

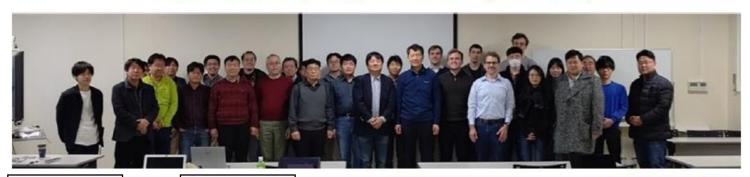
# Status of the JSNS<sup>2</sup>-I/II

ChangDong Shin (KEK)
On behalf of the JSNS<sup>2</sup> I / II collaboration

### JSNS<sup>2</sup> / JSNS<sup>2</sup>-II Collaboration

### (J-PARC Sterile Neutrino Search at J-PARC Spallation Neutron Source)

Collaboration meeting @ J-PARC (2020/Feb)





JAEA KEK Kitasato Kyoto Osaka Tohoku



Chonnam National Jeonbuk National

Dongshin GIST

Kyungbook

Kyung Hee

Seoyeong

Soongsil

Sungkyunkwan
Seoul National of
sci and tech



BNL Florida Michigan Utah



JSNS<sup>2</sup> collaboration (61 collaborators)

- 10 Korean institutions (24 members)
- 6 Japanese institutions (29 members)
- 4 US institutions (7 members)
- 1 UK institution (1 member)



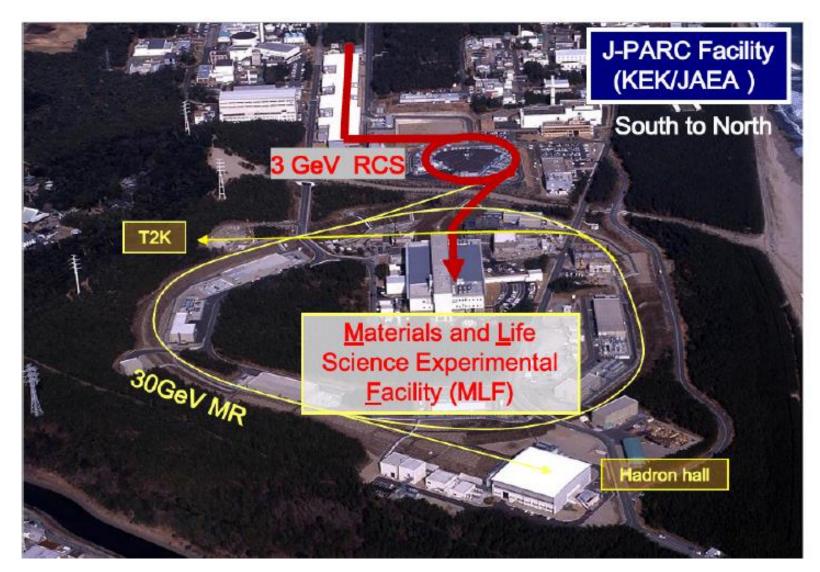
# Indication of a sterile neutrino ( $\Delta m^2 \sim 1 eV^2$ )

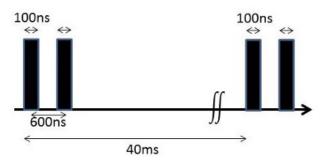
 Anomalies, which cannot be explained by standard neutrino oscillations for ~20 years are shown

Experiments	Neutrino source	Signal	Significance	E(MeV)	L(m)
LSND	μ Decay-At-Rest	$\bar{v}_{\mu}  ightarrow \bar{v}_{e}$	$3.8\sigma$	40	30
MiniBooNE	$\pi$ Decay-In-Flight	$v_{\mu} \rightarrow v_{e}$	$4.8\sigma$	800	600
IVIIIIIBOOINE		$ar{v}_{\mu}  ightarrow ar{v}_{e}$			
BEST	e capture	$v_e \rightarrow v_{\chi}$	$4.2\sigma$	<3	10
Reactors	Beta decay	$\bar{v}_e  ightarrow \bar{v}_{\chi}$	$3.0\sigma$	3	10-100

- JSNS<sup>2</sup> uses the same neutrino source ( $\mu$ ), target (H), and detection principle (IBD) as the LSND
  - Even if the excess is not due to the oscillation, JSNS<sup>2</sup> can catch this directly
  - two advantages : short-pulsed beam and used the gadolinium(Gd)-loaded liquid scintillator(GdLS)

# J-PARC facility





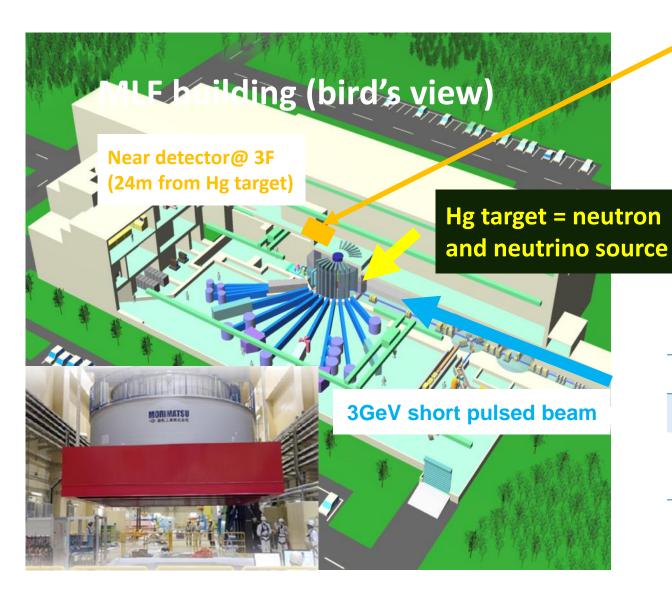
Low duty factor beam (short-pulses + low repetition rate) Gives an excellent signal to noise ratio

