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Energy resolution of CANDLES detector for studying neutrino-less double beta decay of ⁴⁸Ca

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(on behalf of the CANDLES collaboration)

- 1. Introduction
- 2. CANDLES experiment
- 3. $2\nu\beta\beta$ & Energy resolution
- 4. Error in charge measurement
- 5. Photon Counting
- 6. Results and Discussion
- 7. Possibility for Photon Counting

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Double Beta Decay (DBD)

[1] Ann.Rev.Nucl.Part.Sci.52:115







- Obtained in >10 isotopes
- $T_{1/2}^{2\upsilon}$ = 10¹⁸~10²⁰ yr
- Rare, under standard model (SM)
- ***** <u>Physics of **0**νββ decay</u>:
- > Neutrino mass from the $\mathbf{T_{1/2}^{0\nu}}$ $\left(\mathbf{T_{1/2}^{0\nu}}\right)^{-1} = G^{0\nu} \left| \left\langle \mathbf{m_{\beta\beta}} \right\rangle^2 / m_e^2 \right| |M^{0\nu}|^2$

- No observation
- $T_{1/2}^{0\upsilon} > 10^{26} \text{ yr (KamLAND-Zen)}$
- Extremely rare!
- > Nature of neutrino: Majorana or Dirac?
- > Lepton number not conserved (Δ L=2)
 - \Rightarrow New physics beyond SM 4

0vββ experiment with ⁴⁸Ca

✓ Highest $Q_{\beta\beta}$ 4.27 MeV

- Large phase space factor
- Far from BKG (γ: 2.6 MeV; β: 3.3 MeV)
- ⇒Aim for background-free measurement

X Low abundance

- Natural abundance: <0.2 %
- Separate isotopes: expensive
- ⇒Cost-effective enrichment
- □ Energy Resolution $T_{1/2}^{0V} \propto (N_{BKG} \cdot \Delta E)^{-1/2}$ ⇒ Improve sensitivity



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Energy Resolution

 \Rightarrow Improve sensitivity

Future CANDLES



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CANDLES experiment

–To observe $0\nu\beta\beta$ of ^{48}Ca

CANDLES

CAlcium fluoride for studies

of Neutrino and Dark matters

by Low Energy Spectrometer

-Set up @ Kamioka (2700m.w.e depth)





–To observe $0\nu\beta\beta$ of ^{48}Ca

- -Set up @ Kamioka (2700m.w.e depth)
- -CANDLES consists of:
 - 96 CaF₂(nat.): detector + source \Rightarrow 350g ⁴⁸Ca (305kg CaF₂)
 - Liquid scintillator (LS): $2m^3$, 4π active veto









relocate event position

Position X (mm)



CAlcium fluoride for studies of Neutrino and Dark matters by Low Energy Spectrometer CANDLES by Low Energy Spectrometer

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stitute for Laser Technology **T. Shigeki** [**0**νββ] arxiv:2008.09288 (for PRD) [**0**νββ] PRC 78:058501 [**0**νββ] Nucl. Phys. A 730:215

[DAQ] IEEE TNS 66:1174 [DAQ] IEEE TNS 62:1122 [DAQ] IEEE TNS 62:1128

[Detector] NIMA 986:164727 [Detector] Astropart. Phys.100:54 [Detector] NIMA 705:1 [Detector] NIMA 601:282

[Enrich] J. Nucl. Sci. Tech. 55:1473
[Enrich] Austin Chromatogr. 3:1040
[Enrich] J. Chroma. 1415:67
[Enrich] PTEP 2015:053C03
[Enrich] PTEP 2015:033D03

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Background in CANDLES

- Background at $Q_{\beta\beta}$ of ⁴⁸Ca:
- Most background: removed by active shielding
- <u>External (n,γ</u>): passive shielding (Pb,B)
 <u>Impurities background</u>:
- ²¹²Bi²¹²Po sequential decay: pile-up event
- \Rightarrow Waveform analysis
- ²⁰⁸T1 β-decay: remove by tagging preceding α-decay
- \Rightarrow tagging efficiency (DAQ + Analysis)

 $\circ 2\nu\beta\beta$: not reduced by above methods

$2\nu\beta\beta$ in CANDLES

- $2\nu\beta\beta$: irremovable background
- Proportional to M(⁴⁸Ca)
- ~600kg of ⁴⁸Ca in future
- \Rightarrow huge $2\nu\beta\beta$ background
- To distinguish $2\nu\beta\beta$ and $0\nu\beta\beta$

 \Rightarrow Must improve energy resolution.

