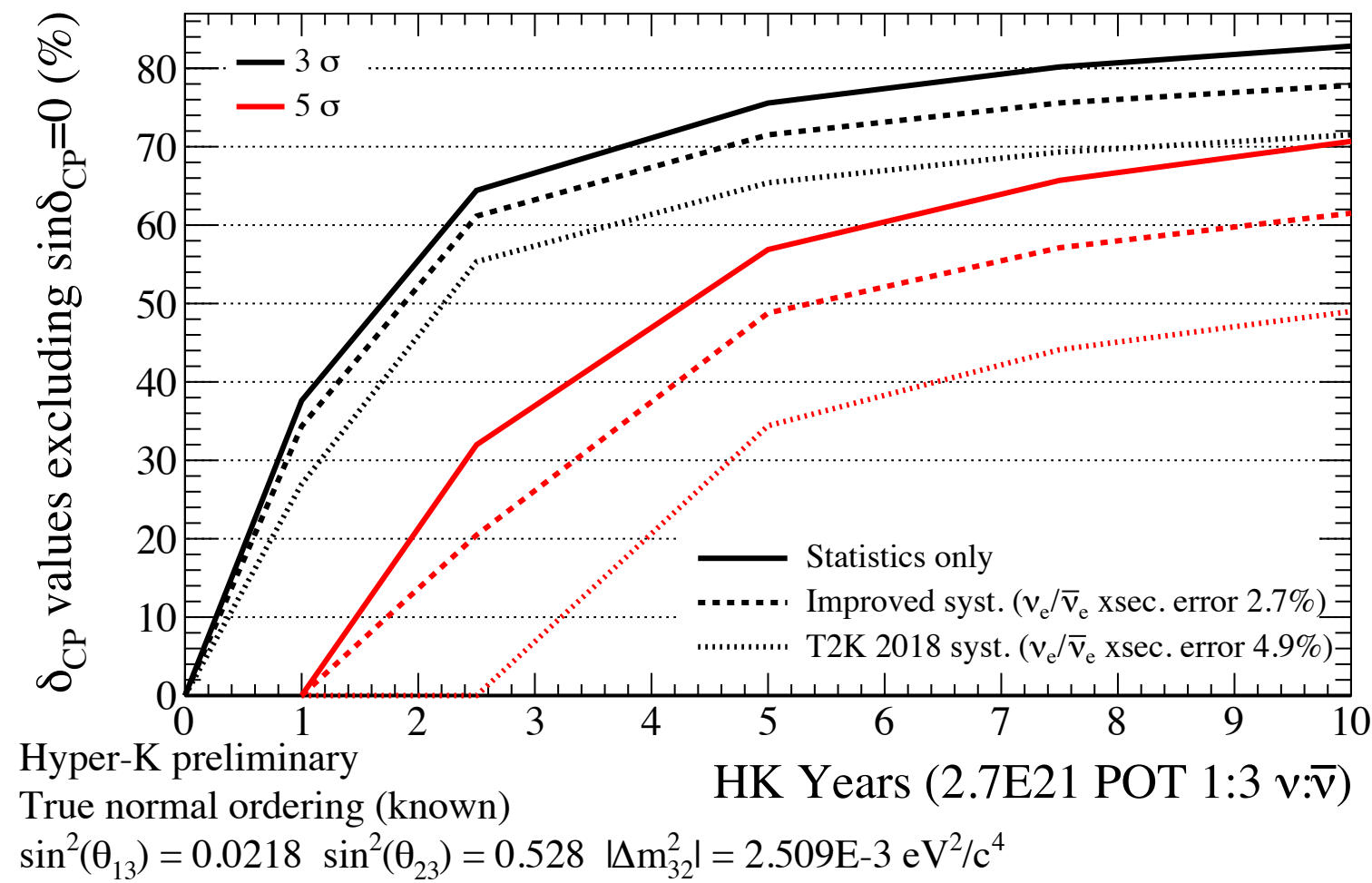
- A few % statistical uncertainties after 10 years operation with >1000 ν_e and $\bar{\nu}_e$ signals

Neutrino oscillation sensitivity

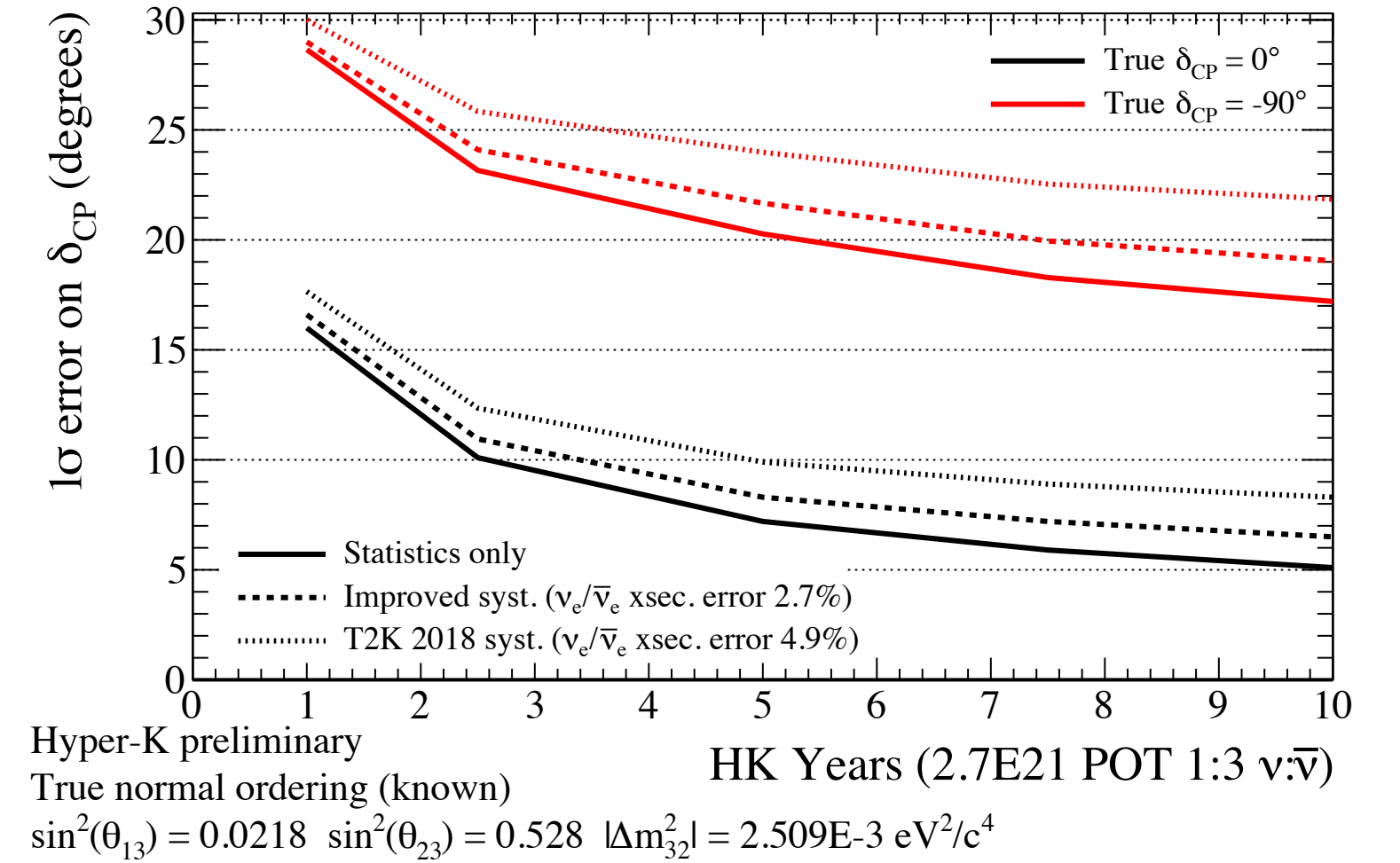
Sensitivity for $\delta_{CP} = -\pi/2$ or $-\pi/4$



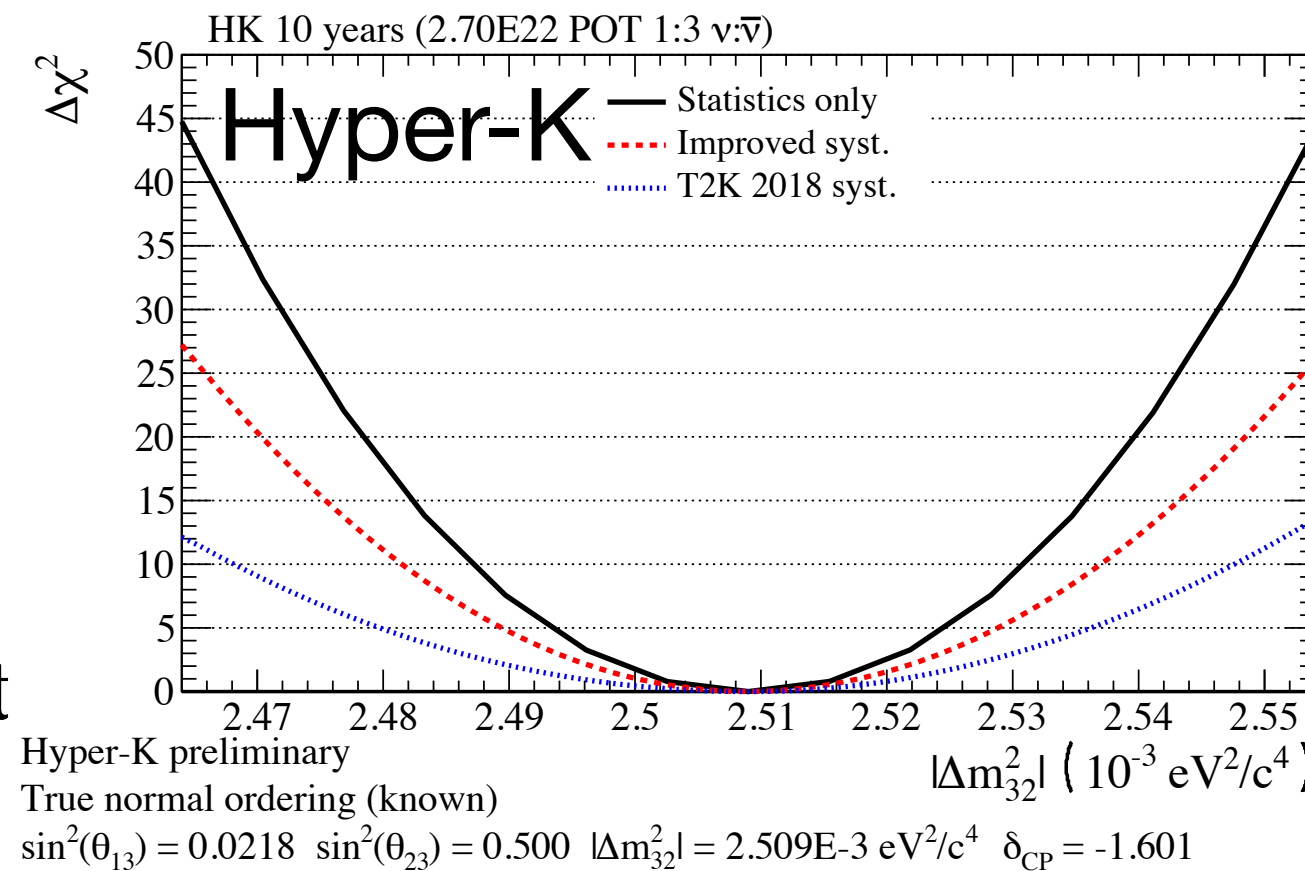
Fraction of δ_{CP} exclusion with 3σ or 5σ



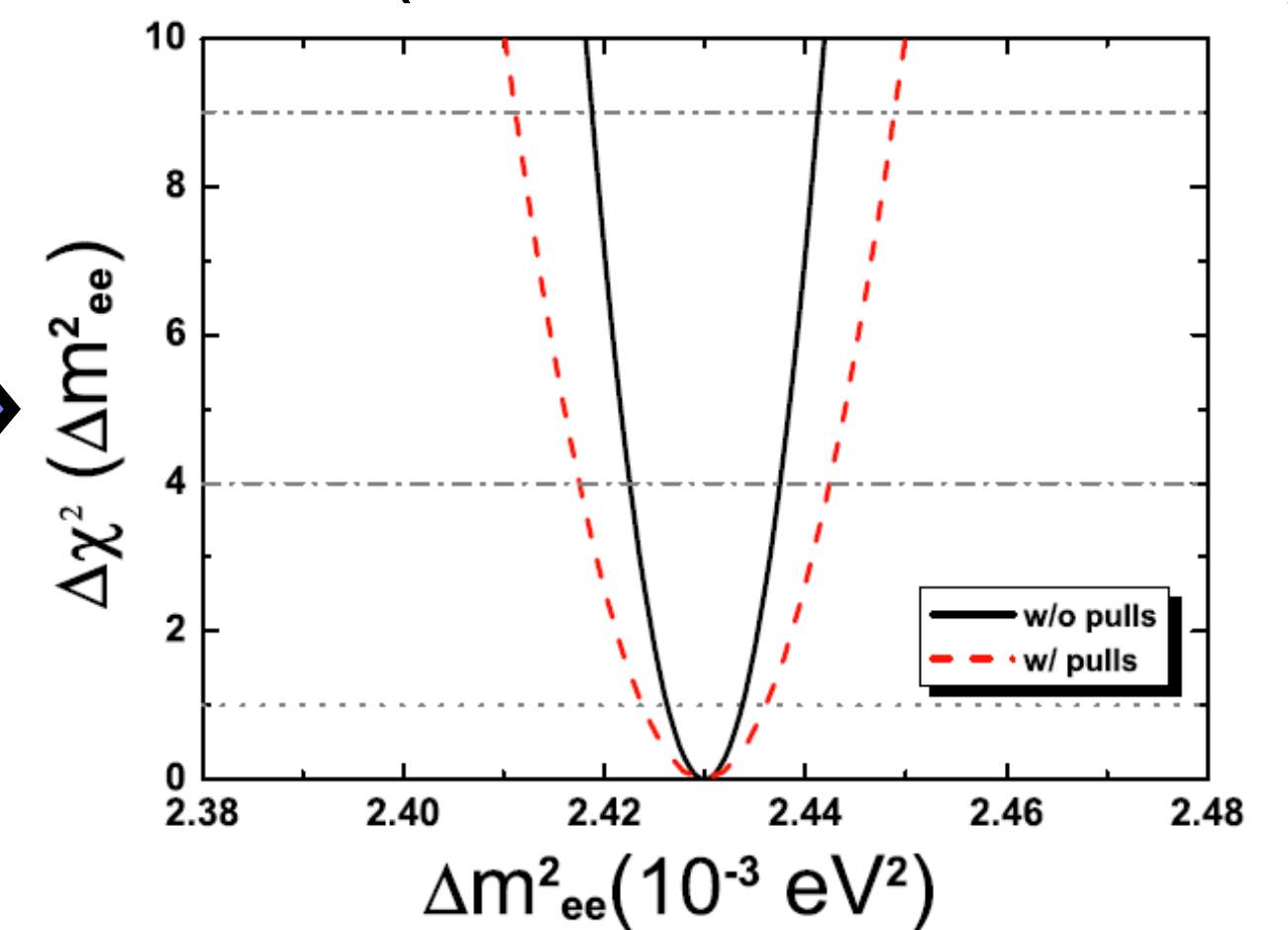
Precision for $\delta_{CP} = 0$ or $-\pi/2$



- Good chance for **discovery of CP violation** with $> 5\sigma$
 - Measurement of δ_{CP} with $\sim 20^\circ$ for $\delta_{cp} = -90^\circ$ / $\sim 7^\circ$ for $\delta_{cp} = 0^\circ$
- Combination of **beam and atmospheric neutrinos** enhances the sensitivity for mass ordering
- Reduction of systematic uncertainty has sizable impact
 - Upgrade of T2K ND280 + a new 1kton scale water Cherenkov (IWCD)



