



Latest Results of PandaX and Prospects

**Xiang Xiao (Sun Yat-sen University)
on behalf of PandaX Collaboration**

July 23, 2025

@ 21st Rencontres du Vietnam, Neutrino Physics

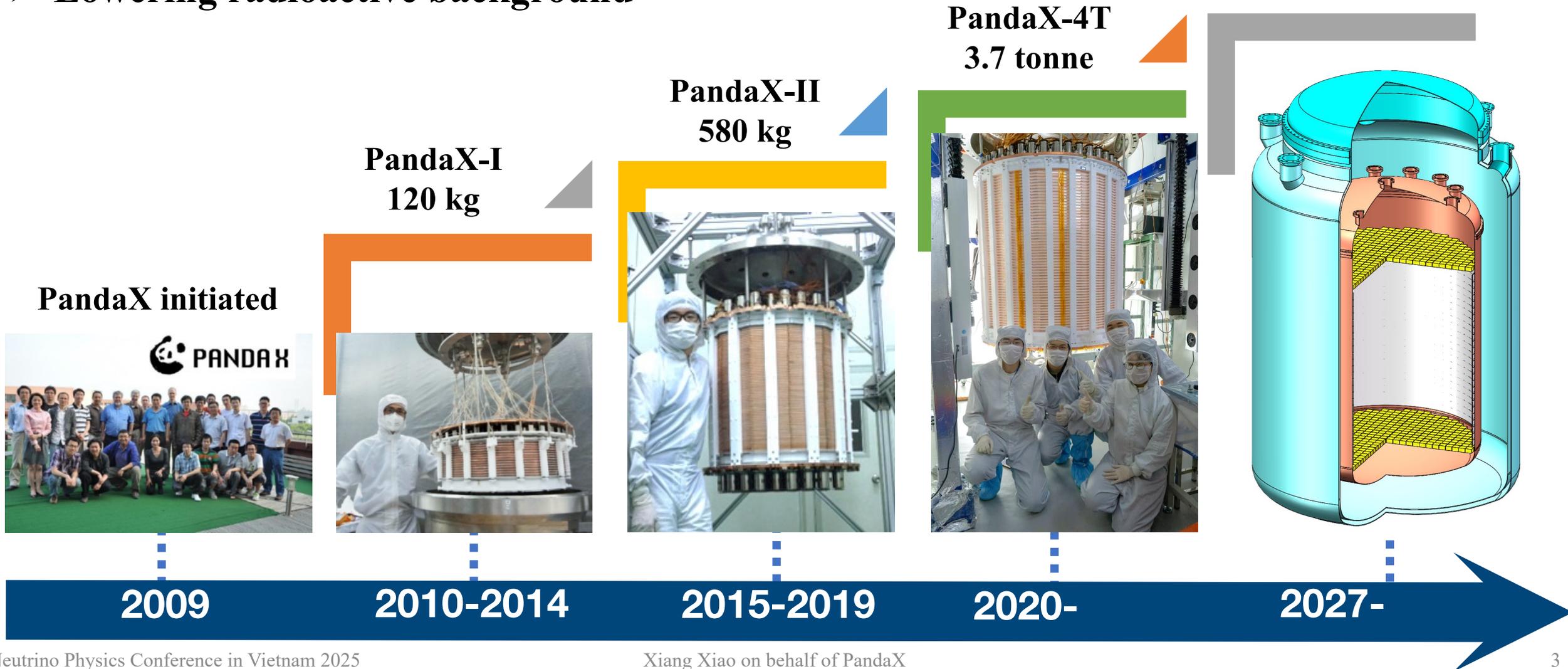
PandaX Collaboration

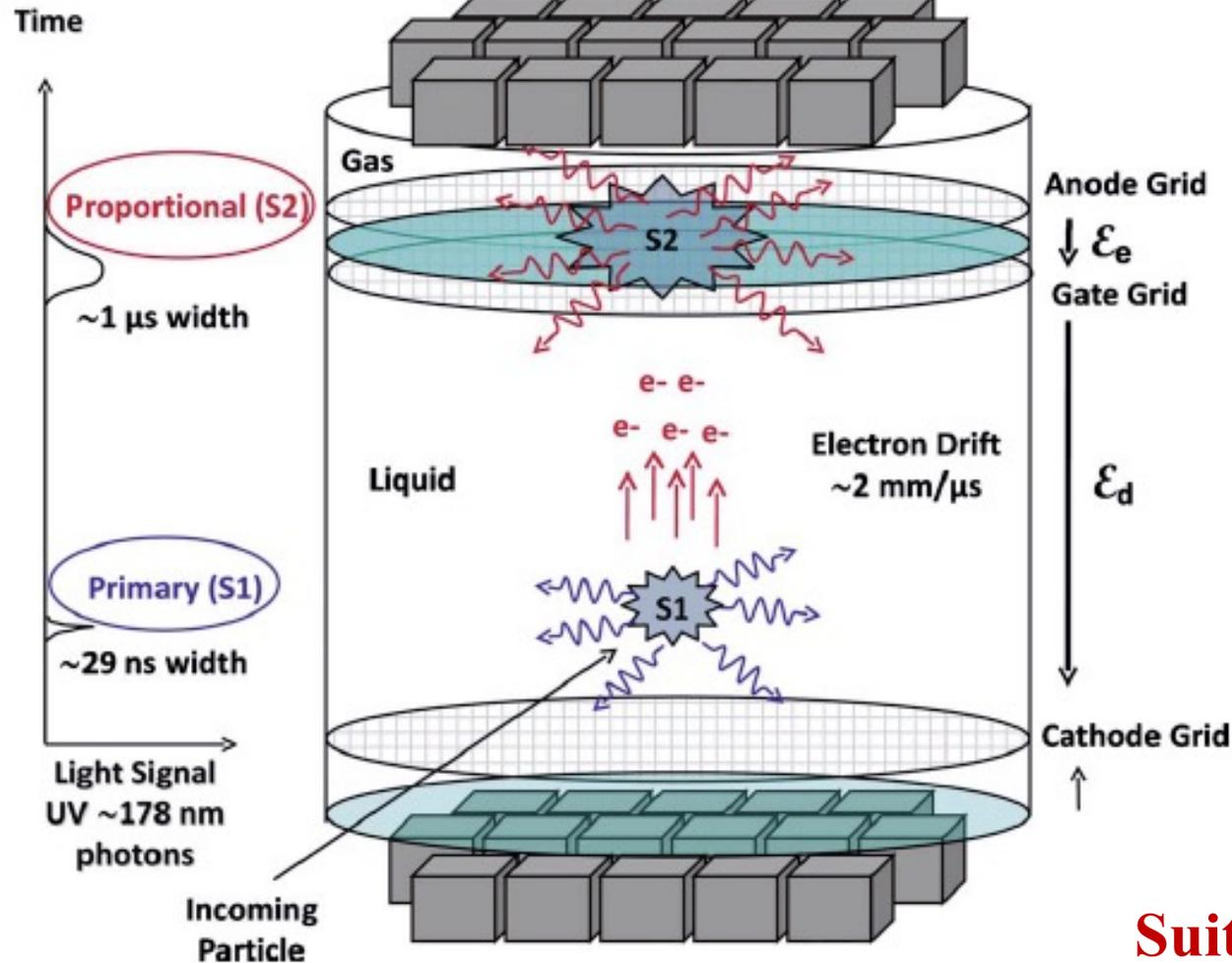
PandaX: Particle and Astrophysical Xenon Experiment



PandaX pathway

- Increasing sensitive target volume
- Lowering radioactive background





➤ 3D position reconstruction

- Fiducialization
- Single-Site (SS) and Multi-Site (MS) discrimination

➤ Particle identification among α , neutron, and γ /electron

➤ Calorimeter from sub keV to a few MeV

➤ Monolithic and scalable



➤ Low background

➤ Large target mass

➤ High detection efficiency

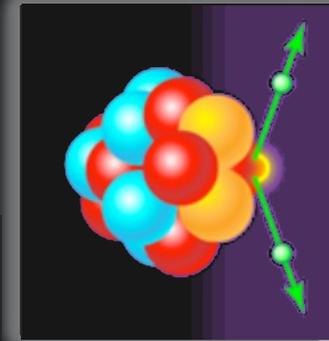


Suitable for detection of dark matter, $0\nu\beta\beta$, and astrophysical neutrinos at the same time!

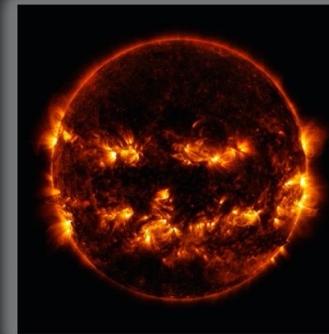
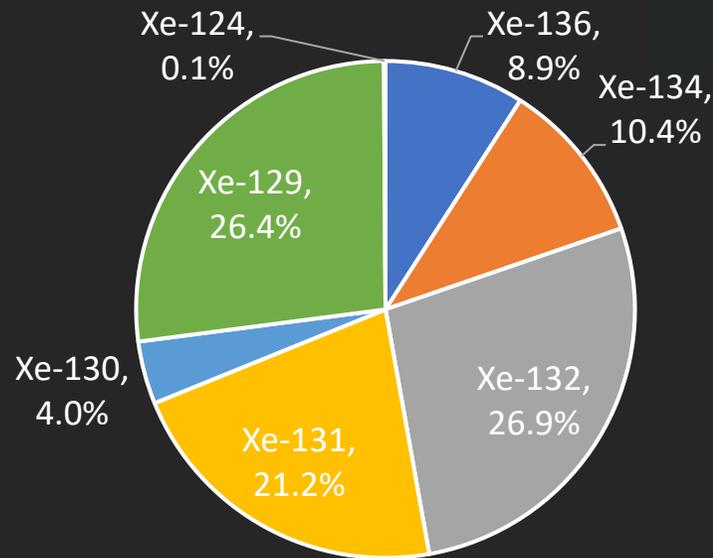
PandaX-4T as multi-physics observatory

- **Wide energy range: sub keV ~ a few MeV**
- **Large volume nature xenon target**
- **Background control and mitigation**

