

# MPPC Characterization Experiments

## Study the Speed of Light in Cable & Muon Rate at Distinct Angles

### Group Members

N.T.M.Hien<sup>\*</sup> N.T.Hang<sup>⊕</sup> T.T.M.Hoan<sup>◇</sup> N.P.Huy<sup>\*</sup> H.Q.Thang<sup>\*</sup>

*Supervisor: Dr. Thanh Dong*

*Supporter: Ph.D. student, Ngoc Tran; Ms. Linh Pham*

<sup>\*</sup>CANTI - Center for Applications of Nuclear Technique in Industry

<sup>⊕</sup>HCM UE - HCM University of Education

<sup>◇</sup>HCM US - HCM University of Science

<sup>\*</sup>USTH - University of Science and Technology of Hanoi

Introduction

Method

Result

Conclusion

Bonus

## Introduction

## Method

Examination of  $V_{\text{breakdown}}$   
Speed of Light in Optical Fibers  
Cosmic Ray Detection (Muon Rate)

## Result

Examination of  $V_{\text{breakdown}}$   
Speed of Light in Optical Fiber  
Cosmic Ray Detection (Muon Rate)

## Conclusion

## Bonus

# etoolbox

Introduction

Method

Result

Conclusion

Bonus

Introduction

Method

Result

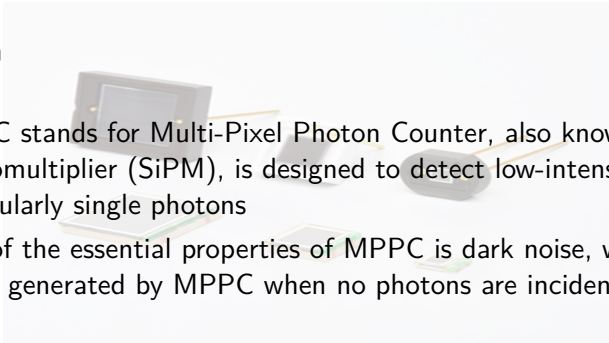
Conclusion

Bonus

## Objectives

- Create a simple yet proper cosmic ray muon detector using scintillators and Multi-Pixel Photon Counter (MPPC).
- Create a simple system to study the speed of light in optical fibers using MPPC.

## Definition

- 
- MPPC stands for Multi-Pixel Photon Counter, also known as Silicon Photomultiplier (SiPM), is designed to detect low-intensity light signals, particularly single photons
  - One of the essential properties of MPPC is dark noise, which is the output signal generated by MPPC when no photons are incident on the detector.

Introduction

**Method**

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fibers

Cosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

Introduction

**Method**

Examination of  $V_{\text{breakdown}}$

Speed of Light in Optical Fibers

Cosmic Ray Detection (Muon Rate)

Result

Conclusion

Bonus

Introduction

Method

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fibers

Cosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

Introduction

Method

Examination of  $V_{\text{breakdown}}$

Speed of Light in Optical Fibers

Cosmic Ray Detection (Muon Rate)

Result

Conclusion

Bonus



Introduction

Method

Examination of  
 $V_{\text{breakdown}}$ Speed of Light in  
Optical FibersCosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

## Setup

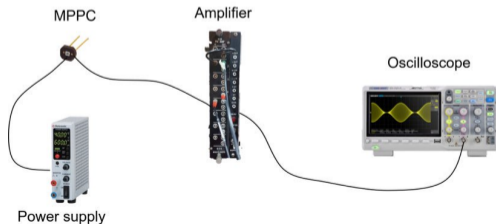
### Instrument

- Power Supply
- MPPC
  - Type: S13360-1325CS
  - Serial: 12482
  - $V_{\text{breakdown}}$  from manufacturer: 51.85V
  - $V_{\text{operator}}$ : 56.85V ( $=V_{\text{breakdown}} + 5V$ )
- Oscilloscope

## Setup

### Setup

- MPPC & a low-pass filter is placed inside a dark box, electrically connected.
- A power supply is linked to the box to generate electricity.
- The circuit will then send signal to an oscilloscope.
- We analyze the data coming on the oscilloscope in this setup.