JUNO (reactor $\bar{\nu}_e$ at $\sim 50\text{km}$)

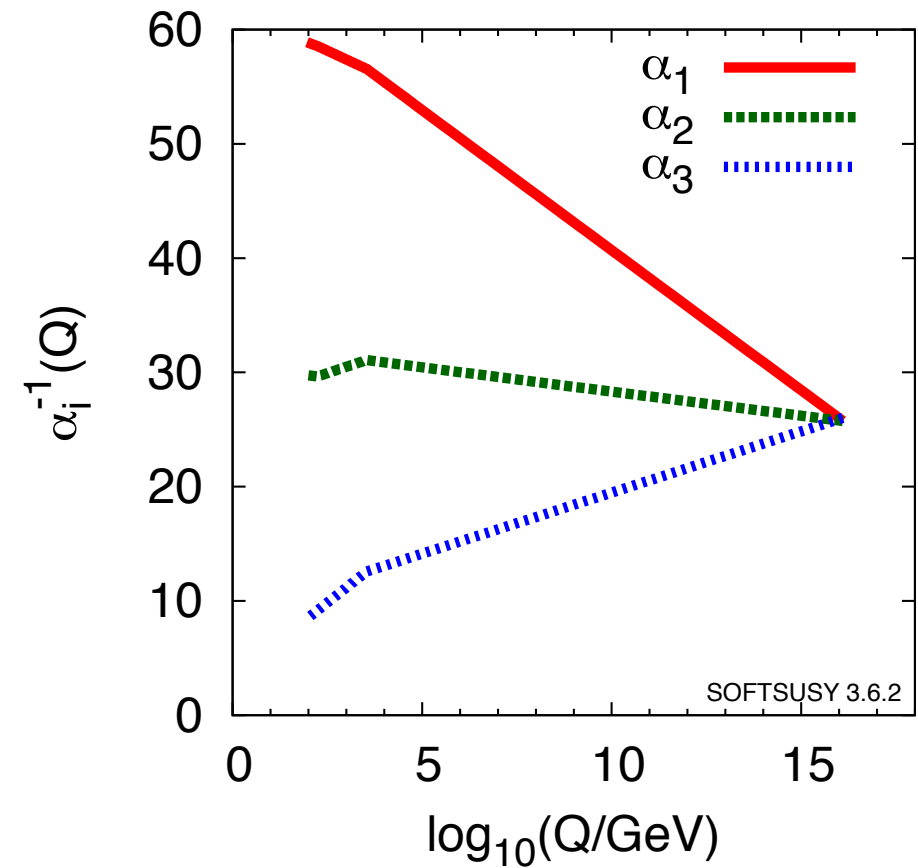


Proton decay Search

Nucleon decay:
clear evidence of GUT

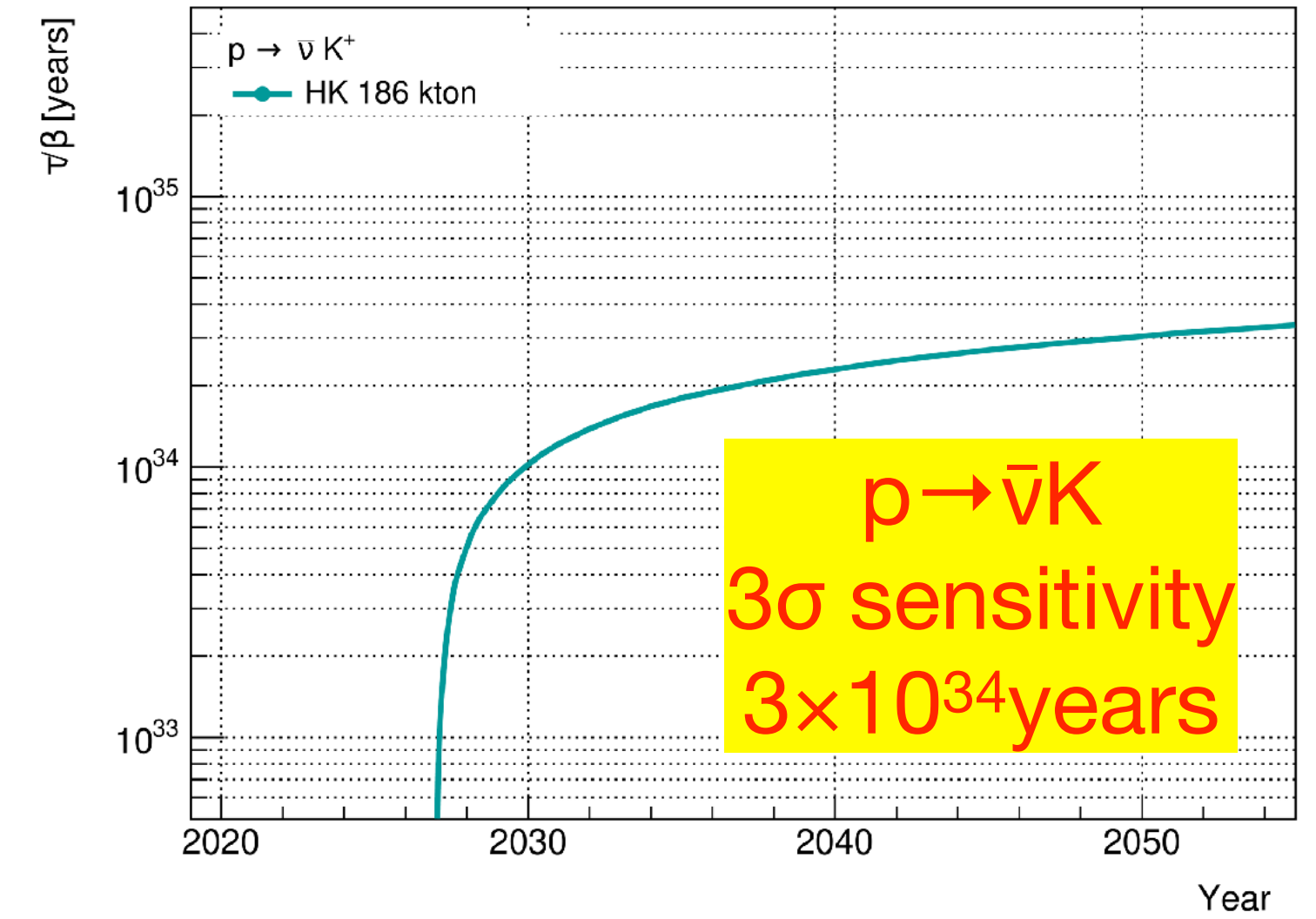
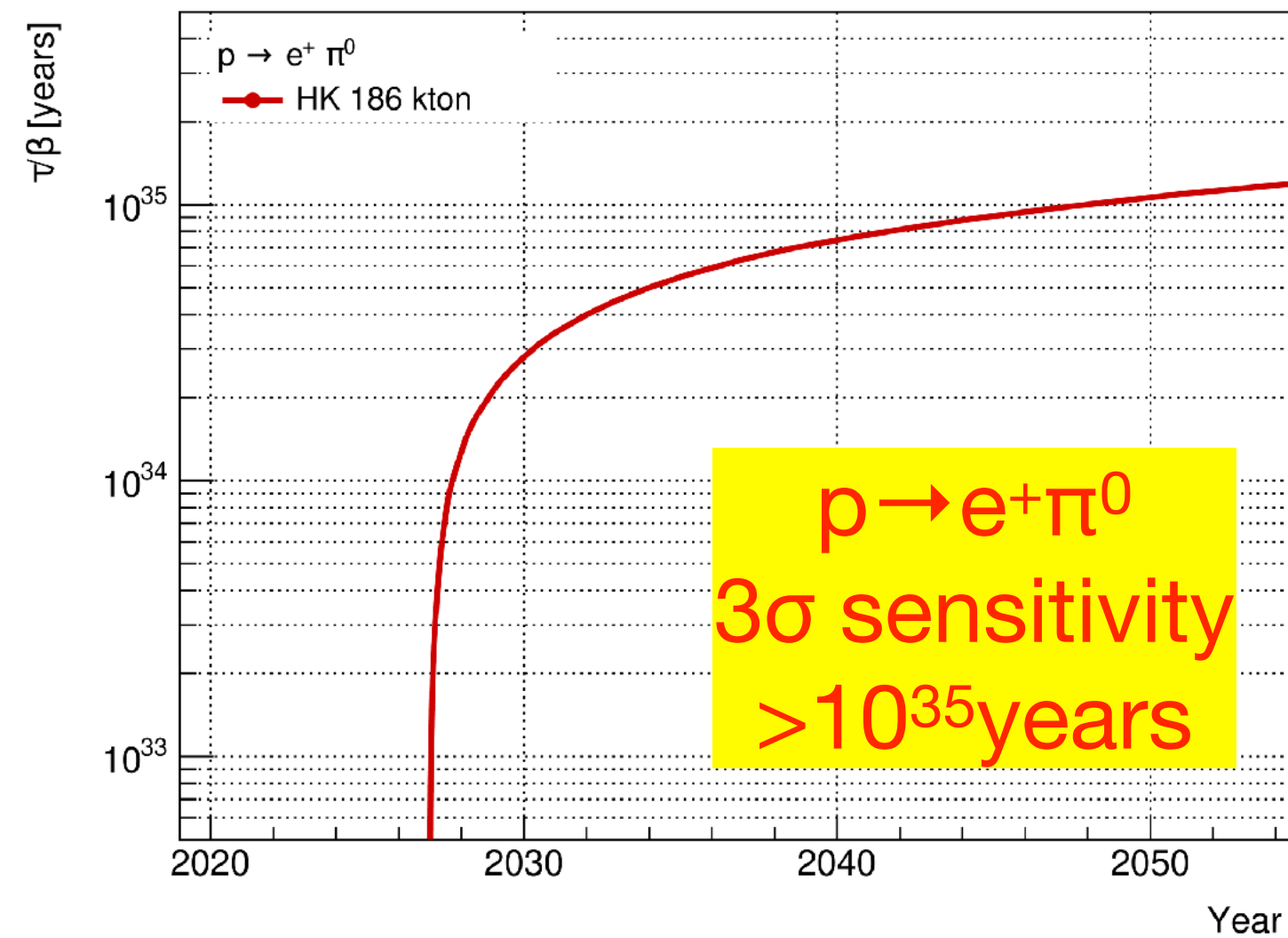
World's best sensitivity for most of possible modes, including $p \rightarrow \bar{\nu}K$, thanks to huge mass and free protons

MSSM: $m_0=M_{1/2}=2$ TeV, $A_0=0$, $\tan\beta=30$

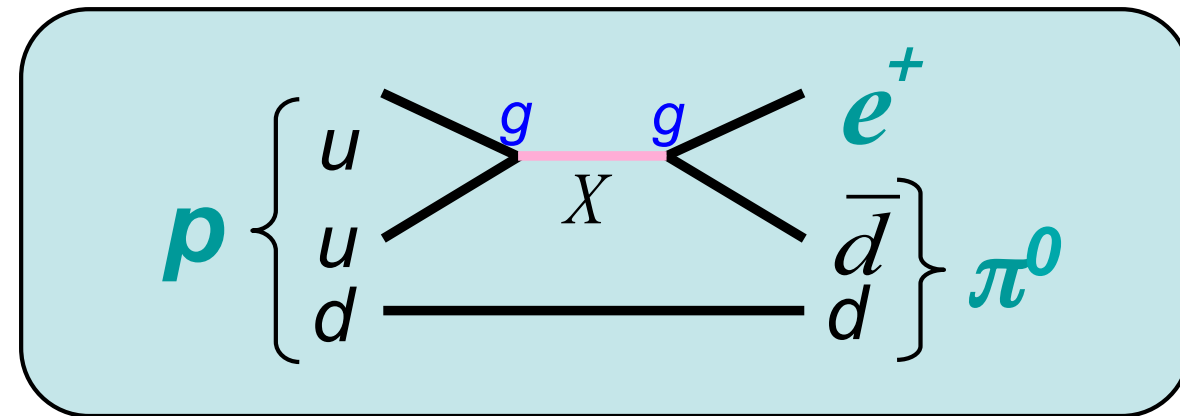


$\sim 10^{16}\text{GeV}$

Two modes as benchmark



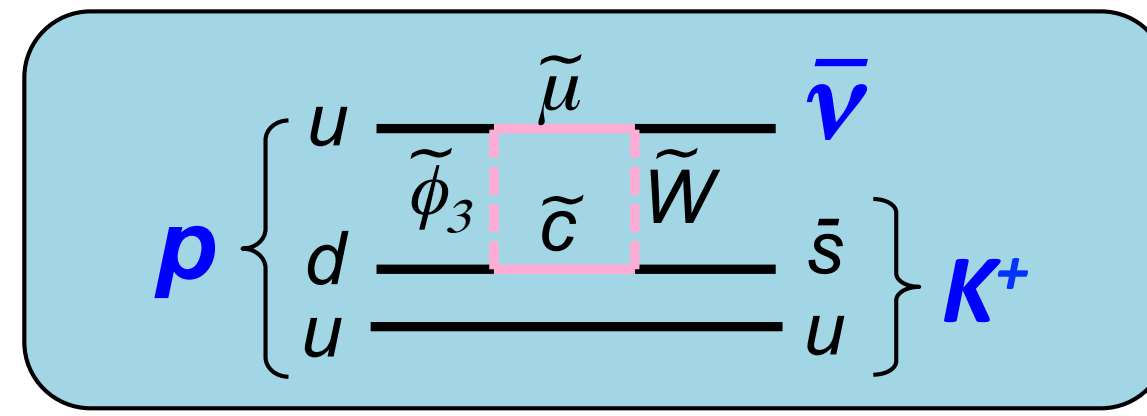
Mediated by gauge bosons



$p \rightarrow e^+ \pi^0$

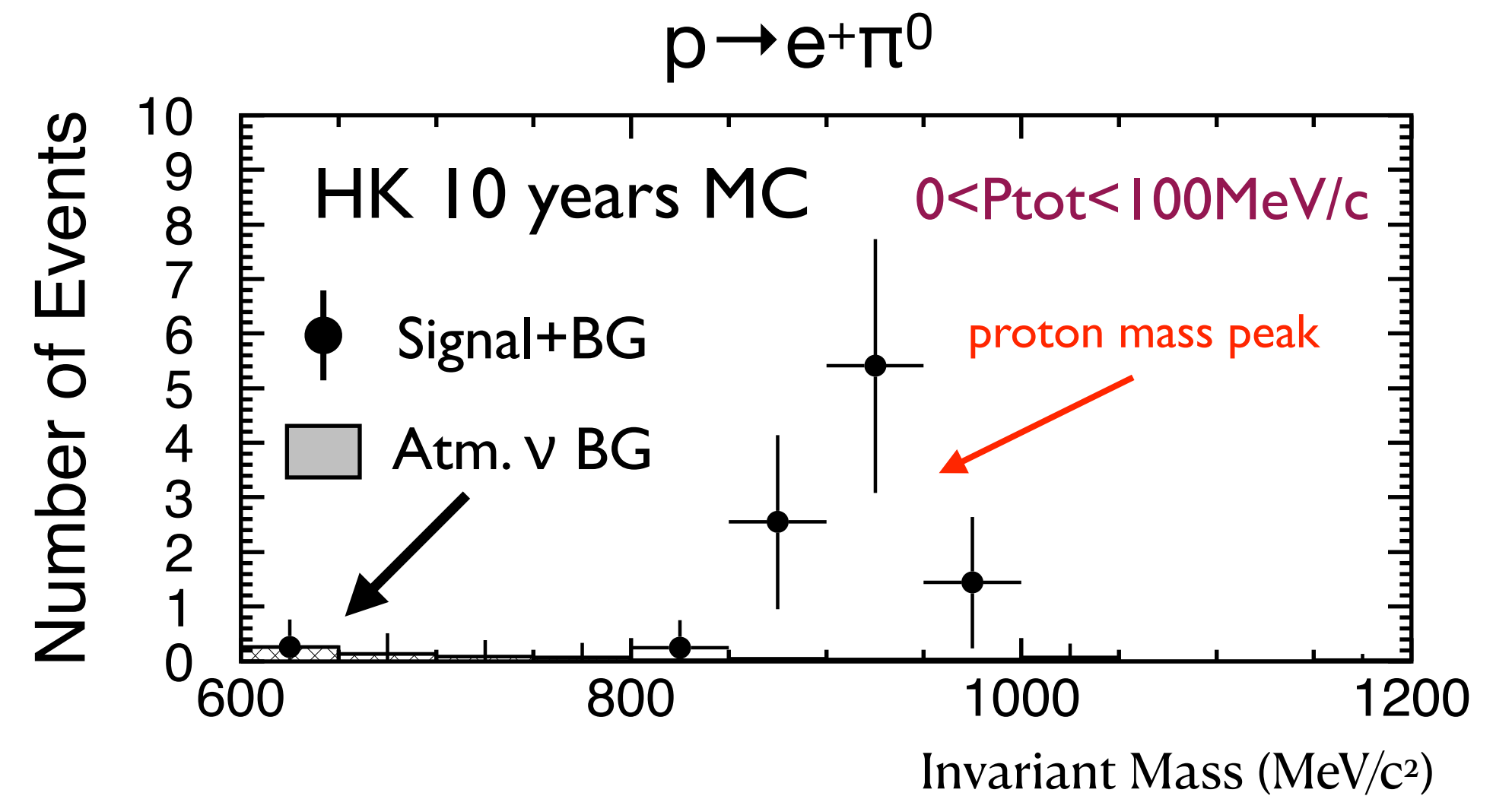
$$\Gamma(p \rightarrow e^+ \pi^0) \sim \frac{g^4 m_p^5}{M_X^4}$$