- Ideal case, resolution = statistical fluctuation of number of p.e.
- Current setup:
- \Rightarrow at Q_{$\beta\beta$}(4272keV): $\sigma_{p.e.} \approx 1.6\%$

• Current resolution:

σ_E=2.6% > σ_{p.e.}
⇒Other fluctuation(s) make energy resolution worse!

 $*\sigma_{\rm E}$ = 2.6%, reported in T. Ohata Dthesis (2018)

Energy Resolution

- Energy resolution depends on: **1.Statistical fluctuation**
- Mainly from fluctuation of number of p.e.
- ⇒• Light yield (cooling detector)
 - Photon collection (light pipes)

2. Detector stabilityTemperature and high-voltage
Monitor during data taking

[Ref] T. Ohata DThesis O.U. (2018)

- 3. Crystal dependence ^[Ref]
 Numbers of scintillation photons
 from different crystals are different.
 ⇒ Calibration for each crystal
 ⇒ Small fluctuation
 (0.3% at 2.6 MeV)
- 4. Error in charge measurement

Research motivation:

 Improve energy resolution.
 ⇒Study the error of charge measurement in CANDLES III

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Error in charge measurement

- $\sigma_{\rm E}$ =2.6% > $\sigma_{\rm p.e.}$ =1.6%
- CaF₂ has a long decay constant 1µsec
- To calculate the energy, signal integration of 4µsec
- \Rightarrow Baseline fluctuation can be accumulated

- Possible fluctuations in a long interval:

 Dark Current in PMTs
 Noises in baseline
 - Digitization error (resolution of FADC)
- \Rightarrow Study the above fluctuations to identify the problem

- Dark Current in PMTs
- Noises in baseline
- Digitization error

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- a. Dark Currentb. Noise in baselinec. Digitization Error

Dark Current (& Small Scint. Light)

- Dark Current affects statistically on the energy resolution.
- Dark Currents in every 100ns of each PMT are counted.
- \Rightarrow Sum dark current in 62 PMTs to

estimate the effect in the CaF₂ waveform.

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- Dark Currents in every 100ns of each PMT are counted.
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- Dark current rate of 10-inch PMTs are the lowest ones.
- The dark current fluc. of 62 PMTs in 4000ns integration: ~1.6 p.e.
- Fluctuation of dark current is negligibly small:

at Q-value, $\sigma_{DC}/Q = 0.04\%$

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a. Dark Currentb. Noise in baselinec. Digitization Error

Noise in Baseline

Noise Amp (fitting) of 62 PMTs

- Fit the baseline of each PMT (sine-wave)
 ⇒get the amplitudes and cycles of noises
 ⇒Highest amplitudes in Five 10-inch PMTs
 ⇒Similar noise cycles in these PMTs
 ⇒Sum all 10-inch PMTs to check the noise
- In sum baseline of 10-inch PMTs (x12):
- Cycle 730 ns (1.3-1.4 MHz)
- Amplitude ~0.73 ADC (3mV)
- \Rightarrow Amplitude < 1 ADC even sum up 12 PMTs

Noise in Baseline

- Noise effect in an interval T is \Rightarrow a function of phase $(\int_0^T \sin(x + \varphi) dx)$
 - Phase factor is random in CaF₂ waveform.
 (difficult to estimate)

 \Rightarrow Estimate the max effect at each interval

 \Rightarrow At T=4000ns, effect is about $\sigma_{noise} \leq 2p.e.$

 $\Rightarrow \sigma_{noise} / Q \le 0.05\%$ (still negligible)

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- a. Dark Current
 b. Noise in baseline
 c. Digitization Error

We are using ADCs: *ADC0850DL, Texas Ins. *ENOB=7.5bits @ 125MHz

Digitization error (DE):

- Difference between measured pedestal and true pedestal (b/c of LSB)
- As a function of pedestal
- Calculate charge:
- \Rightarrow Accumulated in signal integration

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1pe

signal

Digitization error (DE):

- Difference between measured pedestal and true pedestal (b/c of LSB) Digit. Err.
- As a function of measured pedestal
- Calculate charge: $\sum (\text{Ped} \text{Signal}[i])$
- \Rightarrow Accumulated in signal integration

- In 1p.e. signal:
- \Rightarrow DE accumulated by the width
- \Rightarrow Estimate as a function of ped.
- Signal contains many 1p.e.
- \Rightarrow Digitization error in each PMT can be estimated using:
 - Number p.e. & Pedestal⁹