Dark Matter
1 keV – 10 keV



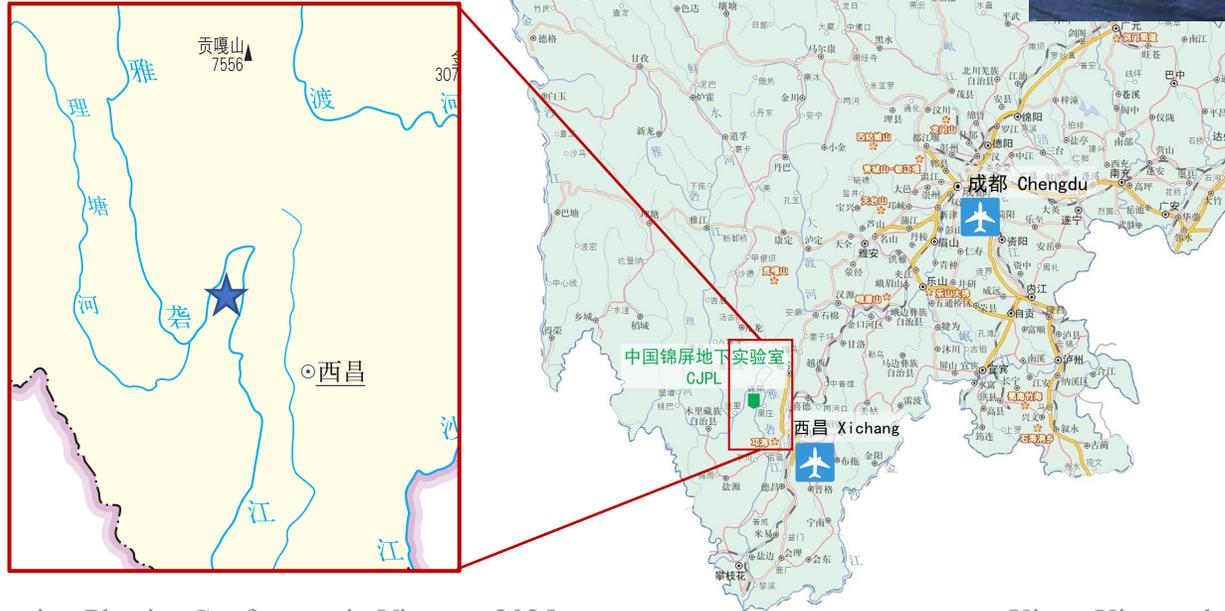
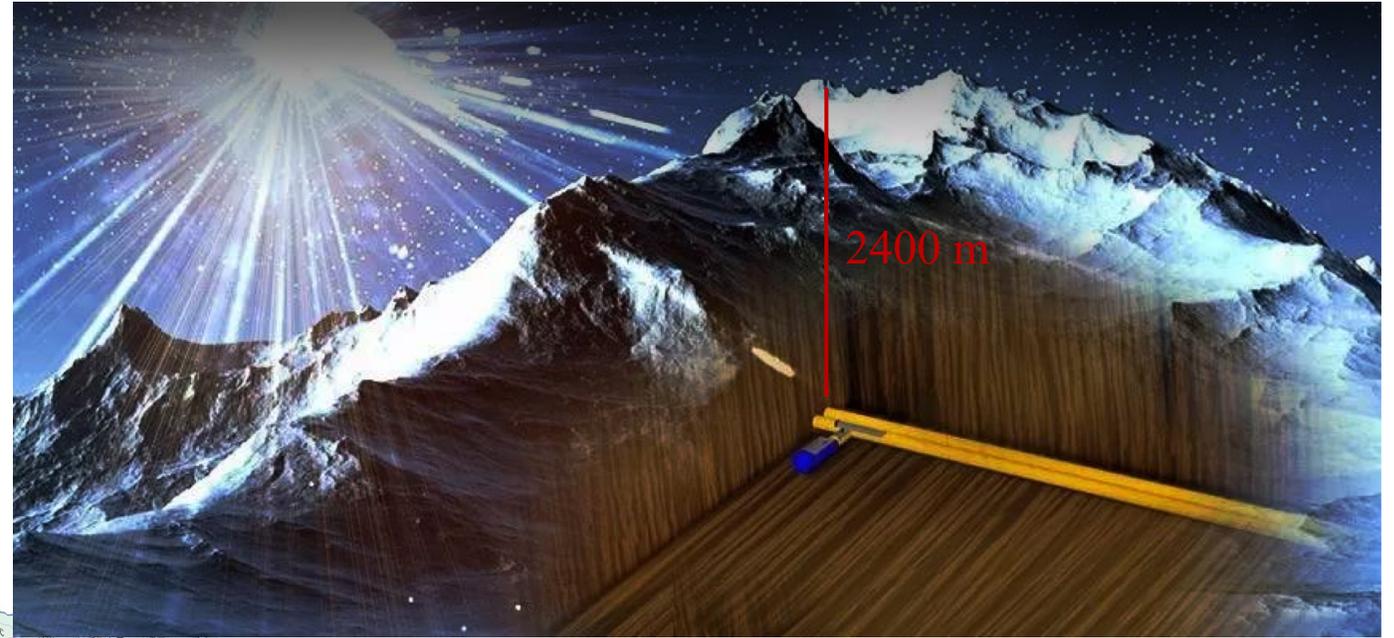
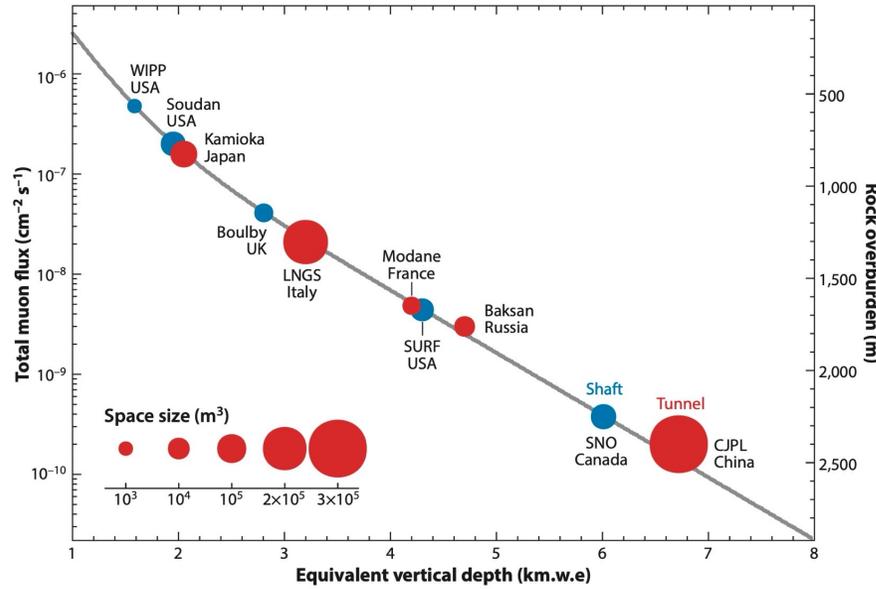
Majorana Neutrino
~2 MeV



Astrophysical Neutrino
~100 keV & sub-keV

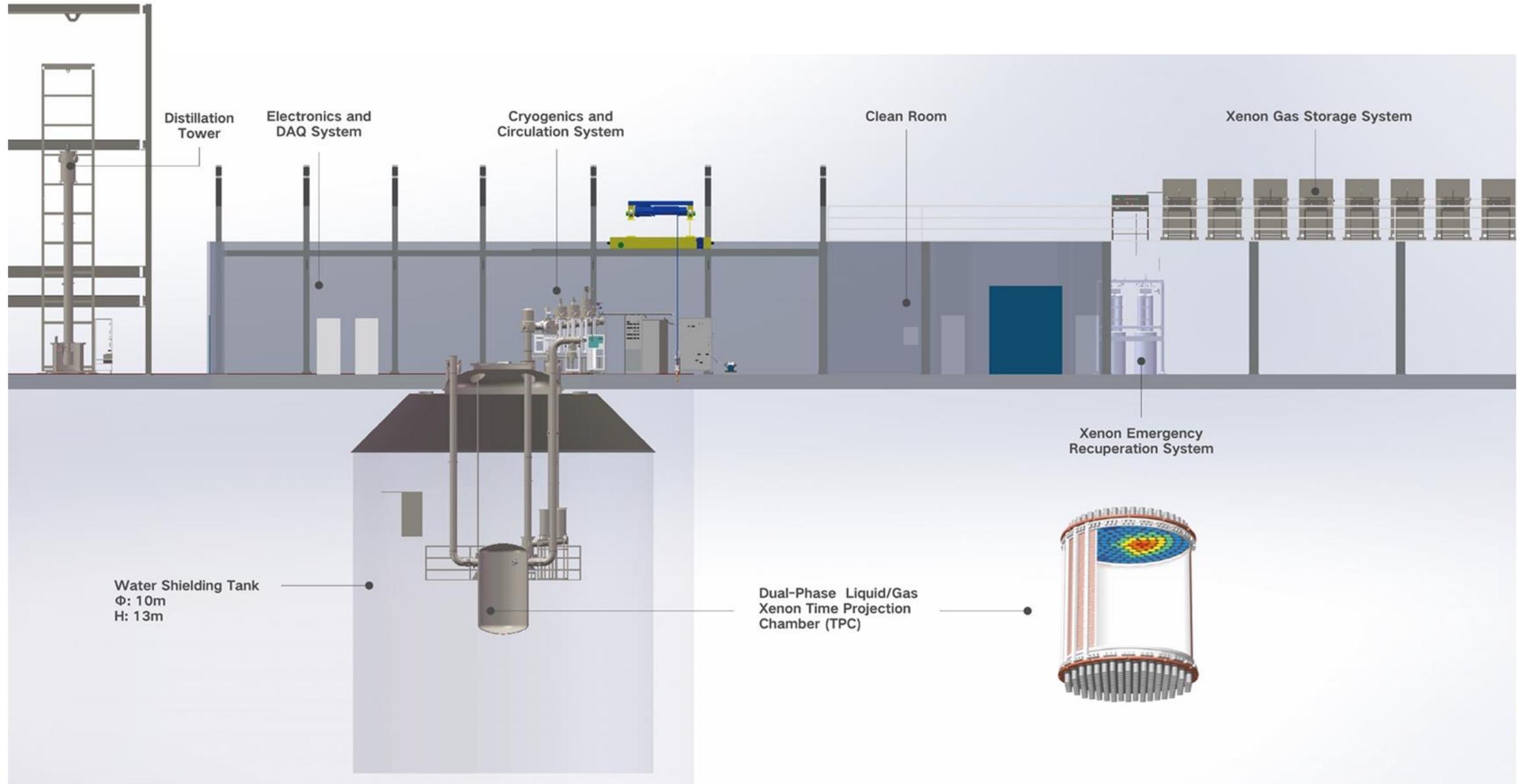


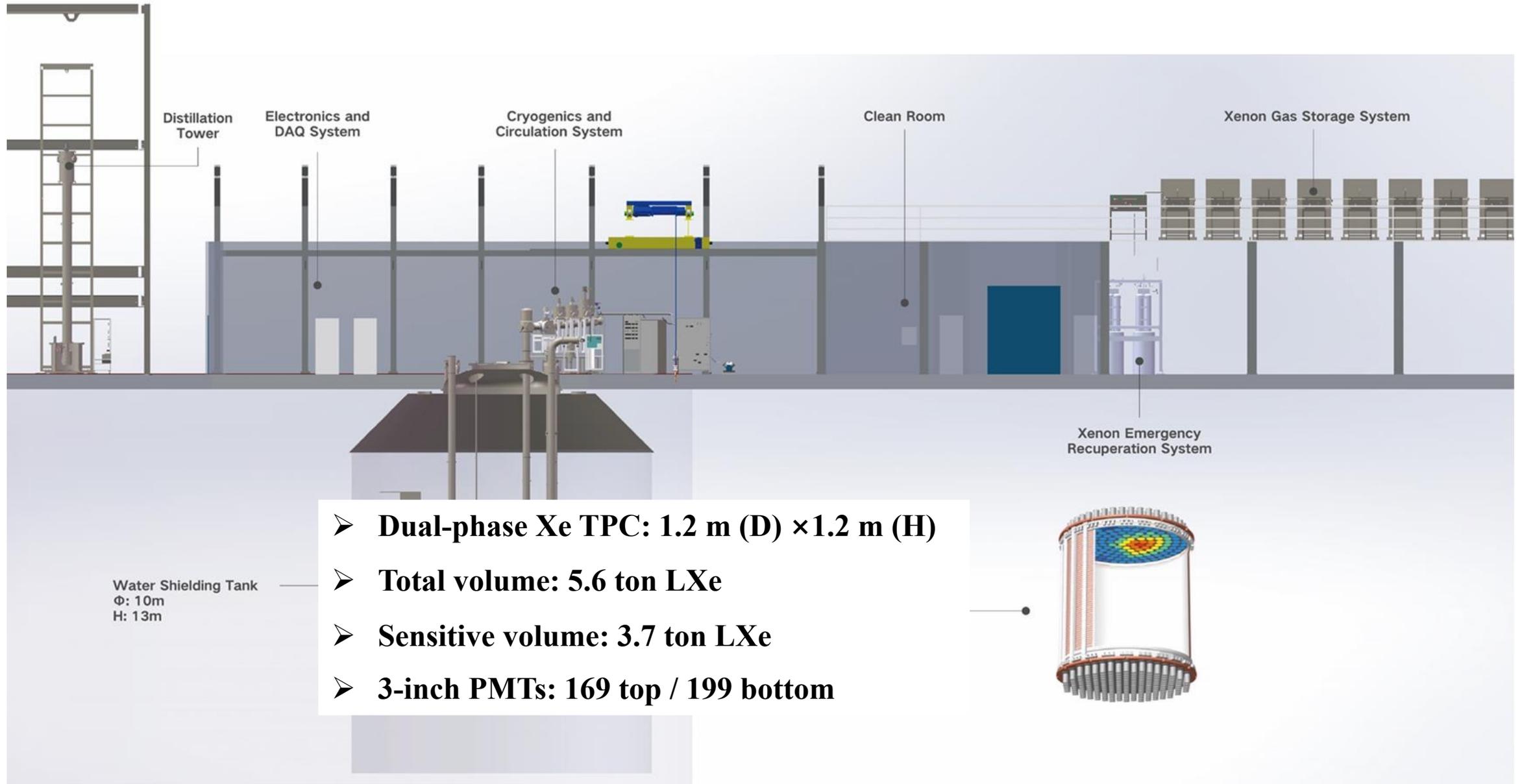
China Jinping Underground Laboratory (CJPL)