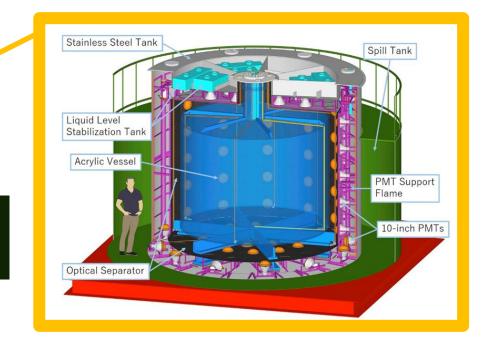
### 1 MW (design)

- 0.6MW (2021 Jan Apr/5)
- 0.7MW (2021 Apr/5 June/22)
- 0.7MW (2022 Jan/28 Apr/6)
- 0.8MW (2022 Apr/7 Jun/3)
- 0.8MW (2023 Apr/15 Jun/2)

### JSNS<sup>2</sup> detector

Nucl. Instrum. Methods A 1014 165742 (2021)





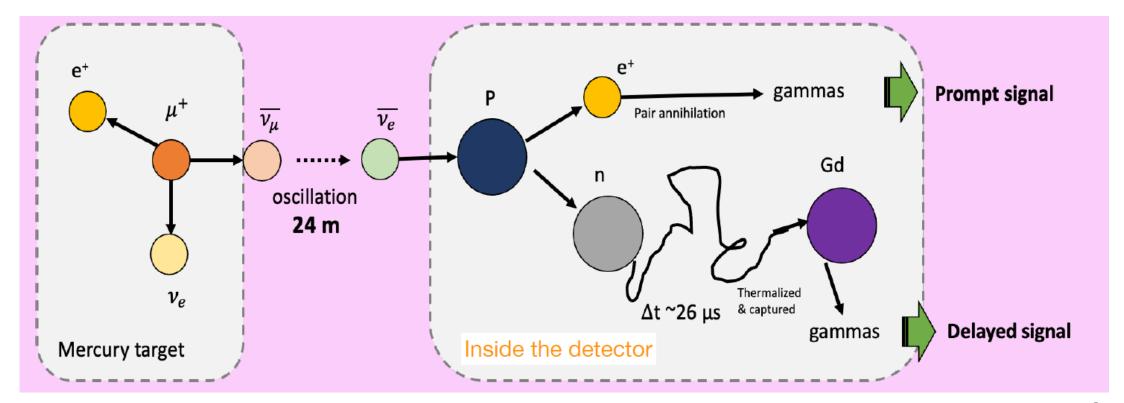
	Liquid	Volume
Target	GdLS + DIN (10%)	17 tons
gamma-catcher and veto	Pure LS	31 tons

- 96, 10-inch PMTs for the inner detector
- 24, 10 inch PMTs for the veto

### **Production and detection**

- If sterile  $\nu$  exist,  $\bar{v}_{\mu} \rightarrow \bar{v}_{e}$  oscillation occurs with 24m.
- Coincidence of Inverse Beta Decay (IBD)
  - Positron annihilation
  - Neutron capture on Gd

	Timing	Energy	
Prompt	$1.5 < T_p < 10$ μs	20 < E < 60 MeV	
Delayed	$\Delta T_{p-d} < 100$ μs	$7 < E < 12 \; MeV$	

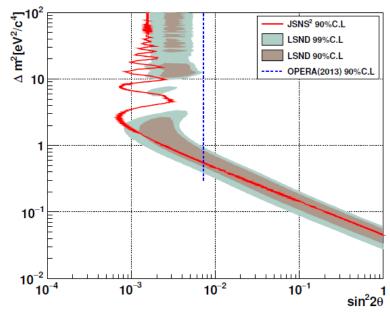


# Expected energy spectrum and sensitivity

- $\bar{v}_e$  follows decay-at-rest  $\bar{v}_\mu$  energy distribution
- Prompt single rate:  $\sim 3.9 \times 10^{-4}$  per spill
- Delayed single rate:  $\sim 4.8 \times 10^{-3}$  per spill
- Spectral fit is sensitive to the difference of energy spectrum

$\times 10^{-3}$	>	$\sin^2 2\theta = 3.0 \times 10^{-3}$
$eV^2$	Events/4MeV	F   10lal
: I	nts/	$\overline{\nabla}_{e}$ from $\mu^{e}$
	š	$\begin{array}{c c} \hline & \overline{\mathbf{v}}_{\mathrm{e}} \text{ from } \overline{\mathbf{v}}_{\mu} \text{ oscillation} \\ \hline & \overline{\mathbf{v}}_{\mathrm{e}} \text{ from } \mu \\ \hline & & 1^{2} \mathrm{C}(\mathbf{v}_{\mathrm{e}}, \mathrm{e}^{\mathrm{i}}) \mathrm{N}_{\mathrm{g.s}} \\ \hline & & - \mathrm{Accidentals} \end{array}$
	ш.	25 Accidentals
		20
		15
		10 1 1
7		5
60		