## Procedure

- Provide power with the source supply.
- Use the *Vertical & Horizontal Scale*, as well as the *Trigger* button to adjust and find the pulse shape.
- Measure the distance from the base to the peak.



- There are mostly pulse of 1 p.e, sometimes 2 p.es, and rarely 3 p.es.

Introduction

Method

Examination of  
 $V_{\text{breakdown}}$ Speed of Light in  
Optical FibersCosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

## Purpose

- Determining the  $V_{\text{BR}}$  of MPPC initially is practically beneficial and recommendable.
- Most MPPC characteristics are dependent on  $V_{\text{over}}$ .

## Note

Recommended to compare or calibrate the characteristics of MPPCs at the same  $V_{\text{over}}$ .

Introduction

Method

Examination of  
 $V_{\text{breakdown}}$

**Speed of Light in  
Optical Fibers**

Cosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

Introduction

Method

Examination of  $V_{\text{breakdown}}$

**Speed of Light in Optical Fibers**

Cosmic Ray Detection (Muon Rate)

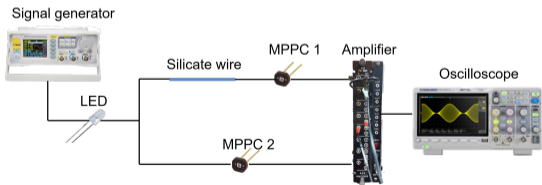
Result

Conclusion

Bonus

## Setup

- A signal generator is connected to an LED.
- The LED then be covered and linked by a cable which will be divided into 2 paths.
- Two MPPCs are joined at the end of two cables, respectively.
- MPPCs are bonded to the amplifier to augment signals.
- Signals then be transmitted to the oscilloscope for analysis purpose.



Introduction

Method

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fibers

Cosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

## Procedure

- Turn on the generator to provide electricity for the LED.
- Observe the received signals on the oscilloscope.
- Change the length of the fiber of one path. Then, observe the alternative signals.

Introduction

Method

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fibers

Cosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

## Purpose

- Determine the speed of light in a given optical fiber.
- From the obtained data, find the *Refractive Index* of that material.



Introduction

Method

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fibers

Cosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

Introduction

Method

Examination of  $V_{\text{breakdown}}$

Speed of Light in Optical Fibers

Cosmic Ray Detection (Muon Rate)

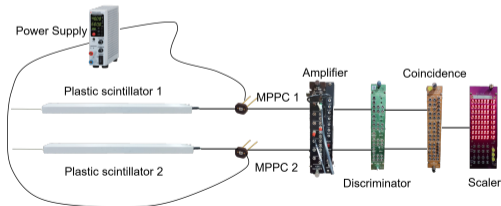
Result

Conclusion

Bonus

## Setup

- MPPC
- Scintillators connected with optical fibers
- Power supply
- Connect the optical fibers to the MPPC



## Setup

### MPPC board:

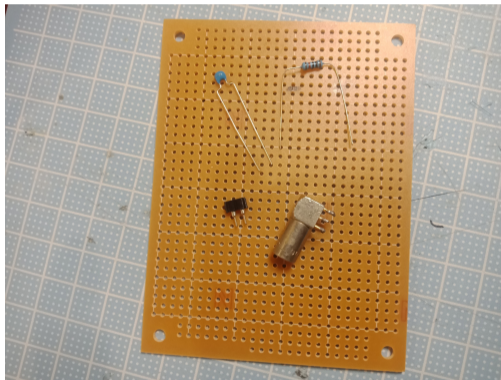


Fig. 1: LED board with components

## Setup

### MPPC board:

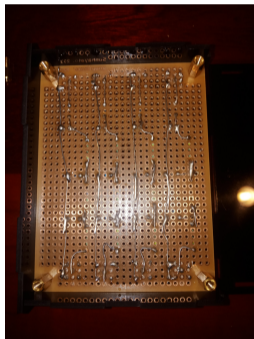


Fig. 2: LED circuit board's front face.