SUSY mediated



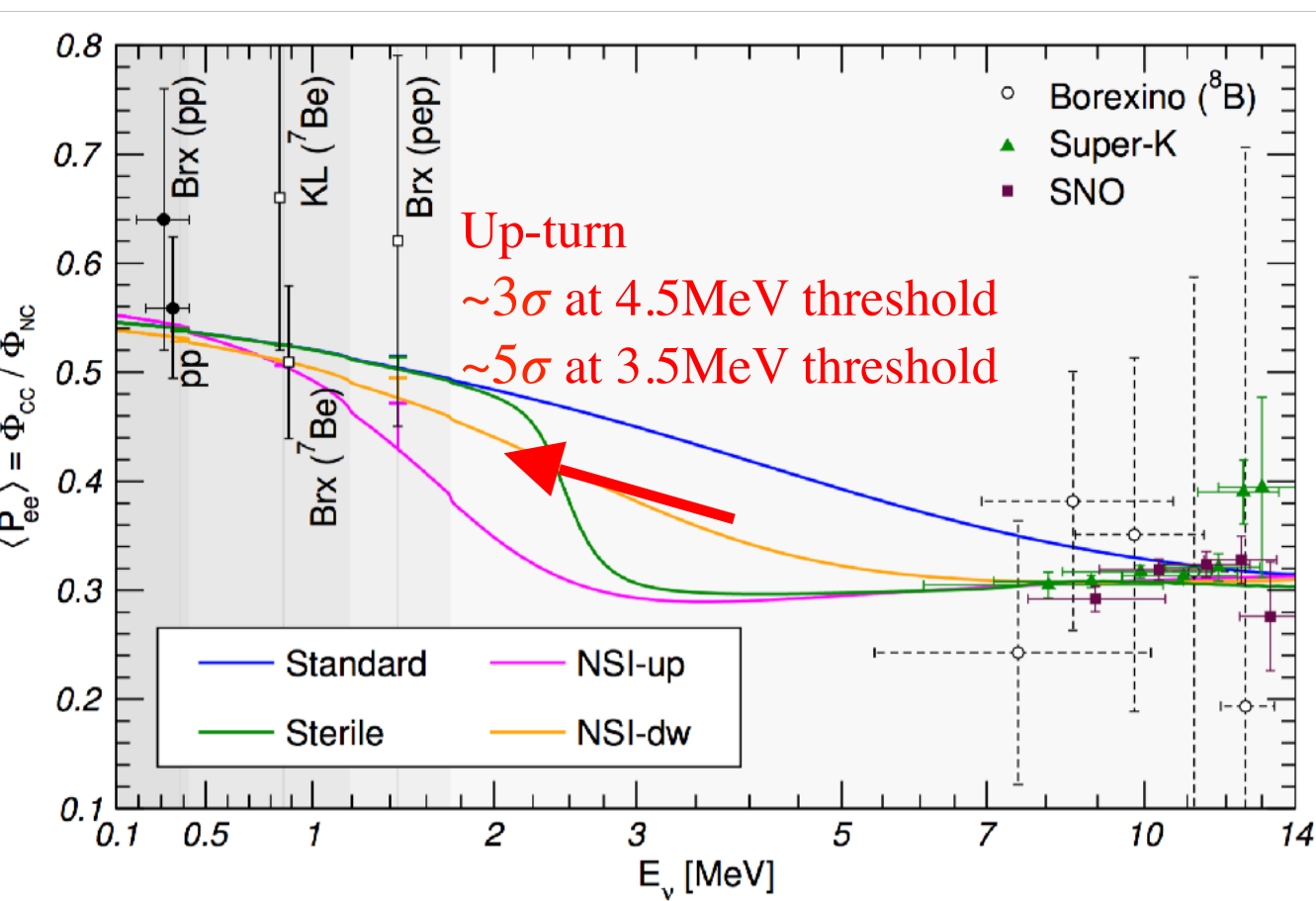
$p \rightarrow \bar{\nu} K^+$

$$\Gamma(p \rightarrow \bar{\nu} K^+) \sim \frac{\tan^2 \beta \times m_p^5}{M_{\tilde{q}}^2 \times M_3^2}$$

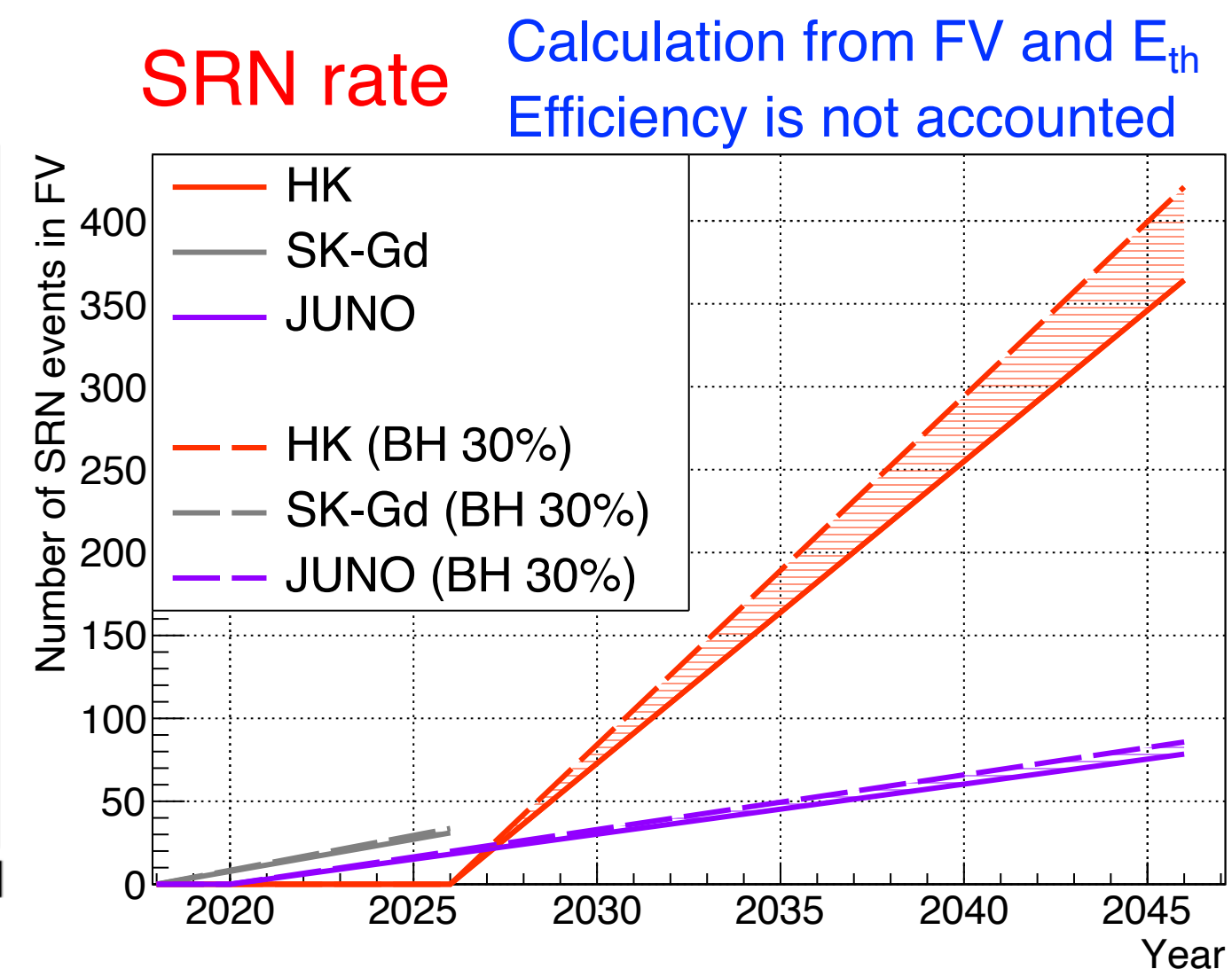
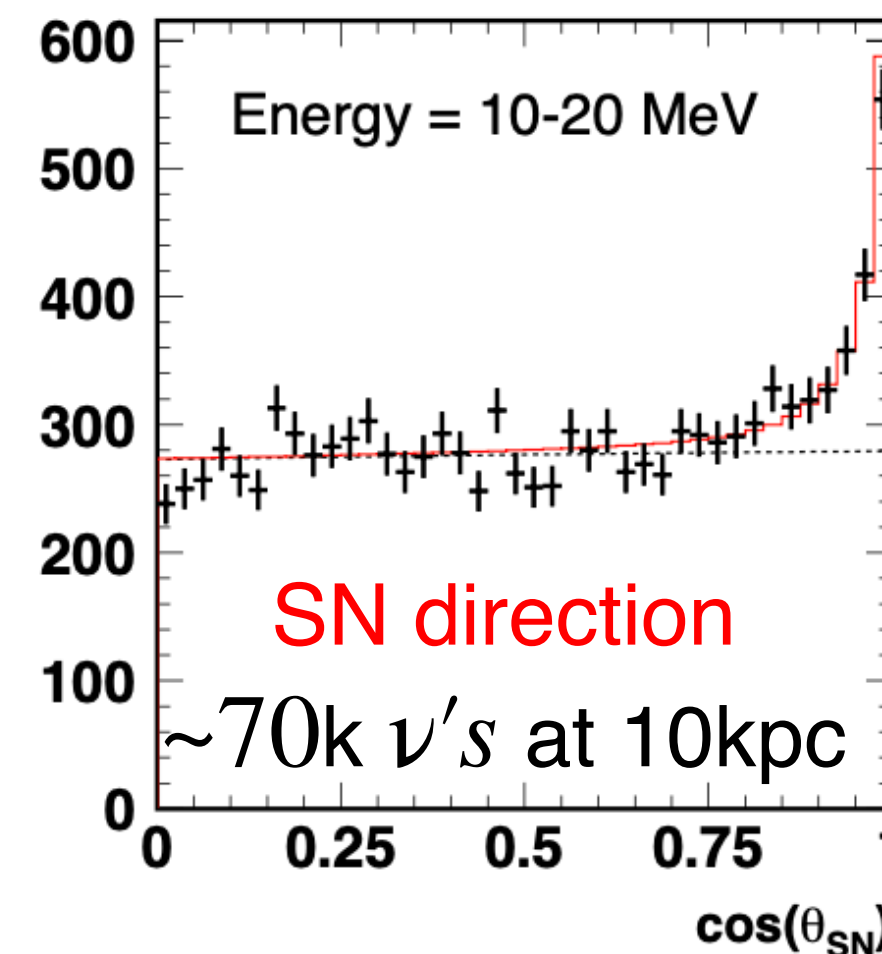
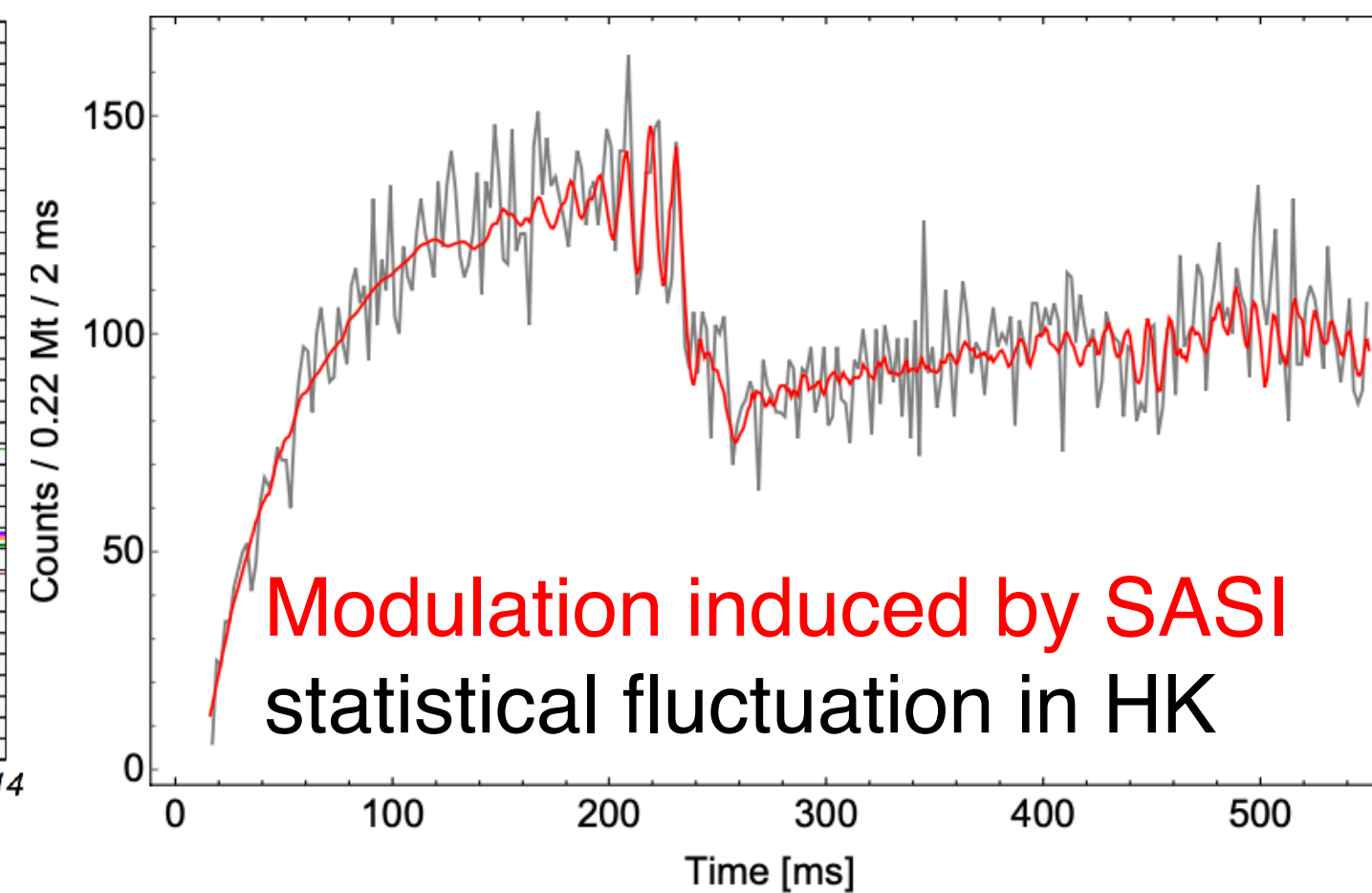