• From experimental data,

 \Rightarrow number of p.e. and pedestal

⇒Estimate the digitization error on

the energy spectrum (40K & 208TI)

• From the Digitization Error dist.,

 $\circ\,\mu_{\text{DE}}$ affects peak position

 $\circ \sigma_{\text{DE}}$ affects energy resolution

•Fluctuations of DE at different peaks are estimated: $\circ {}^{40}$ K (1.46MeV), σ_{DE} /E = 0.6% $\circ {}^{208}$ Tl (2.6MeV), σ_{DE} /E = 0.4% \Rightarrow At Q-value, σ_{DE} /Q is small₃₀

Error in charge measurement (review)

- Several sources of fluctuations were checked
- However,
 - > Dark Current ($\sigma_{DC}/Q_{\beta\beta}$) > negligible
 - > Baseline Noise ($\sigma_{noise}/Q_{\beta\beta}$) negligible
 - > Digitization Error ($\sigma_{DE}/Q_{\beta\beta}$) > small
- \Rightarrow I found the <u>most severe fluctuation in the baseline</u>

is the pedestal uncertainty.

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- a. Dark Current
 b. Noise in baseline
 c. Digitization Error
 d. Pedestal Uncertainty

Pedestal Uncertainty

- In this research, the pedestal of each PMT is calculated using first 40 bins
- Ideally, binomial fluctuation below LSB

- Due to the noise in every PMT
- ⇒The pedestal uncertainty is about 2 times larger than the binomial one.
- ⇒Study pedestal uncertainty at

zero-energy signal

*Signal Integration = N×Ped - Σ_i^N Signal[i]

- In case we sum up many PMTs, it is more severe effect.
- For 62 PMTs, integration interval of 4000 ns, the accumulated fluctuation is $\sigma_{PedErr} = 38.6 \text{ p.e. or } \sigma_{PedErr} / Q_{\beta\beta} = 1\%$
- Compare $\sigma_{p.e.} / Q_{\beta\beta} \approx 1.6\%$ $\Rightarrow \sigma_{PedErr}$ causes a severe fluctuation

Pedestal Uncertainty

- In Signal Integration, the pedestal uncertainty is accumulated at every data point*.
- \Rightarrow Accumulated Fluc = N x σ_{PedErr}
- ⇒The accumulated fluctuation is linearly proportional to number of data points

SUMMARY of fluctuations

Full integral	⁴⁰ K peak	²⁰⁸ Tl peak	⁴⁸ Ca Q-value
(4000 nsec)	(1460 keV)	(2614 keV)	(4272 keV)
σ _{PedErr} : Ped Error	38.6 p.e.	38.6 p.e.	38.6 p.e.
(σ _{PedErr} /N _{pe})	(2.9 %)	(1.6 %)	(1.0 %)
σ _{DE} : Digit. Err.	7.3 p.e.	10.4 p.e.	(small)
(σ _{DE} /N _{pe})	(0.6 %)	(0.4 %)	
σ _{noise} : 730ns noise	<mark>≤2</mark> p.e.	≤2 p.e.	≤2 p.e.
(σ _{HF} /N _{pe})	(≤ 0.15 %)	(≤ 0.08 %)	(≤ 0.05 %)
σ _{DC} : Dark Current	1.6 p.e.	1.6 p.e.	1.6 p.e.
(σ _{DC} /N _{pe})	(0.1 %)	(0.06 %)	(0.04 %)
 Fluctuations from Dark Current, Noise, 			

 Fluctuations from Dark Current, Noise, Digitization Error and Pedestal uncertainty are estimated as functions of integration interval.
 ⇒Fluctuation from Pedestal Uncertainty (σ_{PedErr}) is the most severe

 \Rightarrow Signal integration to get energy \Rightarrow **not good** \Rightarrow Need another method to get energy

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- a. Motivationb. Measurementc. Analysis
Photon Counting: Motivation



SIGNAL INTEGRATION

- The baseline fluctuations: severe
- In current analysis, signal integration is used
 ⇒Baseline fluctuations are accumulated



- Setting threshold
- Count p.e. in each PMT
- \Rightarrow No integration
- \Rightarrow Avoid baseline fluctuation 37

Photon Counting: Measurement



- Waveform of CANDLES:
- First 768 ns, 2ns/sample
- Latter sum every 64 ns.
- The 1p.e. width is short (<50ns)
- If 1p.e. arises after 768 ns, it is difficult to see it.
- \Rightarrow Need to modify the DAQ software
- After modification, data size x3 times larger.