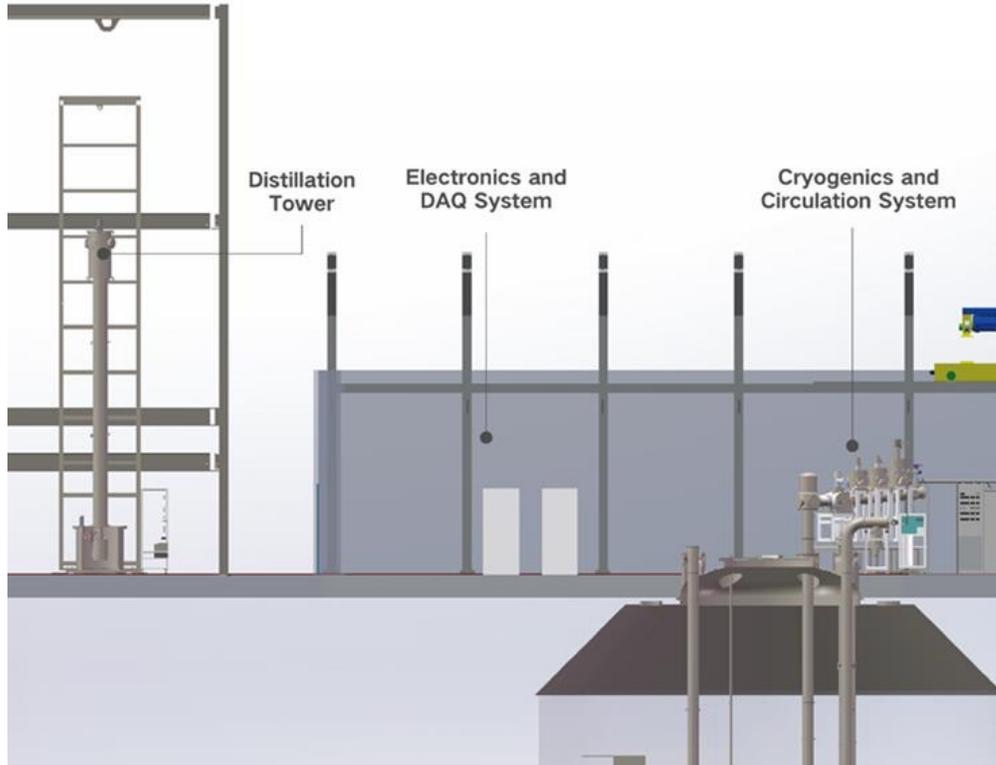


CJPL-II B2 Hall

PandaX-4T @ CJPL-II B2 Hall



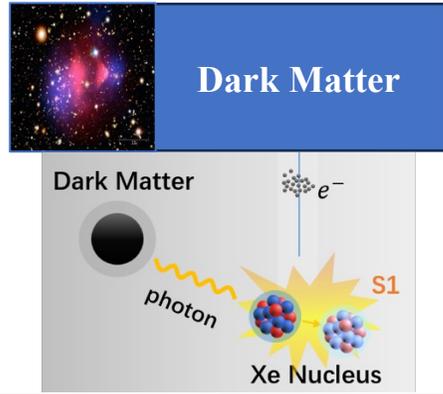




2020/11 – 2021/04	Commissioning run (Run0) 95 days of physics data
2021/07 – 2021/10	Tritium removal xenon distillation, gas flushing, etc
2021/11 – 2022/05	First science run (Run1) 164 days of physics data
2022/09 – 2023/12	CJPL-II B2 hall construction xenon recuperation, detector upgrade
Resume data-taking (Run2)	

- **Dual-phase Xe TPC: 1.2 m (D) × 1.2 m (H)**
- **Total volume: 5.6 ton LXe**
- **Sensitive volume: 3.7 ton LXe**
- **3-inch PMTs: 169 top / 199 bottom**

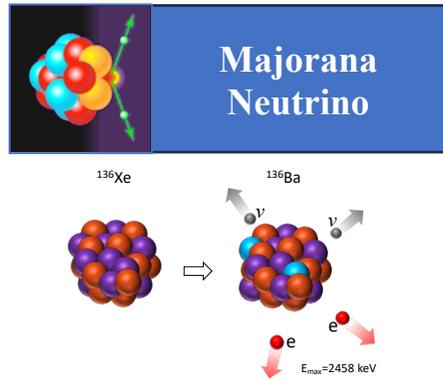




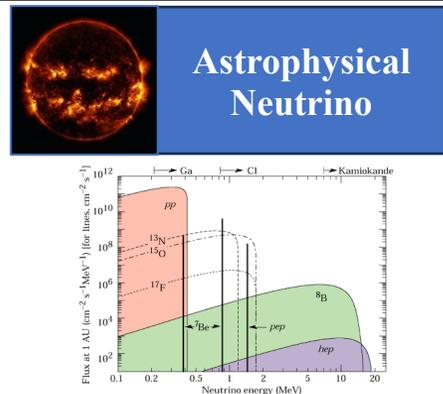
- **Run0+Run1 combined 1.54 tonne-year exposure**
 - Leading constraints for WIMP mass above 100 GeV
 - Leading constraints for ALP and dark photon 150 keV - 1 MeV
 - Competitive constraints on axion, neutrino magnetic moment, etc
- **Limits on the luminance of dark matter**
 - First constraints on DM charge radius

[PRL 134, 011805 \(2025\)](#)
[PRL 134, 071004 \(2025\)](#)
[PRL 134, 041001 \(2025\)](#)

[Nature 618, 47-50 \(2023\)](#)



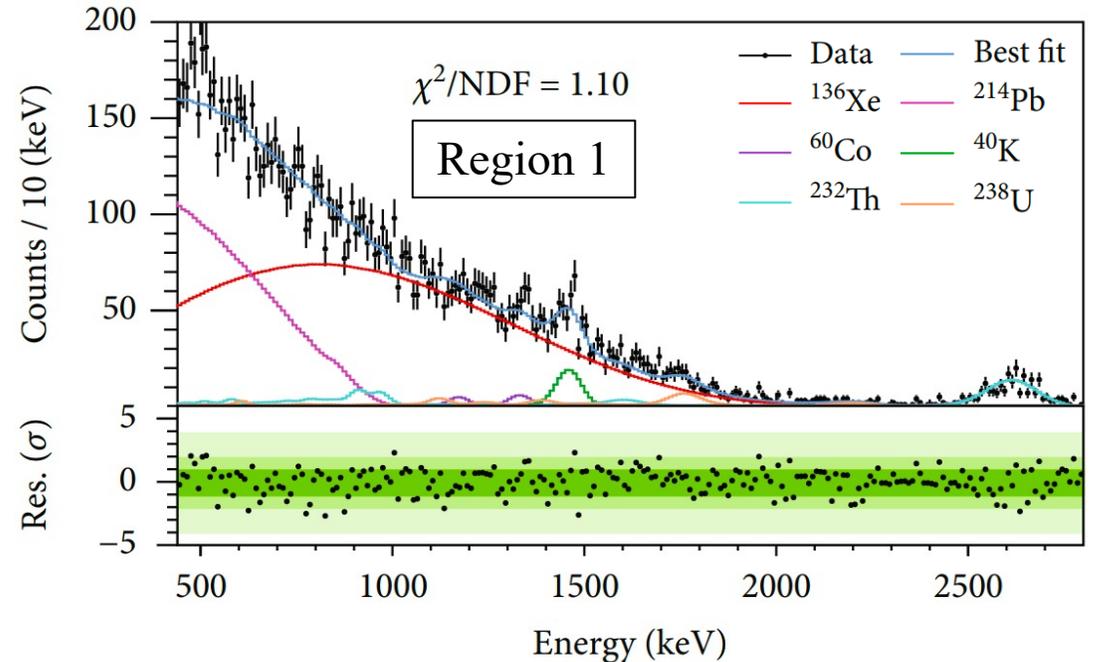
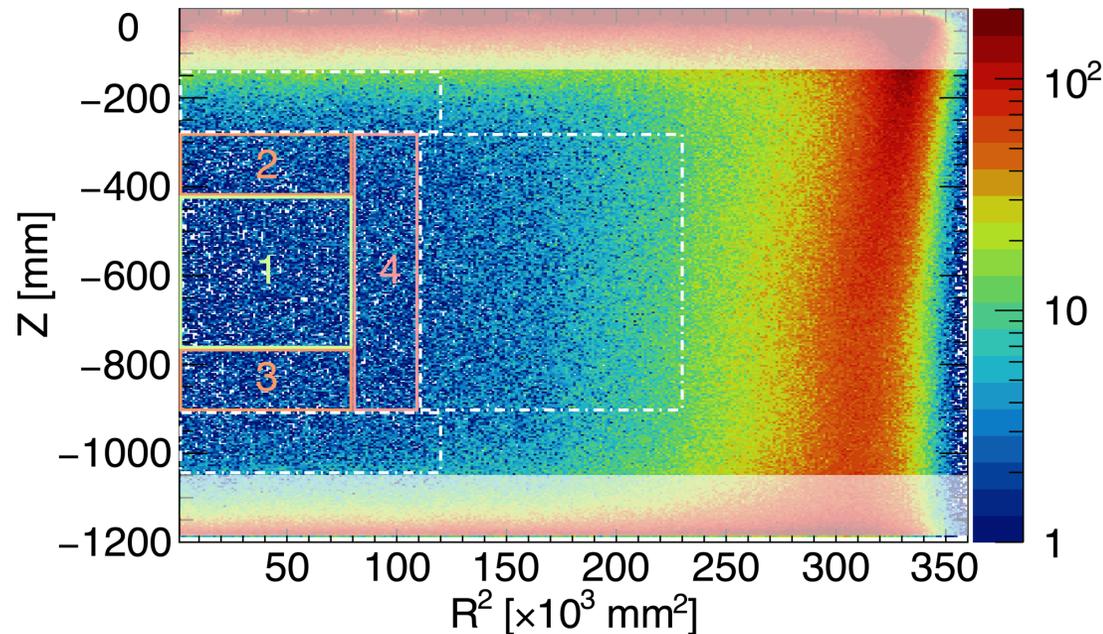
- **First $^{136}\text{Xe } 2\nu\beta\beta$ half-life precise measurement from natural xenon detector**
 - $T_{1/2} = 2.27 \pm 0.03(\text{stat.}) \pm 0.10(\text{syst.}) \times 10^{21}$ yr [Research 2022, 9798721 \(2022\)](#)
- **Leading $^{136}\text{Xe } 0\nu\beta\beta$ half-life constraints from natural xenon detector**
 - 90% CL limits on half-life $T_{1/2} > 2.1 \times 10^{24}$ yr [Science Bulletin 70, 1779-1785 \(2025\)](#)
- **Leading $^{134}\text{Xe } 2\nu\beta\beta$ and $0\nu\beta\beta$ half-life constraints**
 - 90% CL limits on half-life $T_{1/2}^{2\nu\beta\beta} > 2.8 \times 10^{22}$ yr and $T_{1/2}^{0\nu\beta\beta} > 3.0 \times 10^{23}$ yr [PRL 132, 152402 \(2024\)](#)



- **First indication of solar ^8B neutrinos through CEvNS** [PRL 133, 191001 \(2024\)](#)
 - 2.64 σ significance
 - Two ROI regions: paired (3.5 ± 1.3 events) and S2-only (75 ± 28 events)
- **First attempt to detect solar pp neutrinos in liquid xenon detector** [CPC 48, 091001 \(2024\)](#)
 - Flux: $8.0 \pm 3.9(\text{stat.}) \pm 10.0(\text{syst.}) \times 10^{10} \text{ s}^{-1} \text{ cm}^{-2}$

- Dedicated data analysis pipeline is developed for $O(10\text{ keV}) - O(\text{MeV})$ energy range
- Segmented FV to partially include position information
- Binned Poisson likelihood fitting on SS energy spectrum performed simultaneously in four regions
- Outer regions to check background model, consistent at $\sim 1\%$