Signal	$sin^2 2\theta = 3.0 \times 10^{-3}$ $\Delta m^2 = 2.5 eV^2$ (Best fit values of MLF)	87
	$sin^2 2\theta = 3.0 \times 10^{-3}$ $\Delta m^2 = 1.2 eV^2$ (Best fit values of LSND)	62
	$\overline{\nu}_e$ from $\mu^-$	43
	$^{12}C(\nu_e, e^-)^{12}N_{g.s.}$	3
background	beam-associated fast n	$\leq 2$
	Cosmic-induced fast n	negligible
	Total accidental events	20



from v, oscillation

30

40

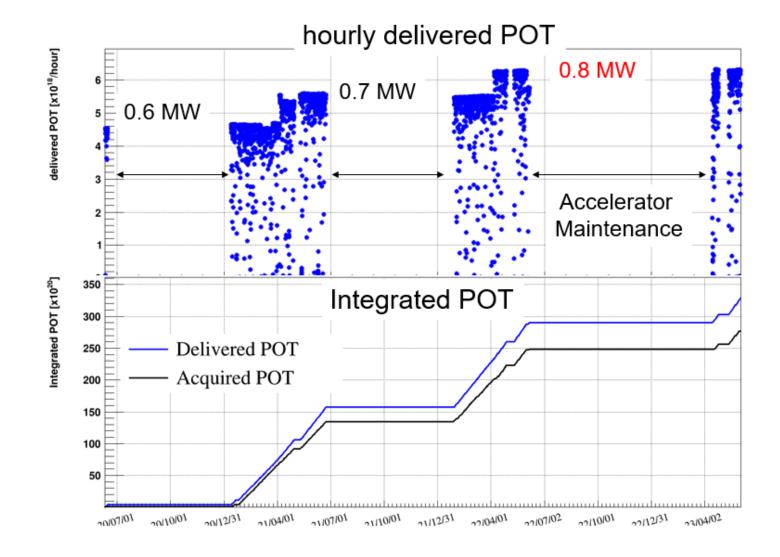
Accidentals

Events/4MeV

15

# **Operation**

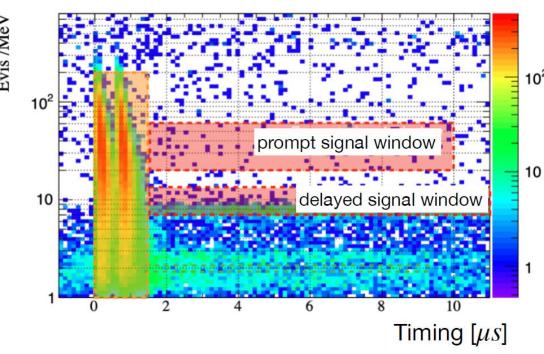
- Now beam power is very close to the design (1MW).
   0.95MW at RCS
- There is an accelerator maintenance period every year.
- 3.28 × 10<sup>22</sup>POT has been delivered. (28.7% of the approved POT of JSNS<sup>2</sup>)

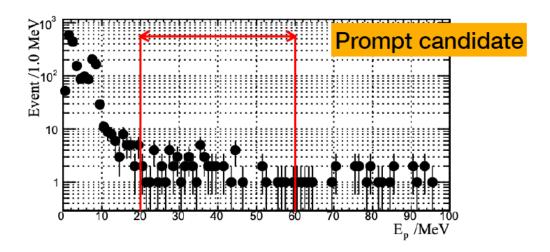


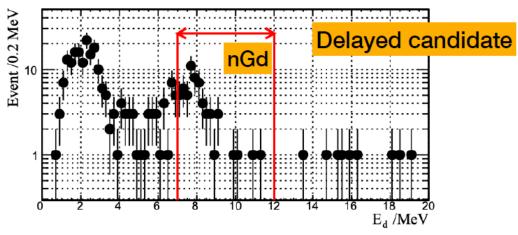
# **Commissioning run**

(Eur, Phys. J.C (2022) 82:331)

- Integrated POT :  $8.9 \times 10^{20}$ 
  - expected IBD << 1 event
- Beam trigger with 25μs width





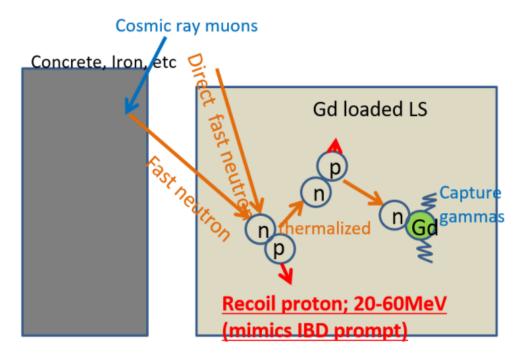


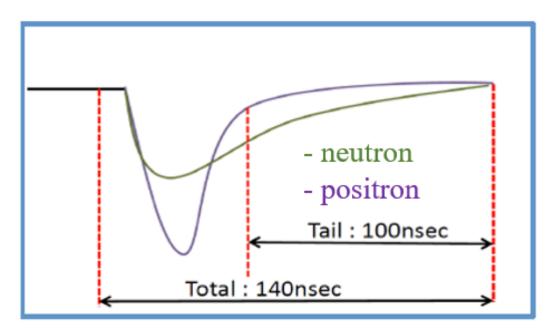
#### Observed correlated event candidates

- 59±8 events / 8M spills
- Cosmic-induced fast neutrons are the dominant background
- expected cosmic-induced fast neutron is 55.9  $\pm$ 4.3 events
- Pulse shape discrimination (PSD) would reject them.