	Physics Run	Photon Count
Size	~50kB/event	~150kB/event
Waveform interval	~9µsec	~4.2µsec



Photon Counting: Measurement



Photon Counting: Measurement

- Maximize 2 nsec means extend a lot of data size in event buffer of FADCs
- Size/event is 3 times larger
 ⇒The readout time/event is ~20msec
 ⇒X2 times longer than the Physics Run
 ⇒Estimate the data taking efficiency.
- Data taking efficiency = N_{record}/N_{incident}
 ⇒Thanks to our development of DAQ system (with 8 event buffers)
 ⇒Acquire ~100% data taking efficiency



Photon Counting analysis: Threshold



 μ_p : mean of 1p.e. pulse height dist. σ_p : rms of 1p.e. pulse height dist.

- In photon counting, a threshold is set to count the photoelectrons.
 - If it is set too low: noises are counted as p.e.
- If it is set too high: we may lose p.e. when counting.
 - Three thresholds μ_p -1 σ_p , μ_p -2 σ_p and μ_p -3 σ_p are used to test
- Check the separation of 1p.e. from baseline \Rightarrow The threshold for photon counting μ_{p} -2 σ_{p}

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- Check the separation of 1p.e. from baseline \Rightarrow **The threshold for photon counting** μ_p -**2** σ_p



Overlap of 1p.e. signals



- Many 1p.e. signals overlap
 ⇒forming into a multi p.e. signal
 (2p.e., 3 p.e., etc.)
- With simple photon counting, a multi p.e. signal is counted as 1 p.e. signal.
 ⇒Miss photoelectrons in counting
- \Rightarrow Bad energy resolution.
- In average, each PMT detects ~62p.e. at $Q_{\beta\beta}$ \Rightarrow Overlap is serious at Q-value!
- The multi p.e. are found at the rising edge of the waveform.

⇒Introduce an alternative method named

"Partial Photon Counting".

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⇒Introduce an alternative method named "Partial Photon Counting". 45

Partial Photon Counting (PPC)



Divide the waveform each PMT into 2 areas: -The 1st area (near rising edge): \Rightarrow many multi p.e. \Rightarrow Signal integral to avoid the lost of p.e. -The 2nd area (near the tail): \Rightarrow not so many multi p.e. \Rightarrow Photon counting to avoid pedestal fluc. - In my research \Rightarrow T_{INTEGRAL} + T_{COUNT} = 4000ns \Rightarrow Check different mixtures of integration and photon counting

Partial Photon Counting (PPC)



Partial Photon Counting (PPC): Result

- Evaluate the energy resolution at 1460keV (40 K) and 2614.5keV (208 Tl) for each histogram \Rightarrow Resolution: a function of signal integral gate \Rightarrow The resolution is improved:
 - σ_E/E(⁴⁰K):~4.5% to ~4.0%
 - σ_E/E(²⁰⁸TI):~3.3% to ~2.9%





Gaussian of ²⁰⁸Tl (impurities in crystal) Gaussian of ²¹⁴Bi (impurities in crystal) Gaussian of ⁴⁰K (impurities in PMTs) Error function as Compton background

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 Study the energy resolution at 1460keV (⁴⁰K) and 2614.5keV (²⁰⁸TI)
 ⇒The baseline fluctuation can explain the tendency of improved resolution

- There is still
 - remaining fluctuation:
 - independent with the integration interval
 - depends on the energy
 (assuming) ⁵³



- Estimate the resolution at Q-value
 ⇒need a fitting equation
- Considering all fluctuations in my study, fluctuations can be divided in 2 groups:
 Depending on the energy
 - Statistical fluctuation, digitization error, remaining fluctuation

oIndependent with the energy pedestal error

- The fitting function is:
- Apply this equation on the energy resolutions of CANDLES ⇒ Good fitting
- The obtained result at ⁴⁸Ca is consistent with the result obtained in prev. researches

(at T_{INTEGRAL} = 4000ns).