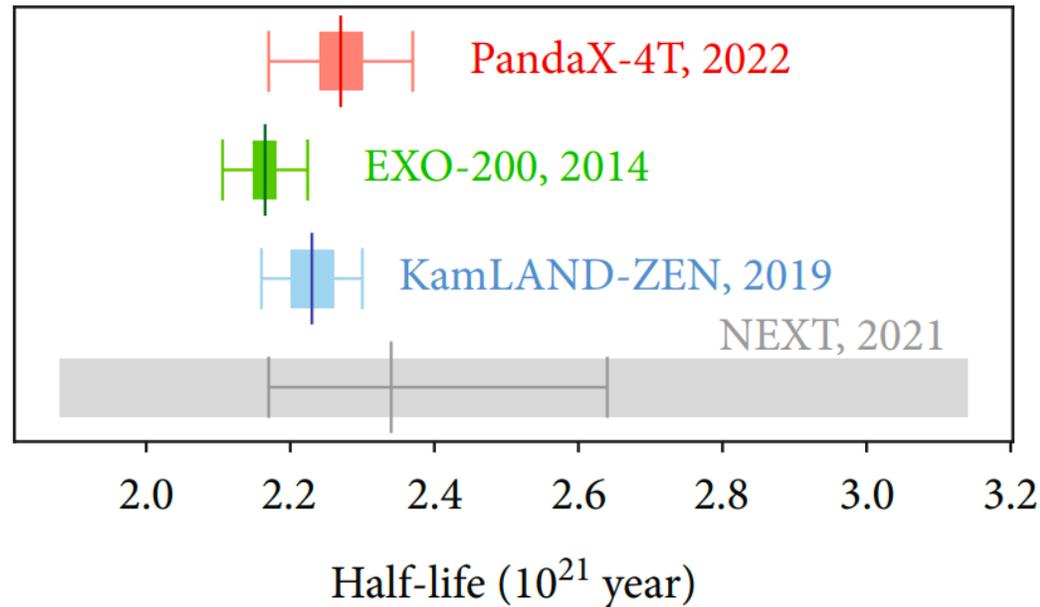
$$L = \prod_{i=1}^{N_R} \prod_{j=1}^{N_{\text{bins}}} \frac{(N_{ij})^{N_{ij}^{\text{obs}}}}{N_{ij}^{\text{obs}}!} e^{-N_{ij}} \prod_{k=1}^{N_{\text{bkgs}}} \frac{1}{\sqrt{2\pi}\sigma_k} e^{-\frac{1}{2}\left(\frac{\eta_k}{\sigma_k}\right)^2}, \quad N_{ij} = n_{\text{Xe}} S_{ij}^{\text{Xe}} + \sum_{k=1}^{N_{\text{bkgs}}} (1 + \eta_k) n_{ij}^k B_{ij}^k$$



^{136}Xe $2\nu\beta\beta$ half-life and background estimation

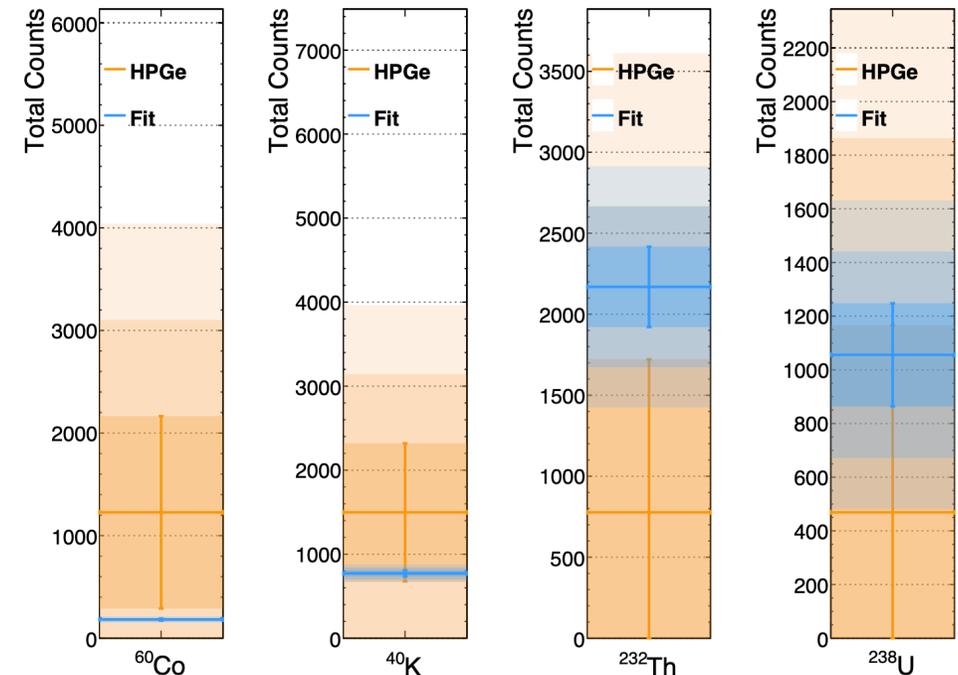
^{136}Xe $2\nu\beta\beta$ half-life measured as: $2.27 \pm 0.03(\text{stat.}) \pm 0.10(\text{syst.}) \times 10^{21}$ year with Run0 data

- First such measurement from natural xenon detectors
- Comparable precision with dedicated ^{136}Xe -enriched $0\nu\beta\beta$ experiments
- Much lower analysis threshold compared with previous measurements
- “*in-situ*” material background fitting results compatible and more precise than HPGe assay



Research 2022, 9798721 (2022)

Material, “Side” category

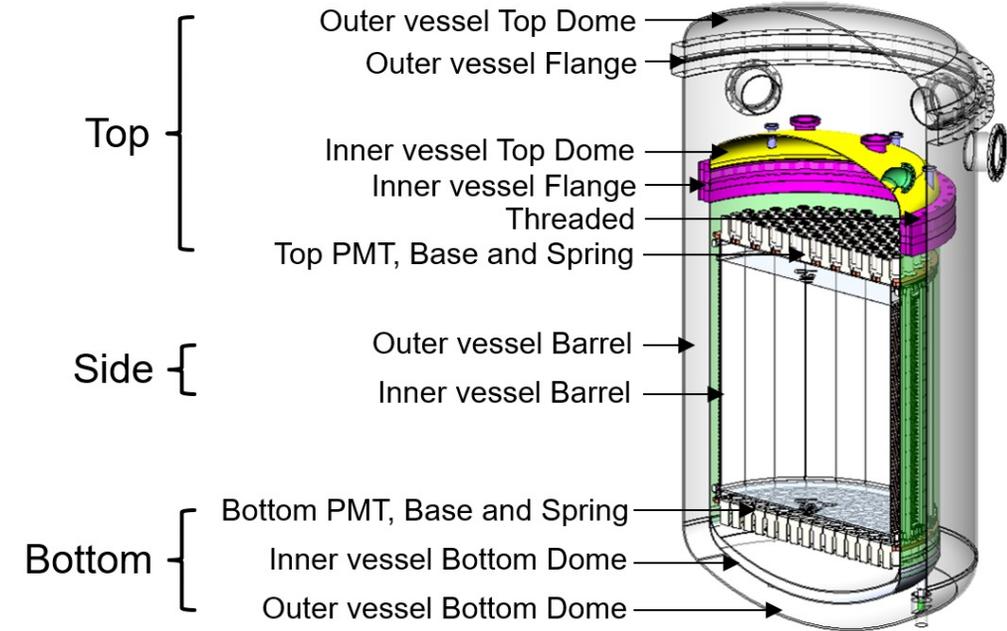
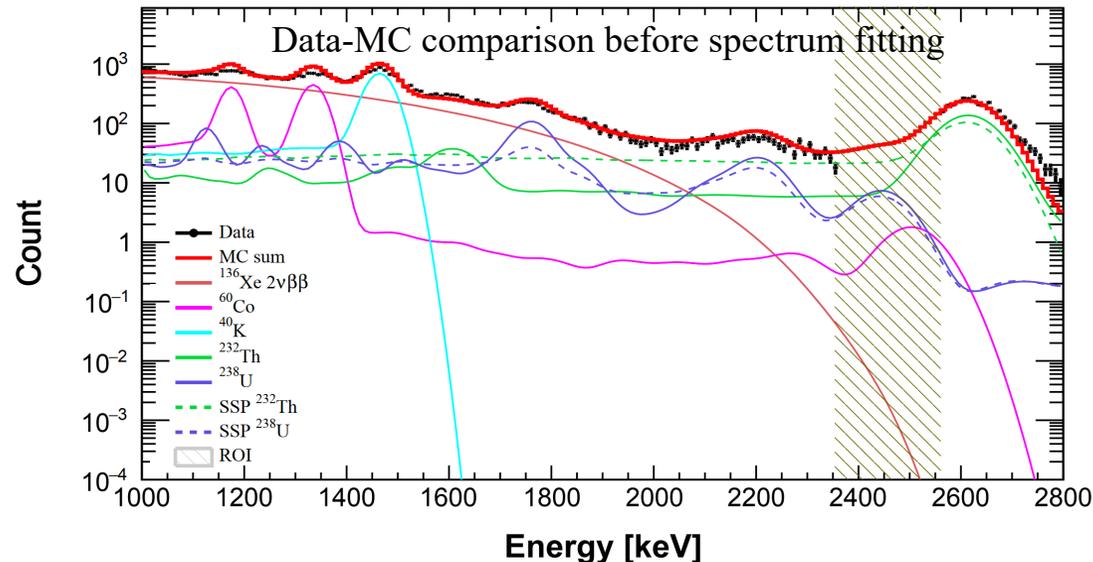


Blind analysis for ^{136}Xe $0\nu\beta\beta$ search

Unified and optimized data reconstruction for Run0+Run1

Background components:

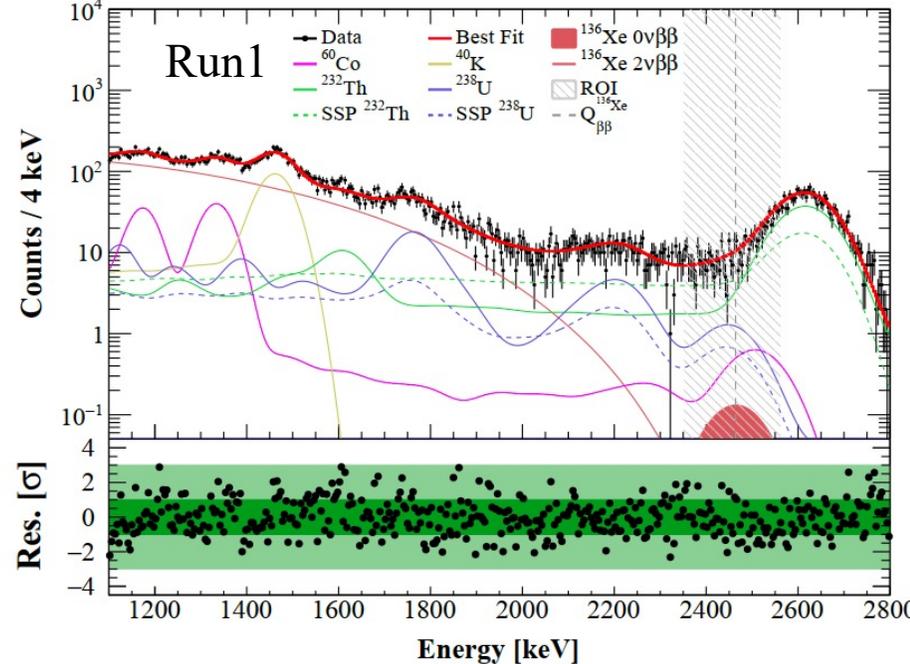
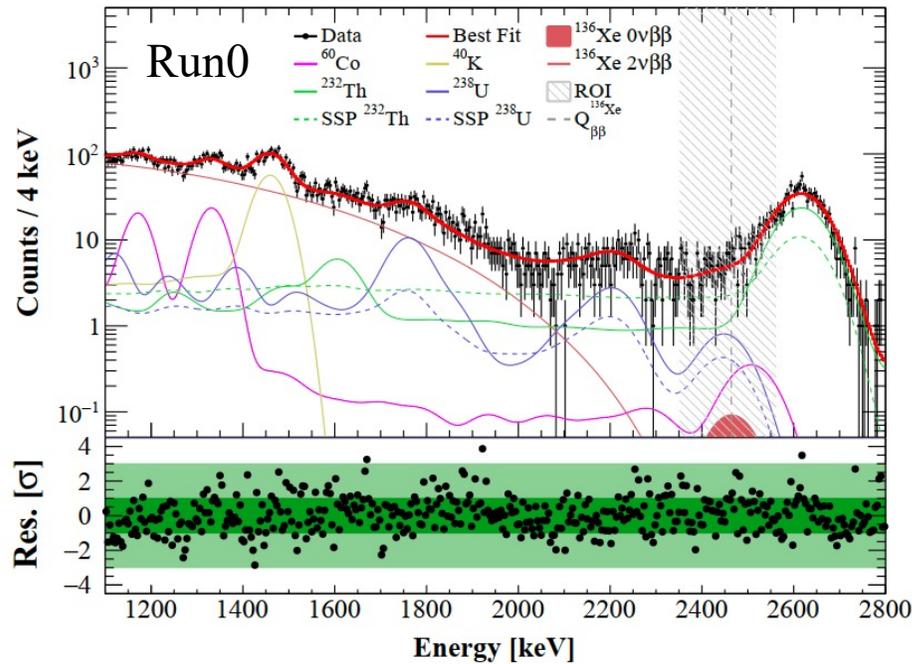
- ^{136}Xe $2\nu\beta\beta$: from ^{136}Xe $2\nu\beta\beta$ half-life measurement)
- Detector material ^{60}Co , ^{40}K , ^{232}Th , and ^{238}U : from HPGe material assay, and grouped into top, side, and bottom parts
- Stainless steel platform (SSP): ^{232}Th , ^{238}U (from MS fitting)



Other background components are checked:

- Residual ^{214}Bi in TPC -> negligible
- Gammas of ^{214}Bi from LXe skin region -> negligible
- 2.5 MeV peak from ^{60}Co cascade gammas -> well modelled

Unblinded fitting and results of $^{136}\text{Xe } 0\nu\beta\beta$



^{136}Xe exposure of
Run0+Run1: 44.6 kg-yr

Energy resolution
@ 2615 keV in FV:
➤ 2.0% in Run0
➤ 2.3% in Run1

$^{136}\text{Xe } 0\nu\beta\beta$ event rate is fitted to be $14 \pm 37 \text{ t}^{-1}\text{yr}^{-1}$,
with a p-value of 0.49 for null results.

$$T_{1/2}^{0\nu\beta\beta} > 2.1 \times 10^{24} \text{ yr at 90\% C.L.}$$

Upward fluctuation, the limit is consistent with
the median sensitivity within 1.1σ .