Neutrino Astrophysics

- Observation of a few ~ 10 MeV neutrinos with time, energy and direction information
 - Unique role in multi-messenger observation
- **Solar neutrinos**: up-turn at vacuum-MSW transition, Day/Night asymmetry, hep neutrino observation
- **Supernova burst neutrinos**: explosion mechanism, BH/NS formation, alert with $\sim 1^\circ$ pointing
- **Supernova Relic Neutrinos (SRN)**: stellar collapse, nucleosynthesis and history of the universe



M. Maltoni et al., Phys. Eur. Phys. J. A52, 87 (2016)

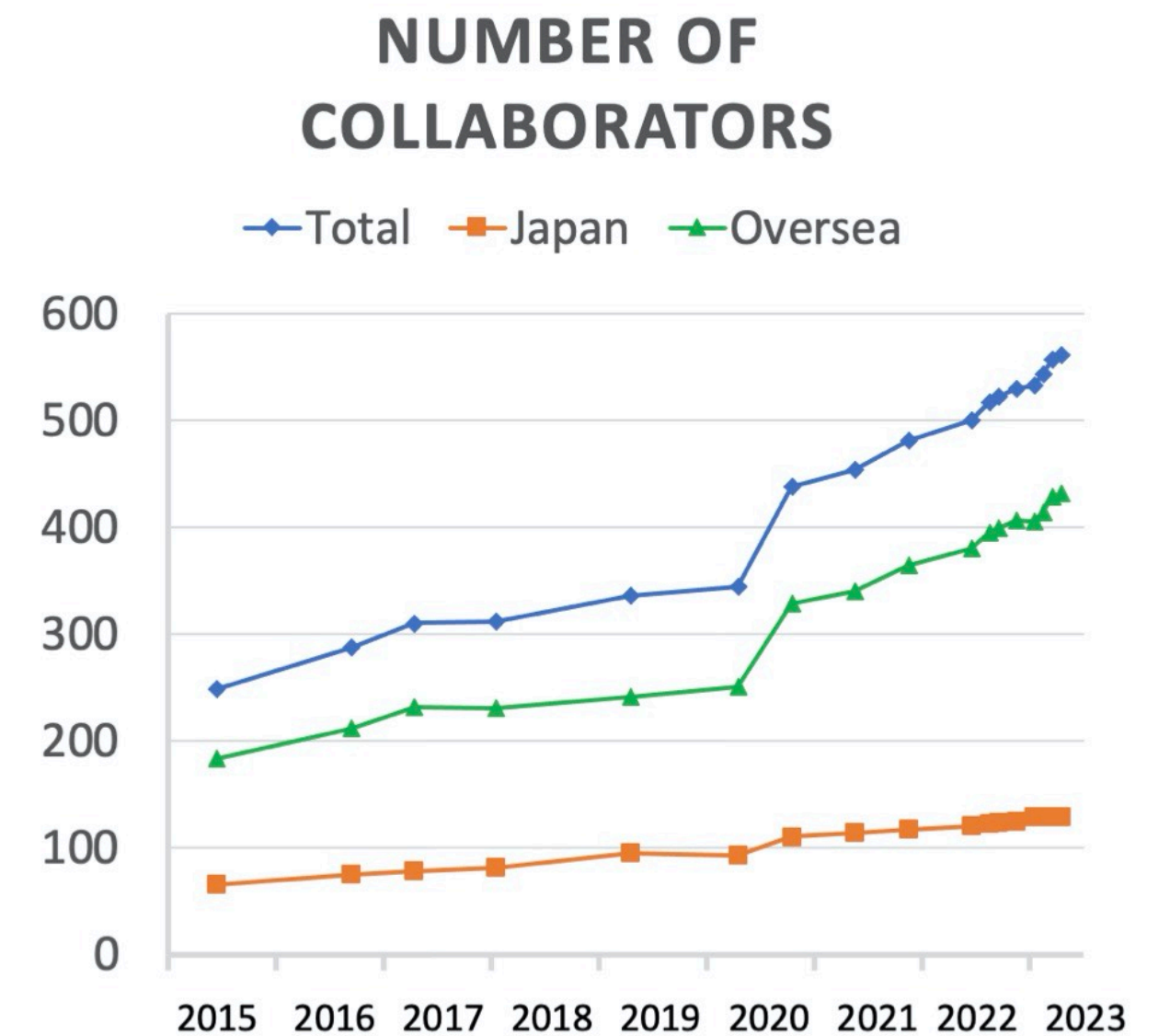


Project Status

Hyper-Kamiokande collaboration

Univ. of Tokyo and KEK host the project

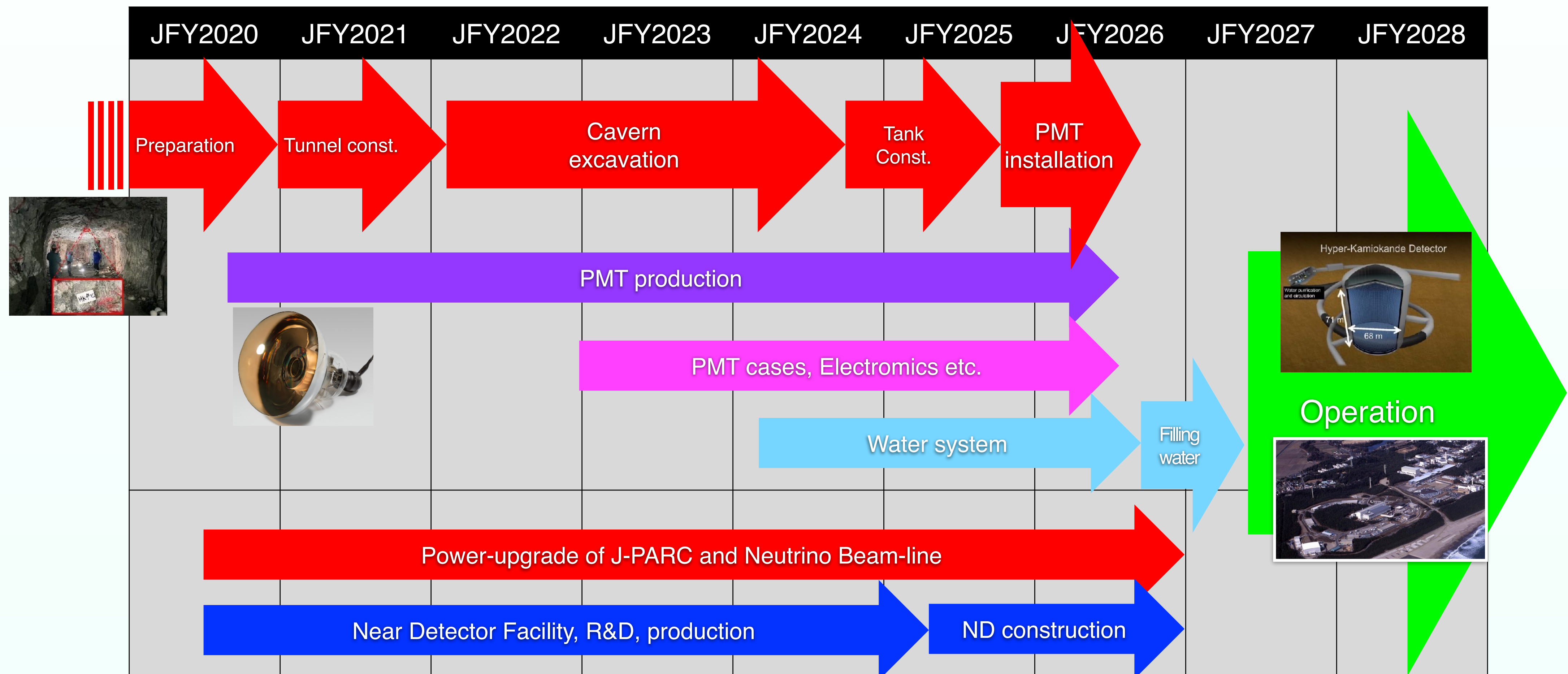
- ~560 people from ~20 countries, ~100 institutions
 - 25% Japanese / 75% non-Japanese
- Recently approved as a recognized experiment (RE45) at CERN



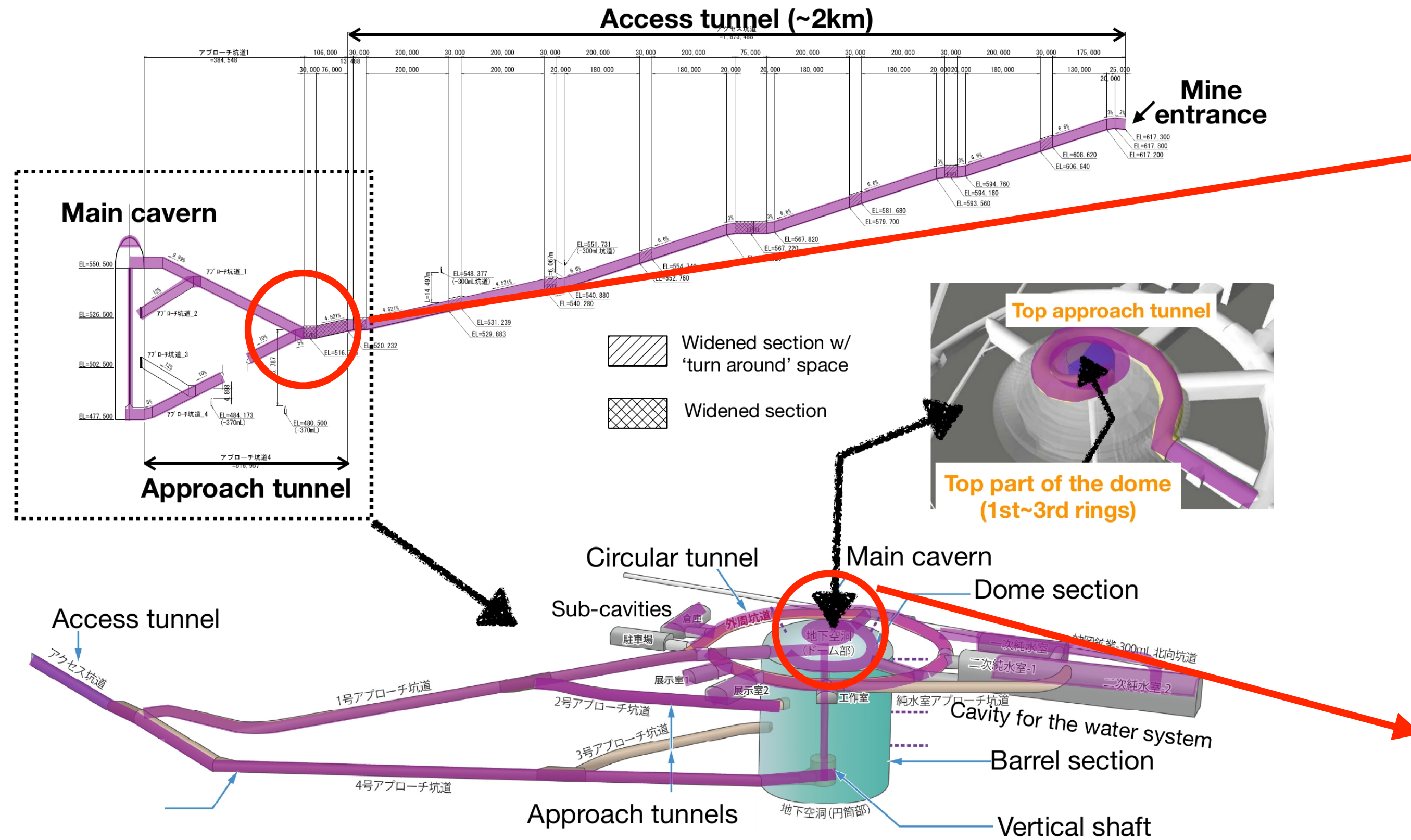
Collaboration Meeting, March 2023 @ Toyama

Timeline

- 2022-2027: Construction, 2027- : Operation
- No change of schedule since the approval of project in 2020



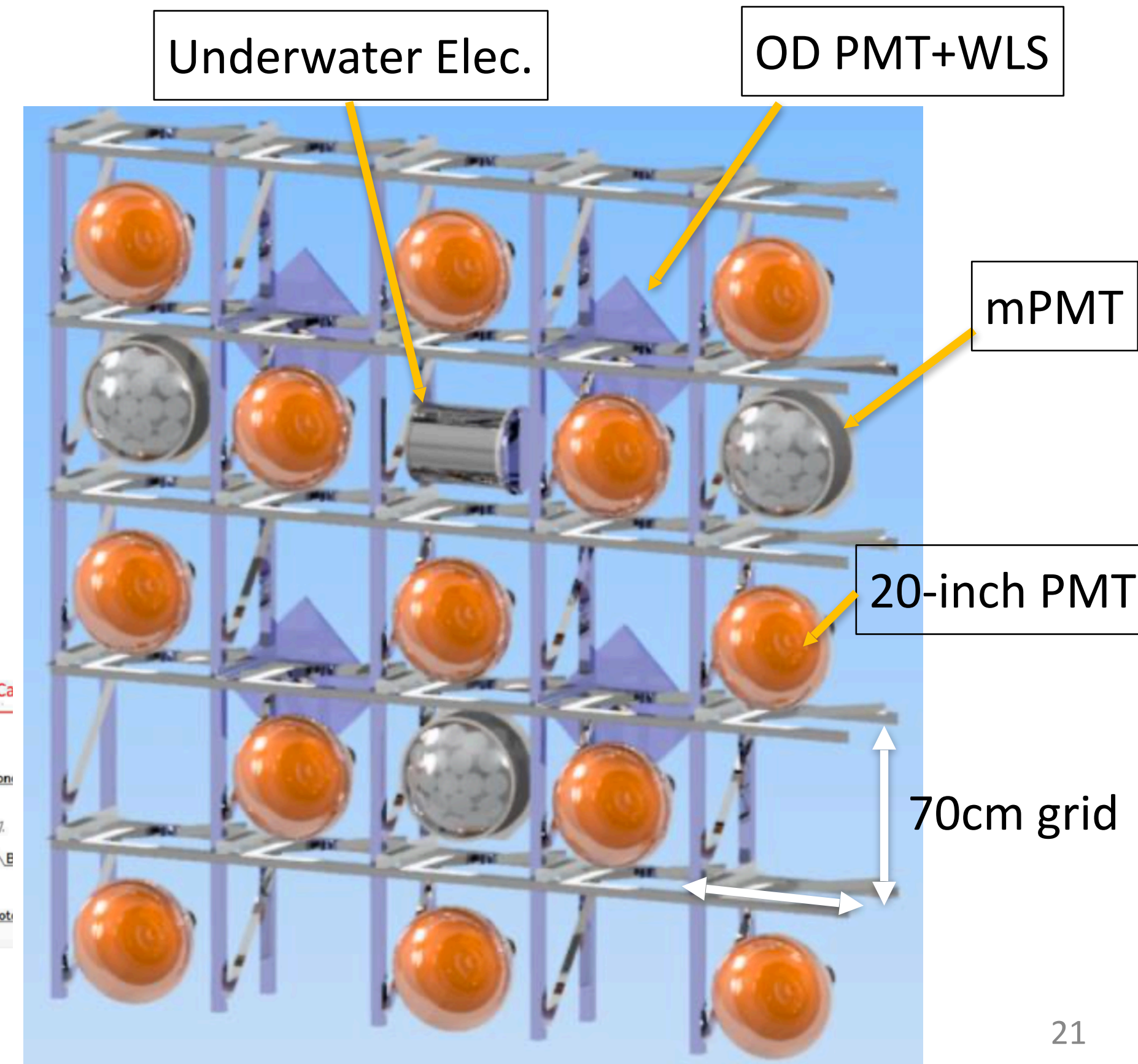
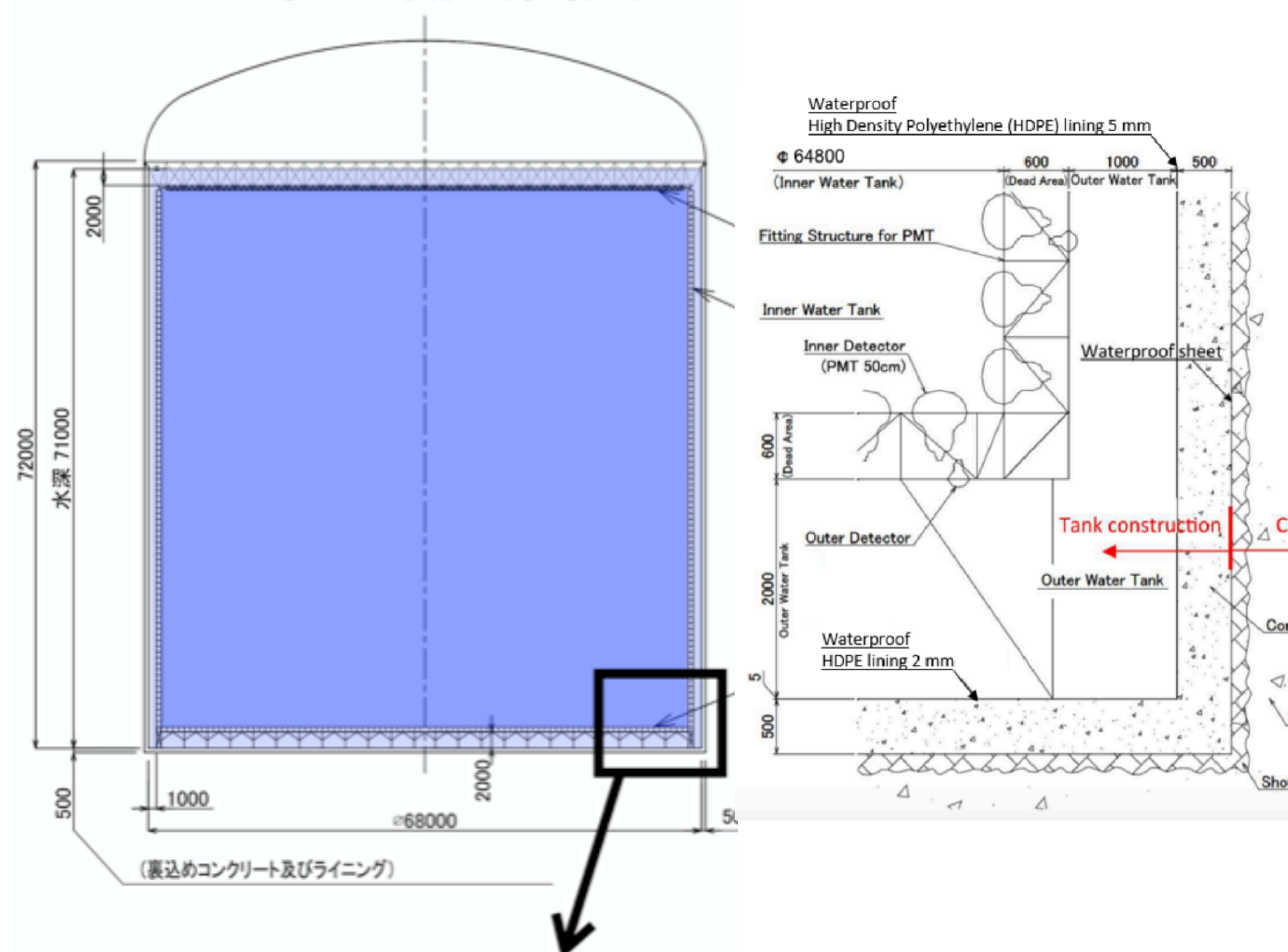
Cavern excavation status