# Pulse Shape Discrimination (PSD)

- Fast neutrons can mimic IBD signals from electron anti-neutrino.
  - correlated background
- PSD can separate the IBD signals and fast neutrons.
  - the goal is to remove 99% of fast neutrons.

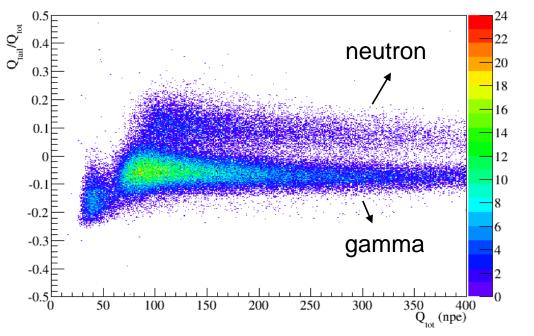


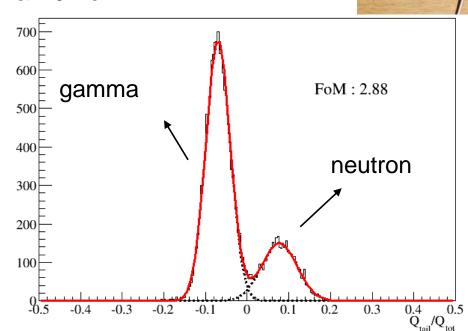


### The conventional PSD method

- The conventional PSD method uses the charge of the tail  $(Q_{tail})$  and total region  $(Q_{total})$  in the waveform.
- Generally, neutrons have large Q<sub>tail</sub>/Q<sub>total</sub> compared to leptons.



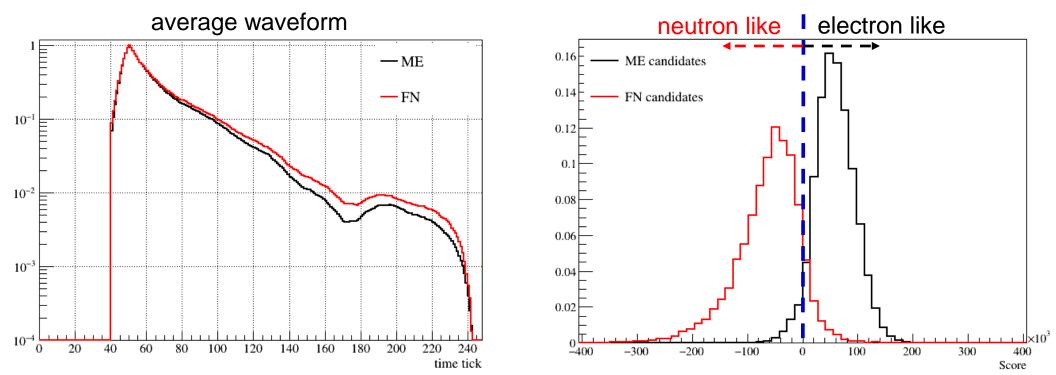




252**C**f

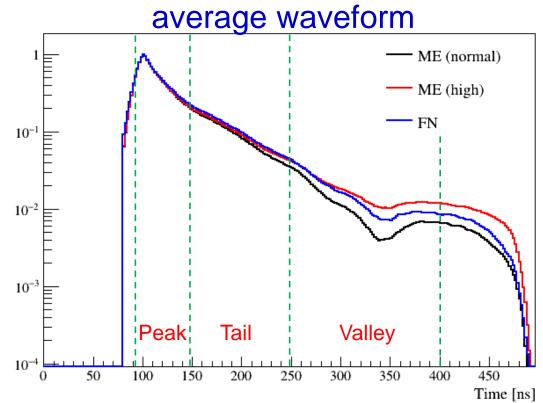
### Likelihood method

- The DAQ of JSNS<sup>2</sup> can measure a waveform every 2ns (500MHz sampling).
- The likelihood judges that all other points look like "neutron" or "electron" after peak normalization as 1.



### 2-Dimensional likelihood method

- More improvement of PSD capability is needed to get better sensitivity for sterile neutrino search.
- There are two types of waveforms in the ME control sample.
  - 2 dimensional likelihood is developed to separate them

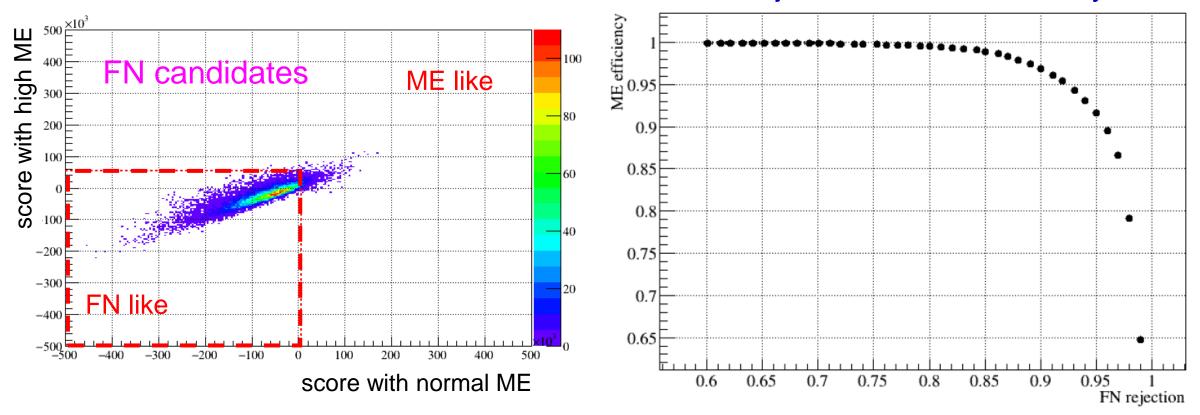


 The after-pulse from the parent muon affects the waveform of the following ME event.

