

# Study Status

Nguyen Thi Kim Ha  
Ho Chi Minh University of Science

NuGroup Meeting, April 13th 2018

# Non-unitary mixing matrix

- 1 Introduction
- 2 The effective Lagrangian
- 3 Neutrino oscillations in vacuum

# Introduction

The  $n \times n$  mixing matrix will indeed be unitary to preserve probability, the same is not true for any given  $m \times m$ , with  $m < n$

$$U_{\text{PMNS}}^{\text{Extended}} = \begin{pmatrix} \overbrace{\begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix}}^{U_{\text{PMNS}}^{3 \times 3}} & \cdots & U_{en} \\ \vdots & \ddots & \vdots \\ U_{s_n 1} & U_{s_n 2} & U_{s_n 3} & \cdots & U_{s_n n} \end{pmatrix}$$

# The effective Lagrangian

The effective low-energy Lagrangian in the mass basis

$$\begin{aligned}
 \mathcal{L}_{eff} = & \frac{1}{2}(\bar{\nu}_i i \partial \nu_i - \bar{\nu}^c_i m_i \nu_i + hc) \\
 & - \frac{g}{2\sqrt{2}}(W_\mu^+ \bar{l}_\alpha \gamma_\mu (1 - \gamma_5) N_{\alpha i} \nu_i + hc) \\
 & - \frac{g}{2 \cos \theta_W}(Z_\mu \bar{\nu}_i \gamma^\mu (1 - \gamma_5)(N^\dagger N)_{ij} \nu_j + hc) + \dots \quad (1)
 \end{aligned}$$

The relation between mass and flavour eigenstates

$$\nu_\alpha = N_{\alpha i} \nu_i \quad (2)$$

Mass eigenstates are orthonormal  $\langle \nu_i | \nu_j \rangle = \delta_{ij}$

Flavour eigenstates are not orthonormal

$$\langle \nu_\beta | \nu_\alpha \rangle = (\tilde{N} \tilde{N}^\dagger)_{\beta\alpha} \neq \delta_{\alpha\beta}$$

where  $\tilde{N}$  is the normalization factor

The effective Lagrangian in the flavour basis

$$\begin{aligned}
 \mathcal{L}_{\text{eff}} = & \frac{1}{2} \left( i\bar{\nu}_\alpha \partial (NN^\dagger)^{-1}_{\alpha\beta} \nu_\beta - \bar{\nu}^c_\alpha [(N^{-1})^t m N^{-1}]_{\alpha\beta} \nu_\beta + hc \right) \\
 & - \frac{g}{2\sqrt{2}} [W_\mu^+ \bar{l}_\alpha \gamma_\mu (1 - \gamma_5) \nu_\alpha + hc] \\
 & - \frac{g}{2 \cos \theta_W} [Z_\mu \bar{\nu}_\alpha \gamma^\mu (1 - \gamma_5) \nu_\alpha + hc] + \dots \quad (3)
 \end{aligned}$$

# Neutrino oscillations in vacuum

We have

$$|\nu_\alpha\rangle = \frac{1}{\sqrt{(NN^\dagger)_{\alpha\alpha}}} \sum_i N_{\alpha i}^* |\nu_i\rangle = \sum_i \tilde{N}_{\alpha i}^* |\nu_i\rangle \quad (4)$$

$$\begin{aligned} |\nu_\alpha(L)\rangle &= \sum_i \tilde{N}_{\alpha i}^* |\nu_i(L)\rangle = \sum_i \tilde{N}_{\alpha i}^* e^{-i\phi} |\nu_i(0)\rangle \\ &= \sum_i \tilde{N}_{\alpha i}^* e^{-i\phi} (\tilde{N}_{\beta i}^*)^{-1} |\nu_\beta\rangle \end{aligned} \quad (5)$$

The probability amplitude of flavour transition  $\nu_\alpha \rightarrow \nu_\beta$  after distance L

$$\begin{aligned}
 A_{\nu_\alpha \rightarrow \nu_\beta} &= \langle \nu_\beta | \nu_\alpha(L) \rangle = \sum_i \tilde{N}_{\alpha i}^* e^{-i\phi} \tilde{N}_{\beta i} \\
 &= \frac{1}{\sqrt{(NN^\dagger)_{\alpha\alpha} (NN^\dagger)_{\beta\beta}}} \sum_i N_{\alpha i}^* e^{-i\phi} N_{\beta i} \quad (6)
 \end{aligned}$$



## The oscillation probability

$$\begin{aligned}
 P_{\nu_\alpha \rightarrow \nu_\beta} &= |A_{\nu_\alpha \rightarrow \nu_\beta}|^2 = \frac{1}{(NN^\dagger)_{\alpha\alpha}(NN^\dagger)_{\beta\beta}} \left| \sum_i N_{\alpha i}^* e^{-i\phi} N_{\beta i} \right|^2 \\
 &= \frac{1}{(NN^\dagger)_{\alpha\alpha}(NN^\dagger)_{\beta\beta}} \left( \sum_i |N_{\alpha i}^*|^2 |N_{\beta i}|^2 \right. \\
 &\quad + 2 \sum_{j>i} \operatorname{Re}[N_{\alpha i}^* N_{\beta i} N_{\alpha i} N_{\beta i}^*] \cos \frac{\Delta m_{ji}^2 L}{2E} \\
 &\quad \left. - 2 \sum_{j>i} \operatorname{Im}[N_{\alpha i}^* N_{\beta i} N_{\alpha i} N_{\beta i}^*] \sin \frac{\Delta m_{ji}^2 L}{2E} \right) \quad (7)
 \end{aligned}$$