- Access tunnel (~2 km) completed in Feb. 2022
- Excavation of the main cavern started in fall 2022 and is proceeding as scheduled

Detector configuration

- 67 m Φ x 66 m Inner Detector (fiducial 190 kt)
 - 20,000 HPK HiQE (x2 SK) 20-inch PMTs will be installed
 - mPMT modules will be integrated as hybrid configuration.
- 1m(wall) or 2m(top/bottom) thick Outer Detector
 - 3" PMTs + WLS boards
- Under-water electronics module
 - Mitigate disadvantage of long cables



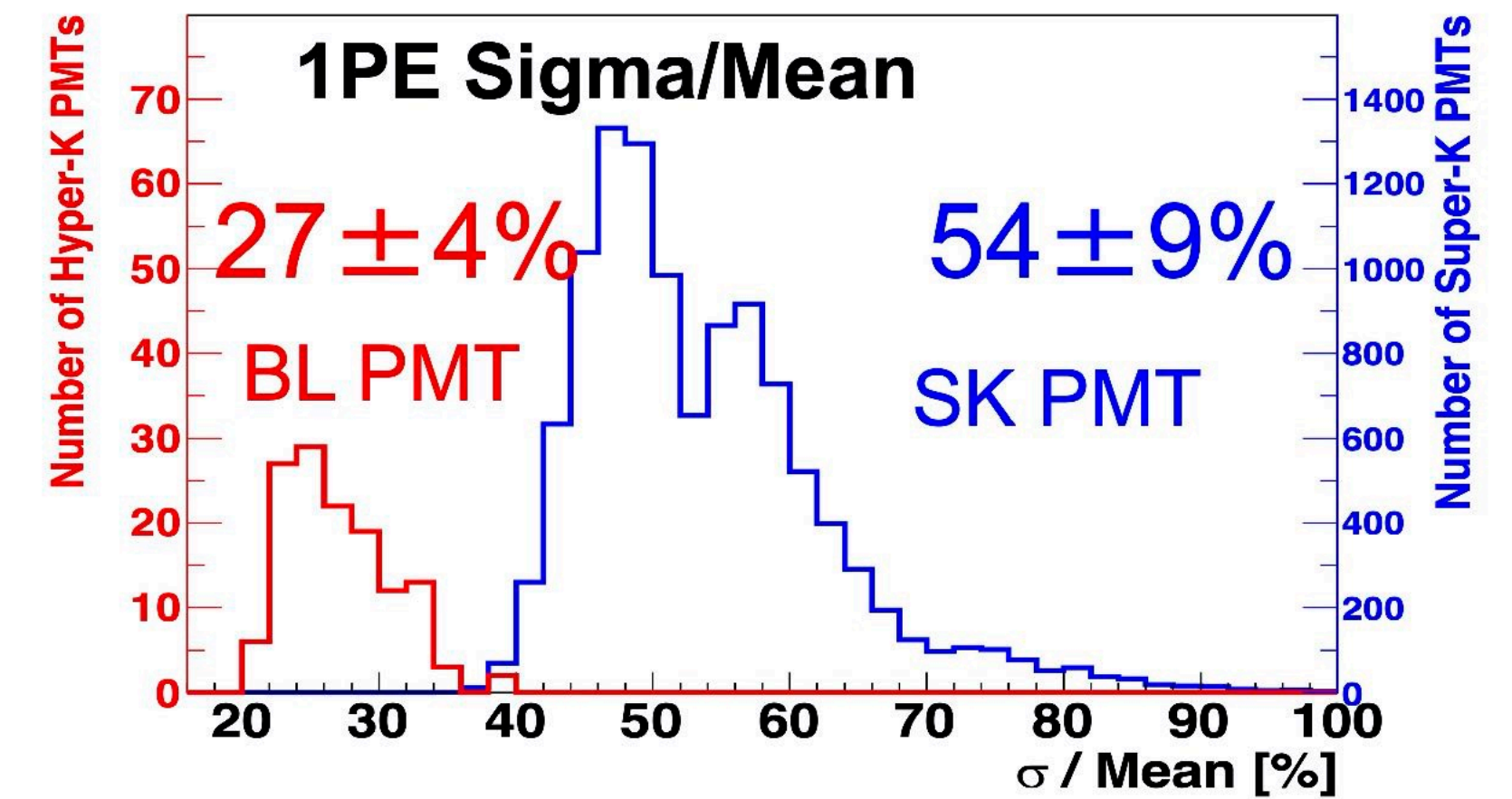
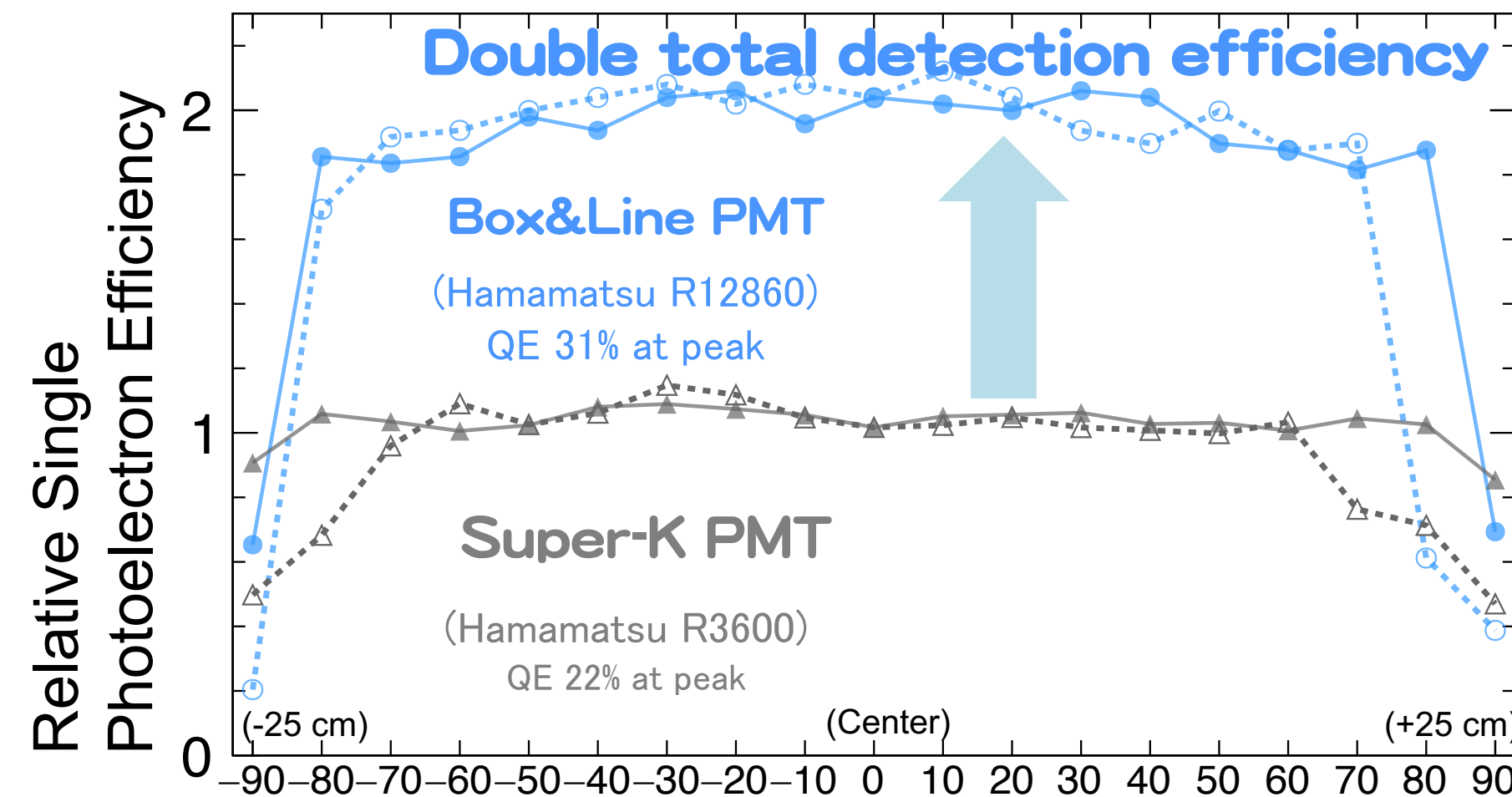
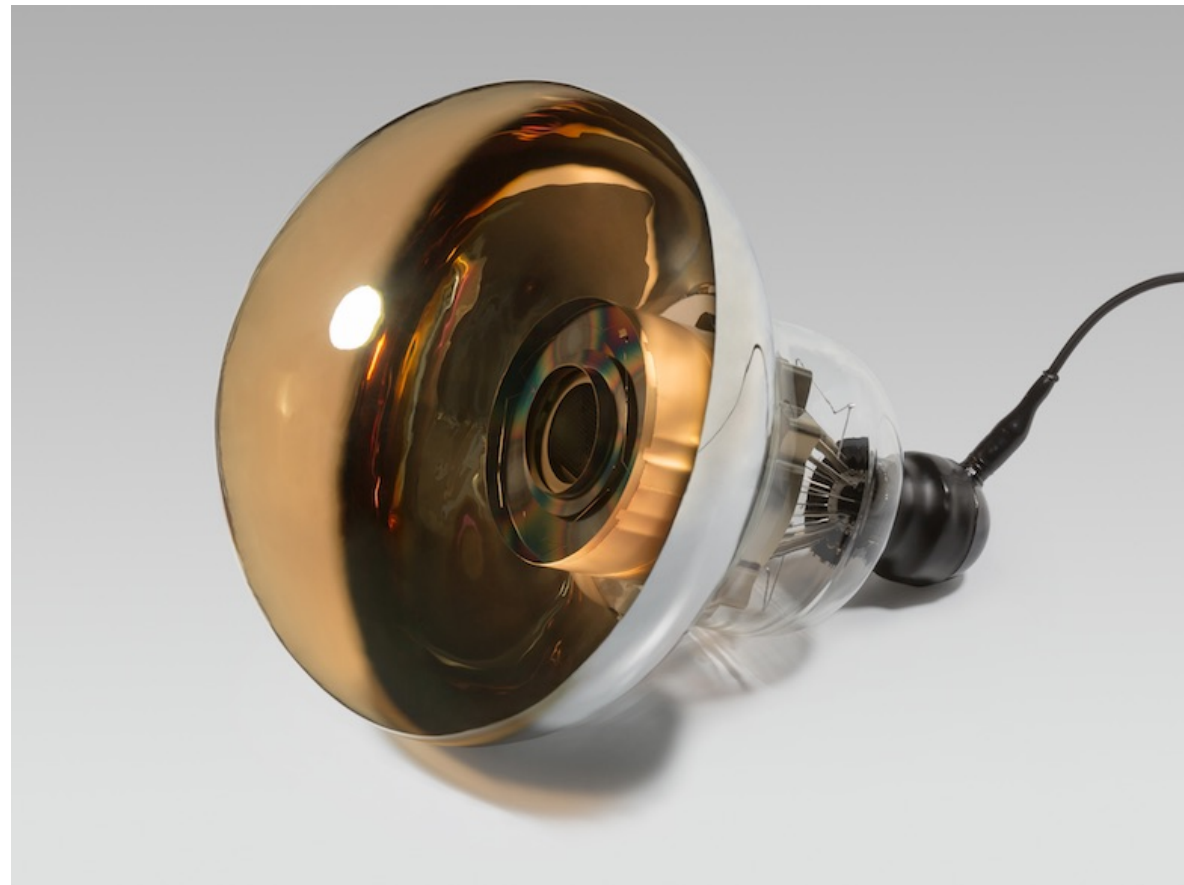
Hyper-K PMT Performance