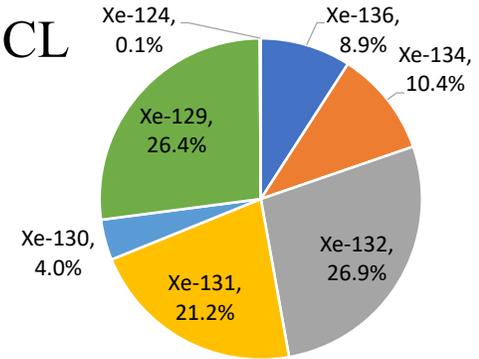
$$\langle m_{\beta\beta} \rangle = (0.4 - 1.6) \text{ eV}/c^2$$

- **Best results among natural xenon detectors so far**
- Improvement to our previous PandaX-II results by an order of magnitude and to the XENON1T results by a factor of 1.8
- Demonstrating the potential of $^{136}\text{Xe } 0\nu\beta\beta$ search with next-generation multi-ten-tonne natural xenon detectors

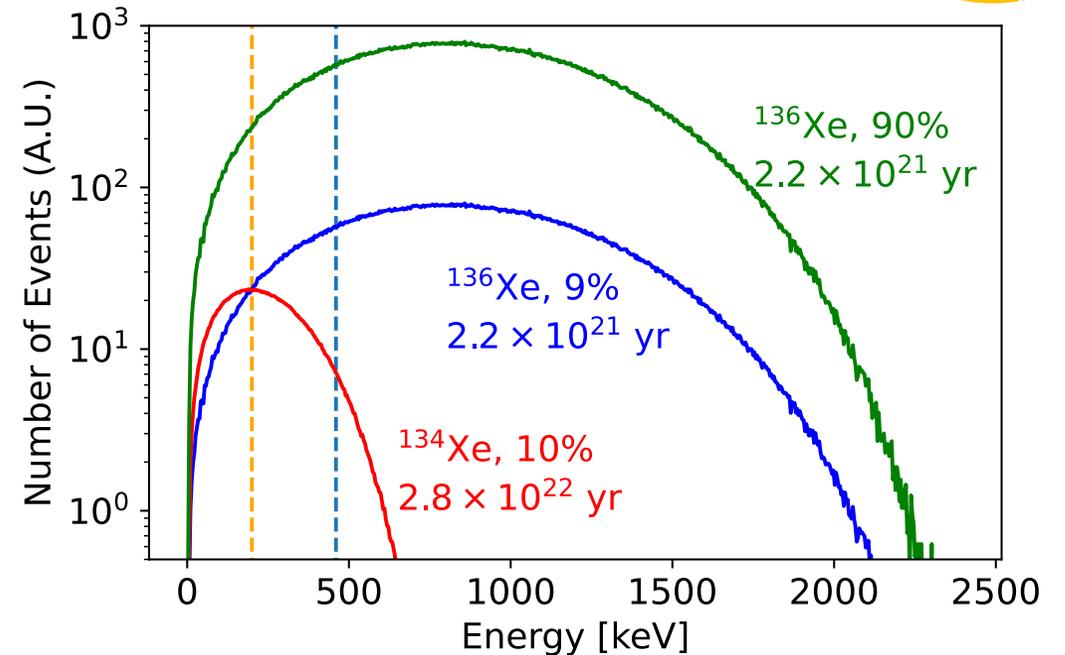
Science Bulletin 70, 1779-1785 (2025)

A blessing: ^{134}Xe $2\nu\beta\beta/0\nu\beta\beta$ searches

- $Q=826$ keV; Half-life from theoretical predictions: 10^{24} - 10^{25} yr; Never been observed yet
- Previous $2\nu\beta\beta$ ($0\nu\beta\beta$) half-life limit from EXO-200: $T > 8.7 \times 10^{20}$ yr (1.1×10^{23} yr) at 90% CL
- **PandaX-4T: more ^{134}Xe ; much less ^{136}Xe ; wider energy range; discovery possible**

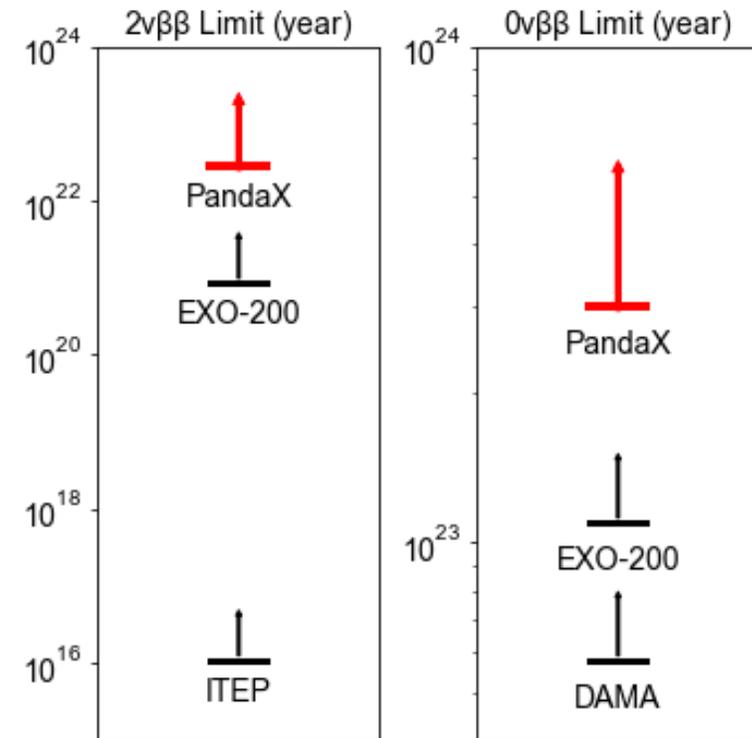
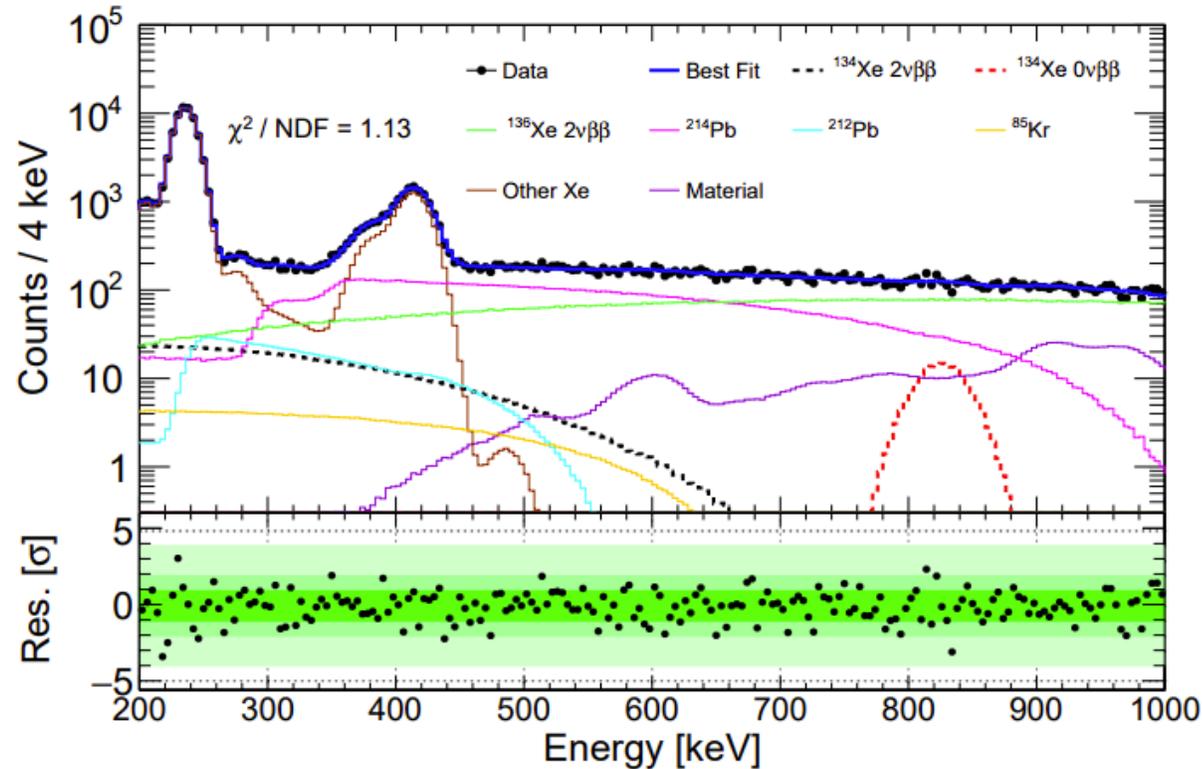


	PandaX-4T	EXO-200
^{134}Xe mass	68.7 kg	18.1 kg
^{136}Xe abundance	8.90%	81%
Analysis threshold	200 keV	460 keV
Live Time	94.9 days	600 days