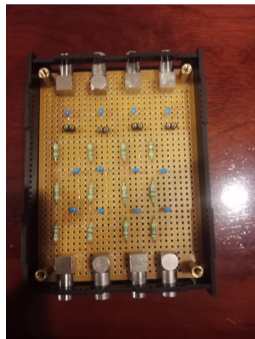


Fig. 3: LED circuit board's back face.

Introduction

Method

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fibers

Cosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

## Setup

### MPPC board



Fig. 4: Complete MPPC Circuit

Introduction

Method

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
 Optical Fibers

Cosmic Ray Detection  
 (Muon Rate)

Result

Conclusion

Bonus

## Setup

The scintillators & MPPC boards are coupled as the following figure:

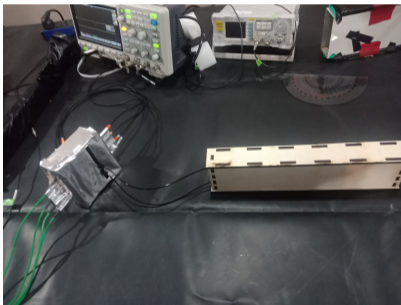


Fig. 5: MPPC & Scintillators setup for Muon Detection

Introduction

Method

Examination of

$V_{\text{breakdown}}$

Speed of Light in  
Optical Fibers

Cosmic Ray Detection  
(Muon Rate)

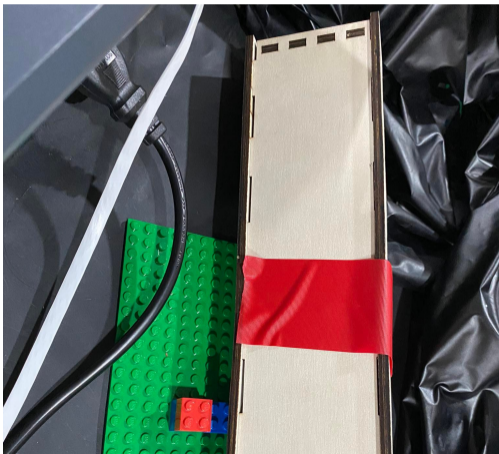
Result

Conclusion

Bonus

## Setup

The scintillators at different angles



## Procedure

- Provide power with the source supply.
- Use the *Vertical & Horizontal Scale*, as well as the *Trigger* button to adjust and set the threshold.
- Cover the scintillators & MPCC boards system to ensure that no incidence external photon could disturb and cause unwanted false signal.
- Record the muon counts for around 15 minutes on a measurement of an angle.
- Analyze the data acquired to have the muon counts distribution.



Introduction

Method

Examination of

$V_{\text{breakdown}}$

Speed of Light in  
Optical Fibers

Cosmic Ray Detection  
(Muon Rate)

Result

Conclusion

Bonus

## Purpose

- Calculate muon rate in different scintillator's angle.
- Obtain the muon counts distribuiton.

Introduction

Method

**Result**

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fiber

Cosmic Ray Detection  
(Muon Rate)

Conclusion

Bonus

Introduction

Method

**Result**

Examination of  $V_{\text{breakdown}}$

Speed of Light in Optical Fiber

Cosmic Ray Detection (Muon Rate)

Conclusion

Bonus

Introduction

Method

Result

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fiber

Cosmic Ray Detection  
(Muon Rate)

Conclusion

Bonus

Introduction

Method

Result

Examination of  $V_{\text{breakdown}}$

Speed of Light in Optical Fiber

Cosmic Ray Detection (Muon Rate)

Conclusion

Bonus

## Outcome

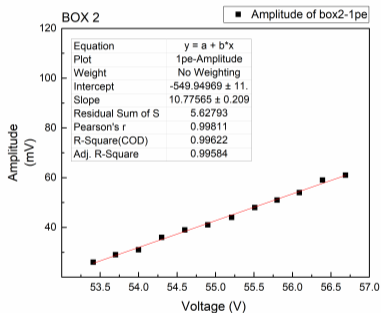


Fig. 6: Box 2 - The Relation between Supply Voltage and Pulse Amplitude - 1 p.e

## Discussion

- Our findings indicate that the breakdown voltage of the MPPC depends on the applied voltage. The higher the applied voltage, the higher the breakdown voltage.
- The result of the breakdown voltage that we obtained has an approximation error of about 3% and the standard deviation is not significant.
- limitations of our study are we only examined the breakdown voltage of one type of MPPC and under a limited range of conditions.
-