### PSD results with 2D likelihood method

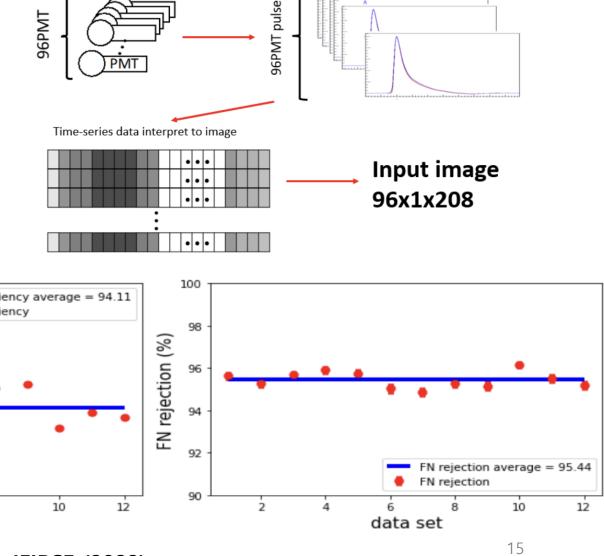
- The ME efficiency with 95% of the FN rejection is 90.98 +/- 0.05%.
- Both ME efficiency and FN rejection depend on a reference point.

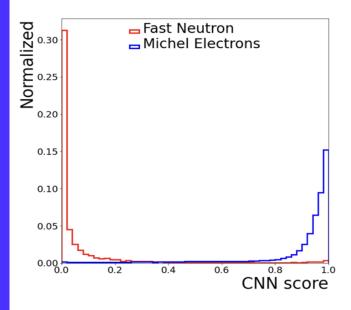


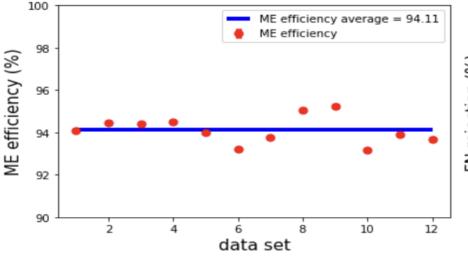


# **Convolutional Neural Network (CNN)**

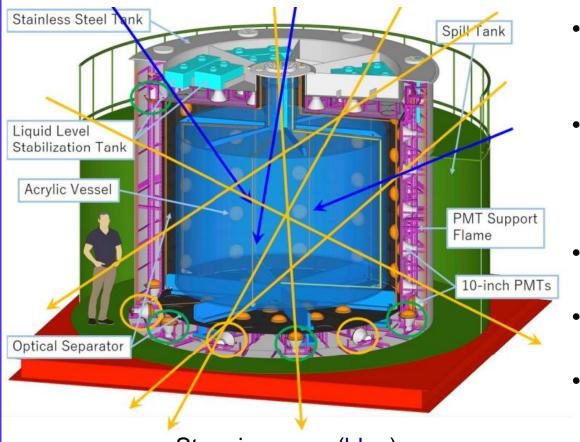
- Treated time-series data from a PMT with image data.
- Two independent efforts show consistent FN rejection results.







### Cosmic muon identification



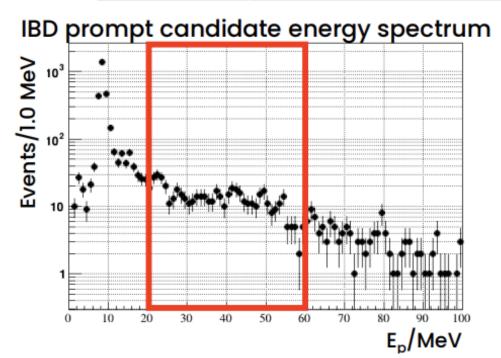
Stopping muon(blue)
Through-going muon (yellow)

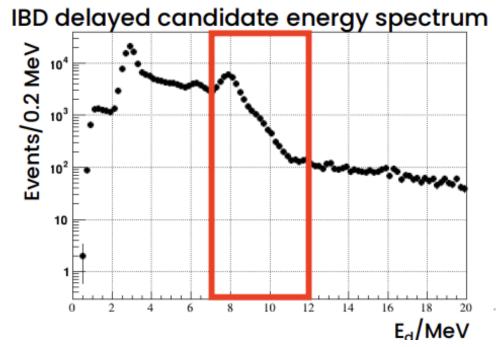
- Michel electron induced by cosmic muon and muon itself are one of backgrounds.
- Cosmic muon can be tagged by Top/Bottom Veto PMTs.
  - 12 PMTs on Top
  - 12 PMTs on Bottom
- Stopping muon candidates rate: 1487.8±0.5 Hz
- Through-going muon candidates rate: 605.4±0.4 Hz
- Michel candidates rate : 110.1±0.2 Hz (10~60 MeV)

# Single rates of the prompt and delayed

- 125  $\mu$ s time window from beam timing.
- External particles and the ME are rejected.