Hamamatsu R12860

(Performance in SK tank, 1.7×10^7 gain)

x2 better photodetection efficiency (QE \times CE)

x2 better charge resolution



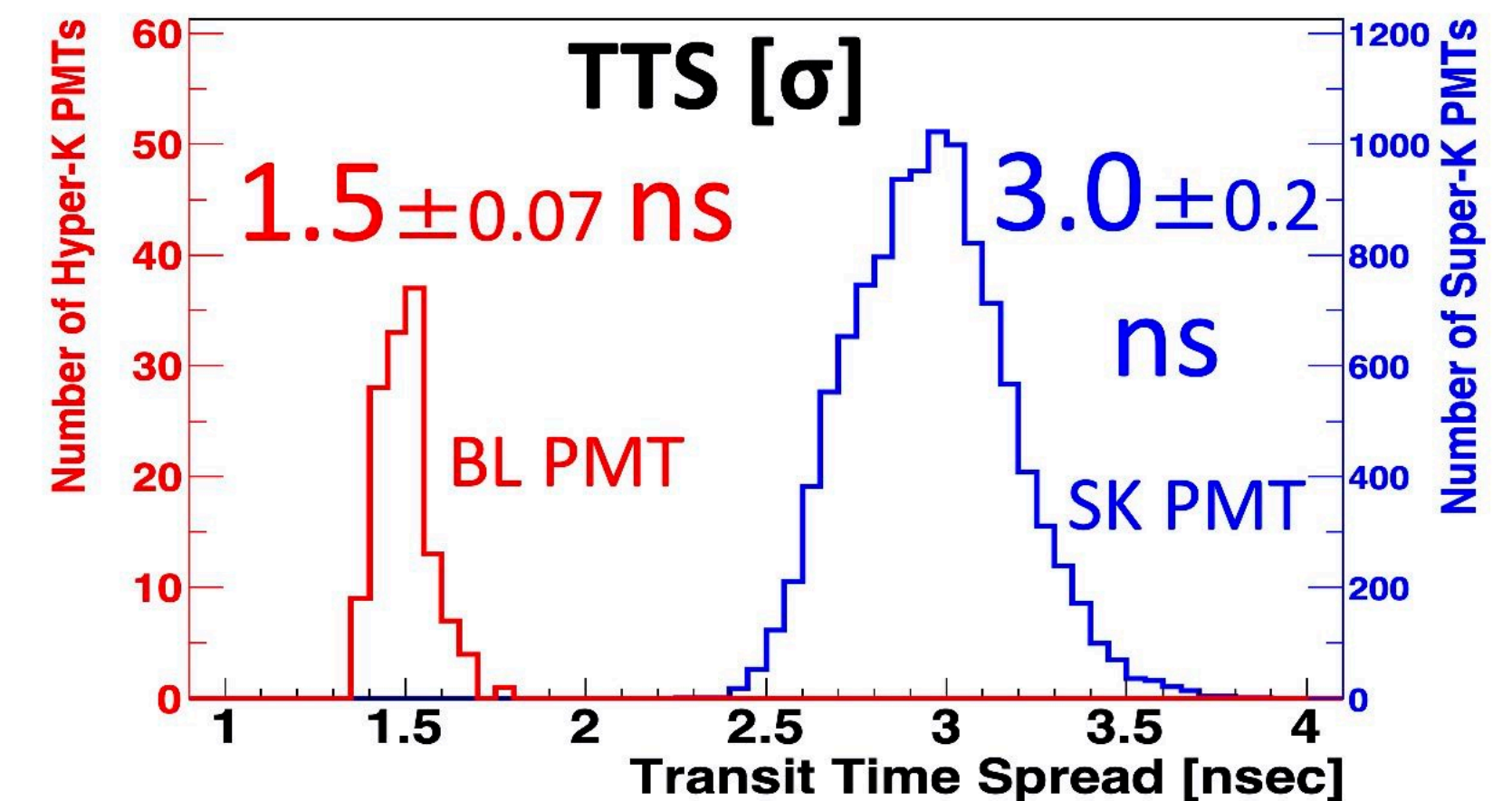
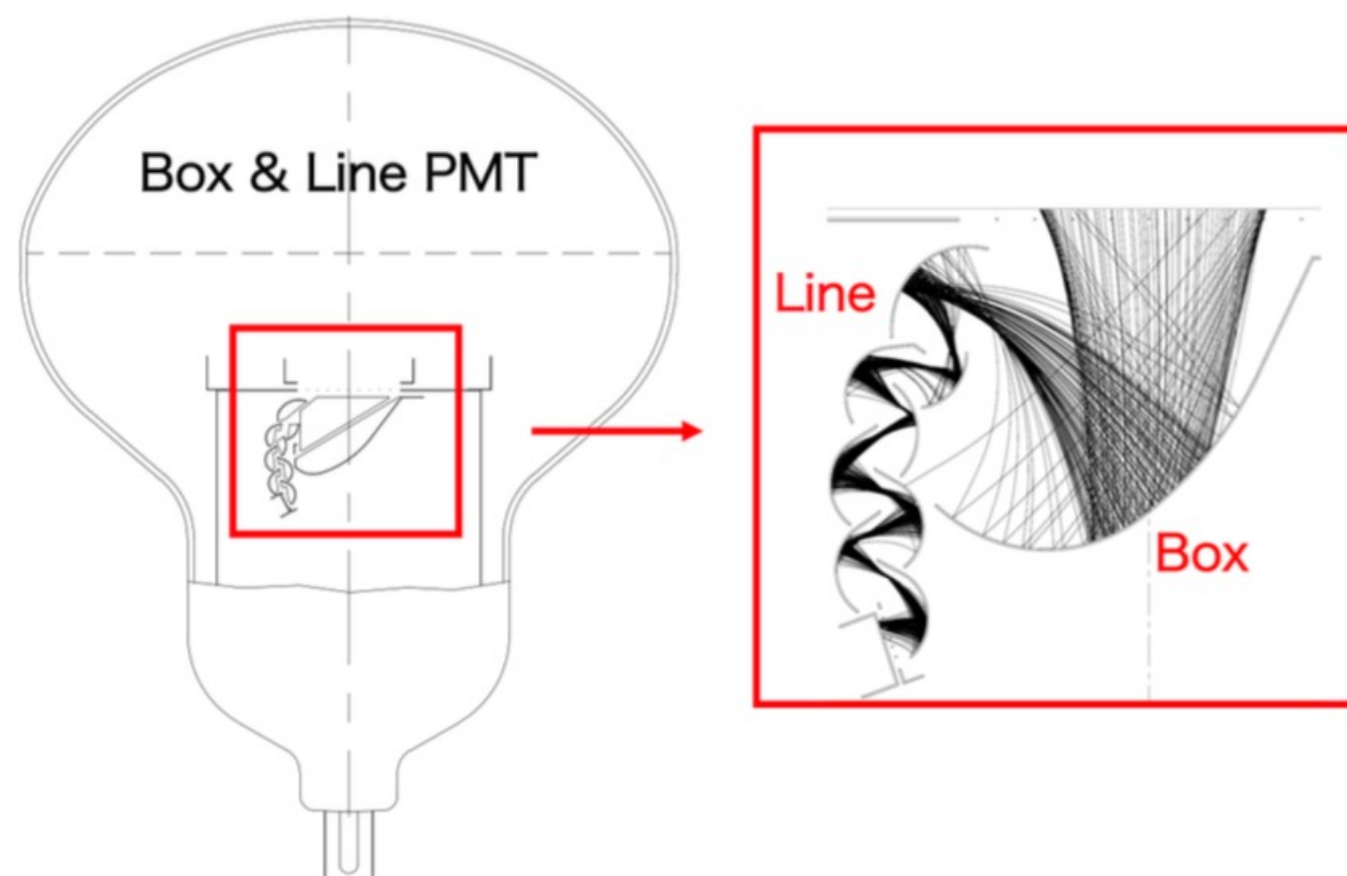
x2 better timing resolution

x2 better pressure tolerance
→ enable deeper tank design,
project cost reduction

All PMTs will be tested >0.85 MPa

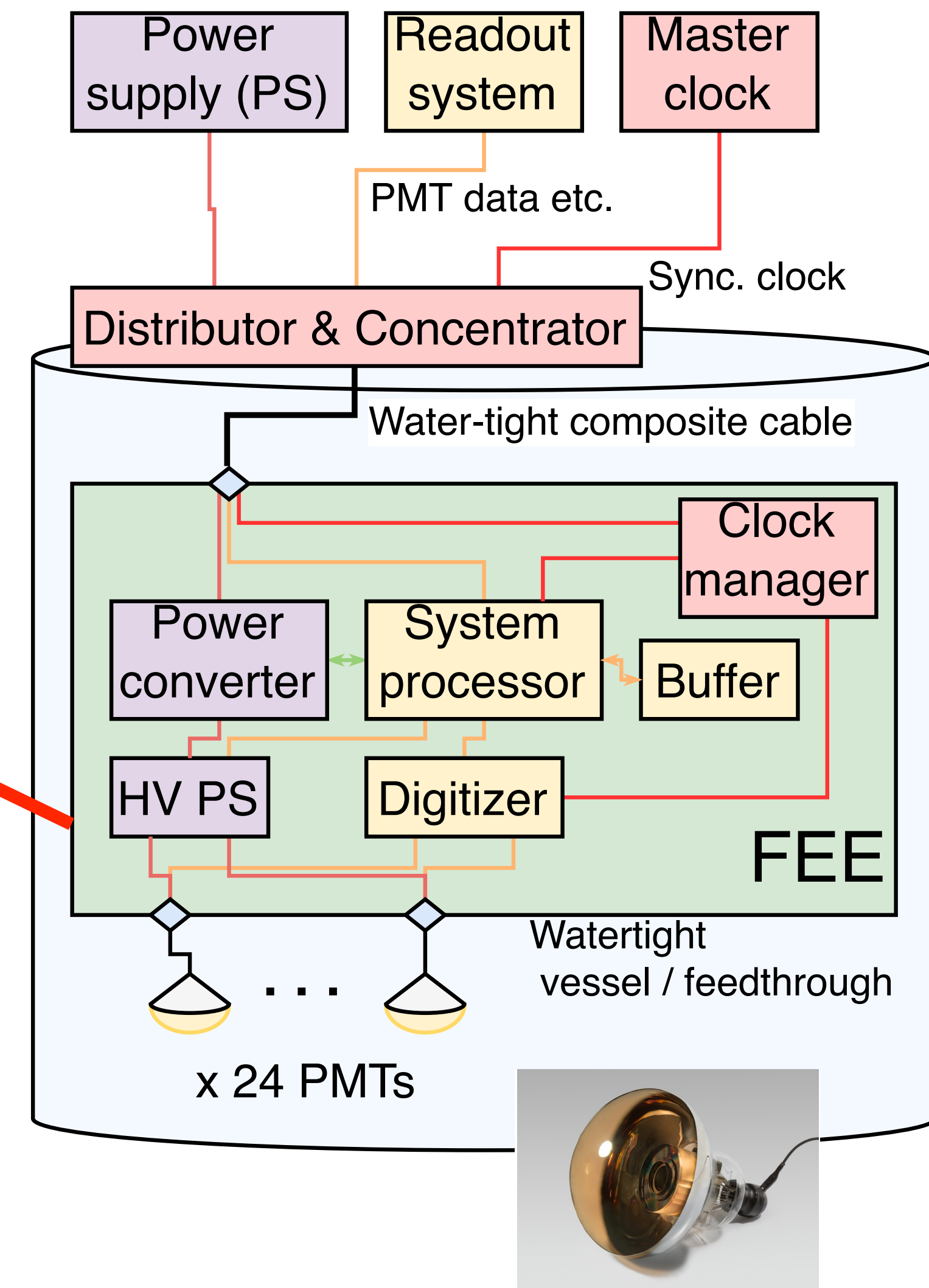
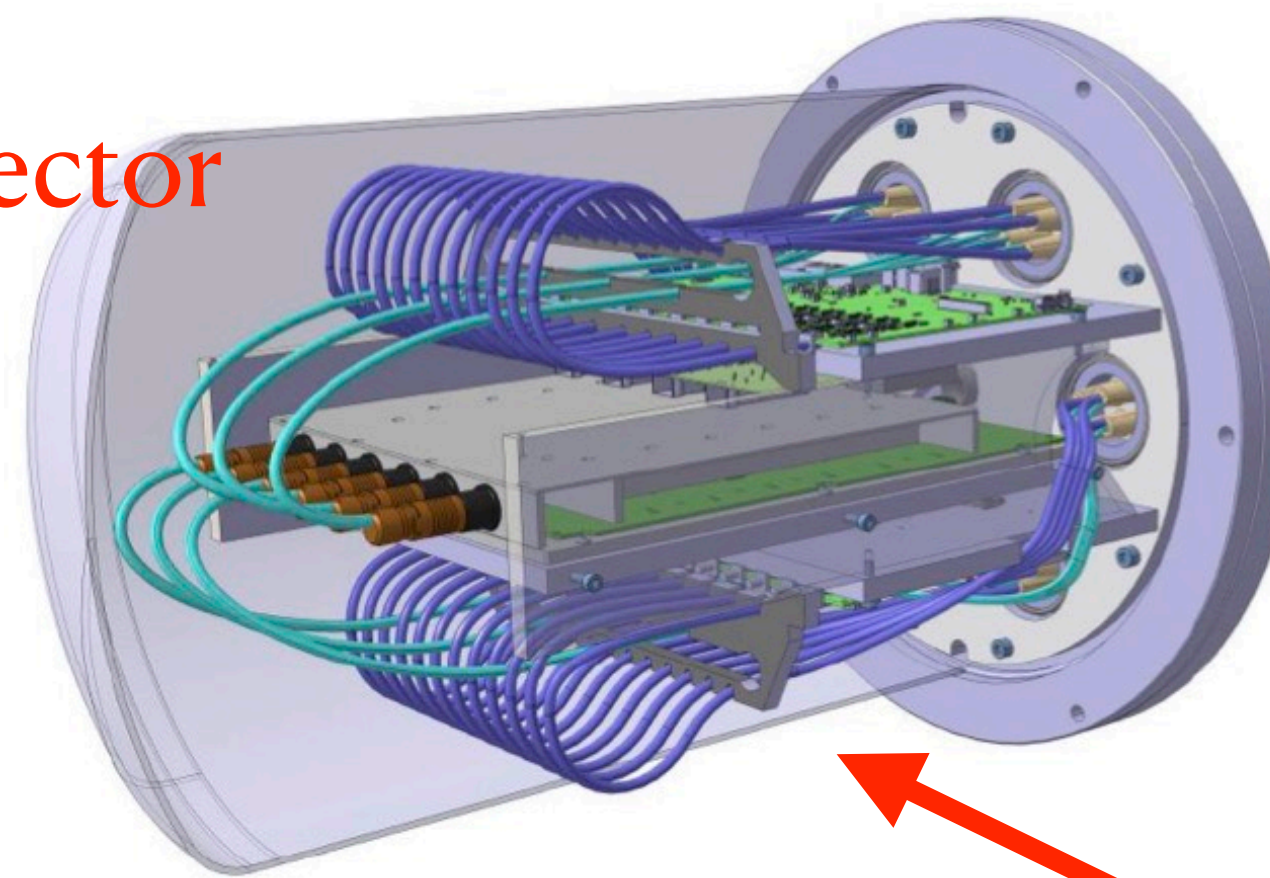
Low dark rate (4kHz) and RI