Introduction

Method

Result

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fiber

Cosmic Ray Detection  
(Muon Rate)

Conclusion

Bonus

Introduction

Method

**Result**

Examination of  $V_{\text{breakdown}}$

**Speed of Light in Optical Fiber**

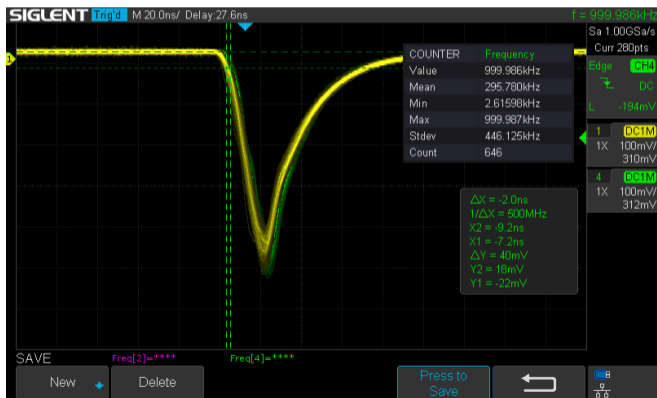
Cosmic Ray Detection (Muon Rate)

Conclusion

Bonus

## Outcome

- For 2 same length of fibers:



- For 2 same length of fibers:

Introduction

Method

Result

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fiber

Cosmic Ray Detection  
(Muon Rate)

Conclusion

Bonus

## Discussion

- The time for light travel is linearly proportional to the optical fiber's length. The slope corresponds to the velocity of light in optical fiber and it is 0.1867 m/ns. From this, we can calculate the refractive index.
- The refraction index is nearly the same in different wire lengths.



Introduction

Method

Result

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fiber

Cosmic Ray Detection  
(Muon Rate)

Conclusion

Bonus

Introduction

Method

Result

Examination of  $V_{\text{breakdown}}$

Speed of Light in Optical Fiber

Cosmic Ray Detection (Muon Rate)

Conclusion

Bonus

## Outcome

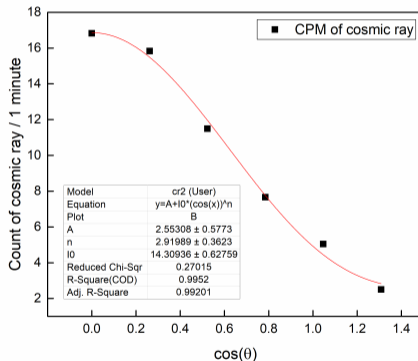


Fig. 13: Relation between the angle of scintillator and counts of cosmic rays per time

## Outcome

$\text{Cos}(\theta)$	Count/mins	Error
1.308	2.50	0.00
1.047	5.04	0.23
0.785	7.67	0.00
0.523	11.49	0.57
0.262	15.83	0.64
0.000	16.82	0.68

Fig. 14: Data table

Introduction

Method

Result

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fiber

Cosmic Ray Detection  
(Muon Rate)

Conclusion

Bonus

## Outcome

- From the plot, we can see that the bigger the angle between the scintillator and the horizontal axis, the larger the recorded cosmic ray count.
- From the experiment, the cosmic ray counts per time along the angle of the scintillator form a Gaussian distribution.

Introduction

Method

Result

Examination of  
 $V_{\text{breakdown}}$

Speed of Light in  
Optical Fiber

Cosmic Ray Detection  
(Muon Rate)

Conclusion

Bonus

## Discussion

- The errors occur due to the fact that we take the mean of measurements over a small time interval. If we consider a longer period of time, the counts of cosmic rays maybe more stable, thus reducing the error.