 \Rightarrow The resolution of ⁴⁸Ca is improved to 2.2⁵%

• In previous research*, the sensitivity of of 48 Ca was estimated as: \square 93 non-enriched crystals of CaF₂(pure) \square T = 129.5 days • In previous research*, the sensitivity of of 48 Ca was estimated as: current sensitivity for $0\nu\beta\beta$ with Full Integration 0.44–0.50 × 10²³ yr,

 $\Box \sigma_E / Q_{\beta\beta} = 2.6\%$ (Full Integration)

- In current CANDLES, $2\nu\beta\beta$ is not dominant
- \Rightarrow The sensitivity in $0\nu\beta\beta$ search is proportional to ($\sigma_{\rm E})^{\text{-1/2}}$
- In the same conditions, if σ_{E} is improved, the sensitivity is increased by a

factor $\sqrt{\sigma_{Full Integration}/\sigma_{Partial Photon Count}}$

 \Rightarrow The sensitivity of CANDLES with $\sigma_{\rm E}/Q_{\beta\beta}$ = 2.2% is improved by = 1.09

If the energy resolution can be improved to 2.2%
 ⇒the sensitivity can be improved to

*T. Ohata, Doctor Thesis, Osaka University (2018)

expected sensitivity for $0\nu\beta\beta$ with **PPC method**

90% C.L.

0.48–0.55 ×10²³ yr, 90% C.L. ⁵⁵

SUMMARY (1)

- Resolution is important in studying $0\nu\beta\beta$ for CANDLES
- The current $\sigma_{\rm E}/Q_{\beta\beta}$ =2.6% > $\sigma_{\rm p.e.}/Q_{\beta\beta}$ =1.6%
- \bullet Pedestal fluctuation (σ_{PedErr}) makes the resolution worse.
- \Rightarrow This fluctuation is negligible with short decay constant (PLS, LS)
- \Rightarrow For CaF₂ in CANDLES, it is large fluctuation ($\sigma_{PedErr}/Q_{\beta\beta}$ =1%)
- \Rightarrow My research goal: reduce $\,\sigma_{\,_{\text{PedErr}}}$ to improve resolution of CANDLES
- \Rightarrow Introduce "partial photon counting".
- \Rightarrow The energy resolution is improved.
- The energy resolution at Q-value is expected to be improved to 2.2% \Rightarrow With this improvement, the sensitivity can be **<u>1.09 times improved</u>**.

SUMMARY (2)

- CANDLES faces a large baseline fluctuation, which can be reduced by "partial photon counting" (PPC)
- PPC requires a lot of data space for daily data taking
- \Rightarrow ~40 GB/day \rightarrow ~120 GB/day
- \Rightarrow In the next step, DAQ improvements (software and firmware) to reduce the data size
- \Rightarrow Energy calibration, resolution at $Q_{\beta\beta}$, long measurement...

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- In the future generations:
- CANDLES plans to introduce the bolometer detector*
- ⇒to improve energy resolution and sensitivity
- Before going to bolometer detector, We still have a possibility to improve detector resolution with CaF₂(pure) at low temperature + photon counting ⇒Discussion in next slides



• At temperature \leq 100 K, Light output

is increased x4 times.

The statistical fluctuation can be

improved x2 times: $\sim 1.6\% \rightarrow \sim 0.8\%$.

 Cooling machines or Liquefied gas (LAr, LN₂, LNe, LH₂)



However, ...

- At cryogenic temperature, the waveform is extended, τ_{decay} : ~1µsec \rightarrow ~40µsec.
- If we use signal integral: a huge σ_{baseline}
- Roughly check p.e. rate (p.e./ns):

○293K, N = 3838, τ = 1µsec ⇒ ~3.8p.e./ns
○<30K, N = 15352, τ = 40µsec ⇒ ~0.4p.e./ns
⇒Less p.e. rate; less overlapping prob.
⇒Photon counting can work!
⇒Avoid the baseline fluctuation
⇒Energy resolution can be much better compared to the current detector 61

- The world best sensitivity is reported KamLand-Zen: $m_{\beta\beta} \approx 61-165 meV^{(*)}$ (above IH region)
- With CaF₂(pure) low temp.
 + photon counting
- We may achieve $\sigma/Q_{\beta\beta} = 0.9\% 1.1\%^{(**)}$
- With this resolution @ $Q_{\beta\beta}$, we may achieve much improved sensitivity: $m_{\beta\beta} \approx 27-118 \text{ meV} (\sigma/Q_{\beta\beta} = 0.9\%)$ $m_{\beta\beta} \approx 80-240 \text{ meV} (\sigma/Q_{\beta\beta} = 1.1\%)$ \Rightarrow Can be close to (or better than) the current world-best limit of $m_{\beta\beta}$