^{134}Xe $2\nu\beta\beta/0\nu\beta\beta$ searches with Run0

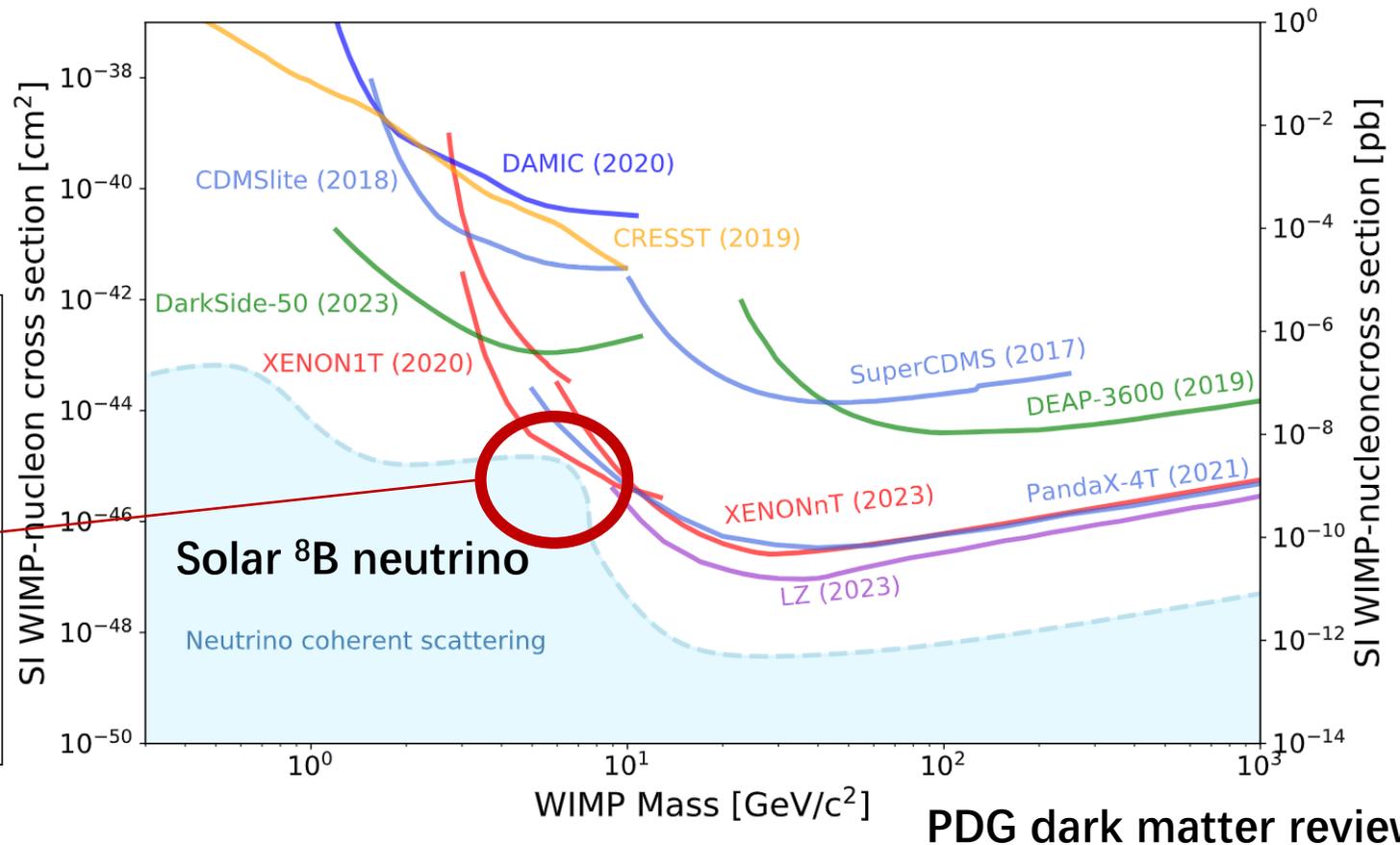
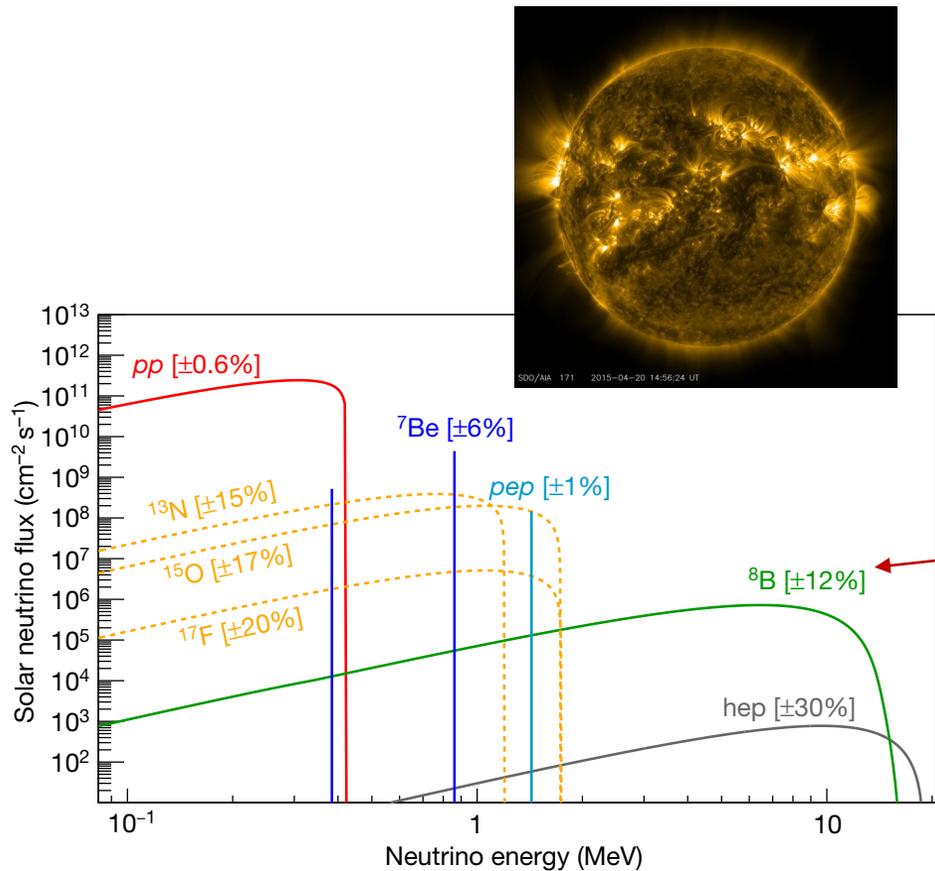
- Simultaneous fit for ^{134}Xe $2\nu\beta\beta$ and $0\nu\beta\beta$ with Run0 data
- Final counts of $2\nu\beta\beta$ and $0\nu\beta\beta$: $10 \pm 269(\text{stat.}) \pm 680(\text{syst.})$ and $105 \pm 48(\text{stat.}) \pm 38(\text{syst.})$
- 90% CL lower limits on the half-life: $T_{1/2}^{2\nu\beta\beta} > 2.8 \times 10^{22}$ yr and $T_{1/2}^{0\nu\beta\beta} > 3.0 \times 10^{23}$ yr



PRL 132, 152402 (2024)

Dark matter search touching neutrino fog...

- In the past decades, the DM search sensitivity improved several orders of magnitude, touching the neutrino fog at ~ 6 GeV DM mass
- Detecting solar ^8B neutrinos through CEvNS becomes feasible with a multi-tonne LXe detector



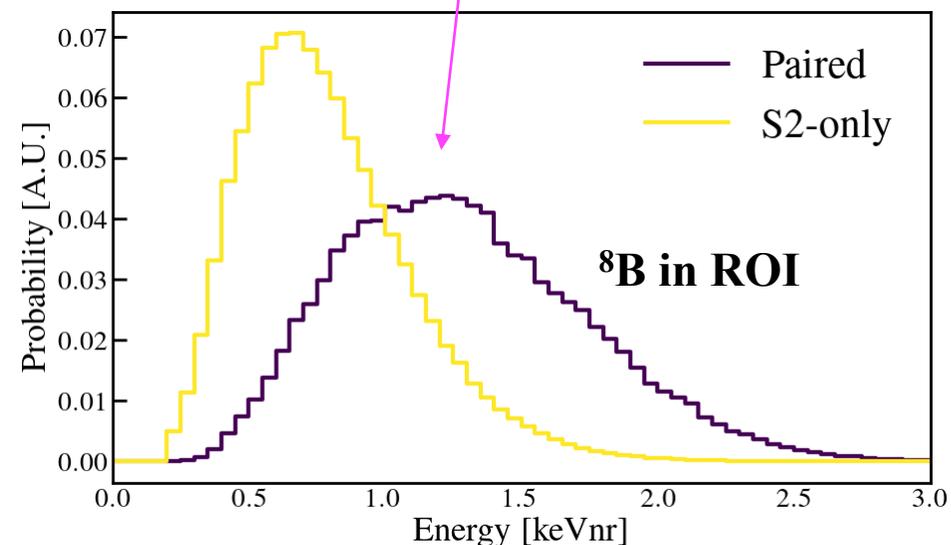
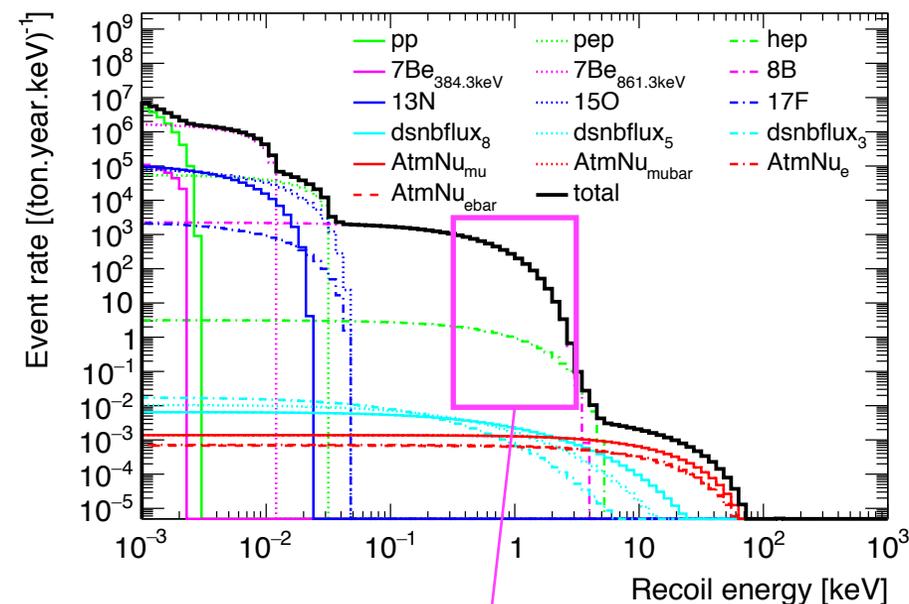
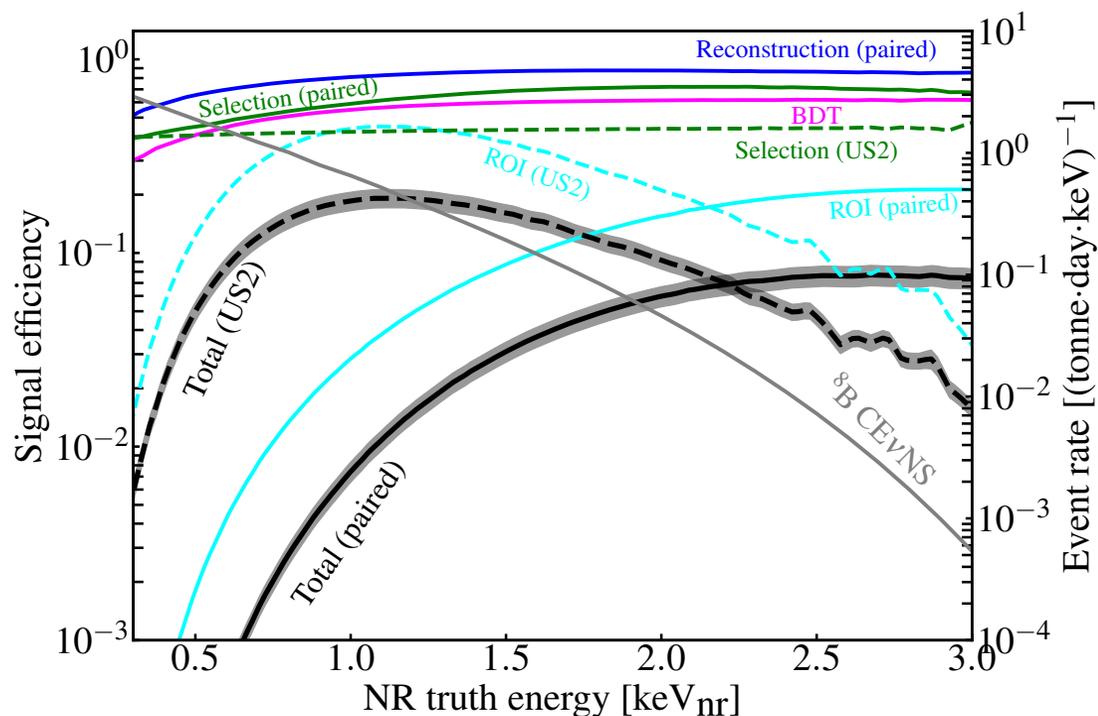
PDG dark matter review

Further lowering energy threshold

Low threshold detection mode:

- low threshold S1+S2 paired ROI (paired)
- Ionization S2-only ROI (US2)

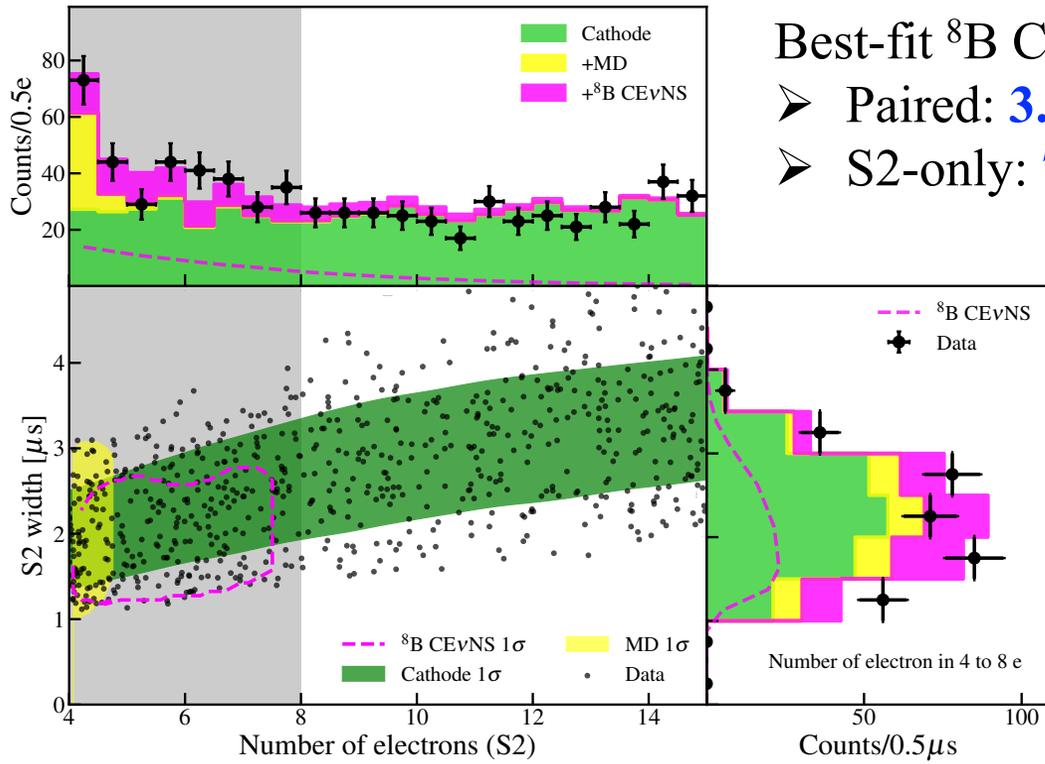
Challenges: background control and signal model at ultralow energy