	Prompt	Delayed
Single rate	$(2.20 \pm 0.09) \times 10^{-4}$ /spill	$(1.80 \pm 0.01) \times 10^{-2}$ /spill

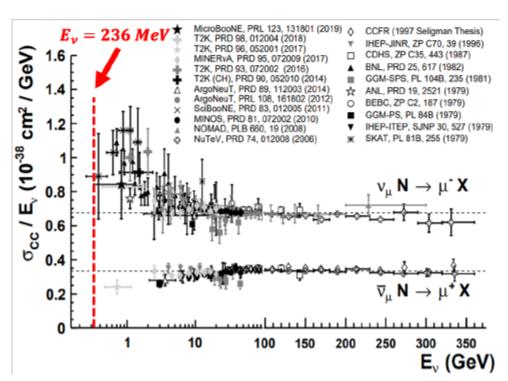


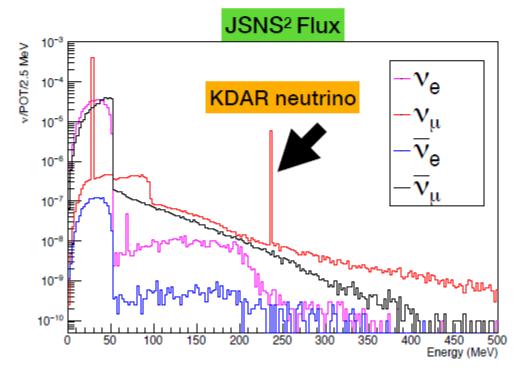


# KDAR neutrino measurement

# Kaon Decay-At-Rest neutrino measurement (KDAR neutrino: 236 MeV mono-energetic)

- Neutrino interaction models are a crucial part of neutrino physics, but poorly known at low energies.
- The JSNS<sup>2</sup> detector has the unique ability to measure the mono-energetic KDAR neutrino.
- Note that it is hard to see a clear energy peak of KDAR neutrinos in the data of the horn focused beam.



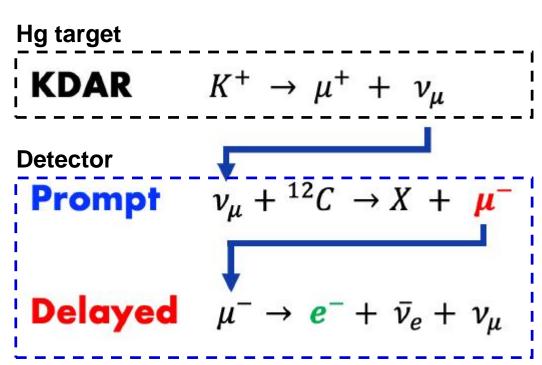


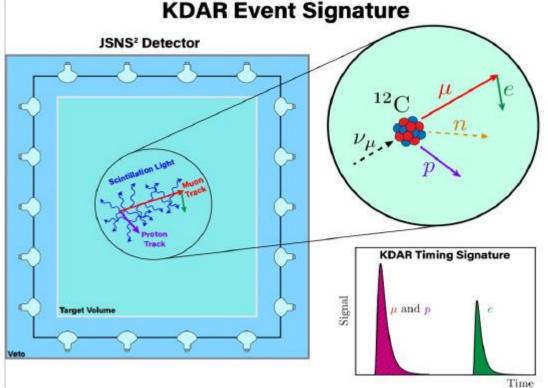
**Neutrino Workshop at IFIRSE (2023)** 

# KDAR signal measurement in JSNS<sup>2</sup>

### A double coincidence between

The initial neutrino interaction products and the subsequent muon decay.





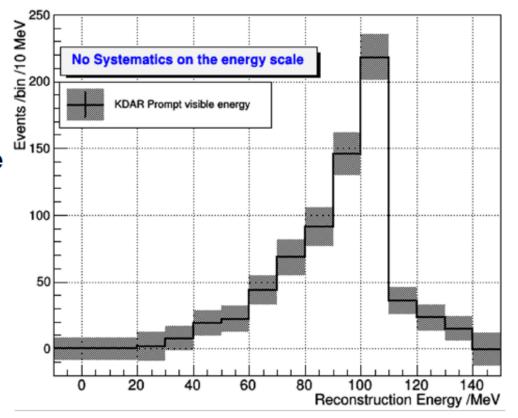
## First clear KDAR signal

### (Toward first precise KDAR measurement)

- KDAR peak is clearly seen
- High purity (95%) KDAR signal
  - Background: 5.2 %

 Note that the systematics on the energy scale are not included yet.