(*) due to different NME values (**) stat. + dark current ^{q2} remain

SUMMARY (3)

- I discussed CaF₂(pure) at low temperature + Photon Counting.
- \Rightarrow Can use liquefied gas or cooling machine
- $\Rightarrow \sigma/E$ at Q-value (expected) ≈ 0.9 -1.1%
- \Rightarrow The resolution is promising to achieve a better sensitivity for
 - CANDLES experiment

CANDLES Collaboration Meeting

2016, 10, 30 at Osaka University

hank vol

BACK UP

0vββ experiment with ⁴⁸Ca

✓ Highest $Q_{\beta\beta}$ 4.27 MeV

- Large phase space factor
- Far from BKG (γ: 2.6 MeV; β: 3.3 MeV)
- \Rightarrow Aim for background-free measurement

X Low abundance

- Natural abundance: <0.2 %
- Separate isotopes: expensive
- ⇒Cost-effective enrichment

Energy Resolution

 \Rightarrow Improve sensitivity



2009

2011 Mar.

- 2012 Mar.
- 2014 Mar.~Sep.

2015 Apr.

: Setup CANDLES-III detector at Kamioka

- : Introduce energy calibration system with ⁸⁸Y source
- : Introduce light pipes \Rightarrow light collection efficiency
- : Magnetic Cancelation coil \Rightarrow photoelectrons collection efficiency & Cooling system \Rightarrow increase light yield
- : Passive shielding (Pb+B) \Rightarrow reduce (n, γ) background (later)



- 2009 2011 Mar. 2012 Mar.
- 2014 Mar.~Sep.
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A box designed for:

- No (contaminated) air leakage inside our detector
- Easily adjust source position vertically to calibrate each crystal
- Small-size source (mm scale)

2009 2011 Mar. 2012 Mar. 2014 Mar.~Sep.

2015 Apr.



- 2009
- 2011 Mar.
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- : Magnetic Cancelation coil \Rightarrow photoelectrons collection efficiency & Cooling system \Rightarrow increase CaF₂ light yield
- : Passive shielding (Pb+B) \Rightarrow reduce (n, γ) background (later)



- Geomagnetic field affects photoelectron collection in PMTs \Rightarrow Magnetic cancellation coil around detector
 - \Rightarrow Photoelectrons x1.29 times increased
 - CaF₂'s light yield increases at low temperature
 - \Rightarrow Cool all crystals to 4°C degree
 - \Rightarrow Light yield x1.33 times increased

- 2009
- 2011 Mar.
- 2012 Mar.
- 2014 Mar.~Sep.
- 2015 Apr.

- : Setup CANDLES-III detector at Kamioka
- : Introduce energy calibration system with ⁸⁸Y source
- : Introduce light pipes \Rightarrow light collection efficiency
- : Magnetic Cancelation coil \Rightarrow photoelectrons collection efficiency & Cooling system \Rightarrow increase CaF₂ light yield
- : Passive shielding (Pb+B) \Rightarrow reduce (n, γ) background

(reduced x100 times for E>5MeV)



Pb blocks for shielding γ -rays



Boron for shielding neutron

4-5mm thickness. Covered 100m² area

B₄C 40% wt silicone rubber (surrounding)

Liquid type for tank's bottom



(a) DAQ Middleware for Physics Run



DAQ-Middleware for Photon Counting

(b) DAQ Middleware for Photon Counting






Fig. 4. Triple-star network in μ TCA crate. The first star is used to connect the trigger module with the AMC-FADC modules. The second star is used for common clock distribution. The final star is used to connect the SpaceWire router with the AMC-FADCs.

●新DAQシステムの使用モジュール

Module	Manufacturer	FPGA	FPGA
		logic	development
μ TCA	Uber Ltd.		
MCH	Shimafuji	SpW-GbE ⁱ interface	by Shimafuji
		Clock distribution	
		SpaceWire Router	Open IP ⁱⁱ
			by Shimafuji
		Trigger Controller	by Osaka
		for CANDLES	University
AMC- FADC	Shimafuji	FADC control	by RCNP ⁱⁱⁱ
			Osaka University
		SpaceWire	Open IP ⁱⁱ
			by Shimafuji



MasterModule



 ⁱ SpW-GbE: The acronyms "SpW" and "GbE" are used in this paper to indicate "SpaceWire" and "Gigabit-Ethernet", respectively.
ⁱⁱ Open IP: The FPGA's intellectual property core is publicly available.

iii RCNP: Research Center for Nuclear Physics