Introduction

Method

Result

**Conclusion**

Bonus

Introduction

Method

Result

**Conclusion**

Bonus

## Project Hindrance:

- Room temperature affects the natural properties of material.
- Noise from the other instruments.
- Could not create a total Faraday cage.
- Other sub-particles than Muon can be detected by scintillator, especially electron.
- Time limited to perform enough long experiments for stable measurements.
- Instrument errors.
- The **Biggest** hindrance is our own inexperience

## What have we learnt?

- How to use oscilloscope system, MPPC and some type of experiments related to it.
- Understand the principles of scintillation detectors, trigger DAQ, cosmic ray detection, etc.
- Know how to solder MPPC circuit, arrange measurement system, and polish optical fibers

Introduction

Method

Result

Conclusion

**Bonus**

Introduction

Method

Result

Conclusion

**Bonus**



As requested, this is our assignment for the homework of Prof. Suzuki.

## Simple experiment (2)

### Exercise

1. Particle decay follows the following formula:

$$N = N_0 e^{-t/\tau},$$

where  $N$  and  $N_0$  are the numbers of events at time  $t$  and 0, respectively, and  $\tau$  is the lifetime.

The right side table is a result of the experiment.

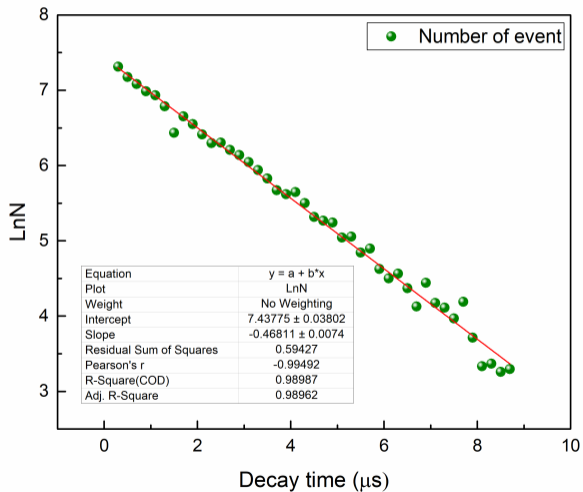
(a) Make a plot of the number of events as a function of the decay time (use log scale as a vertical axis).

(b) Get the muon lifetime.

2. Some events show no 2<sup>nd</sup> signal which corresponds to decay electron. What are those events ?

Decay time $t$ [ $\mu$ sec]	Number of events	Decay time $t$ [ $\mu$ sec]	Number of events
0.3	1501	4.7	194
0.5	1308	4.9	189
0.7	1191	5.1	155
0.9	1082	5.3	157
1.1	1024	5.5	127
1.3	886	5.7	134
1.5	823	5.9	102
1.7	775	6.1	90
1.9	700	6.3	96
2.1	610	6.5	79
2.3	544	6.7	62
2.5	547	6.9	85
2.7	497	7.1	65
2.9	463	7.3	61
3.1	422	7.5	53
3.3	380	7.7	66
3.5	340	7.9	41
3.7	291	8.1	28
3.9	276	8.3	29
4.1	283	8.5	26
4.3	245	8.7	27
4.5	204		

\* ex. "0.3" means  $0.1 \leq t < 0.4$  [ $\mu$  sec]



- For the first question:

$$N = N_0 e^{-\frac{t}{T}}$$

$$\Rightarrow \ln(N) = \ln(N_0) - \frac{t}{T}$$

$$\Rightarrow \ln(N) = \left(-\frac{1}{T}\right)t + \ln(N_0)$$

From the plot, we have the regression line function.

We have

$$\Rightarrow -\frac{1}{T} = -0.47 \Rightarrow T = 2.13 = \text{Muon lifetime}$$

- For the second question:  
The events that show no second signal correspond to muons that did not decay within the detector. Muons are unstable particles and will decay into an electron, a muon neutrino, and an electron antineutrino with a certain probability within their lifetime. However, some muons may travel through the detector without decaying and will not produce a second signal.