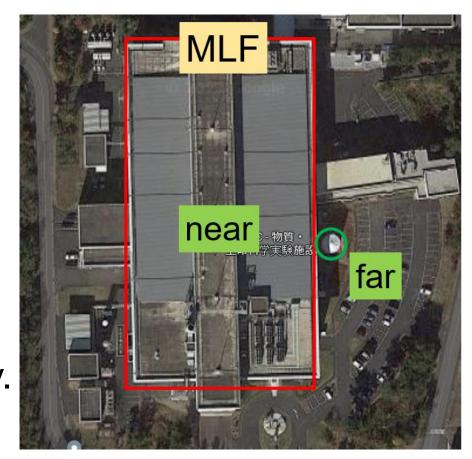
BKG ID	Correlated/ Accidental	BKG (# of events)		
1	Correlated	36.6 +- 34.8	5.0 +- 5.1%	
2	Accidental	1.5 +- 0.1	0.2 +- 0.01%	
KDAR Candidates : 730 events		38.1 +- 38.4	5.2 +- 5.3%	



# JSNS<sup>2</sup>-II

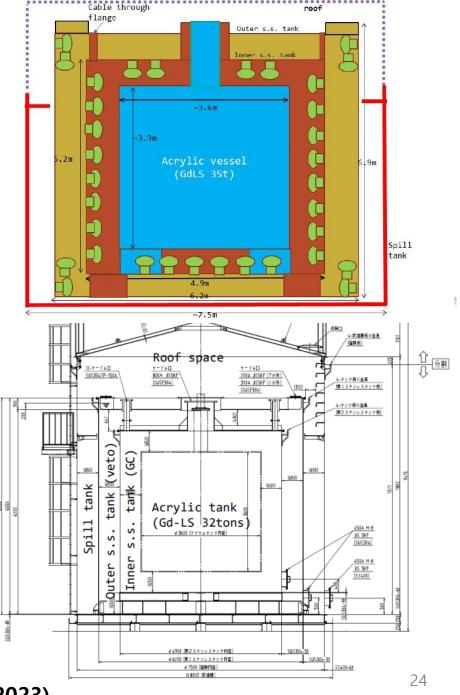
# JSNS<sup>2</sup>-II (Second phase of JSNS<sup>2</sup>)

- New far detector
  - fiducial 32 tonnes and 48 m location
- Two detectors with two different baseline
  - a solid conclusion on LSND anomaly
- The construction is being progressed rapidly.



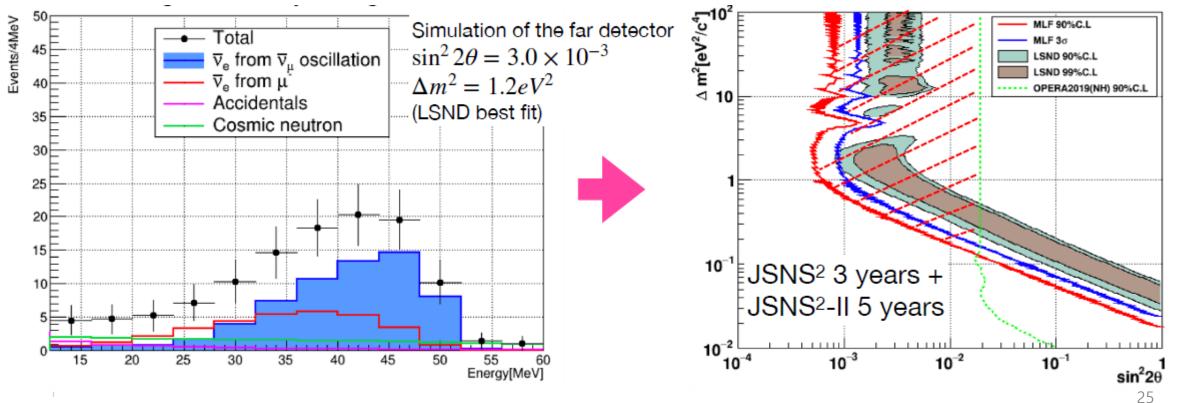
### New far detector

- Almost identical to the existing near detector
  - 37m<sup>3</sup> Gd-LS for the neutrino target
  - 150m<sup>3</sup> no Gd-loaded LS for the veto and gamma catcher.
  - 228 PMTs will be used
- The detector is placed outside of building
  - → Electronics in the "roof space"



# Sensitivity of JSNS<sup>2</sup>-II

- Each background simulation was done based on the JSNS<sup>2</sup> data.
- The sensitivity becomes better in the low  $\Delta m^2$  region.



### **Construction of Stainless Steel Tanks**

- The stainless steel tanks has been constructed.
- The S.S tank was placed at east side of the MLF building.





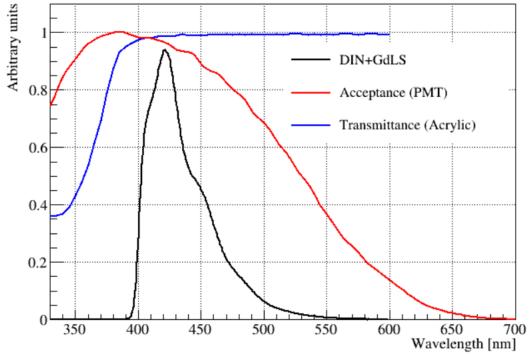
# **Acrylic Tank**

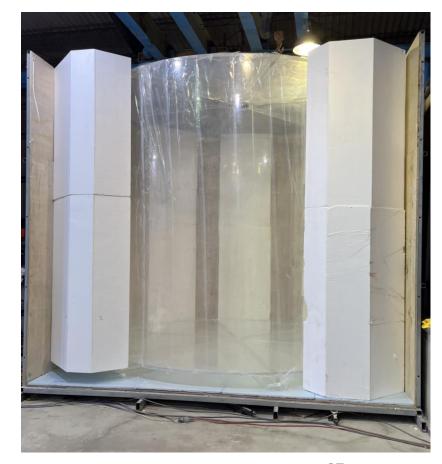
Acrylic tank for the detector target was made by Taiwan company.