Box&Line dynode



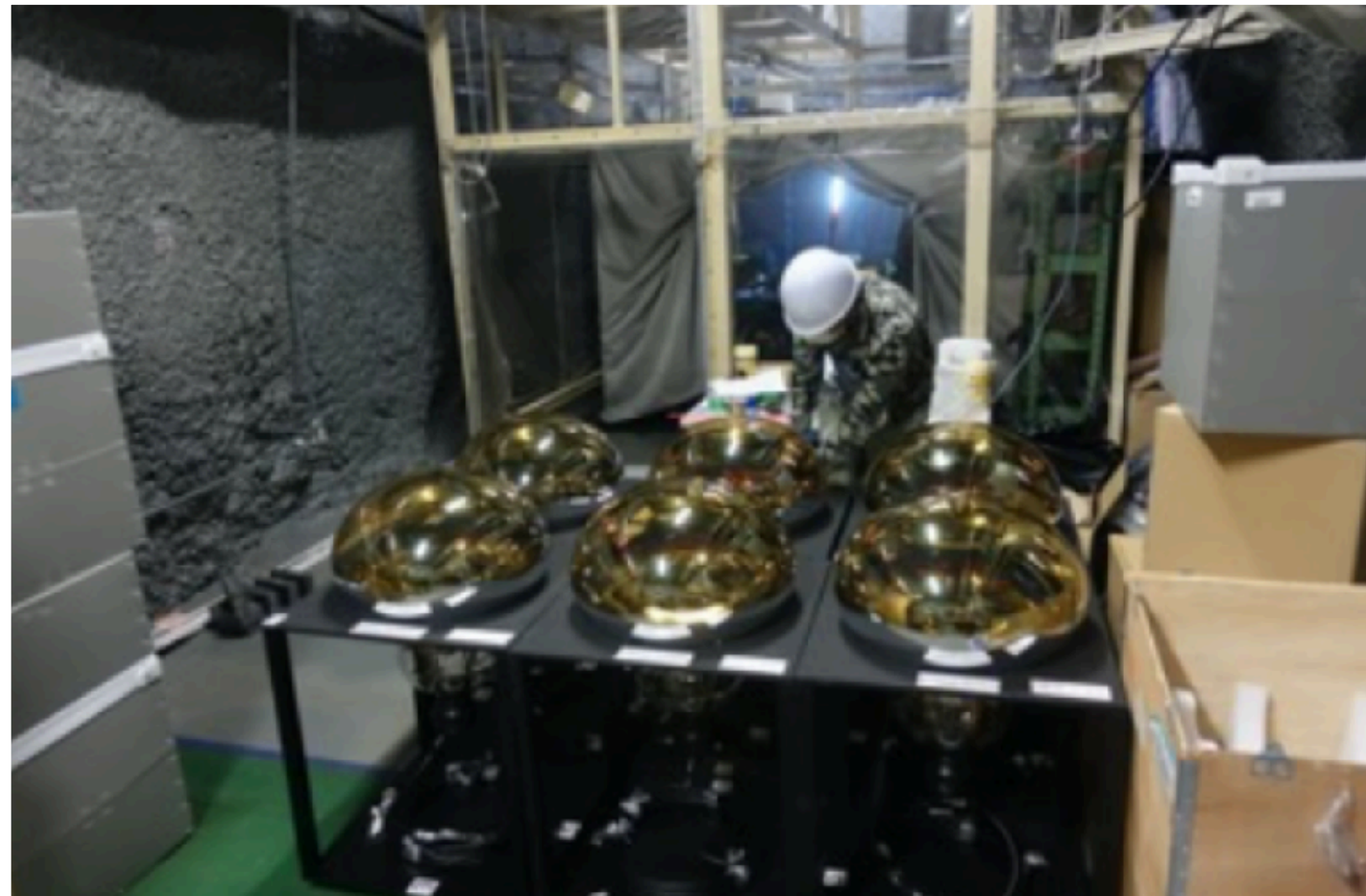
Electronics system

- Developed to maximize the performance of improved PMTs
- Frontend electronics placed underwater
 - Digitizing signals near PMTs
 - Maximize the performance of the detector
- Challenges
 - Everything in water-tight vessels
 - Water-tight connectors and cables
 - Very high reliability required
 - Synchronization of distributed components
- Large international collaboration project by itself
 - Development, production, assembly, testing, calibration, installation, ..
 - Planning assembly/testing at CERN



Detector component (some production starting)

PMTs

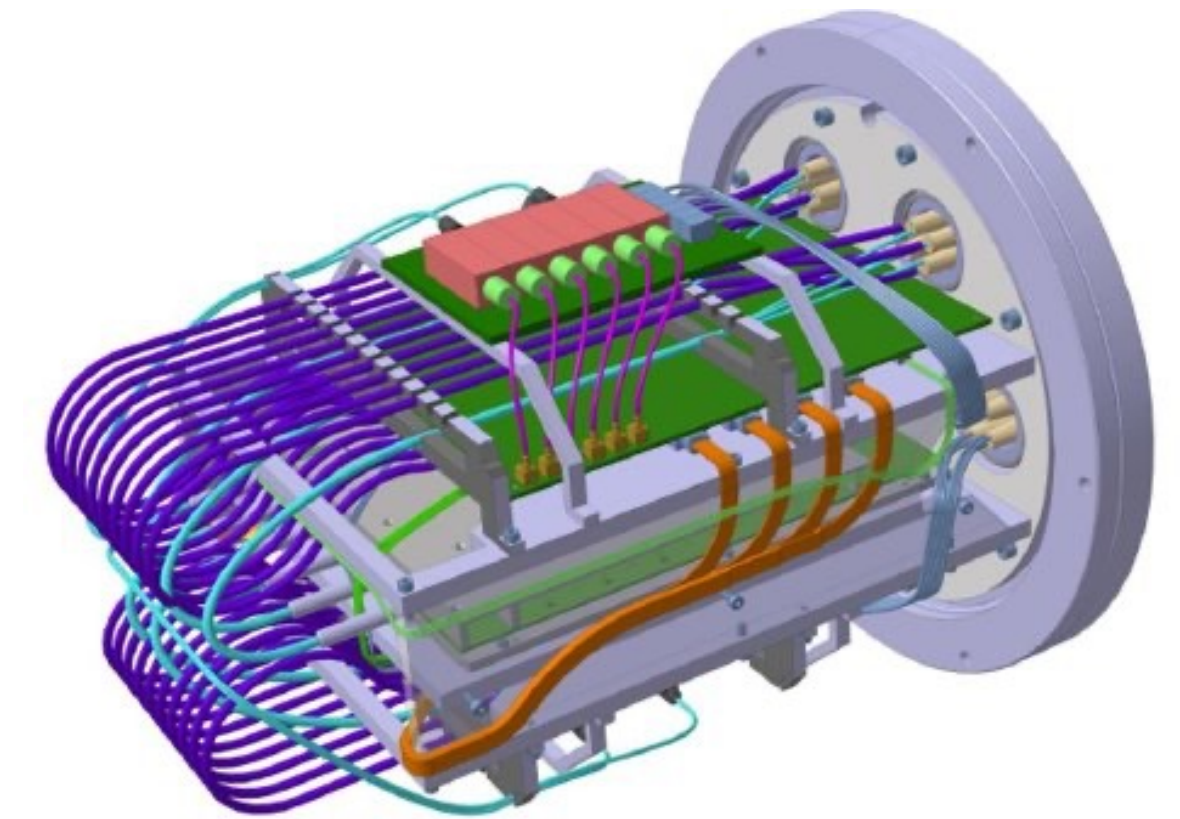


ID mockup



Underwater electronics:

20 x 50 cm ID PMTs + 12 x OD PMTs

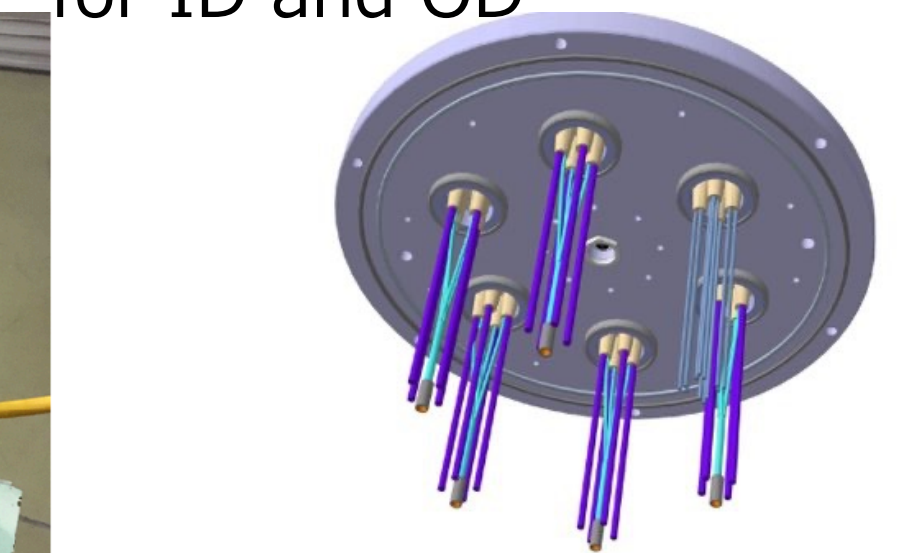
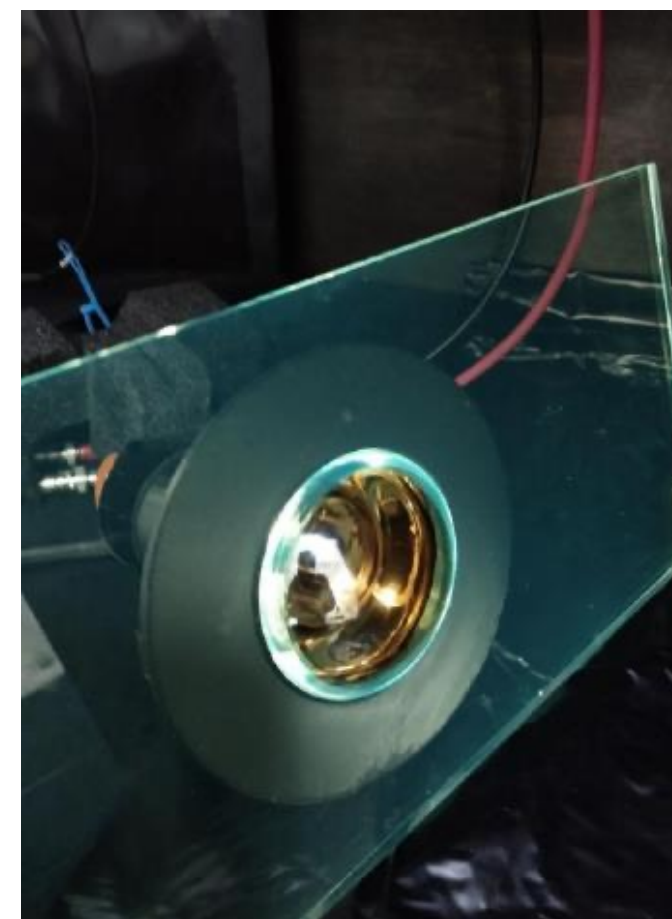


Feedthroughs for ID and OD

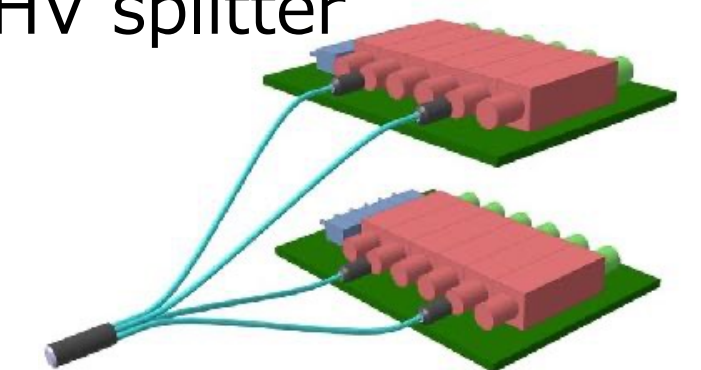
PMT cover



Multi-PMT module: Outer detector: PMT+WLS plate



OD signal + HV splitter



J-PARC Upgrade

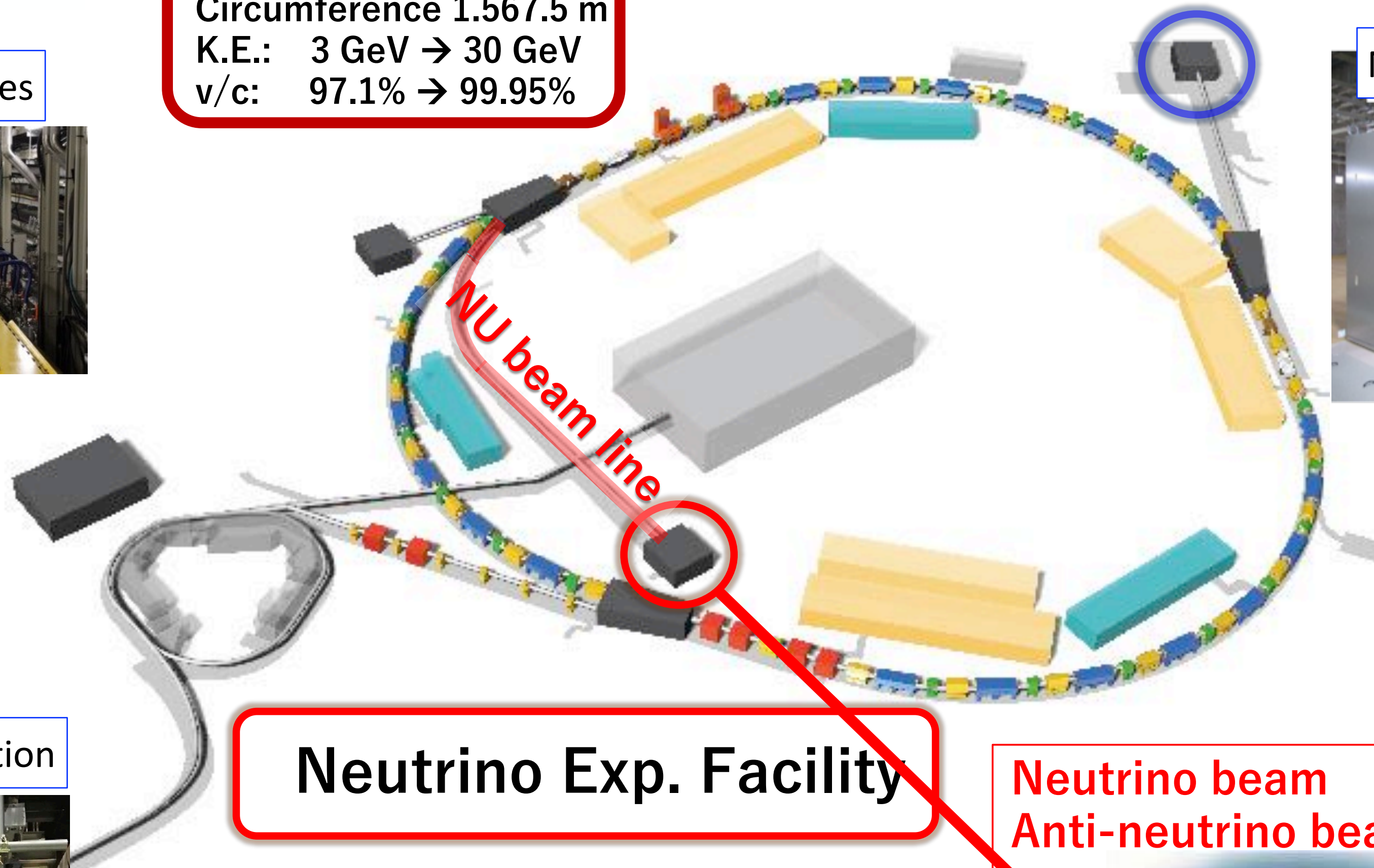


MR-RF cavities



Main Ring

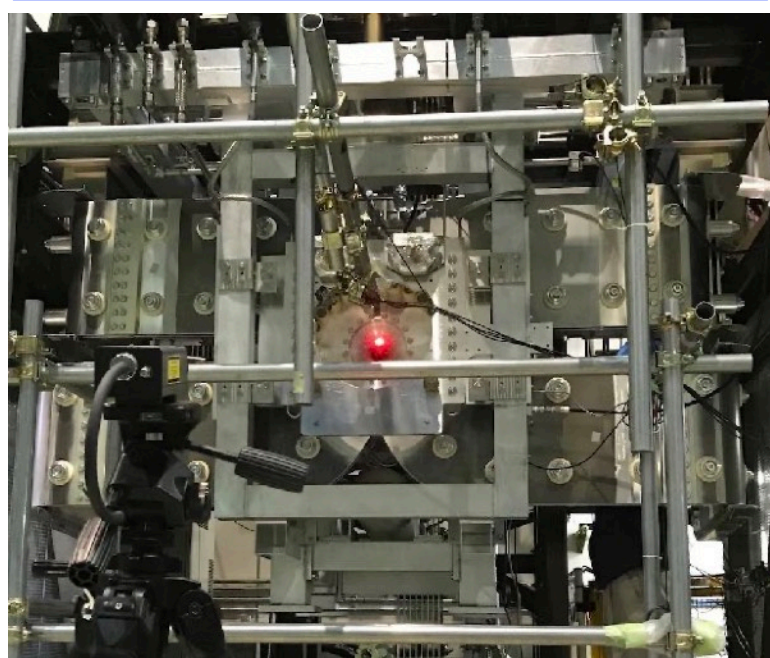
Circumference 1.567.5 m
K.E.: 3 GeV → 30 GeV
v/c: 97.1% → 99.95%