First indication of solar ^8B CEvNS signal

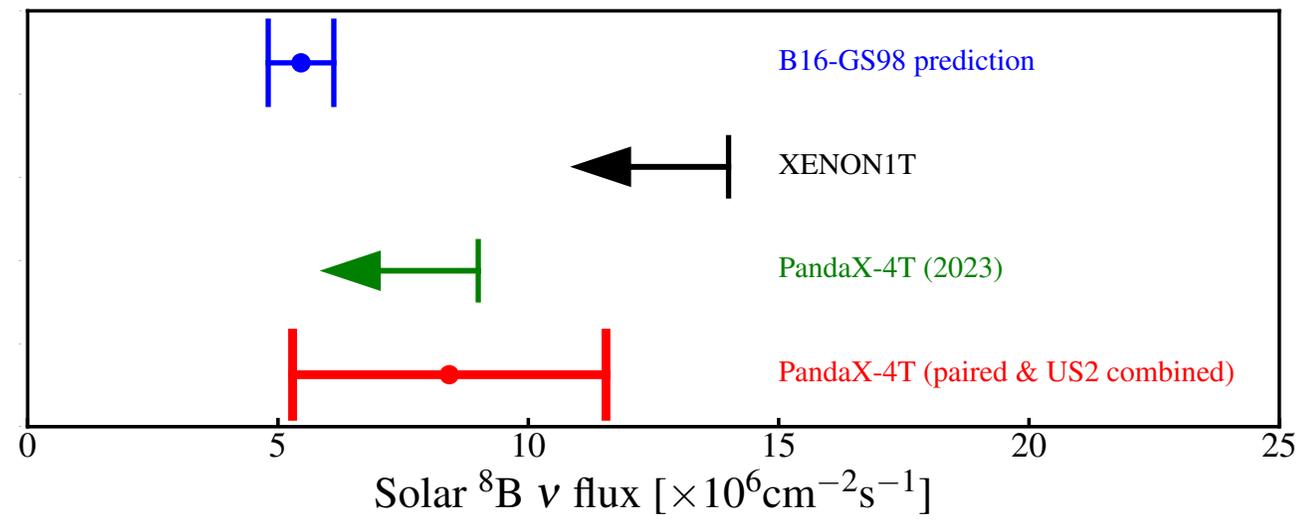
Run0 + Run1 dataset:

- Significance: 2.64σ
- Best-fit ^8B neutrino flux: $(8.4 \pm 3.1) \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$



Best-fit ^8B CEvNS signal:

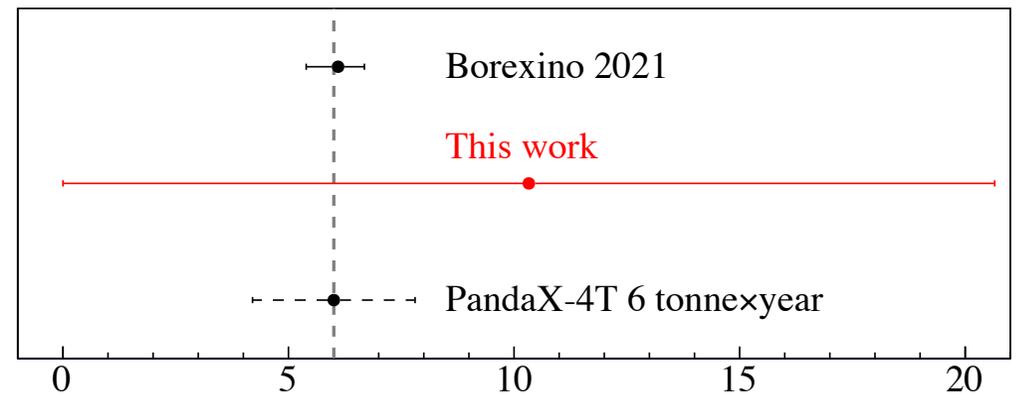
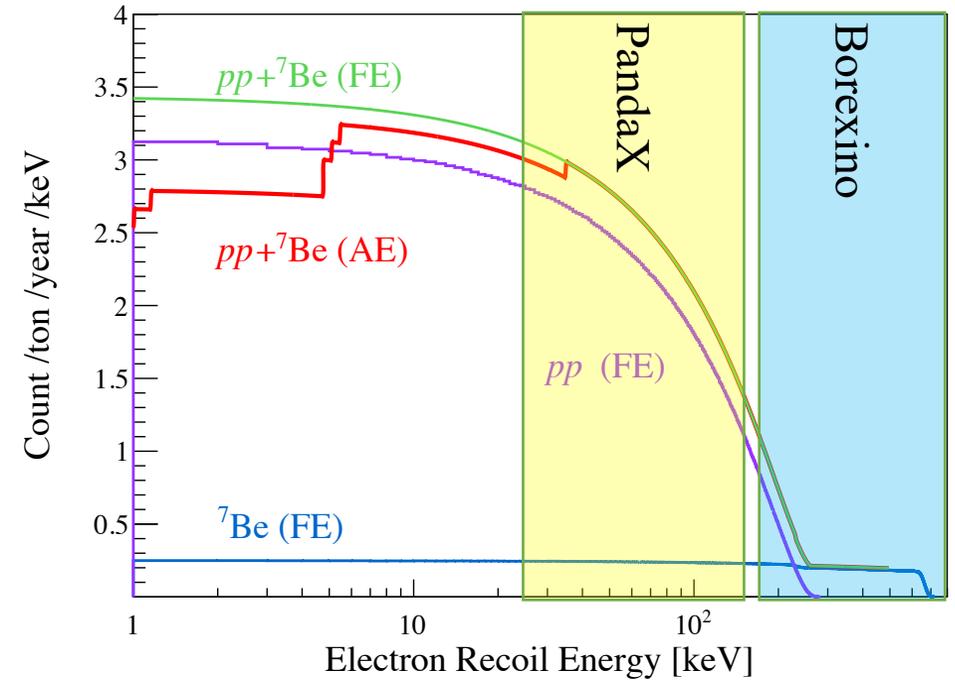
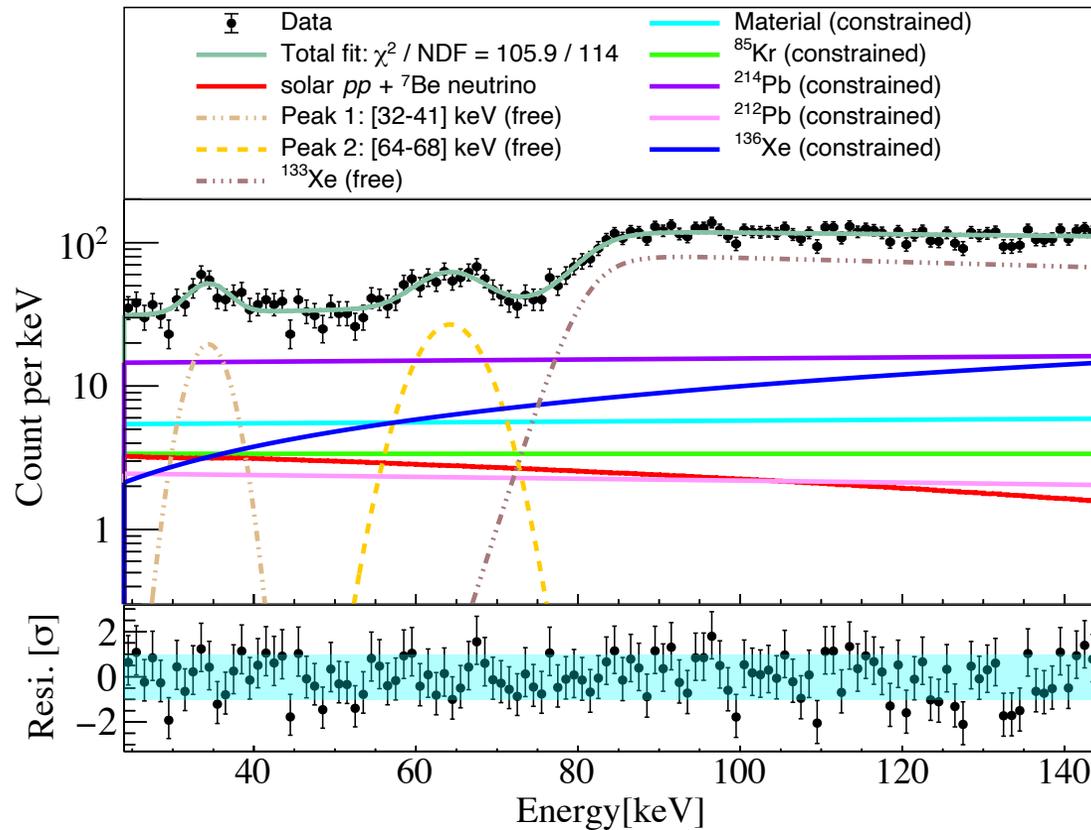
- Paired: 3.5 ± 1.3 evts
- S2-only: 75 ± 28 evts



PRL 133, 191001 (2024)