The transmittance of the acrylic well covers the wavelength region of scintillation light.

→ ~93% @ 400~600nm

The acrylic tank will be installed to the inner S.S tank.

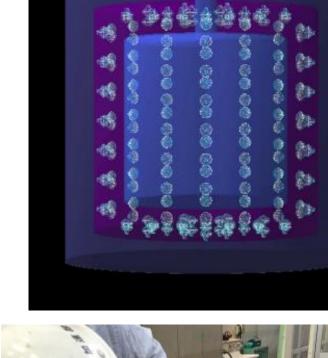




Neutrino Workshop at IFIRSE (2023)

# **PMT** preparation

- 190 PMTs were donated from Double-Chooz
- Dead PMT check was done.
- Our jigs for the far detector were assembled.









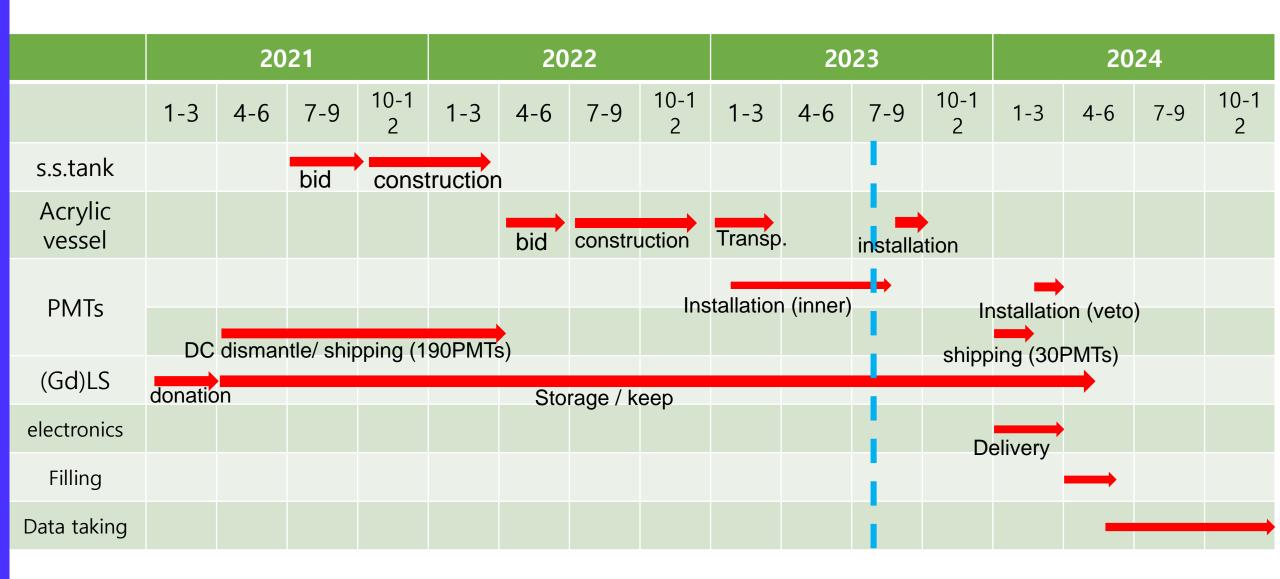


### **PMT** installation

- 142 PMTs were installed.
  - barrel and top lid of the inner detector
- Bottom PMTs will be installed soon
- Calibration system was installed
  - LED
  - temperature sensors



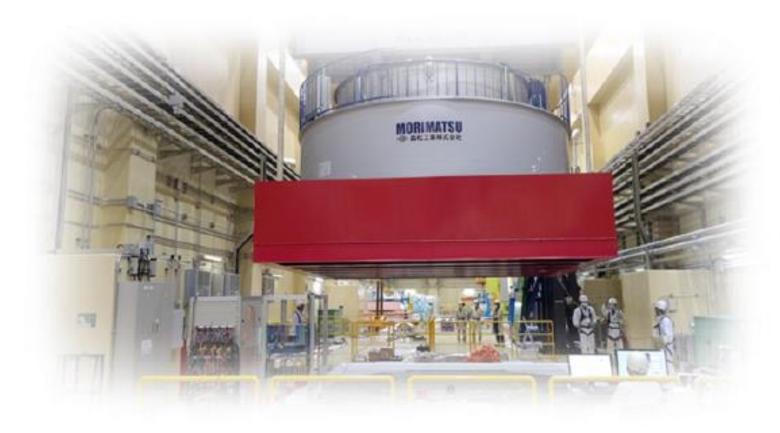
### Construction schedule of JSNS<sup>2</sup>-II



# **Summary**

- There have been 1<sup>st</sup> (2021), 2<sup>nd</sup> (2022) and 3<sup>rd</sup> (2023) long physics runs in JSNS<sup>2</sup>.
- Analyses are ongoing with the data.
  - Has been developing PSD tools
  - Single rates in both prompt and delayed time window were measured
  - Background study is ongoing
- JSNS<sup>2</sup> is working toward the first precise KDAR measurement.
  - clearly see the high purity KDAR signal
- The JSNS<sup>2</sup>-II detector is under construction.
- first data taking on next year

# Thank you for your attention



#### acknowledgements:

- MEXT, JSPS (Japan)
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- DOE, Heising-Simons Foundataion (US)
- · Royal Society (UK)