New main magnet PS for high rep. rate



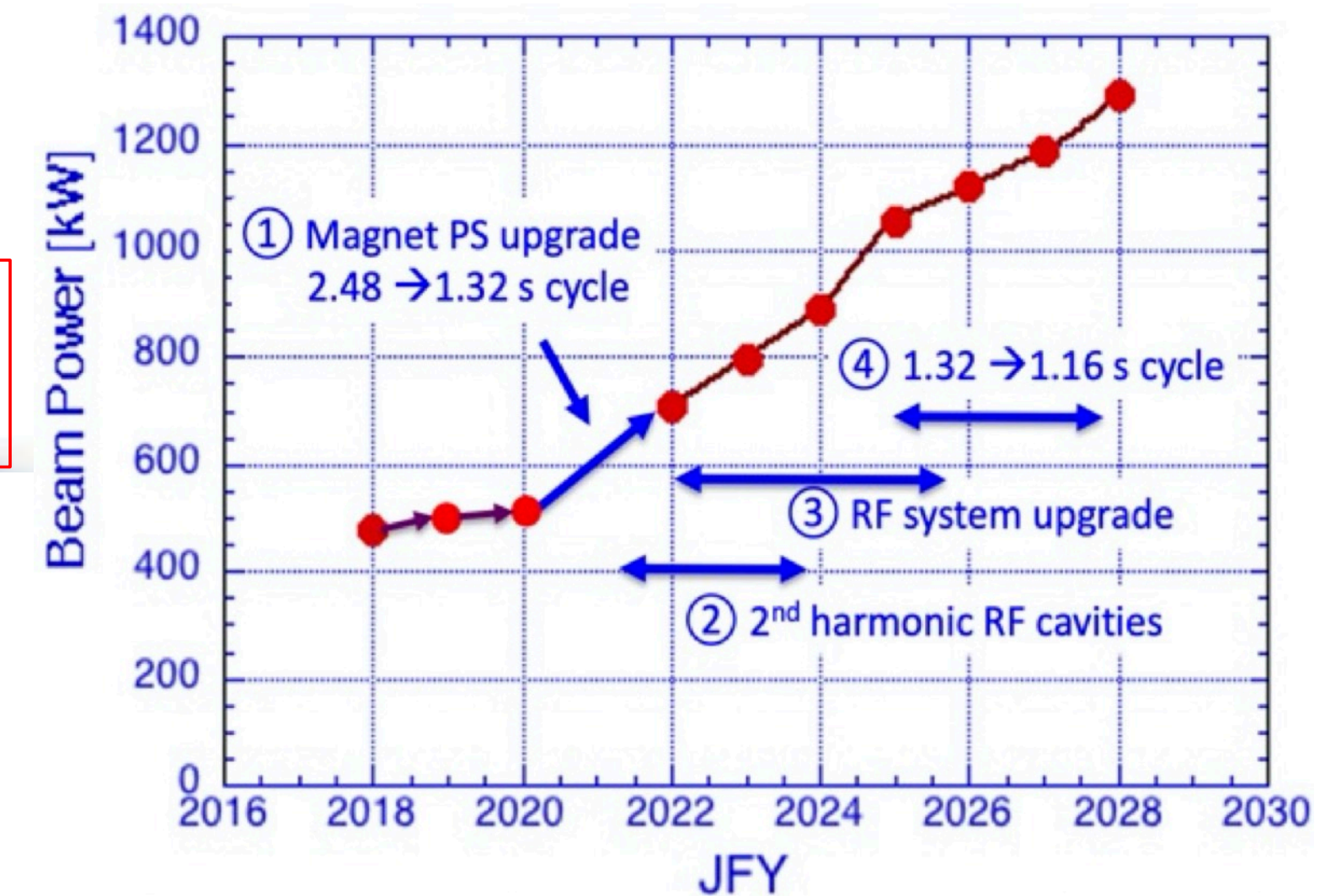
320kA horn operation



Neutrino Exp. Facility

Neutrino beam
Anti-neutrino beam

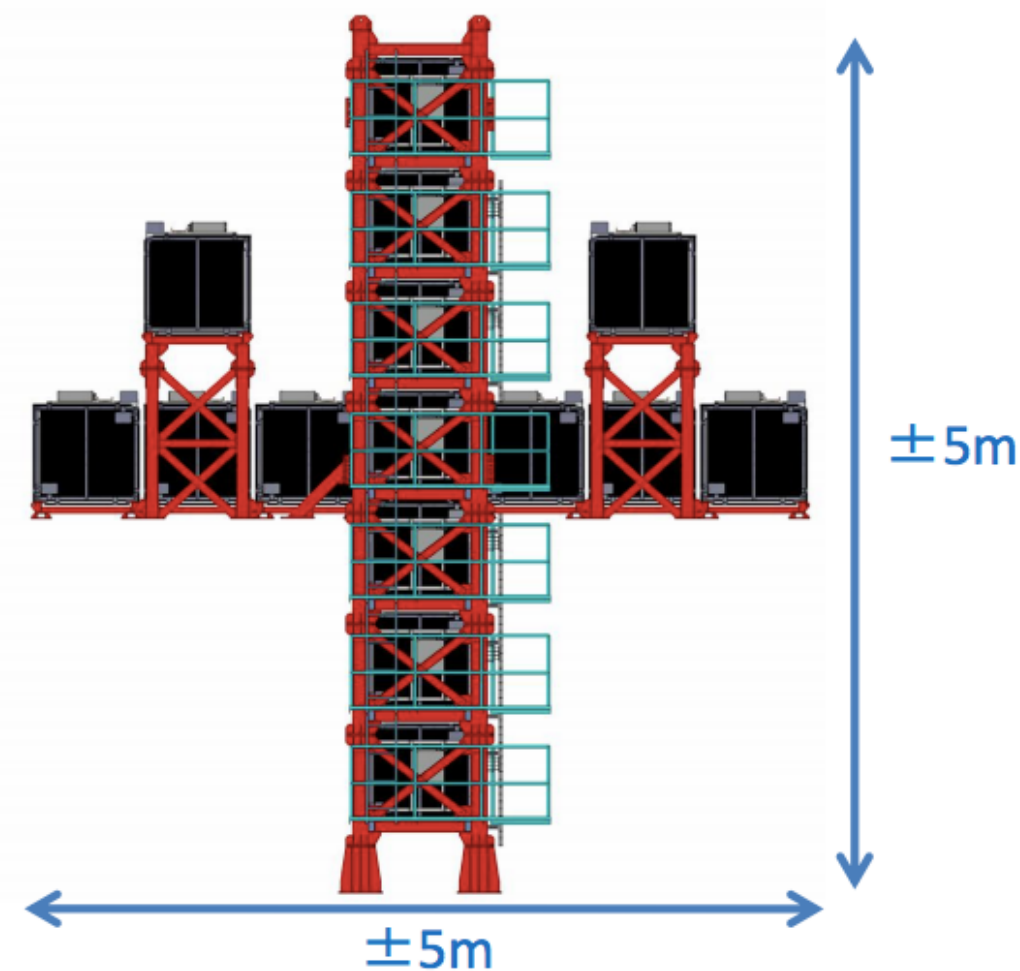
Achieved 515 kW in JFY2020
Aiming 1.3 MW by JFY2028



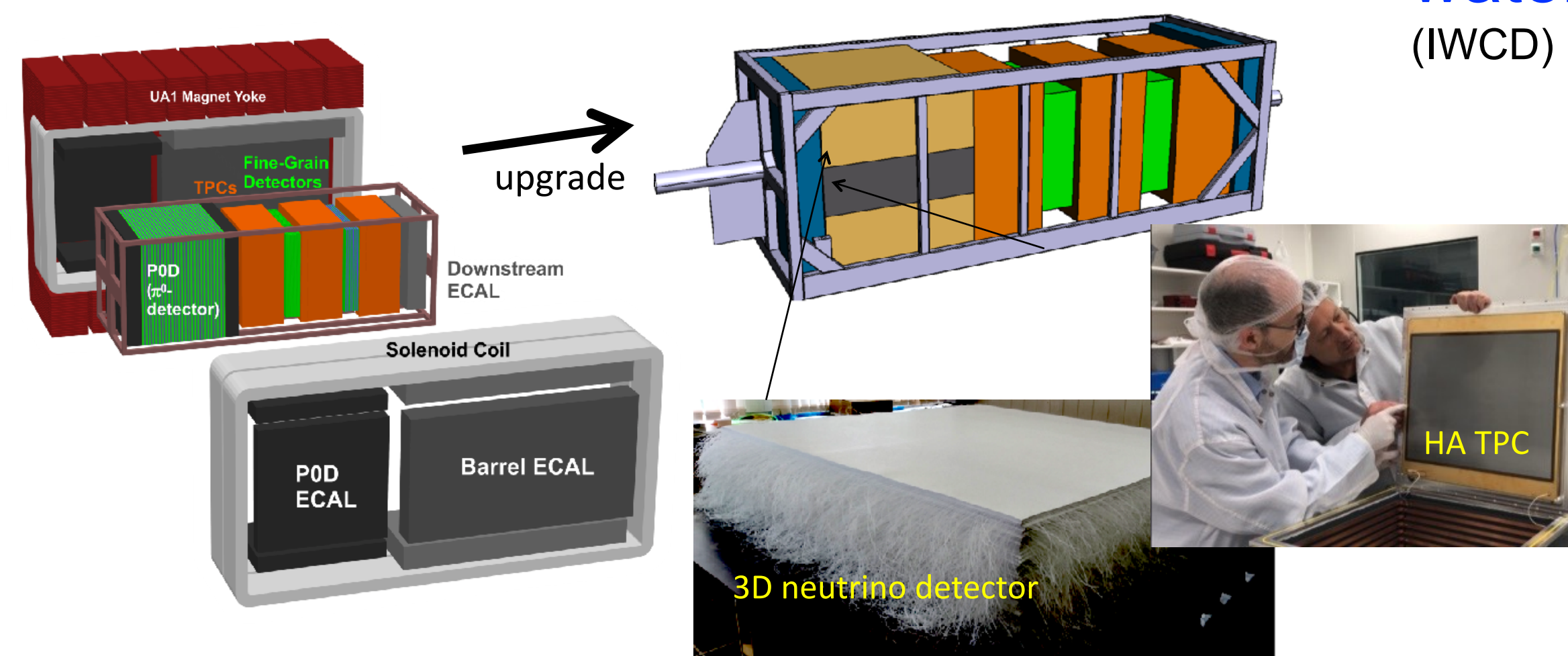
Neutrino detectors at J-PARC

Critical components to precisely understand J-PARC beam and neutrino interactions.

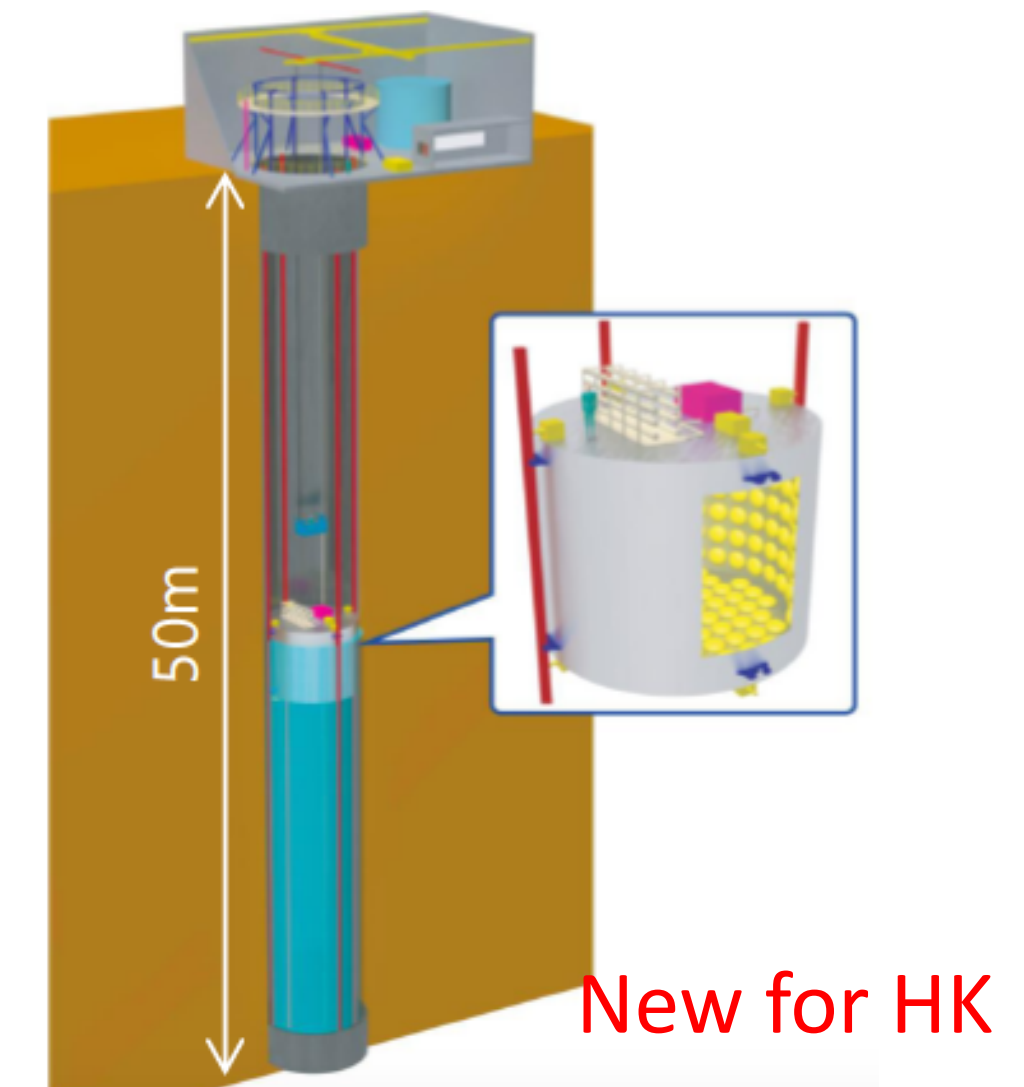
On-axis Detector (INGRID)



Off-axis Magnetized Tracker (ND280 → Upgrade for T2K → Upgrade for HK)



Off-axis spanning Intermediate water Cherenkov detector (IWCD)



- On-axis detector: Measure beam direction and event rate
- Off-axis magnetized tracker: Measure primary (anti)neutrino interaction rates, spectrum and properties. Charge separation to measure wrong-sign background
→ Upgrade by T2K experiment and Intensive discussion for further upgrade in HK-era is on-going.
- Intermediate WC detector: H₂O target with off-axis angle spanning orientation.
→ Detector site investigation and conceptual facility design is on-going.

Connection to FNAL and CERN: Beam test of detectors, Hadron production measurements for J-PARC neutrino beam

Summary

Hyper-Kamiokande with J-PARC

- Important physics targets
 - Neutrino CP violation: Discovery with 5σ for $\sim 60\%$ parameter regions
 - Nucleon Decay Search for testing GUT: $\tau > 10^{35}$ years for $p \rightarrow e^+ \pi^0$
 - Neutrino Astrophysics: Supernova neutrinos
- Big Water Cherenkov detector with 190 kton fiducial mass
 - Facility and Detector construction are on-going for the operation starting in 2027
- J-PARC neutrino beam being upgraded toward 1.3 MW power
- Your support is essential!