An attempt: solar pp neutrino results with Run0

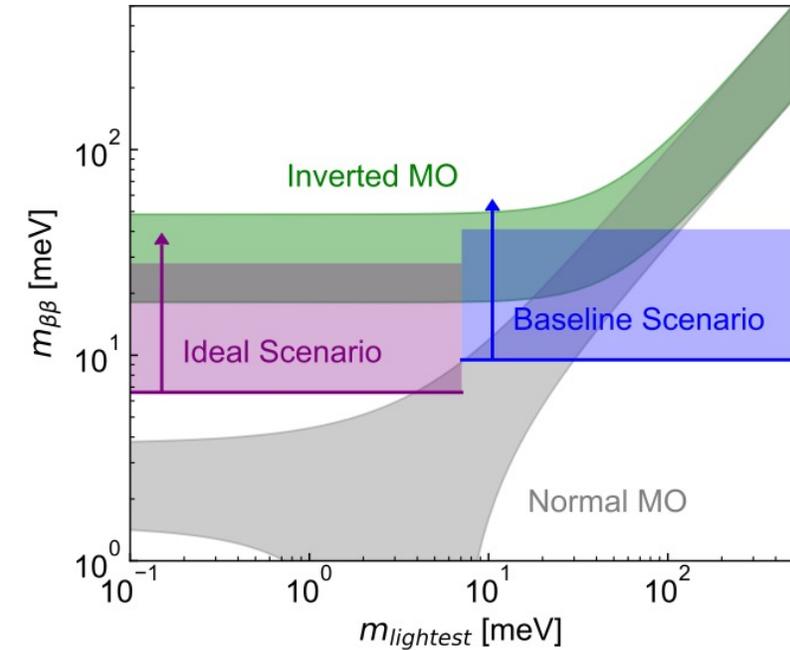
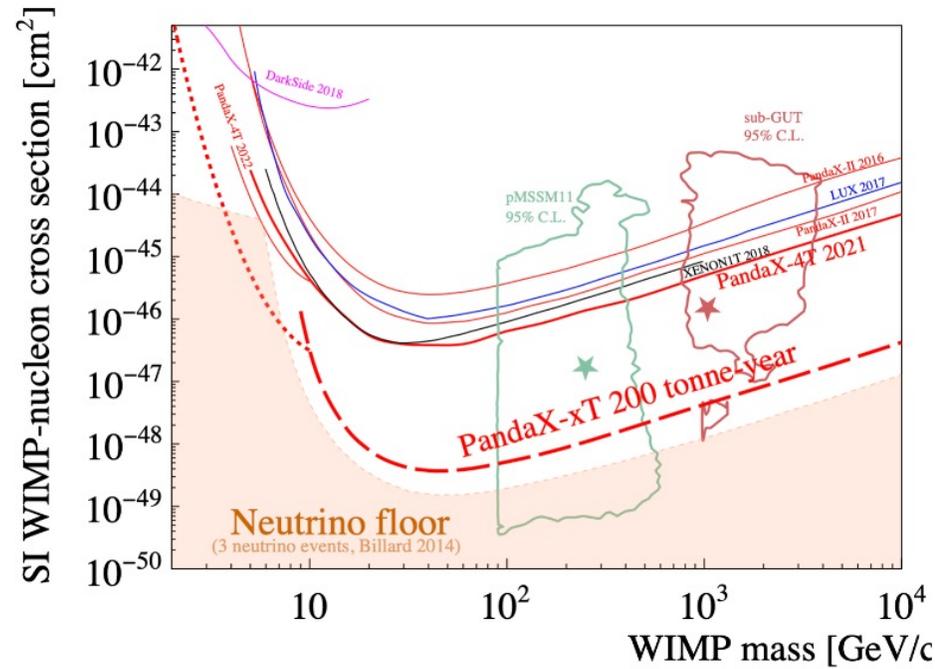
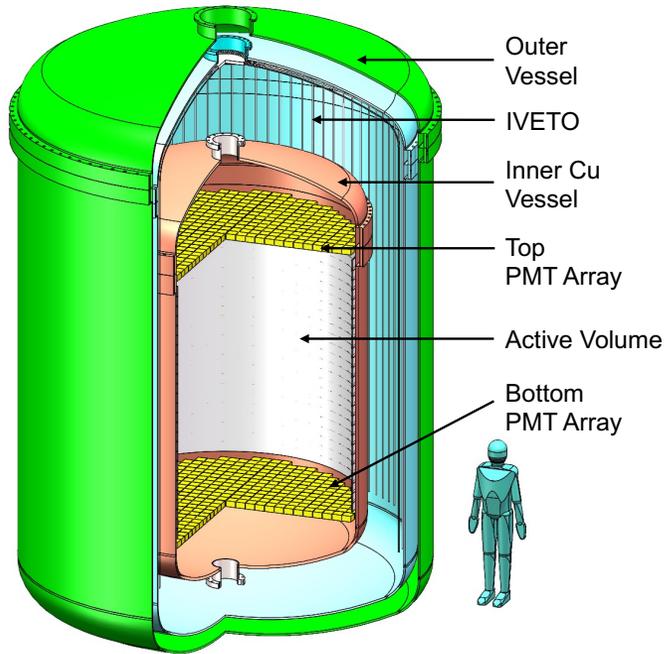
- World's leading result is from Borexino with a recoil energy of >165 keV
- PandaX-4T tried to measure the lower energy spectrum from 24 to 144 keV with a much smaller detector



CPC 48, 091001 (2024) Φ_{pp} ($10^{10} \text{ cm}^{-2} \text{ s}^{-1}$)

Future plan: PandaX-xT

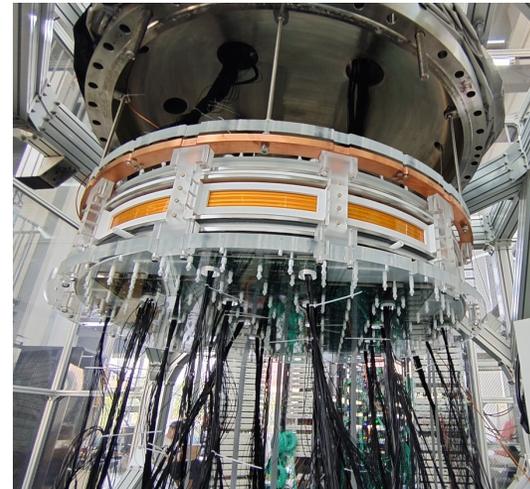
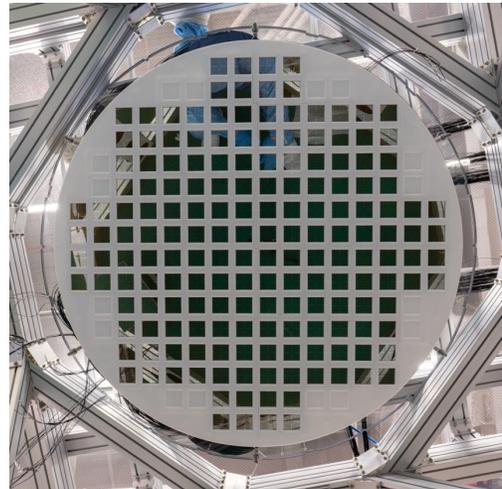
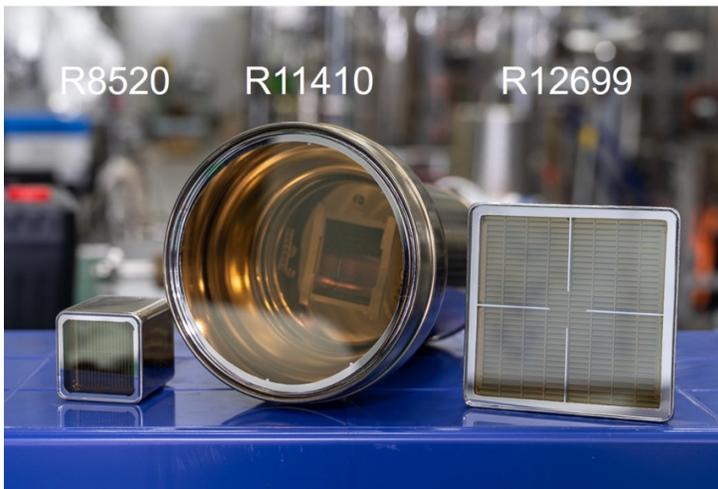
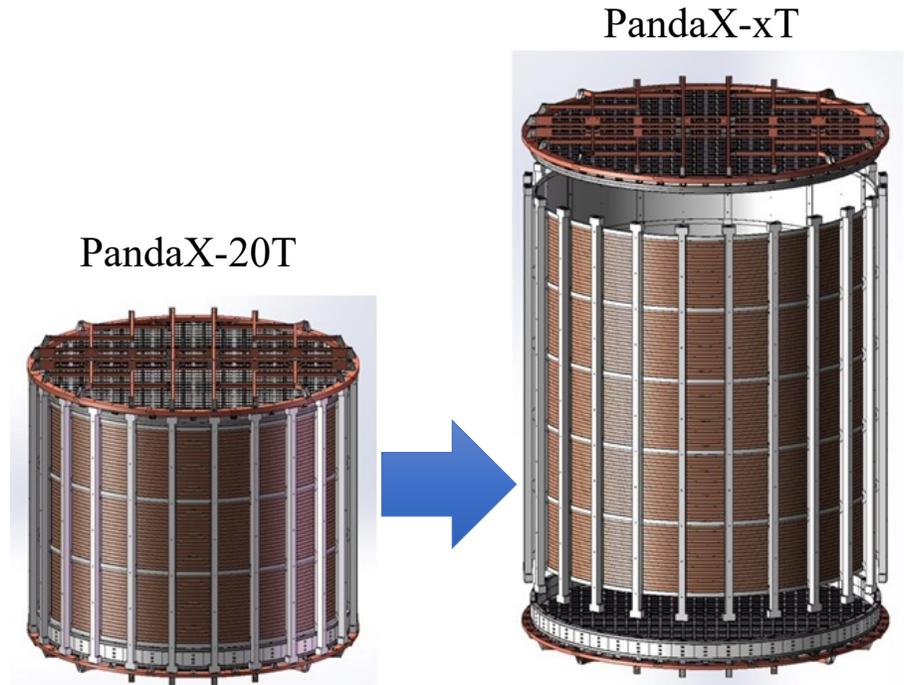
- Staged plan, finally reaching 43 tonne natural Xe in sensitive volume
- Key tests on WIMP and Dirac/Majorana neutrino



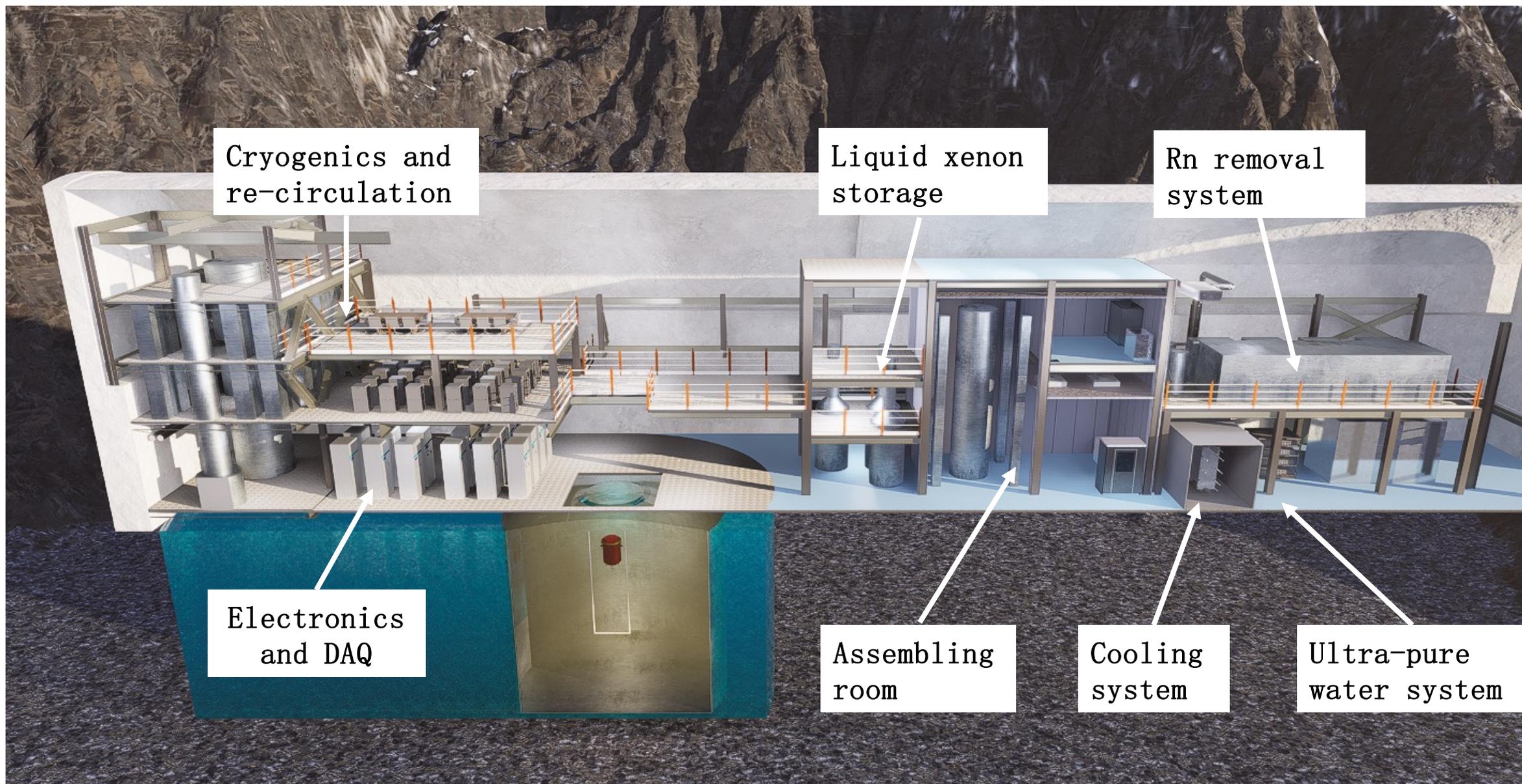
arXiv:2402.03596, SCPMA 68, 221011 (2025)

PandaX-20T: intermediate stage

- R&Ds already started
- Estimated timeline
 - **2026: move to CJPL and assembling**
 - **2027: commissioning**



Planned layout for PandaX-xT at CJPL-II B2 Hall



PandaX-xT First Open Meeting

PandaX-xT First Open Meeting

Apr 10 – 11, 2025
Tsung-Dao Lee Institute
Asia/Shanghai timezone

Enter your search term



- Multi-physics observatory: PandaX-4T → PandaX-xT
- Continue PandaX-4T data-taking in 2025
- Next stage: PandaX-20T is planned to get online in 2027
- Highly welcome new collaborators!

Thank You!

